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Air Force Institute of Technology (ATC) Wright Patterson AFB, OH 45423

THREE BURN INERTIAL UPPER STAGE OPTIMAL ORBIT TRANSFER

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Thesis

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Presented to the Faculty of the School of Engineering of the Air Force Institute of Technology

> Air University In Partial Fulfillment of the Requirements for the Degree of Master of Science

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Mark N. / Brown, B.S. USAF



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Graduate Astronautical Engineering

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List of Symbols

x	-	distance	(x-direction)
у	-	distance	(y-direction)

z - distance (z-direction)

xd - velocity (x-direction)

- yd velocity (y-direction)
- zd velocity (z-direction)

T1 - coast time from parking orbit fix to first stage firing

- Al alpha rotation of first stage alignment
- B1 beta rotation of the first stage alignment
- T2 coast time from first stage firing to second stage firing
- A2 alpha rotation of second stage alignment
- B2 beta rotation of the second stage alignment
- T3 coast time from second stage firing to third stage firing
- A3 alpha rotation of the third stage alignment

B3 - beta rotation of the third stage alignment

 $\Delta V1$ - change in velocity due to first stage firing

AV2 - change in velocity due to second stage firing

 $\Delta V3$ - change in velocity due to third stage firing

G - performance index (Tf)

F - augmented performance index

X - state vector

M - final condition vector

v - Lagrange multiplier vector (final conditions)

V

- A unknown parameter vector
- a da/dt (a arbitrary)
- a^{T} a transposed (a arbitrary)
- μ geocentric gravitational parameter

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Abstract

The Inertial Upper Stage (IUS) being developed for use aboard the Space Shuttle is composed of three solid fuel stages plus a satellite payload. One mission of the IUS system is to launch from a shuttle parking orbit and place the satellite in geosynchronous orbit in minimum time. Actual Space Shuttle parking orbit data and IUS characteristics were used in this study to examine the sequential timing and orientation in inertial space of each stage as it is fired while the spacecraft moves along a transfer orbit to geosynchronous orbit. In addition, the sensitivity of the total transfer time and the final orbital state was found as a result of not meeting one or all of the time and orientation parameters.

This problem is unique in that it considers an optimal orbit transfer problem involving solid fuel stages of fixed thrust and burn time. Previous work with liquid fuel engines examined orbital transfers with the intent of minimizing the amount of propellant or required velocity change needed to accomplish the transfer.

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THREE BURN INERTIAL UPPER STAGE OPTIMAL ORBIT TRANSFER

I. Introduction

Previous work involving optimal orbit transfers was centered around liquid fuel engines. The driving concern with a liquid system is to minimize the amount of propellant to be carried on the spacecraft to keep the overall weight as low as possible. This is especially important for spacecraft of one stage, where the engine is required to be restarted for additional burns. There are a number of studies, of which Reference 9 is an example, that explore the transfer from one orbit to another using two impulsive burns. One study was found which considered three burn optimal transfers (Ref 10), but minimized the total velocity change needed.

The advent of the Space Shuttle made it possible to carry a small spacecraft into near earth orbit inside the cargo bay and then launch it from that parking orbit into another orbit. The storage problems associated with liquid fuel systems made solid fuel engines preferable for this type of spacecraft. The initial configuration for the Inertial Upper Stage consisted of two solid fuel engines. This allowed for the use of previous orbital transfer studies to define the exact capabilities of the configuration. The desire to expand this capability led to the design of the

three stage IU5 which necessitated the need to examine a three burn optimal transfer to geosynchronous orbit in minimum time.

As in past studies, the firing of each stage of the spacecraft is treated as an impulsive burn. This assumption is justified because the burn time of a rocket engine is very small when compared to the total time spent in an orbital transfer. In addition, transfers in the near vicinity of the earth are normally analyzed in three phases which yield increasing degrees of accuracy.

The first phase is to consider the earth and the spacecraft as being alone in inertial space. The earth is considered perfectly spherical and homogeneous, and the effects of the other major bodies such as the sun and moon are ignored. These assumptions allow for a two-body analysis of an object under the influence of a single gravitational field. The equations of motion used to describe the behavior of an object in this environment are accurate enough to allow for initial mission planning. The second phase is to take the two-body equations of motion and add the oblateness and nonhomogeneity effects on the earth. The results are more accurate, but they differ only slightly from those obtained from the two-body equations alone. The last step is to again modify the equations of motion to account for the effects of the sun, moon, and other major bodies. The end product is, indeed, accurate but extremely expensive and tedious to extract.

This study follows the tradition of previous work by

assuming impulsive burns and two-body orbital behavior for this initial mission analysis.

II. Problem Statement

The starting place for this problem is the near earth parking orbit of the Space Shuttle. The IUS System Program office at Patrick AFB (Ref 1) was kind enough to provide the following data describing the shuttle parking orbit at the moment of orbit attainment:

TABLE I

Initial Parking Orbit Data

_		<i> •</i>
	latitude	21.3624°
	longitude	59.2825°
	altitude	7.11693 x 10^5 ft
	inclination	28.78871°
	eccentricity	0.01497
	velocity	2.561851 x 10 ⁴ ft/sec
	$x = 1.029312 \times 10^7$ ft	$xd = -2.248185 \times 10^4 \text{ ft/sec}$
	$y = 1.732354 \times 10^7 \text{ ft}$	yd = 9.356206×10^3 ft/sec
	$z = 7.881747 \times 10^6 \text{ ft}$	$zd = 7.958385 \times 10^3 \text{ ft/sec}$

where x, y, z, xd, yd, and zd are the position and velocity components in the earth-centered inertial (ECI) coordinate frame. In order to describe subsequent positions and times, this initial fix (i.e. initial conditions) was chosen as a starting place for the orbital problem at time equal to zero. For the IUS, the engine burn times are of fixed duration as shown in Table II; therefore, the time of burn initiation and

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Stage	Vehicle Weight (lbs)	Specific Impulse (sec)	Burn Time (sec)	Delta-V (ft/sec)	Propellant Weight (1bs)
1,2, &3	56119	294.332	123.91	4242.175	20263
2&3	33813.6	294.189	140.35	9565.712	21505.6
3	8887.5	302.492	94.49	10997.98	6016.6

IUS Characteristics

the orientation of each hurn in inertial space are the variables of interest.

This problem is worked entirely in an ECI frame of reference. Figure 1 describes the orientation of each stage prior to engine initiation. Since the assumption has been made that each burn will be treated as impelsive, this corresponds to orienting the Delta-V vector in inertial space.

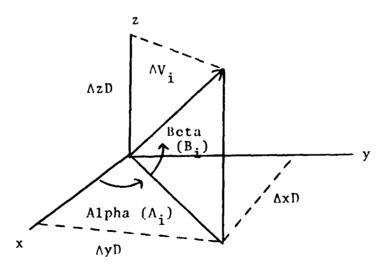


Fig 1. Spherical Coordinates Definition

The angles alpha and beta shown in Fig 1 completely describe the orientation of the spacecraft and its Delta-V vectors in inertial space. A third parameter, time, can then be used to pinpoint the position of the spacecraft along each segment of the transfer orbit. A pictorial representation of this is as follows:

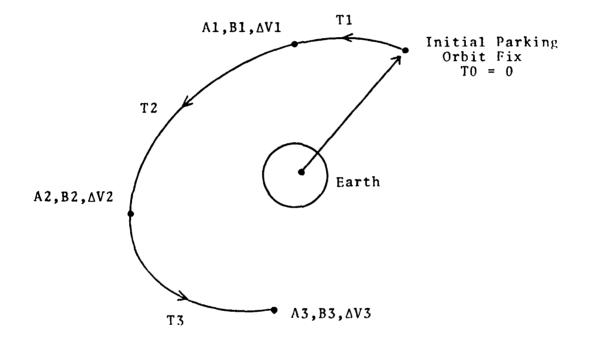


Fig 2. Transfer Orbit

where

T1 = coast time to firing of first stage
A1 & B1 = orientation of first stage at firing
T2 = coast time to firing of second stage
A2 & B2 = orientation of second stage at firing
T3 = coast time to firing of third stage

A3 & B3 = orientation of third stage at firing AV1, Δ V2, Λ V3 = impulsive velocity changes per burn This allows the entire orbit from initial parking orbit to final position to be described by the nine parameters T1, A1, B1, T2, A2, B2, T3, A3, and B3.

As stated before, the calculation of the satellite's position and velocity as it moves through each segment of the transfer orbit will be done using a two-body formulation. Several prediction algorithms are in existance today; all of them solve variations of the differential equation

$$\overline{R} = -\frac{\mu}{R^3} \overline{R}$$
(1)

where μ is the geocentric gravitational parameter, \overline{R} is the vector position of the satellite, \overline{R} is the vector acceleraof the satellite, and R is the scalar distance from the center of the attracting body to the satellite. Two such algorithms were used in this study. The first is an actual series solution of the two-body differential equations known as the f and g series solution to the Kepler problem (Ref 3). The f and g series solution affords a rapid means of determining position and velocity as a function of time and just as importantly, avoids the penalties associated with the numerical integration of differential equations. A second prediction algorithm, known as Hamilton's two body equations (Ref 2), is composed of six first-order differential equations and was used only to check the accuracy of the subroutine written to do the prediction problem.

By using the f and g series in a subroutine, an algorithm defining satellite position and velocity as a function of all nine parameters can be formed.

