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DEPARTMENT OF DEFENSE LAND FALLOUT
PREDICTION SYSTEM. VOLUME II: INITIAL
CONDITIONS, SUPPLEMENT

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Mount Auburn Research Associates, Incorporated

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VOLUME II
INITIAL CONDITIONS
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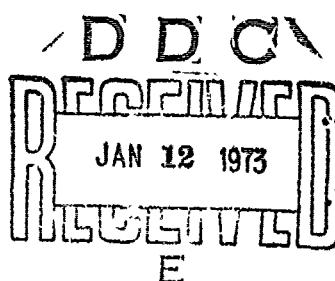
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13. ABSTRACT <p>The DELFIC Initial Conditions Module code has been revised to meet the requirements of the new DELFIC fallout prediction system. This documentation supplement describes the revised code. Discussion of the revised code emphasizes particle size distributions. The code can accept parameters that define lognormal or power-law distributions, or it can accept a distribution in tabular form. Details necessary for use of the code are presented. FORTRAN statement listings of revised subroutines are included.</p> <p style="text-align: center;">Ia</p>		

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II

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III

1. INTRODUCTION

Since publication of DASA-1800-II⁽¹⁾ in 1966, the Defense Land Fallout Interpretative Code (DELFIC) has undergone substantial revision in all of its modules. These revisions have created some new demands on the Initial Conditions Module (ICM) and removed some old restrictions. Of most direct consequence to the ICM are changes in the Cloud Rise Module (CRM)⁽²⁾.

The new CRM accounts for wind shear effects on the cloud rise dynamics. Therefore, shot-time winds above ground zero are input via the ICM rather than via the Cloud Rise-Transport Interface Module (CRTIM) as was done originally. The old CRM could accept no more than forty particle size classes, and the size class structure was rigidly prescribed. These restrictions have been relaxed in the new CRM, and the ICM has been revised accordingly. In addition, the ICM has been given a capability to accept parameters that define a power-law particle size distribution function. From these parameters, it constructs a particle size class table with a user-specified number of entries.

Subroutines LINK1 and DSTBN have been revised, and a new subroutine, SHWIND, which is called by LINK1, has been created. Subroutines MASS, TEMP, TIME, and VAPOR remain unchanged.

Subroutine LINK1 is the ICM executive program. Subroutine DSTBN constructs particle size class tables for lognormal and power-law particle distributions. Subroutine SHWIND reads in the shot-time winds above ground zero.

The logic of the ICM consists of a card input, which is described in Table 1, followed by serial exercise of the subordinate subroutines. Adequate detailed documentation is provided by the FORTRAN statement listings.

Use of the ICM is quite simple with one exception: definition of particle size distributions. Therefore, the bulk of this supplement is devoted to discussions of particle size distributions.

2. THE LOGNORMAL DISTRIBUTION

2.1 Fundamentals

A variable x is said to be normally distributed if the probability of its occurrence in the range x to $x + dx$ is given by

$$dN(x|\mu, \sigma^2) = \frac{1}{\sigma\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{x-\mu}{\sigma}\right)^2\right] dx, \quad (1)$$

where μ is the mean value of x and σ^2 is the variance of x . The square root of the variance, σ , is called the standard deviation.

To define a lognormal distribution, we make the transformation

$$x = \ln(y) . \quad (2)$$

In terms of the variable y Eq. (1) becomes

$$dN(y|\mu, \sigma^2) = \frac{1}{\sigma\sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{\ln y - \mu}{\sigma}\right)^2\right] d(\ln y), \quad (3)$$

and y is said to be lognormally distributed. ⁽³⁾

Some statistical properties of y are as follows:

$$\text{mean}(y) = \exp(\mu + \frac{1}{2} \sigma^2) \quad (4)$$

$$\text{median}(y) = \exp(\mu) \quad (5)$$

$$\text{mode}(y) = \exp(\mu - \sigma^2) \quad (6)$$

$$\text{variance}(y) = [\exp(\sigma^2) - 1] \exp(2\mu + \sigma^2). \quad (7)$$

Let \bar{y} and s be the geometric mean and geometric standard deviation of y . Then

$$\bar{y} = \text{median}(y) = \exp(\mu) \quad (8)$$

and

$$s = \exp(\sigma) . \quad (9)$$

Let λ'_j be the j -th moment of $\Lambda(y|\mu, \sigma^2)$ about the origin. Then by definition

$$\lambda'_j = \int_0^\infty y^j d\Lambda(y|\mu, \sigma^2) , \quad (10)$$

and from the properties of the normal distribution it follows that

$$\lambda'_j = \exp(j\mu + \frac{1}{2} j^2 \sigma^2) . \quad (11)$$

A feature that distinguishes the lognormal distribution from the normal distribution is the existence of moment distributions. The j -th moment distribution is defined as

$$\Lambda(y|\mu, \sigma^2)_j = \frac{1}{\lambda'_j} \int_0^y t^j d\Lambda(t|\mu, \sigma^2) , \quad (12)$$

which can be shown to be⁽³⁾

$$\Lambda(y|\mu, \sigma^2)_j = \Lambda(y|\mu + j\sigma^2, \sigma^2) . \quad (13)$$

The moment distributions provide simple relationships between log-normal distributions of number, surface area, and volume of particles with respect to their diameters.

