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# PROTECTION FROM PENETRATING RADIATION OF A NUCLEAR EXPLOSION By

V. I. KUKHTEVICH, I. V. GORYACHEV, ET AL. - USSR -



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## PROTECTION FROM PENETRATING RADIATION OF A NUCLEAR EXPLOSION

Translation of the Russian-language book by V. I. Kukhtevich, I. V. Goryachev, and L. A. Trykov, edited by Doctor of Physical and Mathematical Sciences V. I. Kukhtevich: Zashchita ot <u>Pronikayushchey Radiatsii Yadernogo Vzryva</u>, Moscow, Atomizdat Press, 1970, 190 pages.

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### FOREWORD

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During the quarter century since the creation of the first atomic bomb, a tremendous number of nuclear and thermonuclear warheads of various sizes have been accummulated in the world. Various transport systems have been created, allowing delivery of nuclear warheads to any point on the earth. Although an agreement was signed in 1963, forbidding the testing of nuclear weapons in the atmosphere, in space and underwater, some countries still continue to perform nuclear weapons tests. American strategic aircraft fly regularly with nuclear bombs onboard.

Some countries, desiring to develop their own nuclear weapons, are delaying in signing the treaty.

Therefore, the problem of protection of the population from the effects of the radiation of nuclear explosions is a problem of prime importance.

The creation of shielding effective against the radiation of nuclear explosions is a difficult and interesting scientific task. Although in principle the design of shielding from the radiation of nuclear explosions is no different from the design of the biological shielding of nuclear reactors, it has a number of peculiarities resulting from the specific geometric conditions and quality of the radiation involved.

Over the past few years, books have been published relating to various aspects of the phenomena accompanying nuclear explosions. The monographs of O. I. Leypunskiy <u>Gamma-Izlucheniye Atomnogo Vzryva</u> [The Gamma Radiation of a Nuclear Explosion] (Atomizdat Press, 1959) and P. A. Yampol'skiy <u>Neytrony Atomnogo Vzryva</u> [The Neutrons of an Atomic Explosion] (Gosatomizdat Press, 1961) are dedicated to description of the physical processes involved in the propagation of gamma radiation and neutrons in a nuclear explosion; the spatial distribution of the doses of neutrons and  $\gamma$  radiation during nuclear explosions are

analyzed. Problems of shielding from radiation are hardly discussed in these books.

The book <u>Deystviye Yadernogo Oruzhiya</u> [The Effects of Nuclear Weapons] (Translated from the English, Voyenizdat Press, 1963) presents a general description of all aspects of nuclear explosions and shielding against damaging effects. However, this book is descriptive in nature and is more an introduction to the problem than a practical guide.

The book of R. V. Petrov et al. Zashchita Ot Radioaktivnykh Osadkov [Protection From Radioactive Fallout] (Medgiz Press, 1963) and that of L. V. Spenser Zashchita Ot Gamma-Izlucheniya Radioaktivnykh Vypadeniy [Protection From the Gamma Radiation of Radioactive Fallout] (Translated from the English, Atomizdat Press, 1965) contain valuable material on protection from the radiation of radioactive fallout. The book of Petrov analyzes the general radiation situation arising when radioactive fallout occurs following an explosion. The book of Spenser presents an approximate method of designing shielding against  $\gamma$  radiation, corresponding to the specific geometric conditions which obtain when areas and structures are contaminated with radioactive fallout.

Thus, as yet no one has discussed the design of shielding from the primary radiation of a nuclear blast, and only one book has presented a method for approximate design of shielding of structures from the radiation of radioactive fallout.

However, many publications have appeared in the periodical literature in the past six or seven years concerning both clarification of the propagation of penetrating radiation in the atmosphere and its transmission through shielding. These data have not yet been systemitized in order to allow their use for the design of shielding from the radiation of nuclear explosions.

The use of these data is important, since correct design requires detailed knowledge of the characteristics of the field of ionizing radiation.