<u>Algorithm</u>

- 1. Given: initial position, velocity, and time (X
 and T0 = 0)
- 2. Coast in parking orbit for time T1 and calculate final position and velocity.
- 3. Align the first burn in inertial space using Al and B1 and calculate its component velocity contribution. Add this to the velocity from (2) to obtain a new state vector.
- 4. Coast in new orbit for time T2 and calculate a final state vector. Repeat the procedures used in (3) for the second stage and find the new orbit parameters.
- 5. Coast in final orbit for time T3, add the vectoral velocity contribution of the third stage, and calculate the final position and velocity.

The requirement that the satellite's final orbit be geosynchronous yields five constraints on the final state vector as defined in the ECI frame.

- 1. The z component of position must equal zero.
- 2. The zd component of velocity must equal zero.
- The satellite's speed must equal geosynchronous speed.
- 4. The satellite's distance must equal geosynchronous

distance.

5. The final orbit must be circular.

The problem then is to find values for the parameters T1, Λ 1, B1, T2, Λ 2, B2, T3, Λ 3, and B3 that allow the satellite to transfer from its initial position in the parking orbit to a final orbit which satisfies the specified end conditions in minimum time. It is readily apparent that this problem is not amenable to closed form analytic treatment, which implies the need to solve it iteratively.

III. The Optimization Problem

A good definition of a parameter optimization problem is how to change parameters in order to satisfy end conditions with the least effort. Optimization theory provides the direction needed to change each parameter to reduce end constraint errors and also offers figures of merit that can be used to tell how well the procedure is progressing. This problem, as stated, involves finding the values of nine parameters that force the final orbit to satisfy five equality constraints and allow the satellite to complete the orbital transfer in minimum time. This defines the performance index as

$$G = T1 + T2 + T3$$
 (2)

which is the total time of flight. The initial conditions for the problem are simply the initial position and velocity components supplied for the parking orbit for reference time equal to zero. The final conditions or end constraints are labeled by the vector M, where

$$M(1) = zd$$

$$M(2) = z$$

$$M(3) = (xd^{2} + yd^{2} + zd^{2})^{1/2} - 1.0096 \times 10^{4}$$

$$M(4) = (x^{2} + y^{2} + z^{2})^{1/2} - 1.3811 \times 10^{8}$$

$$M(5) = x(xd) + y(yd) + z(zd)$$
(3)

where 1.3811×10^8 feet is geosynchronous orbital distance,

and 1.0096 x 10⁴ feet per second is synchronous speed. Each constraint is written so that when its numerical value equals zero, the desired condition has been obtained. By calculating the norm of the M vector, a figure of merit is produced which indicates the degree to which the end constraints are satisfied. The optimization problem is then to drive the norm of the M vector to zero and the performance index to its minimum value.

One procedure for accomplishing this is referred to as suboptimal control. In an optimal control problem, the desired controls, which are functions of time, are directly calculated. For this problem that would mean the alpha and beta parameters are functions of time. The assumption that all three burns are impulsive converts the alpha and beta controls to scalar parameters. The suboptimal control technique approximates these controls using polynomials.

In this problem the controls are the three pairs of angles needed to align the stages for firing and the times of firing. If all the parameters are expressed in one vector A, then

$$\Lambda = [T1 \ \Lambda 1 \ B1 \ T2 \ \Lambda 2 \ B2 \ T3 \ \Lambda 3 \ B3]^{1}$$
(4)

This vector A then contains all the information needed to find the final position and velocity of the satellite.

$$X_{f} = X_{f}(\Lambda)$$
(5)

In addition, the vector Λ also supplies all the information needed to evaluate the performance index and the end condition constraints.

$$G = G(\Lambda)$$

M = M(\Lambda) (6)

If an augmented performance index, F, is defined functionally as

$$F(\Lambda, v) = G(\Lambda) + v^{T}M(\Lambda)$$
(7)

where v represents a vector of Lagrange multipliers, it must satisfy the first variational requirements

$$F_{A}^{T}(A,v) = 0$$

$$F_{A} = \frac{\partial F}{\partial A}$$

$$F_{V}^{T}(A,v) = M(A) = 0$$
(8)

to be a minimum solution. F_A is the partial of the augmented performance index with respect to the parameter vector A. The F_A matrix represents the change in performance resulting from a change in each parameter in A(gradient). Since there are nine parameters in A, F_A is a row vector of nine elements. Hull and Edgeman (Ref 6) describe a second-order parameter optimization technique and algorithm specifically for application to suboptimal control problems. Johnson (Ref 7) and Peterson (Ref 8) used this algorithm to develop a computer program to analyze aircraft time to turn problems. Their program uses three controls, differential equations of motion, and two end condition constraints. The program was modified to accommodate nine parameter controls, a series solution to differential equations of motion, and five end condition constraints. In short, the program which adopts Hull and Edgeman's algorithm uses second-order information to determine how to change the parameters in the vector Λ and the Lagrange multiplier vector, ν , such that the vectors F_{Λ} and M are driven to zero. The vector relationships used to change A and ν are derived from Eq (8). For any guessed values of Λ and ν that are not a solution, the equality will not hold.

$$F_{\Lambda}^{T} \neq 0 \qquad M(A) \neq 0 \qquad (9)$$

If Eq (9) is then linearized about A and v, then

$$\delta F_{A}^{T} = F_{AA} \delta A + M_{A}^{T} \delta v \qquad (10)$$

$$\delta M = M_{\Lambda} \delta A \tag{11}$$

and define

$$\delta F_{A}^{T} = -PF_{A}^{T}$$
(12)

$$\delta M = -QM \tag{13}$$

where P and Q are scaling factors, yields

$$F_{\Lambda\Lambda}\delta\Lambda + M_{\Lambda}^{T}\delta\nu = -PF_{\Lambda}^{T}$$
 (14)

$$M_{A}\delta\Lambda = -QM \tag{15}$$

which can be solved for $\delta \nu$ and $\delta \Lambda.$

$$\delta v = (M_{\Lambda} F_{\Lambda \Lambda}^{-1} M_{\Lambda}^{T})^{-1} (-P M_{\Lambda} F_{\Lambda \Lambda}^{-1} F_{\Lambda}^{T} + QM)$$
(16)

$$\delta \Lambda = -F_{\Lambda\Lambda}^{-1} (PF_{\Lambda}^{T} + M_{\Lambda}^{T} \delta v)$$
(17)

where

$$M_{\Lambda} = \frac{\partial M}{\partial \Lambda}$$
 and $F_{\Lambda\Lambda} = \frac{\partial^2 F}{\partial^2 \Lambda}$

and P and Q are scaling factors which control optimization and end condition satisfaction. $F_{\Lambda\Lambda}$ represents a change in slope and the direction in which it is increasing or decreasing. The algorithm developed to iteratively change Λ and ν follows:

- 1. Guess A and ν
- 2. Use the f and g series solution to find X_{f}
- 3. Compute M, M_A , M_{AA} , F_A , and F_{AA}
- 4. Pick values for P and Q and compute $\delta\nu$ and δA
- 5. Set $\Lambda = \Lambda + \delta \Lambda$ and $\nu = \nu + \delta \nu$
- Check convergence and if unsatisfied, go to step (2)

The computer program uses this algorithm plus a gradient approach that allows the direct calculation of an initial vvector which eliminates guessing five parameters. The procedure for doing this was also developed by Hull and Edgeman and uses first-order information as follows:

$$v = (M_{\Lambda}M_{\Lambda}^{T})^{-1}[(Q/P)M - M_{\Lambda}G_{\Lambda}^{T}]$$
(18)

 $\delta_{\Lambda} = -PF_{\Lambda}^{T}$ (19)

where

 $G_{\Lambda} = \frac{\partial G}{\partial \Lambda}$

For a typical problem, the gradient portion of the program would be used to locate the vicinity of the functional minimum. At this point where gradient methods lose their effectiveness, the second-order algorithm would be used to rapidly converge the problem. As this procedure is followed, the scaling factors P and Q are gradually raised from very small values to final values of one. In the final convergence cycles, P and Q must be equal to one to yield an optimal solution. Although this procedure would appear straightforward, its application to specific problems can, at times, require some finesse as will be shown later.

Numerical Methods (Ref 6:484)

The first- and second-order optimization routines require the matrices M, M_A , M_{AA} , and F_{AA} . These matrices were evaluated using numerical techniques. F_A and F_{AA} are defined as

$$F_{A} = G_{A} + v^{T} M_{A}$$
 (20)

$$F_{\Lambda\Lambda} = G_{\Lambda\Lambda} + \sum_{i=1}^{5} v_i M_{\Lambda\Lambda}$$
(21)

since

$$G = T1 + T2 + T3$$
 (22)

$$G_{\Lambda} = [1 \ 0 \ 0 \ 1 \ 0 \ 0 \ 1 \ 0 \ 0]$$
(23)

and

$$G_{\Lambda\Lambda} = [0]_{9\times9} \tag{24}$$

The only quantities that need to be calculated are M, $M_{\mbox{\scriptsize A}},$ and $M_{\mbox{\scriptsize AA}}.$

The M matrix, or error matrix, is easily evaluated after determining the final position and velocity vectors using the f and g series solution to the Kepler problem. The M_{Λ} and $M_{\Lambda\Lambda}$ matrices are determined using a central differences numerical derivative technique. The technique uses the initial values of the A vector parameters (Λ_n) to calculate an initial M. Each Λ_n is then positively perturbed by

$$A_{n+} = A_n + \delta_n \tag{25}$$

and then negatively by

$$A_{n-} = A_n - \delta_n$$
 (26)

