### UNCLASSIFIED

# Defense Technical Information Center Compilation Part Notice

# ADP012525

TITLE: Magnetic Shear Stabilization of Diocotron Instability

DISTRIBUTION: Approved for public release, distribution unlimited

This paper is part of the following report:

TITLE: Non-Neutral Plasma Physics 4. Workshop on Non-Neutral Plasmas [2001] Held in San Diego, California on 30 July-2 August 2001

To order the complete compilation report, use: ADA404831

The component part is provided here to allow users access to individually authored sections of proceedings, annals, symposia, etc. However, the component should be considered within the context of the overall compilation report and not as a stand-alone technical report.

The following component part numbers comprise the compilation report: ADP012489 thru ADP012577

## UNCLASSIFIED

## Magnetic Shear Stabilization of Diocotron Instability

S. Kondoh, T. Tatsuno, and Z. Yoshida

Graduate School of Frontier Sciences, University of Tokyo, 7-3-1 Hongo, Bunkyo-ku, Tokyo 113-0033, Japan

Abstract. The diocotron instability in a magnetized non-neutral plasma is a close cousin of the Kelvin-Helmholtz instability. A sheared magnetic field brings about coupling between the diocotron modes and the Langmuir waves that propagate along the magnetic field. Motion of electrons parallel to the magnetic field cancels the electric charge produced by the diocotron modes, resulting in stabilization of the diocotron instability.

#### I INTRODUCTION

Recently a variety of new concepts on non-neutral plasma confinement has been proposed [1-3], which significantly differ from the conventional Penning/Malmberg trap [4]. The Prototype Ring Trap (Proto-RT) experiment [1,5] is aimed at pure magnetic confinement of a toroidal non-neutral plasma that is not in a rigid-rotating thermal equilibrium state. In such a system, the plasma flow is generally sheared, and the diocotron instability [6] can be destabilized. Application of magnetic shear is expected to be most effective to stabilize the electrostatic modes. However, exact stability analysis has not been completed, except for the special case of an electron beam with a relativistic speed [7].

The physical mechanism of the diocotron instability is explained as follows [8] (see Fig. 1): When a non-neutral slab plasma has a finite thickness, a perturbation on one of the two plasma surfaces produces surface charges. The resulting perturbed electric field yields an  $\boldsymbol{E} \times \boldsymbol{B}$  flow in the plasma, and the opposite surface is also perturbed. The motion of the opposite surface brings about a reciprocal perturbation, and the waves on the two surfaces couple with each other. Under certain conditions, this coupling yields a positive feedback, and the diocotron instability occurs.

When a non-neutral plasma is confined in a uniform magnetic field, the diocotron modes propagating in the perpendicular direction to the magnetic field are independent of any modes that propagate in the parallel direction. However, if the magnetic field has a shear (see Fig. 2), the wave vector may

CP606, Non-Neutral Plasma Physics IV, edited by F. Anderegg et al. © 2002 American Institute of Physics 0-7354-0050-4/02/\$19.00



**FIGURE 1.** Physical picture of diocotron modes in a uniform magnetic field. Perturbation on one of the two plasma surfaces produces the surface charge and causes the electrostatic field perturbation. This perturbed electric field shakes the body of the plasma itself through  $E \ge B$  drift, and the other surface is also perturbed. The perturbation on the latter surface in turn shakes the former one in the same way. Thus, the waves on the two surfaces couple with each other. Under certain conditions, the diocotron modes can be unstable.



FIGURE 2. Physical picture of stabilizing effect of a sheared magnetic field on diocotron modes. The wave vector almost always has a local parallel component, and the diocotron modes cannot be independent of the parallel modes, such as the Langmuir wave or the plasma oscillation. Therefore, a coupling between them is caused. In a cold non-neutral plasma, the surface charge perturbation produced by the diocotron modes is canceled by the parallel motion of charged particles. Thus, the diocotron instability is stabilized by the sheared magnetic field.

have a local parallel component  $k_{\parallel}(x)$ , and the diocotron modes interact with the parallel modes, such as the Langmuir wave or the plasma oscillation. In a cold non-neutral plasma, the surface charge perturbation produced by the diocotron modes is short-circuited by the parallel motion of charged particles, if the diocotron frequency  $\omega_{\rm D}$  is much smaller than the plasma frequency  $\omega_{\rm p}$ , i.e., when a low density plasma embedded in a strong magnetic field. Therefore, we expect that the diocotron instability is stabilized in a sheared magnetic field.