Actually, the dose within a protective structure can be calculated if we know the spectral and angular distributions of radiation on the outer surfaces of its walls. In a nuclear explosion, the field of radiation consists of radiation scattered in the air and radiation reflected from the earth, with complex spectral and angular characteristics. Equally detailed information is required on the radiation passing through the shielding. In order to prevent this information in a small book, along with characteristics of the attenuation of radiation in shielding and a calculation method, the authors have omitted general definitions and physical interpretation of certain phenomena. The authors assume that the basic concepts of the physics of shielding have been sufficiently fully covered in certain monographs, for example B. Price, K. Horton and K. Spinney Zashchita Ot Yadernykh Izlucheniy [Shielding Against Nuclear Radiation] (Translated from the English, Foreign Literature Press, 1959); G. Goldstein Osnovy Zashchity Reaktorov [Principles of Reactor Shielding] (Gosatomizdat Press, 1961); and Yu. A. Kazanskiy et al. Fizicheskiye Issledovaniya Zashchity Reaktorov [Physical Studies of Reactor Shielding] (Atomizdat Press, 1966).

The principle of the arrangement of the book should be clear from the Table of Contents and requires no explanation.

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Keeping in mind the fact that the planning of shielding structures involves both specialists in the area of nuclear equipment and construction engineers, practical calculations are presented in two versions. A simplified version is presented for an approximate solution, requiring no computer equipment. Precise design of complex structures requires the use of electronic computers. In these cases, calculation methods are presented in this book.

The authors hope that the book will be useful to a broad range of scientific workers and engineers working in the area of radiation protection, as well as persons directly involved in the development of protective structures. CHAPTER 1. A NUCLEAR EXPLOSION AS A SOURCE OF IONIZING RADIATION

§ 1. Brief Description of Phenomena Accompanying a Nuclear Explosion

The explosion of a nuclear warhead is accompanied by liberation of a tremendous quantity of energy, concentrated in a small volume, over an extremely brief period of time. The colossal energy density causes heating of the material in the area of the explosion to temperatures on the order of tens of millions of degrees.

Matter heated to these high temperatures emits energy in the form of electromagnetic radiation over an extremely broad spectrum -- from x-rays to long wave electromagnetic energy.

A large portion of this energy is absorbed by the outer layers of the weapon. These layers, being heated, begin to glow and transmit energy outward. The energy radiated at these high temperatures is absorbed in the air, heating the layers of air adjacent to the bomb. The area of air surrounding the weapon begins to glow, forming the socalled atomic explosion fireball. As the air is increased in temperature, its pressure increases as well. The fireball grows very rapidly due to radiant transfer of energy.

As the fireball grows, the temperature of its outer areas decreases. Whereas at high temperatures the increase in size of the fireball results from radiant heat conductivity, during later stages the transfer of energy is achieved by displacement of the air mass from the area of increased density outward.

In the initial stage following the blast, the leading edge of the shock wave (area of sharp pressure, density and temperature gradient) corresponds with the boundary of the fireball. As the temperature drops, the moment arrives when the shock wave separates from the boundary of the fireball and moves further away from the blast center, while the fireball, expanding slowly, continues to glow for several more tenths of a second. After passage of the shock wave, the pressure behind the wave equalizes; therefore, in the high temperature area adjacent to the blast center, the density of the air is less than atmospheric: an area of low density is formed. This area, due to its low density, rises upward.

The radioactive cloud containing the remains of the bomb rises upward, but soon the heavy particles begin to drop out of the cloud and fall to the ground along the path of movement of the cloud, forming a radioactive wake.