Using Λ_{n+} and Λ_{n-} , M_{+} and M_{-} are calculated. The central differences representation for $M_{\Lambda_{-}}$ is then given by

$$M_{\Lambda_{n}} = \frac{M_{+} - M_{-}}{2\delta_{n}} + \sigma(\delta_{n}^{2})$$
(27)

where $\sigma(\delta_n^{2})$ represents an error term of order δ_n^{2} where δ_n is a very small positive number. The M_A matrix contains five rows. The first row is determined using M₁ in Eq (27), the second row using M₂, and so on. The M_{AA} matrices are determined in a similar manner; however, two elements A_n and A_m, must be perturbed both positively and negatively to obtain M₊₊, M₊₋, M₋₊, and M₋₋. The central differences representation for M_{A_nA_m is}

$$M_{\Lambda_{n}\Lambda_{m}} = \frac{M_{+} - 2M + M_{-}}{\delta_{n}^{2}} + \sigma(\delta_{n}^{2})$$
(28)

for n = m and

$$M_{\Lambda_{n}\Lambda_{m}} = \frac{M_{++} - M_{+-} - M_{-+} + M_{--}}{4\delta_{n}\delta_{m}} + \sigma(\delta_{n}\delta_{m})$$

for $n \neq m$. The M_{AA} matrix is really five matrices. One is determined using M₁ values in the above equations, and then

the others by M_2 through M_5 in turn. The error terms in these equations can be ignored if δ_n and δ_m are quite small. The δ used for the central differences technique is

$$\delta_{n} = DELTA \cdot A_{n}$$
 (29)

and if the absolute value of $\boldsymbol{\delta}_n$ is larger than DELTA, then

$$\delta_n = DELTA$$
 (30)

where DELTA is another small positive number. In this problem, DELTA was initially chosen to equal 1.0 x 10^{-6} .

Selecting Initial Parameter Values

Picking nine parameters to serve as an initial guess for the orbital transfer problem is far from a simple matter. Not only must the parameters be compatible, but they should also provide an initial "miss" relative to the end condition constraints small enough to allow for easy convergence. At this point, some engineering insight was applied with only marginal success. Since the actual parking orbit has an eccentricity of 0.01497, one would think parameters good for a perfectly circular orbit would be good initial guesses for the actual. With this in mind, a test orbit was constructed in an ECI frame with ascending node on the X-axis at 2.16737 x 10⁷ feet, inclined 22.8 degrees at a velocity of 25,618 feet per second. The first parameter, T1, was determined by calculating the time of flight to the descending node. Since the final orbit would lie in the X-Y plane, the first burn

was aligned to eliminate as much inclination as possible using

$$\Lambda V_i = 2V \sin \theta/2 \tag{31}$$

where ΔV_i is the change in velocity due to stage burning, V is the original velocity of the spacecraft at nodal crossing, and θ represents the angular change in the inclination. This allowed Al and Bl to be determined. T2 was chosen to be the time of flight to the next nodal crossing. The second burn was aligned in an attempt to eliminate the remaining inclination yielding A2 and B2. T3 was again chosen to be the time to the next nodal crossing, and A3 and B3 were varied to find the smallest error in the end conditions. This orbit and the actual orbit differ only by a rotation about the Z-axis to align the nodes--ignoring eccentricity. This angle was then added to the three alpha parameters, and T1 was adjusted to reflect the coast time to first nodal crossing from the initial fix. The initial miss vector that this procedure yielded was of magnitude 10¹², largely due to the circularity constraint. Random adjustment of individual parameters proved to be a hopeless means of reducing the error. The circularity constraint proved to be far too sensitive to allow for rapid convergence. To decrease the sensitivity of this constraint, the circularity was multiplied by a scaling factor of 10^{-5} . This allowed the program to drive down the errors in the other four constraints. Each time this occurred, the scaling factor was increased by a factor of ten and the process repeated. This allowed the scaling factor to eventually

be removed once the vicinity of the functional minimum was reached.

IV. Results

Choosing P, Q, and DELTA

The selection of the P and Q scaling factors is unique to every optimization problem. By their selection, one can select several options. The first and most commonly used is to choose P very small and Q large as compared with P. By doing this, the program will attempt to satisfy the end constraints before doing any optimization. A second method is to choose P larger than Q in an attempt to satisfy the end conditions after finding an optimal region. The third common method is to choose P and Q equal and, thereby, optimize while trying to satisfy the end constraints. All three methods were tried on this problem, with only the first being successful. It was found necessary to first satisfy the end conditions and then search for an optimal solution.

Second-order optimization methods are very unstable far from the optimal solution. This requires very small changes in the control parameters through each iteration to prevent divergence. The magnitudes of these "delta" quantities can be controlled using P and Q. Trial and error showed that for P equal to 1.0×10^{-15} and Q equal to 1.0×10^{-6} , sufficiently small parameter changes were generated to allow the program to run for an average of 25 iterations before divergence occurred. This was caused by the numerical

routine generating an excessively large change in one of the parameters. The program was modified to allow continuous operation by examining the norm of these parameter changes. Initially, whenever the norm exceeded 0.38, the divergence was seen to occur. The program was then modified to reinitialize all the Lagrange multipliers and restart itself whenever this was observed. As the program began to approach the vicinity of the functional minimum, the numerical routine became more stable. The limit on the norm of the parameter changes was relaxed to 0.5 and finally to 1.0. This simple procedure proved extremely valuable after unsuccessfully attempting to use a straight gradient method.

A third parameter, DELTA, also showed a large impact on the success of the routine. For a similar problem using differential equations of motion, DELTA would be chosen in the 10^{-4} to 10^{-6} range. This insures the accurate calculation of the M_A and M_{AA} matrices. In addition, the size of the parameter changes (da's) is directly proportional to DELTA, since the "delta" perturbing values equal DELTA times each element of A. In this problem, DELTA equal to 10^{-6} was used based on this previous experience. It was found that the numerical routine stagnated when the solution was being approached. It appears the function contours were steep to the point where the parameter changes equal to 10^{-6} times their present value would not allow the program to move deeper into the contour valleys. At this point, several iterations were run with the same initial parameter values, but with

different DELTAs. By comparing the norms of the constraint errors, it was found that performance improved as DELTA was progressively decreased. Best performance resulted from a DELTA of 10^{-10} , which was then used to find the solution presented here.

Gradient Method

The gradient method, which uses first-order information in an attempt to move toward an optimal solution, proved to be a dismal failure. Because the performance index was defined as the sum of three times, the gradient method tried to drive all three to zero. Gradient information was used, however, to calculate initial values of Lagrange multipliers by Eq (18). These were then used in the second-order method.

Numerical Results

The second-order method used in the computer program, when modified to run continuously, steadily drove toward a solution that satisfied the end constraints and then toward an optimal solution. The parameters that satisfied the end constraints changed very little during the optimization portion of the program, which suggests that this may be a locally unique solution. The optimization iteration process was stopped with P and Q equal to one when the norm of the parameter changes was less than 10^{-9} and the norm of the Lagrange multiplier changes was also very small. The A matrix solution and the sensitivity matrix, M_A, are shown in the following tables.

ТΛ	BL	E	I	I	I

Parameter Solution Set

TABLE IV

End Constraint Errors

٠.

$M(1) = zd = -2.5465 \times 10^{-10} \text{ ft/sec}$
$M(2) = z = 9.5367 \times 10^{-7} ft$
M(3) = synchronous speed = -1.0477 x 10 ⁻⁹ ft/sec
M(4) = synchronous distance = 8.583 x 10 ⁻⁶ ft
$M(5) = circularity (R V) = 1.318 \times 10^{-2}$
Total Transfer Time = 13348.54608 sec = 3.70792 hrs

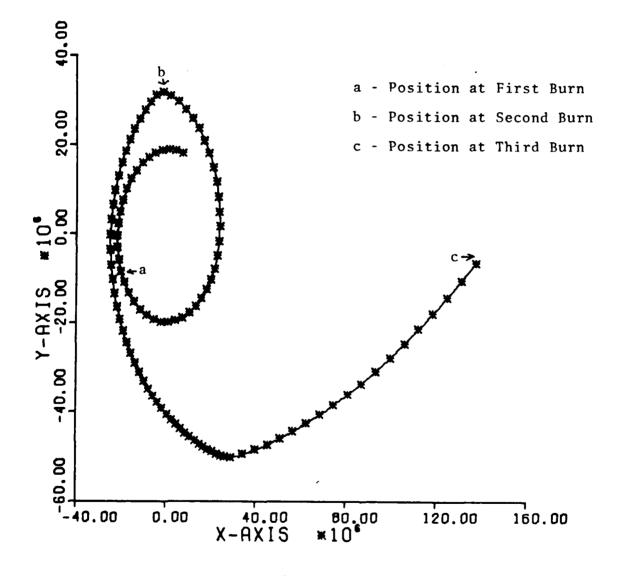
Figures 3 through 8 depict the transfer orbit from several points of view. Figure 5 shows a result that is somewhat surprising. The transfer orbit remains in the same plane as the parking orbit. In other words, the transfer opts to increase the altitude and velocity of the spacecraft early in the transfer and complete the inclination change in the final burn. As Eq (31) points out, the cost to change