2.2 Application to Particle Distributions

In discussions of lognormal particle distributions, confusion frequently arises because distinction is not clearly made between μ and y and between σ and s . Since particular values of μ and σ depend on the base of the logarithms used, we have chosen to confine our discussions in the DELFIC documentation to parameters in the form of y and s .

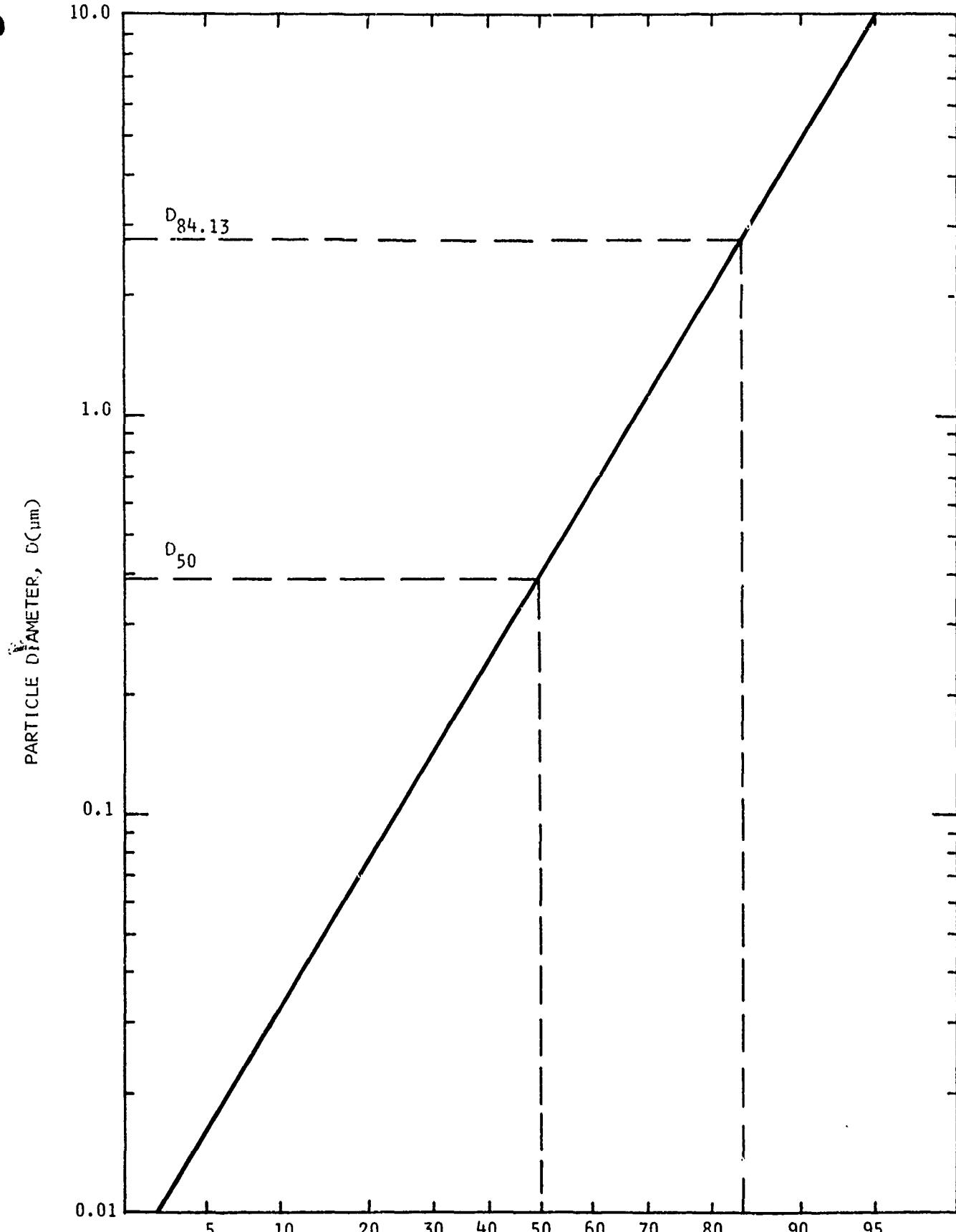
Suppose that we have plotted cumulative numbers of particles versus diameter on log-probability graph paper and have obtained the curve shown in Figure 1. This straight-line curve indicates that the distribution of particle number with respect to diameter, D , is lognormal. Thus, D is equivalent to y in Eq. (3), and from Eqs. (8) and (9) we have

$$y = D_{50}$$

and

$$s = D_{84.13}/D_{50}$$

These are the quantities DMEAN and SD, respectively, that are required as input to subroutine LINK1, and that are printed by LINK1. DMEAN is expressed in units of micrometers. (In words, DMEAN is the median particle diameter in the distribution of numbers of particles with respect to their diameters.) If the user specifies a lognormal distribution but does not input values for DMEAN and SD, the program supplies the values⁽¹⁾:



PERCENT OF TOTAL NUMBER OF PARTICLES HAVING DIAMETERS LESS THAN VALUE INDICATED

FIGURE 1. CUMULATIVE FREQUENCY GRAPH OF LOGNORMALLY DISTRIBUTED PARTICLES

$$DMEAN = \underline{y} = 0.407 \mu\text{m}$$

$$SD = s = 4.0$$

As noted above, the properties of the moment distributions are useful in interrelating distributions of particle number, surface area, and volume with respect to particle diameter. This is because the number distribution is the zeroth moment distribution with respect to diameter, surface area is distributed via the second moment distribution, and volume is distributed via the third moment distribution. Thus, if we assume spherical particles and if the parameters μ and σ are known for either the particle number, or particle area, or particle volume distribution with diameter, then the other distributions can be determined from the equations below. The parameter σ is the same for all three distributions. If we use N, S, and V as subscripts to denote number, surface area, and volume, respectively, we have from Eq. (13)

$$\mu_S = \mu_N + 2\sigma^2$$

$$\mu_V = \mu_N + 3\sigma^2$$

where μ and σ are related to \underline{y} and s by Eqs. (8) and (9).

If base 10 logarithms are used instead of natural logarithms, we distinguish the distribution parameters by use of primes, μ'_N and σ' , and the relations become

$$\mu'_S = \mu'_N + 2\ln(10)(\sigma')^2$$

$$\mu'_V = \mu'_N + 3\ln(10)(\sigma')^2$$

where $\ln(10) = 2.3026$.

The distribution of particle mass with respect to diameter is taken to be equivalent to the volume distribution with respect to diameter. This implies that all particles have the same density.

2.3 Particle Size Class Tables

For computation purposes, the continuous lognormal distribution is replaced by a histogram. The computer program, via subroutine DSTBN, does this automatically by use of the distribution parameters and the number of size classes, NDSTR, which is input by the user.

The user specifies parameters DMEAN, SD, and NDSTR. From these, the parameters μ_N , σ , and μ_V are determined via

$$\mu_N = \ln(DMEAN)$$

$$\sigma = \ln(SD)$$

$$\mu_V = \mu_N + 3\sigma^2 .$$

Define the normal distribution function argument x as

$$x = \frac{\ln(D) - \mu_V}{\sigma}$$

where D is particle diameter. Then

$$N(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^x \exp(-t^2/2) dt$$

Subroutine DSTBN constructs the particle size class table (i.e., histogram) as follows. Each size class contains a constant

volume fraction, ΔN_V , of

$$\Delta N_V = 1/NDSTR .$$

Let D_i , $i = 1, 2, \dots, NDSTR$, be the upper (i.e. the larger particle) boundary diameter of the i -th particle size class. The table is ordered with the largest particles in the first size class, and so on. Then, for the i -th size class

$$N(x_i) = i\Delta N_V$$

and

$$\ln(D_{i+1}) = x_i^\sigma + \mu_V .$$

The upper boundary of the first size class, D_1 , and the lower boundary of the last size class, $D_{NDSTR+1}$, are special cases. These are taken to be the diameters at $\Delta N_V/2$ and $1-\Delta N_V/2$, respectively. That is,

$$N(x_1) = \frac{\Delta N_V}{2}$$

and

$$N(x_{NDSTR+1}) = 1 - \frac{\Delta N_V}{2} .$$

In these calculations x is determined from given $N(x)$ via equation 26.2.23 of Reference 4.

The central particle diameter for the i -th class, d_i , is given by

$$d_i = \sqrt{D_i D_{i+1}} .$$

If NDSTR = 1, a single size class is created with

$$D_1 = (\text{DMEAN}) * (5.0 * \text{SD})$$

$$D_2 = (\text{DMEAN}) / (5.0 * \text{SD})$$

and

$$d_1 = \text{DMEAN}$$

3. THE POWER-LAW DISTRIBUTION

3.1 Fundamentals

Mathematically speaking, power-law distributions are meaningless since distribution functions cannot be defined for them. This is because the power-law function is not properly bounded for zero argument. Freiling has shown that fallout particle distributions that have been represented by power-law functions can equally well be fitted by lognormal distribution functions.⁽⁵⁾ The implication of Freiling's work is that power-law distributions would be more accurately described as truncated lognormal distributions. Nevertheless, power-law distributions frequently are useful in fallout work.

Define the power-law frequency as

$$df(D|k,X) = kD^{-X}dD , \quad (14)$$

where $df(D|k,X)$ is the number of particles in the diameter range D to $D + dD$. If we assume spherical particles with constant density, ρ , we have

$$dF\left(D \left| \frac{\pi \rho k}{6M}, X \right.\right) = \frac{\pi \rho k D^{3-X}}{6M} dD , \quad (15)$$

where $dF\left(D \left| \frac{\pi \rho k}{6M}, X \right.\right)$ is the fraction of the total fallout mass, M , in the diameter range D to $D + dD$.