In this paper, we consider a slab plasma with a flat-top density profile

and show the stabilizing effect of magnetic shear analytically. The diocotron instability is formally equivalent to the Kelvin-Helmholtz instabilities in fluids and plasmas [9]. The magnetic shear stabilization of these instabilities is of the common interest and has a variety of applications (see Sec. IV).

### II EIGENEQUATION FOR DIOCOTRON MODES IN A SHEARED MAGNETIC FIELD

#### A Slab plasma model in a sheared magnetic field

We consider a slab electron plasma embedded in a sheared magnetic field (see Fig. 2). The plasma has a finite thickness  $2\Delta$  in the x-direction. We assume that all equilibrium quantities are functions of only x. We consider a sheared magnetic field such as

$$\boldsymbol{B} = (0, B_y(x), B_z), \tag{1}$$

where  $B_z$  is a constant. Since a non-neutral plasma has a self-electric field, there is a stationary flow that is approximately equivalent to the  $E \times B$  drift for low densities.

The governing equations are

$$\frac{\partial n}{\partial t} + \boldsymbol{v} \cdot \nabla n + n \nabla \cdot \boldsymbol{v} = 0, \qquad (2)$$

$$\frac{\partial \boldsymbol{v}}{\partial t} + (\boldsymbol{v} \cdot \nabla)\boldsymbol{v} = -\frac{1}{s^2}(\boldsymbol{E} + \boldsymbol{v} \times \boldsymbol{B}), \qquad (3)$$

$$\nabla^2 \phi = -n, \tag{4}$$

where n is normalized by the typical electron density  $n_0(0)$ , t by the inverse of the diocotron frequency  $\omega_{\rm D}^{-1} = \varepsilon_0 B_z/en_0(0)$  ( $\varepsilon_0$  is the vacuum dielectric constant and e is the elementary electric charge), the spatial coordinates x, y, zby the half thickness of the slab plasma  $\Delta$ , v by the mean flow velocity at the plasma surface  $|v_0(1)|$ , B by the axial magnetic field  $B_z$ , E by the mean electric field at the plasma surface |E(1)|,  $s \equiv \omega_{\rm p}/\omega_{\rm c} = \omega_{\rm D}/\omega_{\rm p}$  is a dimensionless parameter, and  $E = -\nabla \phi$ . Since we consider a low-density plasma in a strong axial magnetic field, we may assume  $s \ll 1$ . In this limit, we can replace Eq. (3) with

$$\boldsymbol{v}_{\perp} = \frac{-\nabla\phi \times \boldsymbol{B}(x)}{B(x)^2},\tag{5}$$

$$\frac{\partial v_{||}}{\partial t} + (\boldsymbol{v} \cdot \nabla) v_{||} = \frac{1}{s^2} \nabla_{||} \phi, \qquad (6)$$

where  $B(x) = \sqrt{B_y(x)^2 + B_z^2}$ . Linearizing Eqs. (2), (5), and (6), we finally obtain [10]

$$\left(\frac{d^2\phi_1}{dx^2} - k^2\phi_1\right) + \frac{1}{\omega - k_{\perp}v_{\perp 0}}\frac{d}{dx}\left(\frac{k_{\perp}n_0}{B}\right)\phi_1 + \frac{n_0k_{\parallel}^2}{s^2(\omega - k_{\perp}v_{\perp 0})^2}\phi_1 = 0, \quad (7)$$

where we have Fourier-transformed all perturbed variables  $\Psi(x, y, z, t)$  as

$$\Psi(x, y, z, t) = \Psi_0(x) + \Psi_1(x) \exp[i(\omega t - k_y y - k_z z)].$$
(8)