TABL	E	V
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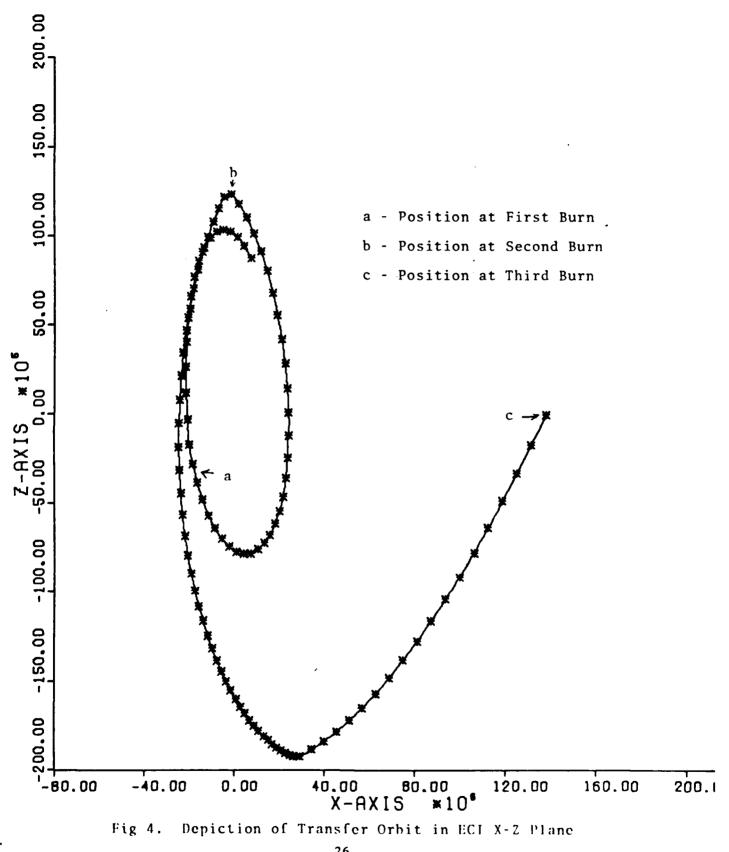
Position and Velocity After Firing of Each Stage

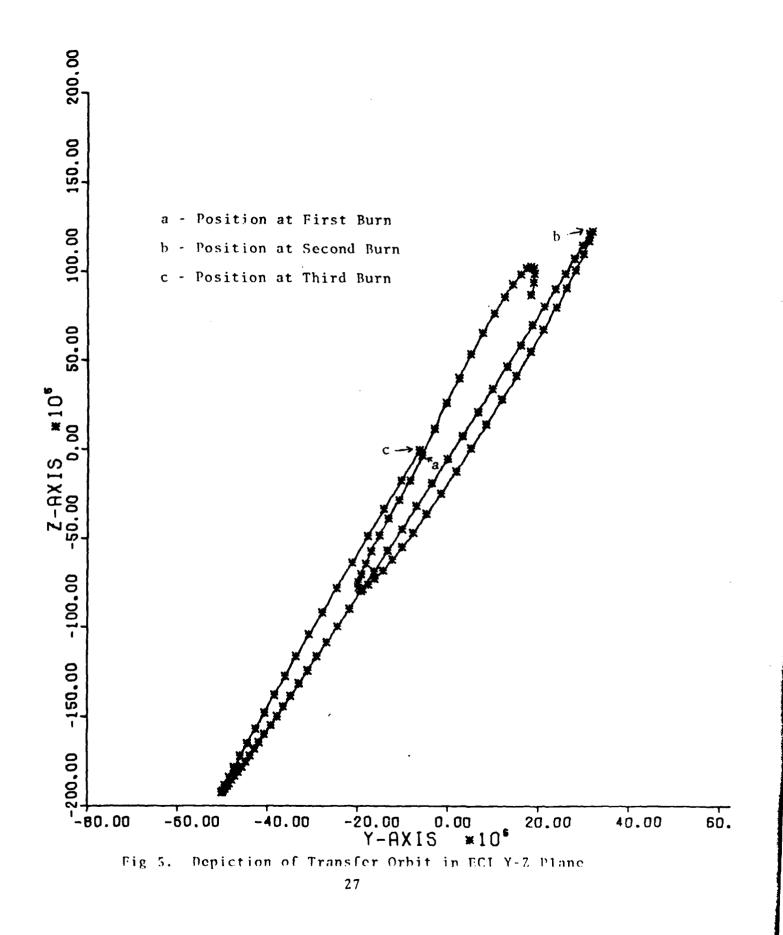
Position and Velocity After	Firing First Stage:
$x = -2.0373124 \times 10^7 \text{ ft}$	xd = 1.11011967 x 10 ⁴ ft/sec
$y = -7.6147903 \times 10^6 \text{ ft}$	$yd = -2.21389677 \times 10^4 \text{ ft/sec}$
$z = -1.3724214 \times 10^6 \text{ ft}$	$zd = -9.3978524 \times 10^3 \text{ ft/sec}$
Position and Velocity After	Firing Second Stage:
$x = -2.885451299 \times 10^6 \text{ ft}$	$xd = -2.05530541 \times 10^4 \text{ ft/sec}$
$y = 3.20133328 \times 10^7 \text{ ft}$	$yd = -1.176502728 \times 10^4 ft/sec$
$z = 1.257707996 \times 10^7 \text{ ft}$	$zd = -5.014671267 \times 10^3 ft/sec$
Final Position and Velocity	after Third Burn:
$x = 1.379638 \times 10^8 \text{ ft}$	xd = 464.30038 ft/sec
$y = -6.351478 \times 10^6 \text{ ft}$	yd = 10085.31809 ft/sec
$z = -8.118152 \times 10^{-5} \text{ ft}$	$zd = 2.8067 \times 10^{-9} \text{ ft/sec}$

inclination is less as the spacecraft velocity decreases. The point at which geosynchronous orbit is obtained is where the spacecraft velocity is lowest. This points out that the assumption used to obtain initial values for the parameters by eliminating the inclination of the transfer orbit was wrong. Figure 6 depicts the entire transfer orbit in three dimensions. Figure 7 is the same as Fig 6, except that the scaling on the Z-axis has been changed to show separation of the near earth portions. Figure 8 is the confirmation that the final orbit is, indeed, circular and geosynchronous.









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3-D VIEW OF TRANSFER ORBIT

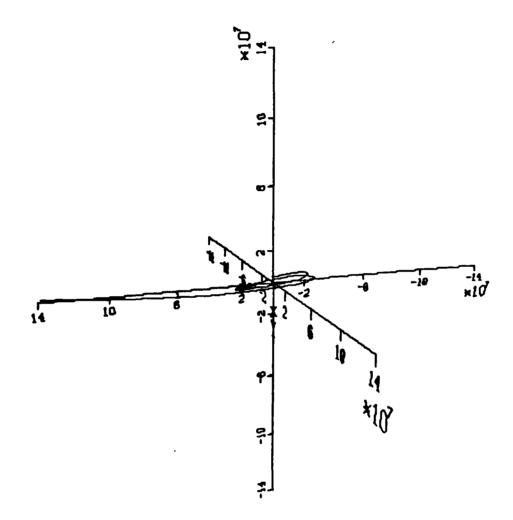


Fig 6. Depiction of Transfer Orbit in 3-D

3-D VIEW OF TRANSFER ORBIT

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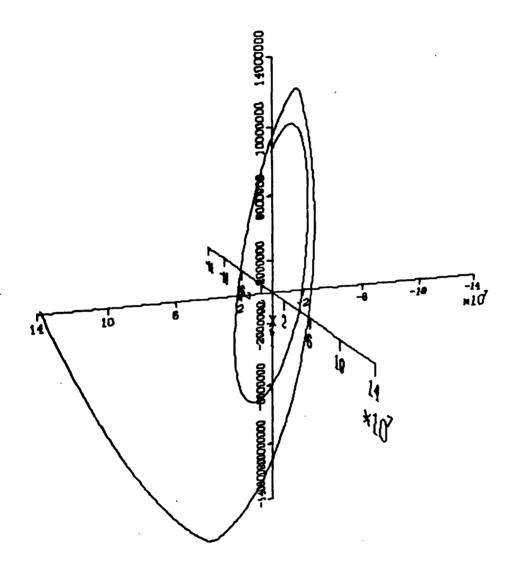
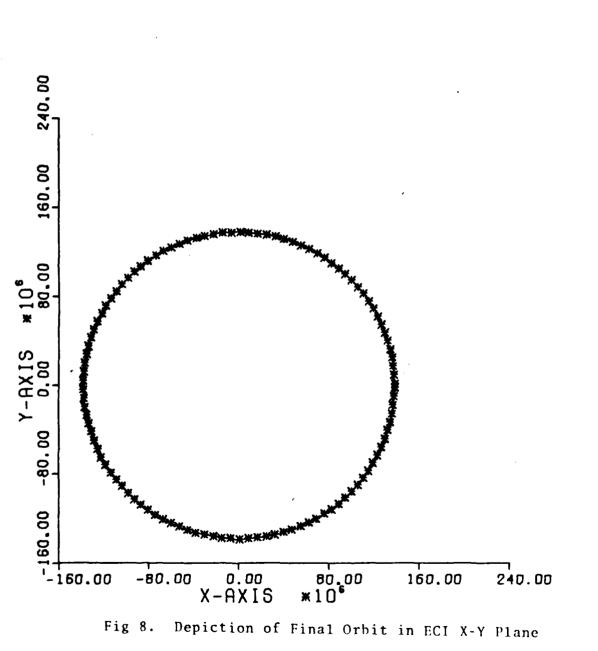


Fig 7. Depiction of Transfer Orbit in 3-D

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The M_{Λ} matrix which is the partial of M with respect to the A matrix indicates the sensitivity of each end condition to a change in any one of the A matrix parameters.