The mass fraction of particles in the macro-range from D_i to D_j is obtained by integration of Eq. (15) between these limits to give

$$\Delta F = \frac{\pi \rho k}{6M(4-X)} \left(D_j^{4-X} - D_i^{4-X} \right), \quad 0 < X < 4 . \quad (16)$$

3.2 Particle Data Analysis

Suppose that we have obtained a sample of fallout particles. We weigh the sample to obtain M (kg), and we size the sample into N fractions, the i -th fraction containing particles in the diameter range ΔD_i centered on D_i (meters). We weigh each fraction and obtain the mass fractions ΔF_i . We determine that the average particle density is ρ (kg/m^3).

To obtain the power law distribution parameters k and X , we plot $\log(\Delta F_i / \Delta D_i)$ versus $\log(D_i)$. A straight line is fitted to the data. From Eq. (15), we see that the intercept and slope are

$$\text{intercept} = c = \log\left(\frac{\pi\rho k}{6M}\right) ,$$

and

$$\text{slope} = m = 3 - X .$$

Then

$$X = 3 - m$$

and

$$k = \frac{6M}{\pi\rho} \log^{-1}(c) .$$

When X and k are determined from M expressed in kilograms, D and ΔD in meters, and ρ in kg/m^3 , they can be input to subroutine LINK1 as EXPO and CAY, respectively.

3.3 Particle Size Class Tables

For use in fallout calculations, subroutine DSTBN creates a histogram representation of the power law distribution. The histogram is comprised of NDSTR particle size classes, where NDSTR is

specified by the user. The mass fraction in each size class, ΔF , is the constant

$$\Delta F = 1/NDSTR .$$

Let D_i be the upper (i.e. larger particle) boundary of the i -th particle size class. The table is ordered with the largest particles in the first class, and so on. Then the smallest particles are contained in the $NDSTR_{th}$ class. If we assume that the smallest particle in this class is much smaller than D_{NDSTR} , we see from Eq. (16) that

$$D_{NDSTR}^{4-X} = \frac{6M(4-X)}{\pi\rho k} \Delta F .$$

By recursive use of this relation with Eq. (16), we find that

$$D_i = (NDSTR - i+1)^{\frac{1}{4-X}} D_{NDSTR} .$$

Size class central diameters, d_i , are

$$d_i = \sqrt{D_i D_{i+1}} .$$

To establish a central and lower boundary diameter for the $NDSTR_{th}$ class, we say that

$$d_{NDSTR} = \left(\frac{1}{2}\right)^{\frac{1}{4-X}} D_{NDSTR}$$

and

$$D_{NDSTR+1} = (d_{NDSTR})^2 / D_{NDSTR} .$$

4. TABULAR DISTRIBUTIONS

4.1 Particle Size Class Tables

If the user so desires, he can input his particle size distribution in histogram form with NDSTR size classes. The table of size classes must be arranged in descending order of particle diameter. Each size class is defined in the input by its upper (i.e. larger particle) boundary diameter, D_i , and mass fraction, ΔF_i . These two data are punched on a separate card for each size class. The last card in the deck contains the lower boundary diameter of the NDSTRth size class. Central particle diameters, d_i , are computed to be

$$d_i = (D_i + D_{i+1})/2 .$$

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5. USER INFORMATION

5.1 Card Input

The ICM card input is described in Table I. This table and the discussions in Sections 2.2, 2.3, 3.2, and 4.1 provide adequate information for use of the code.

5.2 Output

Though the printed output has been modified somewhat, the example output presented in DASA-1800-II is still satisfactory.

Communication with the Cloud Rise Module is via COMMON/SET1/. The contents of COMMON/SET1/ is described in Table 2.2 of DASA-1800-III (Revised).⁽²⁾

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TABLE I
CARD INPUTS TO THE INITIAL CONDITIONS MODULE

<u>Card Number</u>	<u>Contents</u>	<u>Variable Names and FORMATS</u>
1	ICM Run Identifier	DETID(J), J=1,12 (12A6)
2	Control integer to specify particle size distribution type: 1 lognormal 2 power-law 3 tabular	IDISTR (I5)
3	Number of particle size classes	NDSTR (I5)
4(a)*	(For lognormal particle size distribution) Explosion yield (KT), height of burst above GZ(m), soil class indicator: 1.0 for siliceous 2.0 for calcareous, median particle diameter (μm), geometric standard deviation of the particle size distribution, and particle density (g/cm^3). (See Section 2.2.)	W, HEIGHT, USOIL, DMEAN, SD, DNS (6F10.3)
4(b)*	(For power-law particle size distribution) Yield (KT), height of burst (m), soil class indicator (see above), exponent in the particle size distribution frequency function, coefficient in the particle size distribution frequency function, particle density (g/cm^3). (See Section 3.2.)	W, HEIGHT, USOIL, EXPO, CAY, DNS (6F10.3)
4(c)*	(For a tabular particle size distribution) Yield (KT), height of burst (m), soil class indicator (see above), particle density (g/cm^3).	W, HEIGHT, USOIL, DNS (4F10.3)
4(c) ₁ ** .	A table of upper boundary particle diameters (μm) and mass fractions .	DIAM(J), FMASS(J), J=1,NDSTR (2E12.5) .
4(c) _{NDSTR+1} **	The lower boundary diameter (μm) of the last particle size class. (See Section 4.1.)	DIAM(NDSTR+1) (E12.5)

Table I (continued)

Card Number	Contents	Variable Names and FORMATS
5	Number of entries in the wind data table	NHODO (I5)
6***	For each entry in the wind data table: altitude (m, relative to msl), x component of wind (m/sec), y component of wind (m/sec)	ZV(J), VX(J), VY(J), J=1, NHODO (3F12.3)

- * One of the cards 4(a), 4(b), or 4(c) is read according to whether IDISTR is 1, 2, or 3.
- ** These cards are read only for a tabular distribution.
- *** These cards are read only if NHODO > 0.