The most important factors in the damaging effects of a nuclear blast are the shock wave, light radiation and ionizing radiation. The quantity of energy in these damaging forms extending for various distances from the center of the blast depends on the type of nuclear warhead and the medium in which the blast occurs. If the blast occurs in the atmosphere at an altitude of less than 30 km, 35 % of the total blast energy is expended in the form of light radiation. Under these conditions, about 50% of the energy of nuclear fission is expended in the formation of the shock wave. The remaining 15% is liberated in the form of ionizing radiation. Of this radiation, about 5% is the initial radiation, acting during the first one minute following the moment of the blast, while 10% of the total energy of the nuclear explosion makes up the residual radiation, which appears over an extended period of time. The residual radiation results almost entirely from radioactive fission products. In a thermonuclear explosion, in which only about one-half of the total explosive energy is liberated as a result of the fission reaction, residual radiation amounts to only 5% of the energy liberated by the blast.

The portion of the energy of ionizing radiation is a small fraction of the total energy liberated in a nuclear blast; however, due to the great penetrating capacity and difficulty of detecting ionizing radiation, this radiation may have a damaging effect even at great distances from the radiation source and even in shelters. Therefore, ionizing radiation is an important damaging factor of a nuclear weapon.

Depending on the nature of the object and the purpose of the atomic attack, a nuclear blast may be performed in the air, at the surface of the ground (water) or underground (underwater). Therefore, atmospheric, surface and underground (underwater) explosions are differentiated.

Usually, an atmospheric nuclear explosion is used to destroy city or industrial buildings, to damage personnel and equipment on the field of battle. Under these conditions, it is set off at an altitude of several hundreds or thousands of meters over the ground, depending on the power of the blast. A surface blast is designed to destroy strong surface structures, railroad junctions, heavy shelters, airfields, and to damage surface ships. It may be performed at an altitude of a few tens of meters over the surface or directly at the surface (contact explosion).

An underground nuclear blast may be used to damage especially strong underground structures. In certain cases, a surface or underground explosion may be used for radioactive contamination of the terrain in the enemy rear area as well.

The external picture of the blast, as well as the effects of its damaging factors, depend on the position of the blast center in relationship to the surface of the earth. During an atmospheric blast, the fireball does not touch the surface of the earth at the moment of its maximum brightness, and its dimensions depend on the power of the blast. The maximum radii of the fireball for atmospheric and surface explosions of the same power do not differ very greatly. The mean radius of the fireball can be described approximately by the following expression:

$$r_{fb} = 33.6q^{0.4}m,$$
 (1.1)

where q is the power of the explosion, kt.

The maximum size of the fireball is not immediately achieved, although the rate of growth of the fireball is rather great, so that it can be considered that it does not rise to any significant height above the point of the explosion during the time of its formation. Therefore, the height of a blast at which the fireball will not contact the surface of the earth is determined by its radius [see (1.1)].

If the distance between the blast center and surface of the earth (water) is not over  $r_{fb}$ , the external picture of the explosion is essentially the same as for an explosion occurring directly on the surface. Therefore, blasts occurring both directly on the surface and at a slight altitude above the surface are both considered surface blasts. With increasing altitude of the explosion, an intermediate position develops, in which the fireball does not contact the surface of the earth, but the column of dust and soil rising from the earth catches up with the rising cloud of fission products and is connected to it. This type of explosion is looked upon as a low altitude atmospheric blast.

During the period when the fireball glows, its temperature is so high that, at least in its interior, all of the matter of which the nuclear weapon is made is in the vapor state. This matter includes the radioactive fission products, uranium (or plutonium) which has not fissioned, the envelope around the charge and other parts of the nuclear weapon.

At the moment of an explosion, a mass of gas flies outward from the explosion center. A hollow sphere is formed in the center, the dimensions of which depend on the power of the explosion. After the departure of the shock wave, the pressure within the fireball returns to normal, but the density of the gas becomes near zero, since the temperature within the fireball is extremely high.

It is this significant difference between the densities within and without the fireball which leads to the rise of the radioactive cloud.

### § 2. The Neutrons of a Nuclear Explosion

During a nuclear explosion, the primary source of neutrons consists of the prompt fission neutrons emitted in the first few tenths of a microsecond.