### III STABILIZING EFFECT DUE TO PARALLEL MOTION

#### **A** Non-resonant frequency regime

First, we show that the diocotron modes are stabilized ( $\omega_i = 0$ ) for wave numbers without resonance between the phase velocity and the plasma flow, that is,  $\omega_{\rm r} - k_{\perp} v_{\perp 0} \neq 0$  for all x. Multiplying Eq. (7) by  $\phi^*$  and integrating it over  $(-\infty, \infty)$ , we obtain from the imaginary part

$$\omega_{\rm i} \int_{-\infty}^{\infty} \left[ \frac{1}{|\omega - k_{\perp} v_{\perp 0}|} \frac{d}{dx} \left( \frac{k_{\perp} n_0}{B} \right) + \frac{2n_0 k_{\parallel}^2 (\omega_{\rm r} - k_{\perp} v_{\perp 0})}{s^2 |\omega - k_{\perp} v_{\perp 0}|^4} \right] |\phi_1|^2 dx = 0.$$
(9)

Here we used the boundary condition

$$\phi_1(\pm\infty) = 0. \tag{10}$$

Since  $s^2 \ll 1$  and  $\omega_{\Gamma} - k_{\perp} v_{\perp 0} \neq 0$  at any point in the plasma region, we obtain  $\omega_{i} = 0$ , which means stability. This mathematical treatment is same as the standard Rayleigh's analysis [11].

#### **B** Dispersion relation with resonances

If the plasma has a resonant point, the analysis in Sec. III A does not apply to check whether the eigenvalues  $\omega$  for Eq. (7) are real or not. In this case, we need to solve Eq. (7) directly. The eigenfunction determined by Eq. (7) is oscillatory, because the sign of the last term, which we assumed to be very large ( $s \ll 1$ ), is positive. If  $\omega$  is not real, the real and imaginary parts of the eigenfunction have a relative phase angle of about  $\pi/2$ . When we consider a density profile with a sharp boundary, we have to connect both the real and imaginary parts of the eigenfunction at the plasma surfaces using the same boundary condiction. If both of them have a different phase angle, this process fails, which implies that  $\omega$  must be real. The essential characteristic of this eigenvalue problem is well understood by the following simplified model. First, we neglect the second term  $k^2\phi_1$  in the bracket of the first term of Eq. (7) in the plasma region (-1, 1), since it is much smaller than the last term when  $n_0 \simeq 1$ . We also assume that  $k_{\perp}n_0/B$ jumps at  $x = \pm 1$  and its variation is negligible anywhere else, i.e.

$$\frac{d}{dx}\left(\frac{k_{\perp}n_0}{B}\right) = f(x)\left[\delta(x+1) - \delta(x-1)\right],\tag{11}$$

where f(x) is a given finite function. Furthermore, we assume

$$\frac{n_0(x)k_{\parallel}(x)^2}{s^2} \equiv a^2 = \text{const.} \gg 1,$$
(12)

$$k_{\perp}(x)v_{\perp 0}(x) = x.$$
(13)

Under these assumptions, we can solve Eq. (7) analytically, and the dispersion relation is given by [10]

$$\left[ (a+k\omega_{i})^{2}-k^{2}\omega_{r}^{2}+\left(k+\frac{1}{2}\right)^{2}+f(-1)f(1)-\left(k+\frac{1}{2}\right)(f(-1)+f(1))\right. \\ \left.+k\omega_{r}(f(-1)-f(1))+i(a+k\omega_{i})(f(-1)-f(1)-2k\omega_{r})\right](\omega+1)^{2ai} \\ = \left[ (a-k\omega_{i})^{2}-k^{2}\omega_{r}^{2}+\left(k+\frac{1}{2}\right)^{2}+f(-1)f(1)-\left(k+\frac{1}{2}\right)(f(-1)+f(1))\right. \\ \left.+k\omega_{r}(f(-1)-f(1))-i(a-k\omega_{i})(f(-1)-f(1)-2k\omega_{r})\right](\omega-1)^{2ai}, (14)$$

where  $\omega = \omega_r + i\omega_i$ . We can show that  $\omega$  is real for Eq. (15). where  $\omega_r = \text{Re } \omega$ and  $\omega_i = \text{Im } \omega$ . Taking the absolute number of Eq. (15) gives

$$\frac{|A_1|}{|A_2|} = \exp\left[a \arg\left(\frac{\omega+1}{\omega-1}\right)\right],\tag{15}$$