6. FORTRAN STATEMENT LISTINGS

Complete FORTRAN statement listings are given for the following subroutines. These subroutines are operational on the UNIVAC 1108.

<u>SUBROUTINE</u>	<u>Page</u>
LINK1	22
DSTBN	27
SHVIND	29

The machine used to prepare these listings prints a # symbol to represent a 4-8 punch; this symbol should be an apostrophe ('). In FORMAT and DATA statements, the apostrophe is used to define Hollerith character fields.

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SUBROUTINE LINK1
 INITIAL CONDITIONS (FIRBALL) MODULE
 MT. AURURN RESEARCH ASSOCIATES JANUARY 1972

**** * PROGRAM TO DETERMINE THE INITIAL CONDITIONS SPECIFICATIONS OF TIME, TEMPERATURE, TOTAL SOIL MASS, FRACTION OF THE SOIL PURUFN IN THE VAPOR PHASE, AND THE SIZE FREQUENCY DISTRIBUTION OF THE CONDENSED PHASE SOIL

THE FIRST CARD CONTAINS ANY ARBITRARY ALPHANUMERIC IDENTIFICATION.

OTHER INPUT PARAMETERS ARE - TEST PARAMETER (IDISTR) TO DETERMINE IF THE PARTICLE SIZE FREQUENCY DISTRIBUTION IS LOG-NORMAL, POWER LAW, OR TABULAR, YIELD IN KILOTONS, HEIGHT(DEPTH) OF BURST IN METERS, A SOIL TYPE INDICATOR, FALLOUT PARTICLE DENSITY(GM/CM**3), MEAN(MICROMETERS) AND STANDARD DEVIATION FOR A LOG-NORMAL PARTICLE SIZE FREQUENCY DISTRIBUTION, THE NUMBER OF PARTICLE SIZE CLASSES IN THE PARTICLE SIZE FREQUENCY DISTRIBUTION. IF EITHER A TABULAR OR POWER LAW DISTRIBUTION IS USED, THE MEAN AND STANDARD DEVIATION ARE NOT CALLED FOR SINCE THEY DO NOT APPLY. IF A LOG-NORMAL DISTRIBUTION IS TO BE SUPPLIED BY THE PROGRAM, THE MEAN AND STANDARD DEVIATION FIELDS ARE LEFT BLANK.
 SHOT TIME WINDS ABOVE GZ ALSO ARE INPUT. THESE ARE USED TO COMPUTE WIND SHEAR EFFECTS ON CLOUD RISE AND FALLOUT ADVECTION DURING THE CLOUD RISE TIME INTERVAL.

FOR UNDERGROUND BURSTS INPUT DEPTH OF BURST AS A NEGATIVE NUMBER

THE OUTPUT UNITS ARE MASS IN KILOGRAMS, LENGTH IN METERS, TIME IN SECONDS, TEMPERATURE IN DEGREES KELVIN, YIELD IN KILOTONS, DISTRIBUTION PARAMETERS IN MICRONS

**** * GLOSSARY **** *

CAY	COEFFICIENT OF THE FREQUENCY FUNCTION FOR THE POWER LAW PARTICLE SIZE FREQUENCY DISTRIBUTION	LINK1
DETID(I)	INITIAL CONDITIONS IDENTIFICATION ARRAY	LINK1
DIAM(I)	ARRAY(201), UPPER BOUNDARY OF THE I-TH PARTICLE SIZE CLASS. THE LAST ENTRY IN THE DIAM ARRAY IS THE LOWER BOUNDARY OF THE LAST(SMALLEST) PARTICLE SIZE CLASS. THE LENGTH OF THE DIAM ARRAY IS ALWAYS ONE GREATER THAN THE NUMBER OF SIZE CLASSES(MICROMETERS)	LINK1
DMEAN	MEDIAN DIAMETER (MICROMETERS) OF LOGNORMAL PARTICLE SIZE DISTRIBUTION	LINK1
DNS	FALLOUT PARTICLE DENSITY (GM/CM**3)	LINK1
EXPO	EXPONENT OF THE FREQUENCY FUNCTION FOR THE POWER LAW PARTICLE SIZE FREQUENCY DISTRIBUTION	LINK1
FMASS(I)	ARRAY OF FRACTION OF TOTAL PARTICULATE MASS IN I-TH PARTICLE SIZE CLASS. MAXIMUM LENGTH OF ARRAY = 200	LINK1
HEIGHT	HEIGHT OF BURST (METERS) ABOVE GROUND ZERO	LINK1
IDISTR	CONTROL INTEGER FOR PARTICLE SIZE DISTRIBUTION 1 - LOGNORMAL DISTRIBUTION 2 - POWER LAW DISTRIBUTION 3 - TABULAR DISTRIBUTION READ IN ON CARDS (ARRAY WHY)	LINK1
IS	CONTROL INTEGER SPECIFIES WHETHER LOGNORMAL DISTRIBUTION IS SPECIFIED BY THE USER OR BY THE	LINK1