Upon fission of nuclei of  $U^{235}$  or other fissionable material, the energy spectrum of prompt neutrons can be described by the following formula [1]:

$$\Phi_0(E_n) = \sqrt{E_n} e^{-\frac{3}{2} \frac{E_n}{\epsilon}}.$$
 (1.2)

In this expression, parameter  $\varepsilon$  depends on the energy of excitation of the fission fragments which is related in turn to the mean number of neutrons v per fission event. This relationships is expressed by the following dependence:

$$\varepsilon \approx 0.78 + 0.621 (v + 1)^{1/3} \text{ MeV}.$$
 (1.3)

The value of v for fission by thermal neutrons is well known for most fissionable nuclei:

$$U^{235} - v = 2,47 \pm 0,03;$$
  

$$U^{233} - v = 2,51 \pm 0,03;$$
  

$$P_{11}^{239} - v = 2,90 \pm 0,04.$$

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This indicates that when  $U^{233}$  is split by thermal neutrons, the value of  $\varepsilon$  is 0.02 MeV higher than for the fission of  $U^{235}$ ; for  $Pu^{239}$ , the increase is 0.07 MeV. Thus, for thermal neutron fission, the spectrum of  $Pu^{239}$  contains 30% more neutrons with energies of 10 MeV than the spectrum accompanying fission of  $U^{235}$ .

It is known that v increases approximately linearly with the energy of the neutrons causing fission, and that the increase for  $U^{235}$  is 0.15 per 1 MeV. This change in v, together with the relationship (1.3) between  $\varepsilon$  and v presented above indicates that the mean energy of fission neutrons for  $U^{235}$  increases by 0.04 MeV as the energy of the neutrons causing fission is increased by 1 MeV. This slight change in the energetic characteristics of the fission neutrons with a change in the energy of neutrons causing fission or the energy of the nuclei undergoing fission can be ignored.

During a thermonuclear explosion based on reactions with lithium deuteride and the products of this reaction, the following main reactions occur [2]:

 $D + D \rightarrow He^{3} + n + 3,25 \text{ MeV};$   $D + D \rightarrow T + p + 4,0 \text{ MeV};$   $n + Li^{6} \rightarrow He^{4} + T + 4,8 \text{ MeV};$  $T + D \rightarrow He^{4} + n + 17,6 \text{ MeV}.$ 

The probability of the last two reactions is approximately 100 times higher than the probability of the first two reactions; therefore, explosions of this type characteristically emit neutrons with energies of around 14 MeV. These neutrons interact with the structural elements of the bomb and the unreacted nuclear fuel. The higher energy of neutrons in this case results in the nuclear reaction (n, 2n). Furthermore, the reaction of inelastic scattering of neutrons on the heavy nuclei of the structural elements of the weapon is quite probable, resulting in the formation of neutrons with most probable energy around 4 MeV [3].

The spectrum of neutrons formed upon explosion of a bomb based on the fission-synthesis-fission principle, ignoring the moderation of neutrons within the material of the bomb, should be approximately as shown on Figure 1.1. This hypothetical spectrum consists of three main components: fission neutrons with most probable energy around 800 KeV (curve 1), neutrons developing as a result of inelastic scattering of 14 MeV neutrons, with most probable energy about 4 MeV (curve 2) and the 14 MeV neutrons (curve 3). The relationship between the various groups of neutrons can be significantly changed, depending on the share of power in the energy balance of the explosion attributed to fission and to thermonuclear reactions, as well as the design of the bomb.



Figure 1.1. Probable Spectrum of Neutrons Formed Upon Explosion of a Weapon Operating According to the Fission-Synthesis-Fission Principle (Ignoring Moderation in the Shell):

1, fission neutrons; 2, neutrons resulting from inelastic scattering of neutrons with energy 14 MeV; 3, neutrons with energy 14 MeV developing in the synthesis reaction

The actual spectrum of neutrons formed by explosion of a bomb operating according to the fission-synthesis-fission principle will be softer and smoother, as a result of moderation of neutrons in the materials and structure of the warhead.