TABLE V	I
---------	---

М	Matrix
---	--------

	T1	Λ1	B1
M1	9.9414	-11045.0	-17707.0
M2	-78377.0	83614230.0	147191047.0
M3	26.306	-28542.0	-45384.0
M4	-497674.0	473346710.0	756855918.0
M5	9.9097	-7.73×10^{12}	-1.215×10^{13}
	Τ2	Α2	B 2
M1	-10.353	3133.023	1377.528
M2	-248.21	4481256.0	60289542.0
M3	-26.293	7204.65	6391.623
M4	304516.131	-103134155.0	-82096056.0
M 5	1.596×10^9	1.3178×10^{12}	3.077×10^{11}
	Т3	A3	ВЗ
М1	-0.802865	0	10995.048
М2	14068.347	0	0
M3	-2.0329	2449.40	-239.582
М4	50201.67	0	0
М5	-646330649.0	-1.48×10^{12}	-7.567×10^{9}

The M_A matrix is extremely useful when one is interested in how changes in one of the parameters in the A matrix will effect the end conditions. As an example, consider the last column, which represents the change in end conditions due to

TABLE VII

 M_{Λ} Values for B3

	B3
М1	10995.048
M2	0
M3	-239.582
м4	0
M5	-7.567 x 10 ⁹

a change in B3. These numbers indicate that if the third stage was misaligned by plus one radian, it would increase the zd component of velocity by 10995.048 feet per second. Since B3 has no effect on position, the numbers that correspond to z and geosynchronous distance are both zero. The satellite's speed would decrease by 240 feet per second, and the circularity of the orbit by 7 x 10^9 . This clearly illustrates the use of the M_A matrix as well as the wide differences in magnitudes of the individual errors.

V. Application

The application of the information found in this study is both simple and straightforward. The parking orbit and the transfer orbit are fixed in inertial space due to the choice of an ECI reference frame. The earth, on the other hand, spins relative to this frame. To place a satellite in geosynchronous orbit over a specific spot on the equator, one has basically two choices. The first is to time the launch of the satellite from the earth into its parking orbit and then along the transfer orbit such that the final position of the transfer orbit coincides with the desired equatorial position. A second and more practical approach would be to launch into the parking orbit whenever convenient. The period of the parking orbit is roughly 90 minutes, which allows for a transfer every hour and a half.

Because this problem was worked in an ECI frame of reference, the position of the satellite at the end of the transfer relative to its position at the beginning is always constant. Therefore, the angular difference in their positions, as seen in the X-Y plane and measured about the Zaxis, is constant. By allowing for the rotation of the earth, a transfer to a given position over the equator can be calculated in terms of longitude. Optimal launch conditions occur when the longitude of the parking orbit initial fix is

114.90 degrees greater than the longitude of the desired geosynchronous orbital position. For example, if the initial orbit was obtained at longitude 23.90°E and the transfer begun, the final position would be at longitude 91.0°W, which corresponds to the Galapagos Island chain, after a flight time of 3.70792 hours.

VI. Recommendations

One of the major assumptions made in this study was that the three engine burns were impulsive. As was shown, this reduced the alpha and beta controls to scalar parameters. In reality, the engines operate for specific lengths of time as shown in Table II. Since the computer program has the capability to determine the alpha and beta controls as polynomial functions of time, it would be interesting to remove the impulsive burn assumption and rework this problem using the solution found here as a good initial guess.

This study found only one solution to this transfer problem and suggests that it may be unique. If the original parking orbit were equatorial, there would exist two mirror image solutions. It would be interesting to see if the methods used in this study would yield both solutions.

The suboptimal control technique, as applied to this problem, did yield a valid optimal solution, but the method did prove expensive in terms of computer time. It would be worthwhile to apply one or two other second order or quasisecond order optimization methods to this problem to see if a more efficient method exists.

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APPENDIX A

Computer Program SubOpt

PROGRAM ORBIT(INPUT;OUTPUT) 100= 110= DIMENSION DA(9),MAA3(9,9),MAA4(9,9),MAA5(9,9),WZ(9) DIMENSION A(9), AA(9), GA(9) 120= DIMENSION M(5), MA(5,9), MAA1(9,9), MAA2(9,9) 130 DIMENSION B(6),DB(6),BB(6),BP(6),BP1(6),BP2(6),BP3(6),BP4(6) 140= DIMENSION DEL(9), MAMAT(5,5), MAMATI(5,5), MADAT(5) 150= 160≕ DIMENSION FAT(9), FAA(9,9), FAAI(9,9), MAFAAI(5,9) DIMENSION W(5,5),WI(5,5),C(5),P(9),PNU(5),DPNU(5),PNUMA(9) 170= DIMENSION E(9),FA(9),FFAA(9,9),FFAAI(9,9) 180= 190≔ DIMENSION GAA(9,9) DIMENSION MTXFXT(12,12),FFAAB(22,22) 200= BIMENSION MTX(12,12), MTXT(12,12), CONA(12), CONB(12), MTXF(12,12) 210= 220= COMMON/MISC/NPH,NPI,NA,TF,DT,N,S,LL,PP,TOT(8),FIXT 230= COMMON X,Y,Z,XD,YD,ZD 240= COMMON/QUES/ALFHA 250= REAL M+M1+M2+M3+M4+MM1+MM2+MM3+MM4+MM5+MM6+MM7+MM8+MA 260= REAL M5, M6, M7, M8, M9, M10 270= REAL MM9, MM10, MM11, MM12, MM13, MM14, MM15 280≈ REAL MM16, MM17, MM18, MM19, MM20, MAA1, MAA2, MAA3, MAA4, MAA5 290= REAL MAMAT, MAMATI, MAGAT, MAFAAI, NORMA, NORMA, NORMA 300≈C PROGRAM SUBOPT IS A SUBOPTIMAL CONTROL TECHNIQUE USED TO FIND AN 310≈C 320≈C APPROXIMATE SOLUTION TO AN OPTIMAL CONTROL PROBLEM. THE SOLUTION TO MOST OFTIMAL CONTROL PROBLEMS IF THE CONTROLS CANNOT BE SOLVED 330=C 340×C ANALYTICALLY IS TO GUESS THE CONTROLS AND THE LAGRANGE MULTIPLIERS AND SEE IF END CONDITIONS ARE MET AND THE PERFORMANCE INDEX MINIMIZED 350=C 340≈C THE SUBOPTIMAL CONTROL TECHNIQUE ASSUMES THE CONTROLS ARE A LINEAR 370=0 COMBINATION OF POLYNOMIALS WITH UNKNOWN COEFFICIENTS. IN SO DOING 380≈C THE PROBLEM IS CHANGED FROM A FUNCTIONAL MINIMIZATION PROBLEM TO A 390=C PARAMETER OPTIMIZATION PROBLEM. 400=C IN THIS EXAMPLE THE CONTROLS ARE NOT FUNCTIONS OF TIME. THIS SIMPLIFIES THE PROGRAM TO ONE OF PARAMETER OPTIMIZATION. 410=C 420=C THIS VERSION ADOPTS THE ORBITAL TRANSFER PROBLEM 430=C 440=C 450=C A ==MATRIX OF PARAMETERS(T1,A1,B1,T2,A2,B2,T3,A3,B3) 460=C DA =MATRIX OF PARAMETER CHANGES (NP X 1) 470=C NE =NUMBER OF STATE VARIABLES 480=C NP =TOTAL NUMBER OF PARAMETERS NC =NUMBER OF END CONDITIONS TO BE SATISFIED 490=C 500=C P =SCALING FACTOR PERFORMANCE INDEX 510=C Q =SCALING FACTOR FOR END CONDITION CONSTRAINTS (M) = SCALAR PERFORMANCE INDEX 520=C G 530=C GA = PARTIAL DERIVATIVES OF G WRT A'S (1 X NP) 540=C GAA = SECOND PARTIAL DERIVATIVES OF G WRT A'S (2 - NP X NP) 550=C Μ = MATRIX OF PRESCRIBED FINAL CONDITIONS (NC X 1) 560≈C = PARTIAL DERIVATIVES OF M WRT A'S (NC X NP) MA MAA = SECOND PARTIAL DERIVATIVES OF M WRT A'S (2 - NP X NP) 570=C 580=C В =STATE VARIABLES (X,Y,Z,XD,YD,ZD) 590=C F = AUGMENTED PERFORMANCE INDEX (G + PNUT*M) 3=005 FA = PARTIAL DERIVATIVE OF F WRT A'S (1 X NP)

610≈C FAT = PARTIAL DERIVATIVES OF F WRT A'S TRANSPOSED (NP X 1) FAA = SECOND PARTIAL DERIVATIVE OF F WRT A'S (NP X NP) 620≈C PNU = LAGRANGE MULTIPLIERS (NC X 1) 630=C DPNU= CHANGE IN LAGRANGE MULTIPLIERS (NC X 1) 640=C 650=C THE CIRCULAR PARKING ORBIT IS INCLINED 22.8 DEGREES 660≈C 670=C AT AN ALTITUDE DF 2,1637E07 FT 680=C 690= T=0.0 700≈ X=1.029312E07 710= Y=1.732354E07 720= Z=7.881747E06 `0≔ XD=-2.248185E04 , 40= YD=9.356206E03 750= ZD=7.958385E03 760=C 770=C THE THREE BURNS ARE ORIENTED IN TIME AND SPACE 780=C BY NINE PARAMETERS 790=C 800≔ T1=2030.244999504 810= A1=-.4353610070055 820= B1=,7111815869911 T2=3847.461750268 830= 840= A2=.7323419833131 850= B2=-2,731217478098 =068 T3=7470.839340005 870= A3=1,300130746286 880= B3=.0223640901197 890=C 900=C THE FINAL ORBIT IS CALCULATED USING THE F & G SERIES 910=C SOLUTION TO THE TWO BODY PROBLEM. 920=C 930= R = (X * X + Y * Y + Z * Z) * * 0.5940= V=(XD*XD+YD*YD+ZD*ZD)**0.5 950= * • X PRINT*, X(0) =• , Y 960= FRINT*, Y(0)= 970= FRINT*, "Z(0)= • , Z 980= PRINT*, *XD(0)= • , XD 990= PRINT*, *YD(0) =",YD *,ZD 1000 =PRINT**TI(0) =1010= PRINT*, *RANGE= * , R 1020 =PRINT*, VELOCITY= .V PRINT*,* 1030= 1040 =NF=9 1050 =NC=51060= NE=61070= ITER=0 90= MAX=100 ≕090 ⊾ ZED=1. 1100= DELTA=1.E-08 1110= IMET=2 1120 =Q=1. 1130= P=1. FNORMA=10.0 1140= 1150 =FGNORM=10.0 CC1=1.0E-04 1160 = CC2=1.0E-02 1170= 1180= CC3=1.0E-08 1190= CC4=1.0E-04 1200= CC5=1.0E-04

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. .