C	PROGRAM	LINK1 59
C	0 - PROGRAM SPECIFIED LOG-NORMAL DISTRIBUTION	LINK1 60
C	1 - USER SPECIFIED LOG-NORMAL DISTRIBUTION	LINK1 61
C	ISIN	LINK1 62
C	ISOUT	LINK1 63
C	NDSTR	LINK1 64
C	NHODO	LINK1 65
C	PS(I)	LINK1 66
C	SD	LTK1 67
C	SSAM	LINK1 68
C	TME	LINK1 69
C	TMP1	LINK1 70
C	TMP2	LINK1 71
C	T2M	LINK1 72
C	USOIL	LINK1 73
C	VPR	LINK1 74
C	VX(I)	LINK1 75
C	VY(I)	LINK1 76
C	W	LINK1 77
C	ZSCL	LINK1 78
C	ZV(I)	LINK1 79
C	*****	LINK1 80
C	COMMON /SET1/	LINK1 81
C	1CAY ,DETID(12) ,DIAM(201) ,DMEAN ,DNS ,EXPO ,LINK1 91	LINK1 92
C	2FMASS(200) ,IDISTR ,IEXEC ,IRISE ,ISIN ,ISOUT ,LINK1 93	LINK1 94
C	3NDSTR ,PS(200) ,SD ,SSAM ,TME ,TMP1 ,LINK1 95	LINK1 96
C	4TMP2 ,T2M ,USOIL ,VPR ,W ,HEIGHT ,LINK1 97	LINK1 98
C	5ZSCL ,NHODO ,ZV(200) ,VX(200) ,VY(200) ,LINK1 99	LINK1 100
C	*****	LINK1 101
1	FORMAT(12A6)	LINK1 102
2	FORMAT(/3X,60HTHE SPECIFIED STANDARD DEVIATION IS NEGATIVE HENCE INCORRECT//)	LINK1 103
3	FORMAT(7F10.3)	LINK1 104
4	FORMAT///25X28H**** INPUT PARAMETERS ****/20X,5HYIELD,40X,E12.5LINK1 105	LINK1 106
5	1,2X,2HKT/20X,24HHEIGHT OR DEPTH OF BURST,21X,E12.5,2X,6HMETERS/20XLINK1 107	LINK1 108
6	2,13HSOIL CATEGORY	LINK1 109
7	FORMAT(1H+,65X,9HSILICEOUS)	LINK1 110
8	FORMAT(1H+,65X,10HCALCAREOUS)	LINK1 111
9	FORMAT(/20X, 36HPARTICLE SIZE FREQUENCY DISTRIBUTION/LINK1 112	LINK1 113
10	125X32HA LOG-NORMAL DISTRIBUTION WITH -/30X,15HMEDIAN DIAMETER,20 ,LINK1 114	LINK1 115
	2E12.5,2X,11HMICROMETERS/30X,28HGEOMETRIC STANDARD DEVIATION, 7X, LINK1 116	LINK1 117
	3E12.5/25X, 34HTHIS DISTRIBUTION WAS SPECIFIED BY)	
	FORMAT(1H+,65X,11HTHE PROGRAM)	
	FORMAT(1H+,65X,8HTHE USER)	
	FORMAT(15I)	


```

220 READ(ISIN,3)W,HEIGHT,USOIL,EXPO,CAY,DNS           LTNK1175
      GO TO 23
211 READ(ISIN,3)W,HETGHT,USOIL,DNS                LINK1176
      READ(ISIN,195)(DIAM(I),FMASS(I),I=1,NODSTR)
      LD=NODSTR+1
      READ(ISIN,195)DIAM(LD)                         LINK1177
C
C   CHECK ORDERING OF THE HISTOGRAM TABLE           LTNK1178
      DO 215 I=2,LD
      IF(DIAM(I) .LT. DIAM(I-1)) GO TO 215          LINK1179
      WRITE( ISOUT,198)
      GO TO 190
215 CONTINUE                                         LINK1180
C
C   23   CONVERT HOB - DOB FROM METERS TO FEET       LTNK1181
      23 HEIGHT=HEIGHT/0.3048                         LINK1182
C   ZSCL IS THE SCALED HOB - DOB                   LINK1183
      60 ZSCL=HEIGHT/((W)**(1.0/3.4))               LINK1184
C
C   TEST THE DATA TO SEE IF THE MODEL IS APPROPRIATE
      IF(HEIGHT)66,66,63                            LINK1185
      63 IF(ZSCL-190.0)170,70,150                  LINK1186
      66 IF(ZSCL+20.0)143,70,70
70   CALL TIME                                         LTNK1187
      CALL TEMP                                         LINK1188
      CALL MASS                                         LINK1189
      CALL VAPOR                                         LINK1190
      GO TO (90,95,95),IDISTR                         LINK1191
C
C   TEST FOR ACCEPTABLE SPECIFICATIONS OF PRE-SHOT PARTICLE SIZE
      FREQQUENCY DISTRIBUTION.                      LINK1192
      90 IF(SD)91,92,92                                LINK1193
      91 WRITE (ISOUT,2)                               LINK1194
      GO TO 190