The distribution of neutrons leaving the shell of a nuclear weapon by energies is determined by the structural characteristics of the weapon, i. e. the quantities characterizing the permeability of the outer portions of the weapon. The problem of transformation of the spectrum of fission neutrons as they penetrate through the shell of a weapon was discussed in detail in [4]. Calculation of the spectrum of neutrons (density  $2 \text{ g/cm}^3$ ) is presented in [4] and its results are shown on Figure 1.2.



Figure 1.2. Spectrum of Neutrons Formed . Upon Explosion of Weapon Based on Fission-Synthesis-Fission Principle:

1, spectrum of fission neutrons; 2, spectrum
of neutrons leaving shell with radius of 25
cm; 3, spectrum of neutrons leaving shell
with radius 50 cm

As we can see, the spectrum of neutrons leaving the shell is quite different from the spectrum of the primary fission neutrons. The spectrum of the scattered neutrons includes a significant fraction of neutrons at low energies. Figure 1.2 also indicates that the forms of the spectra of neutrons leaving shells of various thicknesses correspond, differing only in the absolute quantity of neutrons emitted during the process of moderation.

In addition to the fission neutrons, a small quantity (less than 1%) of so-called delayed neutrons is formed in the chain reaction of fission of uranium or plutonium; these neutrons are emitted from the strongly excited nuclei with excess neutrons. Table 1.1 shows the basic data characterizing the delayed neutrons which develop upon fission of uranium and plutonium [5]. The yields of neutrons for each group are presented in the form of the fraction of the total number of delayed neutrons represented by the delayed neutrons emitted in each group.

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d	Life,	Energy,	Share o	of Delay	ved Neu	trons,	$a_i / \Sigma a_i$	
Grou	Half sec	Mean KeV	Th <sup>232</sup>	U111	Uz 3 2	U238	Pu <sup>za</sup>	Pu:40
1 2 3 4 5 6	54,5 21,8 6,0 2,2 0,50 0,18	250 460 405 450 520	0,034 0,150 0,155 0,446 0,172 0,043	0,086 0,274 0,228 0,316 0,073 0,023	0,038 0,213 0,188 0,407 0,128 0,026	0,013 0,137 0,162 0,388 0,225 0,075	0,038 0,280 0,216 0,328 0,103 0,035	0,025 0,270 0,184 0,359 0,135 0,027
Tota per even	l yié fissi t	eld .on	0,0496 ±0,0020	0,0070 ±0,0004	0,0165 ±0,0005	0,0412 ±0,0017	0,0063 ±0,0003	0,0088 ±0,0006
Tota per emit fiss	l yie neuti ted c ion	eld ron luring		0,0029	0,0067	- ,	0,0021	

### DELAYED NEUTRONS EMITTED DURING FISSION

For the fissionable elements  $U^{235}$  and  $Pu^{239}$ , the half life periods of delayed neutrons correspond within the measurement errors, since the delayed neutrons are formed from the same fragments. The yields of individual groups of delayed neutrons differ, since the curve of the distribution of fragments is different for the different isotopes.

The total yield of delayed neutrons increases with increasing mass number of fissionable material. The greater yield of delayed neutrons in  $U^{238}$  and  $Th^{232}$  apparently results from the presence of longer decay chains.

The mean energy of delayed neutrons is on the order of 0.5 MeV or less (significantly lower than the energy of prompt fission neutrons). However, we must also note for further analysis that in contrast to prompt neutrons, the delayed neutrons are formed after the shell has been blown away and therefore do not experience its moderating influence as they fly outward.

In addition to the sources of neutrons during fission and synthesis which we have mentioned, two more should be noted: photoneutrons and activation neutrons.