1210=	B(1)=X
1220=	B(2)=Y
1230=	B(3)=Z
1240=	B(4)=XD
1250≈	B(5)=YD
1260=	B(6)=ZD
1270≔	A(1)=T1
1280=	A(2)=A1
1290=	A(3)=B1
1300=	A(4)=T2
1310=	A(5)=A2 ·
20=	A(6)=B2
≖30د ـ	A(7)=T3
1340≈	A(8)=A3
1350=	A(9)=B3
1360=555	ITER=0
1370≔	PNU(1)=00001
1380=	PNU(2)=8.0E-08
1390≈	PNU(3)=0000537
1400=	PNU(4)=-1.757E-07
1410=	PNU(5)=-1.7E-11
1420=	DPNU(1)=DFNU(2)=0.0
1430=	DPNU(3)=DPNU(4)=DPNU(5)=0.0
1440≈52	FORMAT(1X,5E15.7)
1 450 ≈50	FORMAT(1X,8E15.7)
1460=C	
	TERMINING M MATRIX BY INTEGRATING DIFFERENTIAL CONSTRAINTS
1480≔C	
1490=	DO 1 I=1, NF
1500=1	DA(I)=0.0
1510 =1000	DO 2 I=1,NP
1520=	A(I)=A(I)+DA(I)
1530=2	_ AA(I)=A(I)
1540=	PRINT*, * *
1550=	PRINT*,"ITERATION NUMBER ",ITER," P= ",P," Q= ",Q
1560=	PRINT*," "
1570=	DO 3 I=1,NE
1580≈3	BB(I)=B(I)
1590=	S=1.0
1600=	CALL TRANS(AA, BB)
1610=	M(1)=ZD
1620=	M(2)=Z
1630=	M(3)=((XD*XD+YD*YD+ZD*ZD)**0.5)~1.0096E04
1640=	M(4)=((X*X+Y*Y+Z*Z)**0.5)-1.3811E08
1650≖	M(5)=X*XD+Y*YD+Z*ZD
1660=	M(5)=M(5)/ZED
<70≔	G=AA(1)+AA(4)+AA(7)
- 780=	ITER=ITER+1
1690=	PRINT*, *
1700=	PRINT*, "A MATRIX"
1710=	DO 666 I=1,NP P
1720=666	$PRINT*_*A(*_*I_*) = *_*A(I)$
1730=	
1740=	DO 4 I=1,NC
1750	FNU(I)=FNU(I)+DFNU(I)
1780-1	PRINT*, "M(",I,")= ", M(I)
1770	
1780=	DO 5 I=1,NC
1790	
1800=	NORMM=SQRT(NORMM) 40
•	

1810=	PRINT*, "
1820=	$PRINT*_{,}"NORM OF M = ",NORMM," P.I. = ",G$
	PRINTAG NURPEUP PERSONNAL PAIN - 90
1830=C	
1840=C	DETERMINING MA AND MAA BY CENTRAL DIFFERENCES
1850=C	
18 60≔	S=0.0
1870=	DO 100 I=1,NF
1880=	DO 6 J=1,NP
1890=6	AA(J)=A(J)
1900=	DEL(I)=DELTA*A(I)
1910=	IF(ABS(DEL(I)).LE.DELTA) DEL(I)=DELTA
20=	AA(I)=A(I)+DEL(I)
1930=	DO 7 J=1+NE
1940=7	BP(J) = B(J)
1950=	CALL TRANS(AA,BP)
1960=	M1=ZD
1970=	M2=Z
1980=	M3=((XD*XD+YD*YD+ZD*ZD)**0.5)-1.0096E04
1990=	M4=((X*X+Y*Y+Z*Z)**0,5)~1.3811E08
2000=	MS=XxXD+YXYD+ZXZD
2010=	M5=M5/ZED
2020=	G1=AA(1)+AA(4)+AA(7)
2030=	AA(I)=A(I)-DEL(I)
2040=	DO 8 J=1,NE
2050=8	BF(J)=B(J)
2030=	CALL TRANS(AA, BP)
2070=	M6=ZD
2080=	M7=Z
2090=	M8=((XD*XD+YD*YD+ZD*ZD)**0.5)-1.0096E04
2100=	M9=((X*X+Y*Y+Z*Z)**0.5)-1.3811E08
2110=	M10=X*XD+Y*YD+Z*ZD
2120=	M10=M10/ZED
2130=	G2=AA(1)+AA(4)+AA(7)
2140=	MA(1,I)=(M1-M6)/(2,0*DEL(I))
2150=	MA(2,I)=(M2-M7)/(2.0*DEL(I))
2160=	MA(3,I)=(M3-M8)/(2,0*DEL(I))
2170=	MA(4,I) = (M4-M9)/(2,0*DEL(I))
2180=	MA(5,I)=(M5-M10)/(2.0*DEL(I))
2190=	GA(I)=(G1-G2)/(2.0*DEL(I))
2200=	DO 100 K=1,I
2210=	IF(K.EQ.I)GO TO 707
2220=	IF(IMET.EQ.1)G0 T0 100
2230=	DO 9 J=1,NP
2240=9	
2250=	AA(I) = A(I) + DEL(I)
2260=	DEL(K)=DELTA*A(K)
70=	IF(ABS(DEL(K)).LE.DELTA) DEL(K)=DELTA
2280=	AA(K) = A(K) + DEL(K)
2290=	DO 10 J=1,NE
2300=10	BP1(J)=B(J)
2310=	CALL TRANS(AA, BF1)
2320=	MM1=ZD
2330=	MM2=Z
2340=	MMZ=((XD#XD+YD#YD+ZD#ZD)##0.5)-1.0096E04
2350	MM4=((X*X+Y*Y+Z*Z)**0.5)-1.3011E08
2360=	MMS=XXXD+YXYD+ZXZD
2370=	MM5=MM5/ZED
2380	GG1 = AA(1) + AA(4) + AA(7)
2390=	DO 11 $J=1+NP$ 41
2400=11	AA(J) = A(J)
•	

2410=	AA(I) = A(I) + DEL(I)
2420=	AA(K)=A(K)-DEL(K)
2430=	DO 12 J=1,NE
2440=12	BP2(J)=B(J)
2450=	CALL TRANS(AA, BP2)
2460=	MM6=ZD
2470=	
2480=	MM8=((XD*XD+YD*YD+ZD*ZD)**0.5)-1.0096E04
2490=	MM9=((X*X+Y*Y+Z*Z)**0.5)-1.3811E08
2500=	MM10=X*XD+Y*YD+Z*ZD
2510=	MM10=MM10/ZED
2520=	GG2=AA(1)+AA(4)+AA(7)
: '0=	DO 13 J=1+NP
2040=13	
2550=	AA(I)=A(I)-DEL(I)
2560=	AA(K) = A(K) + DEL(K)
2570=	DO 14 J=1+NE
2580=14	BP3(J)=B(J)
2590=	CALL TRANS(AA, BF3)
2600=	MM11=ZD
2610 **	MM1.2=Z
2620=	MM13=((XD*XD+YD*YD+ZD*ZD)**0.5)-1.0096E04
2630=	MM14=((X*X+Y*Y+Z*Z)**0.5)-1.3811E08
2640=	MM15=XXXD+YXYD+ZXZD
2650=	MM15=MM15/ZED
2650=	GG3=AA(1)+AA(4)+AA(7)
2670=	DO 15 J=1,NP
2680=15	
2590=	AA(I)=A(I)-DEL(I)
2700=	AA(K) = A(K) - DEL(K)
2710=	DO 16 J=1,NE
2720=16	BP4(J) = B(J)
2730=	CALL TRANS(AA,BF4)
2740=	MM16=ZD
2750=	
2760=	MM18=((XD#XD+YD#YD+ZD#ZD)##0.5)-1.0096E04
2770=	MM19=((X*X+Y*Y+Z*Z)**0.5)-1.3811E08
2780= 2790=	MM20=X*X0+Y*Y0+Z*Z0
2800=	MM20=MM20/ZED CC4=AA(1)\AA(A)\AA(7)
2810=	GG4=AA(1)+AA(4)+AA(7) MAA1(I,K)=MAA1(K,I)=(MM1-MM6-MM11+MM16)/(4.0%DEL(I)%DEL(K))
2810=	MAA2(I,K)=MAA2(K,I)=(MM2-MM7-MM12+MM17)/(4.0%DEL(I)%DEL(K))
2830=	MAA3(I;K)=MAA3(K;I)=(MM3-MM8-MM13+MM18)/(4.0*DEL(I)*DEL(K)) MAA3(I;K)=MAA3(K;I)=(MM3-MM8-MM13+MM18)/(4.0*DEL(I)*DEL(K))
2840=	MAA4(I,K)=MAA4(K,I)=(MM4-MM9-MM14+MM19)/(4.0*DEL(I)*DEL(K))
2850=	MAA5(I,K)=MAA5(K,I)=(MM5-MM10-MM15+MM20)/(4.0%DEL(I)*DEL(K))
2860=	GAA(I,K)=GAA(K,I)=(GG1-GG2-GG3+GG4)/(4.0*DEL(I)*DEL(K))
2870=	GO TO 100
`B0=707	MAA1(I,I)=(M1-2.0*M(1)+M6)/(DEL(I)**2)
ະຟ90≕	$MAA2(I + I) = (M2 - 2 \cdot 0 \times M(2) + M7) / (DEL(I) \times 2)$
2900=	MAA3(I+I)=(M3-2+0*H(3)+M8)/(DEL(I)**2)
2910=	MAA4(I+I)=(M4-2.0*M(4)+M9)/(DEL(I)**2)
2920=	MAA5(I,I)=(M5-2.0*M(5)+M10)/(DEL(1)**2)
2930=	GAA(I,I)=(G1-2.0*G+G2)/(DEL(I)**2)
2740=100	CONTINUE
2750=	IF(IMET-1)759,759,747
2960 =0	max - maximum - maximum Alexandria Alexandria Alexandria Alexandria Alexandria Alexandria Alexandria Alexandria
	NDING INITIAL LAGRANGE MULTIPLIERS AND DA'S (OFAUIENT TECH)
2980=C	renewerse werne ewenne wernersensene endewerken werde tekt warkenste mitter AMME OF COPHIULLICE EDUULL.
2990=759	DD 18 I=1,NC
3000	DO 18 J=1,NC η_D
•	