      92 IF(DMEAN)94,95,95                           LINK1195
      94 WRITE (ISOUT,17)                           LTNK1196
      GO TO 190
C
      95 CALL DSTBN                                     LINK1197
C
C   CONVERT HOB - DOB PACK TO METERS FROM FEET
      HEIGHT=HEIGHT*0.3048                         LINK1198
C
C   CONVERT VPR AND SSAM FROM GRAMS TO KILOGRAMS
      VPR=VPR/1000.0                                 LINK1199
C
C   DURING COMPUTATION SSAM CONTAINS THE VALUE OF THE TOTAL MASS OF
C   GAS AND CONDENSED PHASE MATERIAL IN THE CLOUD.
      SSAM=SSAM/1000.0-VPR                          LINK1200
C
C   WRITE INITIAL CONDITIONS RESULTS
      WRITE( ISOUT,4)W,HEIGHT                        LINK1201
      IF(USOIL-1.0)301,301,302                     LINK1202
301  WRITE (ISOUT,5)                               LINK1203
      GO TO 305
302  WRITE (ISOUT,6)                               LINK1204
305  GO TO (309,310,311),IDISTR                 LINK1205
309  WRITE( ISOUT,7)DMEAN,SD
      IF (IS)102,103,102                           LINK1206

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103 WRITE (ISOUT,4) LINK1233
    GO TO 315 LINK1234
102 WRITE (ISOUT,9) LINK1235
    GO TO 315 LTNK1236
  311 WRITE (ISOUT,18) NDSTR LTNK1237
C           PRINT FINAL PARTICLE SIZE CLASS LINK1238
C
  315 WRITE (ISOUT,193) LTNK1239
    DO 602 J=1,NDSTR LINK1240
    J0=J+1
    DM1=DIAM(J0)*1.0E-6
    DM2=DIAM(J)*1.0E-6
  602 WRITE (ISOUT,194) J,PS(J),DM1,FMASS(J),DM2
    GO TO 106 LINK1241
  310 WRITE (ISOUT,197) NDSTR,CAY,EXPO
    GO TO 315 LINK1242
106 WRITE (ISOUT,13) TME,TMP1,TMP2,VPR,SSAM
118 WRITE (ISOUT,192) LTNK1243
  200 CALL SHWIND
    WRITE (ISOUT,15) LINK1244
    RETURN
  143 WRITE (ISOUT,11) LINK1245
    GO TO 190
  150 WRITE (ISOUT,12) LINK1246
  190 CALL EXIT
    END

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SUBROUTINE DSTBN
COMMON /SET1/
1CAY      ,DEFID(12) ,DIAM(201) ,DMEAN      ,DNS      ,FXPO      ,DSTBN    1
2FMASS(200),IDISTR   ,IEXEC      ,IRISE      ,ISIN      ,ISOUT      ,DSTBN    2
3NDSTR    ,PS(200)   ,SD        ,SSAM       ,TME       ,TMPI      ,DSTBN    3
4TMP2     ,T2M       ,USOIL     ,VPR       ,W         ,HEIGHT     ,DSTBN    4
5ZSCL     ,NH000     ,ZV(200)   ,VX(200)   ,VY(200)   ,          ,DSTBN    5
                                         ,          ,          ,          ,          ,DSTBN    6
                                         ,          ,          ,          ,          ,DSTBN    7
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C
C LOGNORMAL DISTRIBUTION TO 100
C POWER FUNCTION DISTRIBUTION TO 200
C TABULAR DISTRIBUTION TO 300
C

C EQUATION 26.2.23 OF NBS-AMS 55 HANDBOOK IS USED TO COMPUTE THE
C PROBABILITY FUNCTION ARGUMENT FROM THE RATIONAL POLYNOMIAL
C APPROXIMATION TO THE NORMAL PROBABILITY FUNCTION.
C TA(X)=SQT(ALOG(1.0/X**2))
C APX(X)=TA(X)-(2.515517+0.802853*TA(X)+0.010323*TA(X)**2)/
C (1.0+1.432799*TA(X)+0.139269*TA(X)**2+0.001308*TA(X)**3)
C LD=NDSTR+1
C GO TO (100,200,300),IDISTR
100 IF(DMEAN)111,111,112
111 DMEAN=0.407
SD=4.0
112 IF(NDSTR-1)101,101,102
101 PS(1)=DMEAN*1.0E-6
C5=SD**5
DIAM(1)=DMEAN*C5
DIAM(2)=DMEAN/C5
FMASS(1)=1.0
GO TO 400
102 BARMU=ALOG(DMEAN)
SIGMA=ALOG(SD)
BARMU=BARMU+3.*SIGMA**2