Photoneutrons are formed by the reaction  $(\gamma - n)$ , when the energy of photons exceeds the binding energy of a neutron. For most nuclei, the binding energy is greater than 7 MeV. Significant formation of photoneutrons can be expected only in the nuclei D, Be<sup>9</sup>, C<sup>13</sup> and Li<sup>6</sup>, the threshold energies of photons of which are 2.23, 1.67, 4.9 and 5.3 MeV respectively. The values of photoneutron cross sections can be found in [6].

The only reaction in which activation neutrons can arise is the reaction  $0^{17}$  (n, p)N<sup>17</sup>. The nucleus of the isotope N<sup>17</sup> decays with a half life of 4.14 sec, forming the isotope  $0^{17}$ . Most of the neutrons emitted by  $0^{17}$  have energies of 1±0.2 MeV.

The reaction threshold of  $0^{17}(n, p)N^{17}$  is about 8.0 MeV, and its cross section for neutrons in the fission spectrum is on the order of 3.2 µbarn. Therefore, the contribution of this reaction to the total yield of delayed neutrons is apparently quite slight and will be ignored in this work.

Let us now determine the total number of neutrons leaving a nuclear weapon at the moment of the explosion.

As we know, upon fission of nuclei of the material used as nuclear fuel  $(U^{235} \text{ or } Pu^{239})$ , each fission event liberates approximately  $2.8 \cdot 10^{-4}$  erg of energy. As we know, detonation of one g TNT liberates an energy of  $4.18 \cdot 10^{10}$  erg (1000 cal); this indicates that a nuclear explosion acccompanied by the liberation of energy equivalent to the energy of detonation of 1 Kt of TNT involves N  $\approx 1.5 \cdot 10^{23}$  fission events. If we consider the number of neutrons required to maintain the chain reaction, the fission of N nuclei liberates N( $\nu$  - 1) free neutrons. For example, for a plutonium bomb, assuming  $\nu \approx 3$ , the number of neutrons liberated upon explosion of a weapon with a power of 1 Kt will be approximately  $3 \cdot 10^{23}$ .

Thus, the number of neutrons liberated in the fission chain reaction is determined by the power of the explosion, i. e. the quantity of fissioning material; however, the number of neutrons flying outward depends also on the design of the weapon. The number of neutrons liberated upon explosion of a purely thermonuclear bomb corresponds approximately with the number of neutrons liberated upon explosion of a bomb of the same power, based on the fission reaction.

When a three-phase bomb is exploded, the thermonuclear reaction is used as an intermediate process, which initiates high energy neutrons, causing fission of  $U^{238}$  atoms included in the composition of the shell. Therefore, a significant portion of the energy of the explosion of this type of bomb is the energy of fission of the uranium nuclei. As studies performed in [7] have shown, the mean number of neutrons v produced by fission of  $U^{238}$  by neutrons with an energy of 14 MeV is  $4.5\pm0.32$  per fission event. Knowledge of the value of v for  $U^{238}$  allows us to estimate the number of neutrons liberated upon explosion of a three-phase thermonuclear bomb. It corresponds approximately to the yield of neutrons by a weapon operating according to the fission principle and is approximately  $(2-2.5)\cdot10^{26}$  n/Mt.

### § 3. γ Radiation of a Nuclear Explosion

The  $\gamma$  radiation observed upon explosion of a nuclear weapon consists of many components differing in their origin and the time of radiation, energy of quanta and radiated energy, all related to the origin.

The  $\gamma$  radiation formed in a nuclear explosion can be subdivided into primary and secondary as to its origin. Primary  $\gamma$  radiation is emitted in each fission event with the products (fragments). Secondary  $\gamma$  radiation includes the  $\gamma$  quanta developing upon interaction of neutrons with the materials of the charge and the matter making up the surrounding medium (such as air or soil).