3010=	MAMAT(I,J)=0.0
3020=	DO 18 K=1,NP
3030=18	MAMAT(I,J)=MAMAT(I,J)+MA(I,K)*MA(J,K)
3040=	CALL GAUSD(NC+1+0E-30+MAMAT+MAMATI+IER+NC)
3050=	DO 19 I=1,NC
3060=	MAGAT(I)=0.0
3070=	DO 19 J=1,NP
3080=19	MAGAT(I)=MAGAT(I)+MA(I,J)*GA(J)
3090=	00 20 I=1,NC
3100=	FNU(I)=0.0
3110=	DO 20 J=1,NC
5150=50	FNU(I)=FNU(I)+MAMATI(I,J)*(M(J)-MAGAT(J))
	PRINT*, *
3140=	DO 21 I=1,NC
3150=21	FRINT*,"PNU(",I,")=",FNU(I)
3160=	PRINT*," •
3170=	DO 34 I=1,NP
3180=	E(I)=0.0
3190=	DO 34 J=1,NC
3200=34	E(I)=E(I)+MA(J+I)*PNU(J)
3210=	DO 35 I=1,NP
3220=	FA(I)=GA(I)+E(I)
3230=35	DA(I) = -P*FA(I)
3240=	F=0.0
3250=	DO 36 I=1,NP
3260=36	F=F+DA(I)**2
3270=	GRNDRM=SQRT(F)
3280=	DIFF=(FGNORM-GRNORM)/FGNORM
3290=	FONORM=GRNORM
3300=	PRINT*,* *
3310=	PRINT*, "FA MATRIX"
3320=	PRINT 50,(FA(I),I=1,NP)
3330=	FRINT*, * *
3340=	PRINT*, "GRADIENT METHOD DA'S"
3350=	PRINT 50,(DA(I),I=1,NP)
3360=	PRINT*, *
3370=	IF(DIFF.LT.CC2.AND.F.EQ.1.0) V=1.0
3380=	IF(V.EQ.1.0)PRINT*, GRADIENT METHOD CONVERGENCE
3390=	PRINT*,* *
3400=	IF(V.EQ.1.0) P=.1
3410=	
3420=	GU TO 1000
3430=C 3440=C	
3450=C	
3460=C	
3480≕C 7470=C	
80=C	
3490=C	
3500=C	
3510=C	
	LCULATING DPNU AND DA (SECOND ORDER TECH)
-3530≈C	COCHIING DEND HAD DA (SECUND URDER IECH)
3540=747	DO 22 I=1,NP
3550%	PNUMA(I)=0.0
3560 *	00 23 J=1+NC
3570=23	FNUMA(I)=FNUMA(I)+FNU(J)*MA(J,I)
3580=22	FAT(I) = GA(I) + PNUMA(I)
3590~	PRINT*** *
3600=	PRINT*, "PNU'S" 43
•	
-	

	3610=	PRINT 50,(PNU(I),I=1,NC)
	3620=	PRINT*, "
	3630=	PRINT*, "FA MATRIX"
	3540=	PRINT 50,(FAT(I),I=1,NP)
	3620=	DO 24 I=1,NP
	3660=	DO 24 J=1+NP
	3670=	FAA(I,J)=PNU(3)*MAA3(I,J)+PNU(4)*MAA4(I,J)+PNU(5)*MAA5(I,J)
	3680=24	FAA(I,J) = FAA(I,J) + PNU(1) + MAA1(I,J) + PNU(2) + MAA2(I,J) + GAA(I,J)
	3690=	PRINT#, FAA= ', FAA(1,J)
	3700=	
	3710=7007	CALL GAUSD(NF,1.0E-30,FAA,FAAI,JER,9)
	720= 3730=	DO 25 I=1,NC DO 25 J=1,NF
	3740=	MAFAAI(I;J)=0.0
	3750=	DO 25 $K=1$ · NP
	3760=25	$MAFAAI(I_{J}) = MAFAAI(I_{J}) + MA(I_{J}K) + FAAI(K_{J})$
	3770=	$DO = 26 I = 1 \cdot NC$
	3780≔	DO 26 J=1,NC
	3790=	$W(1,\mathbf{J})=0.0$
	3800=	DO 26 K#1+NP
	3810=26	W(I,J)=W(I,J)+MAFAAI(I,K)*MA(J,K)
	3820=	CALL GAUSD(NC,1.0E-30,W,WI,KER,NC)
	3830=	DO 27 I=1,NC
	3840=	C(I)=0.0
	3850=	DO 27 J=1,NF
	3860=27	C(I)=C(I)+MAFAAI(I,J)*FAT(J)
	3870=	DO 28 I=1,NC
	3880=	DPNU(I)=0.0
	3890=	
	3900=28 3910=	DFNU(I)=DFNU(I)+WI(I,J)*(-F*C(J)+Q*M(J))
	3920=	PRINT*," " PRINT*,"DPNU'S"
	3930=	PRINT 50, (DPNU(I), I=1, NC)
	3940=	$DO 29 I=1 \cdot NP$
	3950=	D(I) = 0.0
	3960=	DO 29 J=1,NC
	3970=29	D(I)=D(I)+MA(J,I)*DFNU(J)
	3980=	DO 30 I=1,NP
	3990=	DA(I)=0.0
	4000=	DO 30 J=1,NP
	4010=30	DA(I)=DA(I)+FAAI(I,J)*(-P*FAT(J)-D(J))
	4020=	NORMP=0.0
	4030=	DO 31 I=1,NC
	4040=31	NORMF=NORMP+DPNU(I)**2
	4050= 4060=	NORMP=SQRT(NORMP) PRINT*, * *
)70=	PRINT#, NORM OF DPNU'S= ', NORMP
	4080=	PRINT#9 "
	4090=	PRINT#, DA MATRIX"
	4100=	$PRINT 50 \cdot (DA(I) \cdot I = 1 \cdot NP)$
	4110=	NORMA=0.0
	4120=	DO 32 I=1,NP
	4130=32	NORMA=NORMA+DA(I)**2
	4140=	NORMA=SORT (NORMA)
	4150-	DIF=(FNORMA-NORMA)/FNORMA
	4160=	FNORMA=NORMA
	4170=	PRINT*, *
	4180=	PRINT*, NORM OF DA'S = ', NORMA
	4190=	IF(ITER.GE.10.0)GO TO 555
	4200=	IF(NORMA.GE.1.0) GO TO 4371 μμ
•	•	

4210=	Q=1.
4220≕C 4230≔C CO	INVERGENCE CRITERIA
4240=C	
4250=200	IF(NORMM.LE.CC3.AND.NORMA.LE.CC4.AND.F.EQ.1.0) GO TO 201
4260= 4270=	IF(F.GT.1.0) F=1.0 GO TO 1000
4280=4371	Q=Q/100.
4290	GO TO 555
4300=201	PRINT*," "
4310=	PRINT*, **CONVERGENCE* PERFORMANCE INDEX= *,G,* P= *,P
120=757	STOP
30=	END
4340=C	
4350=C	
	TRIX INVERSION SUBROUTINE
4370=C	
4380=C 4390=	SUBROUTINE GAUSD(M,EFS,B,C,KER,LAY)
4390	DIMENSION B(LAY,LAY), C(LAY,LAY),A(20,20), X(20,20)
4410=	DOUBLE PRECISION Z,A,X,S,RATIO,EP
4420=	EP = EPS
4430=	N = M
4440=	DO 100 J = 1,N
4450=	DO 100 K = $1 \cdot N$
	$A(J_{F}K) = B(J_{F}K)$
4470=	DO 1 I=1,N
4480≔ 4480= 1	$DO 1 J=1 \cdot N$
4490= 1 4500=	X(I,J) = 0.0D0 D0 2 K=1,N
	$X(K_{*}K) = 1.000$
4520= 10	
4530=	KP=0
4540=	Z = 0.0D0
4550=	DO 12 K=L+N
4560=	IF(Z-DABS(A(K,L))) 11,12,12
	Z = DABS(A(K,L))
4580=	
4590= 12 4600=	CONTINUE IF(L-KF)13,20,20
	DO 14 J=L;N
4620=	Z=A(L,J)
4630=	A(L,J) = A(KP,J)
4640= 14	A(KP,J)=Z
4650=	DO 15 J=1,N
4660=	Z=X(L,J)
170= 00	X(L,J) = X(KP,J)
80= 15 4690= 20	- X(KP;J)=Z - IF(DABS(A(L;L))-EP)50;50;30
4700= 30	
4710= 31	
4720=	DO 36 K=LP1,N
4730=	IF(A(K+L))32+36+32
4740= 32	
475 0 =	DO 33 J=LP1+N
4760 - 33	
4770-	DO 35 J=1+N
- 4780≕ - 35 - 4790≕ - 36	
4790°° 38 4800= 34	
	Lic.