FRAC=1.0/FLOAT(NDSTR)
DO 103 ND=1,NDSTR
103 FMASS(ND)=FRAC
NH=NDSTR/2
DO 104 I=1,NH
PRR=FLOAT(I)*FRAC
DIAM(I+1)=BARMU+APX(PRR)*SIGMA
J=NDSTR-I+1
104 DIAM(J)=BARMU-APX(PRR)*SIGMA

C
C FOR THE 2 EXTREME INTERVALS THE AVERAGE DIAMETER IS
C ASSUMED TO BE AT HALF A MASS FRACTION FROM ZERO AND ONE
C
PRB=FRAC/2.0
PS(1)=BARMU+APX(PRB)*SIGMA
PS(NDSTR)=BARMU-APX(PRB)*SIGMA
DIAM(1)=2.*PS(1)-DIAM(2)
DIAM(LD)=2.*PS(NDSTR)-DIAM(NDSTR)

C
C CALCULATE MEAN DIAMETERS FROM BOUNDARY VALUES.
C
J=NDSTR-1
IF(J-1)107,107,105
105 DO 106 I=2,J
106 PS(I)=0.5*(DIAM(I)+DIAM(I+1))


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107 DO 108 I=1,NDSTR          DSTBN 59
  DIAM(I)=EXP(DIAM(I))        DSTBN 60
108 PS(I)=EXP(PS(I))*1.0E-6   DSTBN 61
  DIAM(LD)=EXP(DIAM(LC))     DSTBN 62
  GO TO 400                  DSTBN 63
200 IF(EXPO<4.0)201,202,202  DSTBN 64
202 WRITE(ISOOUT,2001)        DSTBN 65
  CALL EXIT                  DSTBN 66
2001 FORMAT('#1#,1X,*EXONENT OF POWER LAW POWER LAW PARTICLE SIZE FREQUENCIES') DSTBN 67
  1ENCY DISTRIBUTION GT. OR EQ. 4.0#)
201 IF(NDSTR>1)203,204,204  DSTBN 68
203 NDSTR=10                 DSTBN 69
204 AN=FLCAT(NDSTR)          DSTBN 70
  FRAC=1.0/AN                DSTBN 71
  DO 205 I=1,NDSTR          DSTBN 72
205 FMASS(I)=FRAC           DSTBN 73
  POW=1.0/(4.0-EXPO)         DSTBN 74
  DMIN=(6.0*SSAM*FRAC/(POW*CAY*DMS*3.14159E6))**POW  DSTBN 75
  DO 206 IJ=1,NDSTR         DSTBN 76
  AJ=FLOAT(IJ)-1.0           DSTBN 77
206 DIAM(IJ)=(AN-AJ)**PCW*DMIN  DSTBN 78
  PS(NDSTR)=DMIN*0.5**POW    DSTBN 79
  DIAM(LD)=PS(NDSTR)**2/DIAM(NDSTR)  DSTBN 80
  ND=NDSTR-1                 DSTBN 81
  DO 207 IJ=1,ND             DSTBN 82
207 PS(IJ)=SQRT(DIAM(IJ)*DIAM(IJ+1))  DSTBN 83
  DO 208 IJ=1,L0             DSTBN 84
208 DIAM(IJ)=1.0E+6*DIAM(IJ)  DSTBN 85
  GO TO 400                  DSTBN 86
300 DO 301 I=1,NDSTR         DSTBN 87
301 PS(I)=0.5*(DIAM(I)+DIAM(I+1))*1.0E-6  DSTBN 88
400 RETURN                  DSTBN 89
  END                         DSTBN 90
                                DSTBN 91

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C
C          SUBROUTINE SHWIND
C
C          READS IN SHOT TIME WIND DATA ABOVE GROUND ZERO
C
C          COMMON /SET1/
C          1CAY      ,DETID(12) ,DIA1(201) ,DMEAN      ,DNS        ,EXPO      ,SHWND 1
C          2FMASS(200),IDISTR   ,IFXEC      ,IRISE      ,ISIN       ,ISOUT     ,SHWND 2
C          3NDSTP    ,PS(200)   ,SD         ,SSAM       ,TME        ,TMP1      ,SHWND 3
C          4TMP2      ,T24        ,USOIL      ,VPR        ,W          ,HEIGHT    ,SHWND 4
C          5ZSCL     ,NHODO     ,ZV(200)    ,VX(200)    ,VY(200)    ,SHWND 5
C          READ(ISIN,1)NHODO
C          IF(NHODO)100,100,200
100 WRITE(ISOUT,5)
C          GO TO 300
200 READ(ISIN,2)(ZV(J),VX(J),VY(J),J=1,NHODO)
C          WRITE(ISOUT,3)NHODO
C          WRITE(ISOUT,4)(ZV(J),VX(J),VY(J),J=1,NHODO)
300 RETURN

C
1 FORMAT(I5)           ,SHWND 19
2 FORMAT(F12.3, 2F12.3) ,SHWND 20
3 FORMAT(1F12.3,1X,1F12.3,1X,1F12.3,1X,1F12.3,1X,1F12.3) ,SHWND 21
4 FORMAT(1F12.3,1X,1F12.3,1X,1F12.3,1X,1F12.3,1X,1F12.3) ,SHWND 22
5 FORMAT(1F12.3,1X,1F12.3,1X,1F12.3,1X,1F12.3,1X,1F12.3) ,SHWND 23
C          END

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