where

$$A_{1} = (a + k\omega_{i})^{2} - k^{2}\omega_{r}^{2} + \left(k + \frac{1}{2}\right)^{2} + f(-1)f(1) - \left(k + \frac{1}{2}\right)(f(-1) + f(1)) + k\omega_{r}(f(-1) - f(1)) + i(a + k\omega_{i})(f(-1) - f(1) - 2k\omega_{r}),$$
(16)

$$A_{2} = (a - k\omega_{i})^{2} - k^{2}\omega_{r}^{2} + \left(k + \frac{1}{2}\right)^{2} + f(-1)f(1) - \left(k + \frac{1}{2}\right)(f(-1) + f(1)) + k\omega_{r}(f(-1) - f(1)) - i(a - k\omega_{i})(f(-1) - f(1) - 2k\omega_{r}).$$
(17)

If  $\omega_i > 0$ , the left-hand side of Eq. (16) is greater than unity, while the righthand side is less than unity. Therefore,  $\omega_i > 0$  cannot be satisfied. If  $\omega_i < 0$ , the left-hand side of Eq. (16) is less than unity, while the right-hand side is greater than unity. Therefore,  $\omega_i < 0$  cannot be satisfied. Thus  $\omega_i = 0$ , which means stability. If  $\omega_i \neq 0$ , the eigenfunctions of the three regions  $\phi_I$ ,  $\phi_{II}$ , and  $\phi_{III}$  cannot be connected properly at  $x = \pm 1$ .

#### IV SUMMARY

We have shown that the magnetic shear has a strong stabilizing effect on the diocotron instability. The fluid motion parallel to the magnetic field shortcircuits the charge perturbation of the diocotron modes. The scaling parameter is  $s \equiv \omega_D/\omega_p$ . Since the time scale of the parallel motion is  $\sim \omega_p^{-1}$ , the condition  $s \ll 1$  enables the parallel motion of the plasma to cancel the perturbed charge sufficiently. Typical non-neutral plasmas in laboratories satisfy this condition.

Mathematically the last term of the eigenequation (7) prohibits non-real eigenvalues, because the last term makes the eigenfunction oscillatory. If  $\omega \notin \mathcal{R}$ , the relative phase angle between the real and imaginary parts of the eigenfunction is about  $\pi/2$ . This phase angle disables both the real and imaginary parts of the eigenfunction to be connected properly at the plasma surfaces.

We note that our analysis is based on a modal approach which, however, may not be complete for non-Hermitian systems [12–14]. There remains a possibility of secular algebraic behavior, although we have shown that there are no exponentially unstable modes. This problem will be discussed elsewhere.

#### REFERENCES

- Z. Yoshida, Y. Ogawa, J. Morikawa, et al., Non-Neutral Plasma Physics III, AIP Conference Proceedings 498 (American Institute of Physics, Princeton, 1999) p.397.
- 2. S. Robertson and B. Walch, Phys. Plasmas 7, 2340 (2000).
- 3. K. Avinash, Phys. Plasmas 1, 3731 (1994).
- 4. J. H. Malmberg and J. S. deGrassie, Phys. Rev. Lett. 35, 577 (1975).
- 5. S. Kondoh and Z. Yoshida, Nucl. Instrum. Meth. Phys. Res. A 382, 561 (1996)
- R. C. Davidson, *Physics of Nonneutral Plasmas* (Addison-Wesley, California, 1990).
- 7. E. Ott and J. M. Wersinger, Phys. Fluids 23, 324 (1980).
- 8. W. Knauer, J. Appl. Phys. 37, 602 (1966).
- 9. C. F. Driscoll and K. S. Fine, Phys. Fluids B 2, 1359 (1990).
- 10. S. Kondoh, T. Tatsuno, and Z. Yoshida, Phys. Plasmas 8, 2635 (2001).
- 11. S. Chandrasekhar, Hydrodynamic and Hydromagnetic Stability (Clarendon Press, Oxford, 1961).
- L. N. Trefethen, A. E. Trefethen, S. C. Reddy, and T. A. Driscoll, Science 261, 578 (1993).
- 13. G. D. Chagelishvili, A. D. Rogava, and I. N. Segal, Phys. Rev. E 50, R4283 (1994).
- 14. S. M. Mahajan and A. D. Rogava, Astrophys. J. 518, 814 (1999).