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4910=	40	
4820=		II=N+1-I
4830=		DO 43 J=1,N
4840=		S = 0.0D0
4850=		IF(II-N)41,43,43
4860=	41	
4870=		DO 42 K=IIF1,N
4880=	42	
4890=	-43	
4900=		KER=1
4910=		10200 J = 1.N
4920=		DD 200 K = 1 N
1930=	200	$C(J_*K) = X(J_*K)$
/40==		GO TO 75
4950=	50	KER=2
4960=	70	PRINT 71
4970=	71	FORMAT(1X,*MATRIX SINGULAR IN GAUSD*)
4980=	75	CONTINUE
49 90≕		RETURN
5000=		END
5010=C		
5020≖C		
5030=C		
5040=		SUBROUTINE CHEBY(T)
5050=		DIMENSION Z(7)
5060=		COMMON/MISC/NPH,NPI,NA,TF,N,S,LL,PP,TOT(8),FIXT
5070=		COMMON/QUES/ALPHA
5080=		Z(1)=T
5090=		1011=2,7
5100=		K=I-1
5110=1		Z(I) = Z(K) * T
5120=		TOT(1)=1.0
5130=		$TOT(2) = 2.0 \times Z(1) - 1.0$
5140=		$TOT(3) = 8.0 \times Z(2) - 8.0 \times Z(1) + 1.0$
5150=		$TOT(4) = 32.0 \times Z(3) - 48.0 \times Z(2) + 18.0 \times Z(1) - 1.0$
5160=		$TOT(5) = 128.0 \times Z(4) - 256.0 \times Z(3) + 160.0 \times Z(2) - 32.0 \times Z(1) + 1.0$
5170=		$TOT(6) = 512.0 \times Z(5) - 1280.0 \times Z(4) + 1120.0 \times Z(3) - 400.0 \times Z(2) + 50.0 \times Z(1) - 1.0$
5180≖		$TOT(7) = 2048.0 \times Z(6) - 6144.0 \times Z(5) + 6912.0 \times Z(4) - 3584.0 \times Z(3) + 840.0 \times Z(2)$
5190=		1 -72.0*Z(1)+1.0
5200=	•	TOT(8)=8192.0*Z(7)-28672.0*Z(6)+39424.0*Z(5)-26880.0*Z(4)
5210=		1 +9408.0*Z(3)-1568.0*Z(2)+98.0*Z(1)-1.0
5220=		RETURN \$ END
5230=		SUBROUTINE TRANS(A,B)
5240=		DIMENSION A(9),B(6)
5250=		COMMON X,Y,Z,XD,YD,ZD
5260=		X=B(1)
5270=		Y=₿(2)
5280=		Z≈B(3)
290≈		XD=B(4)
5300=		YD=B(5)
5310=		ZD=B(6)
5320≈		T1=A(1)
5330=		A1=A(2)
5340=		B1=A(3)
5350=		T2=A(4)
5360=		A2=A(5)
5370=		B2=A(3)
5390 5390		T3=A(7)
		A3=A(8)
5390= 5400=		A3=A(8) B3=A(9)
		46

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E A A A.	705-T1
5410=	TOF=T1 CALL FNG(TOF,X,Y,Z,XD,YD,ZD)
5420=	CALL FROCIUFIXITYZIXDITUIZDI
5430=C	
5440=C	VELOCITY CHANGE DUE TO FIRING OF FIRST STAGE
5450=C	
5460=	DXD=4242.175*COS(B1)*COS(A1)
5470=	DYD=4242.175*COS(B1)*SIN(A1)
5480=	DZD=4242.175*SIN(B1)
5490=	XD=XD+DXD
5500=	YD=YD+DYD
5510=	ZD=ZD+DZD
5520=	TOF=T2
-30=	CALL FNG(TOF,X,Y,Z,XD,YD,ZD)
ວ540=C	
5550=C	VELOCITY CHANGE DUE TO FIRING OF SECOND STAGE
5560≖C	
5570=	DXD=9565.712*COS(B2)*COS(A2)
5580=	DYD=9565.712*COS(B2)*SIN(A2)
5590=	DZD=9565.712*SIN(B2)
5600=	X0=XD+DXD
5610=	YD≕YD+DYD
5620=	
5630=	
5640=	CALL FNG(TOF,X,Y,Z,XD,YD,ZD)
5650=C	
5660=C	VELOCITY CHANGE DUE TO FIRING OF THIRD STAGE
5670=C	
5680=	DXD=10997.798*COS(B3)*COS(A3)
5390=	DYD=10997.798*COS(B3)*SIN(A3)
5700=	DZD=10997.798*SIN(B3)
5710=	XD≈XD+DXD
5720=	ϒϼ϶ϒϼͱϧϒϼ
5730=	ZD=ZD+DZD
5740=	R=(X*X+Y*Y+Z*Z)**0.5
5750=	V=(XD*XD+YD*YD+ZD*ZD)**0.5
5760=	RETURN
5770=	END
5780=	SUBROUTINE FNG(TOF,X,Y,Z,XD,YD,ZD)
5790=	AMU=1.4076468E16
5800 =	E=1.0E-06
5810=	RD=(X*X+Y*Y+Z*Z)**0.5
5820≈	EPS=(XD#XD+YD#YD+ZD#ZD)/2.0-AMU/RD
5830≈	A=-AMU/(2.0*EPS)
5840=	XN=1.0
5850=1	ZZ=XN*XN/A
5860=	C=(1.0~COS(ZZ**0.5))/ZZ
5870≈	S=((ZZ**0.5)-SIN(ZZ**0.5))/(ZZ**1.5)
380=	TN=(X*XD+Y*YD+Z*ZD)*C*(XN*XN)/(AMU**0.5)
J890=	TN=TN+(1.0-R0/A)*S*(XN**3.0)
5900=	
5910=	TN=TN/(AMU**0.5)
5920=	RN=(XN*XN*C)+(X*XD+Y*YD+Z*ZD)*(1.0~ZZ*S)/(AMU**0.5)
5930=	RN=(\N\#\N\#C)+(\A*\D+[*+D+2*2D)*(1,0~22*5)/(AMD**0.5) RN=RN+RO*(1.0~ZZ*C)
5930= 5940=	NT=TOF-TN
	DI=IDF-IN IF(DT.LE.E)60 TO 99
5950=	
5930-	XN=XN+DT*(AMU**0.5)/RN
5970= 5900 00	
5980=99	F=1.0-(XN*XN*C)/(RO)
5990=	G=TOF-(XN**3.0)*S/(AMU**0.5)
6000 m	₽≖X
•	<u>Ц7</u>

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6010=	Q=Y
6020=	V=Z
6030=	X≕F*X+G*XD
6040=	Y=F*Y+G*YD
6050=	Z=F*Z+G*ZD
6060=	GD=1.0-(XN*XN)*C/((X*X+Y*Y+Z*Z)**0.5)
6070=	FD=(AMU**0.5)*(ZZ*S-1.0)/(RD*((X*X+Y*Y+Z*Z)**0.5))
<080=	FD=FD*XN
J90 =	XD=FD*P+GD*XD
6100=	YD=FD*Q+GD*YD
6110=	ZD=FD*V+GD*ZD
6120=	RETURN
6130=	END
6140=*EOR	
6150=*EOF	

<u>Vita</u>

Mark N. Brown was born on 18 November 1951 in Valparaiso, Indiana. He graduated from high school in Valparaiso in 1969 and attended Purdue University, from which he received the degree of Bachelor of Aeronautical and Astronautical Engineering in May 1973. Upon graduation, he received a commission in the USAF through the ROTC program. He attended Undergraduate Pilot Training at Laughlin AFB and was assigned to the 87th Fighter Interceptor Squadron at K.I. Sawyer AFB, Michigan. He served initially as a T-33 pilot and then as an F-106 pilot until entering the School of Engineering, Air Force Institute of Technology, in June 1979.

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parking orbit data and IUS characteristics were used in this study to examine the sequential timing and orientation in inertial space of each stage as it is fired while the spacecraft moves along a transfer orbit to geosynchronous orbit. In addition, the sensitivity of the total transfer time and the final orbital state was found as a result of not meeting one or all of the time and orientation parameters.

This problem is unique in that it considers an optimal orbit transfer problem involving solid fuel stages of fixed thrust and burn time. Previous work with liquid fuel engines examined orbital transfers with the intent of minimizing the amount of propellant or required velocity change needed to accomplish the transfer.

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