As to time,  $\gamma$  radiation from a nuclear explosion is divided into prompt and delayed. Prompt (short-lived)  $\gamma$  radiation includes that emitted by the material of the bomb before it evaporates and flies outward (about 10<sup>-5</sup> sec). In this classification, prompt  $\gamma$  radiation includes: the  $\gamma$  quanta produced directly in the process of fission and with slight delays, a portion of the radiation from short-lived fission products, the photons formed upon capture and inelastic scattering of neutrons by the materials of the nuclear charge. The delayed  $\gamma$  radiation results from short-lived (with half life

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>10<sup>-5</sup> sec) and long-lived decay products, and capture of moderated neutrons in the surrounding medium. Let us study the characteristics of these types of radiation in greater detail.

Figure 1.3 shows the energy spectrum of prompt fission  $\gamma$  quanta [8]. In the range from 1 to 7 MeV, this spectrum can be represented by the simple exponential rule  $\Phi(E_{\gamma}) = 8.0 \cdot e^{-1.1E\gamma} \text{ MeV}^{-1}$  with an accuracy of  $\pm 40\%$ , while in the area of energies from 1 to 4.5 MeV, the error is less,  $\pm 15\%$ . Below energies of 1 MeV the experimental data are higher than what would be indicated by this formula; in this area, the formula  $26.8e^{-2.30E\gamma}$  provides a better fit. However, the  $\gamma$  quanta at low energies are easily absorbed by the shell of the weapon and are therefore insignificant radiation sources.

It was experimentally established in [9] that the form of the  $\gamma$  spectrum accompanying the fission of nuclei of  $U^{2\,33}$  and  $Pu^{2\,39}$  does not differ significantly from the  $\gamma$  spectrum formed upon fission of ,,235 As concerns the total yield of energy of prompt  $\gamma$  quanta upon fission, in the interval of energies from 0.3 to 10 MeV it is 7.2±0.8 MeV, and in this area of energies 7.4±0.8 quanta are emitted per fission event [7]. In other words, the mean energy of instantaneous  $\gamma$  quanta from fission is approximately 1 MeV. The data presented relate to the case of fission of nuclear fuel by thermal neutrons. In [10, 11], the dependence of the total energy of  $\gamma$  radiation of fission on the energy of excitation of fissionable nuclei is studied. The comparison of the spectrum and number of  $\gamma$  quanta upon fission of  $U^{235}$  by thermal neutrons with energies of 2.8 and 14.7 MeV performed in these works showed that the energy of  $\gamma$  radiation and the yield of  $\gamma$  quanta are retained within limits of 15%. For comparison, Figure 1.3 shows data characterizing the energy composition of  $\gamma$  radiation arising in  $U^{2\,38}$  under the influence of neutrons with an energy of As we can see, the spectrum is almost the same as the spec-14 MeV. trum of  $\gamma$  radiation developing upon fission of  $U^{235}$  by thermal

neutrons. The instantaneous  $\gamma$  radiation is significantly attenuated in the materials of which the nuclear charge is made, since its time of luminescence is approximately 100 times less than the time required for the bomb to burst. The degree of attenuation of the prompt  $\gamma$ radiation within the weapon can be estimated by the following method.

Let us analyze a nuclear charge in the form of a sphere of radius  $r_0$  at the moment of the explosion. In order to simplify our calculations, we will make the following assumptions: the sources of  $\gamma$  quanta (i. e. the fissioning nuclei) are distributed evenly throughout the entire volume of the charge; the instantaneous  $\gamma$  radiation is monoenergetic; the absorption of  $\gamma$  radiation in the material of the charge occurs exponentially. These assumptions are apparently physically justified in this analysis.



Figure 1.3. Spectrum of  $\gamma$  Radiation Arising in U<sup>238</sup> Under the Influence of Neutrons With 14 Mev Energy (Curve 1), and Energy Spectrum of Prompt  $\gamma$  Quanta Emitted Upon Fission of U<sup>235</sup> or Pu<sup>239</sup> by Thermal Neutrons (Curve 2)

We would like to know the number of times by which the dose of  $\gamma$  radiation at some distance from the weapon is decreased due to selfabsorption in the material of the charge in comparison to the case when a source of the same total intensity is concentrated at the center of the charge.

Let us look at Figure 1.4. Distance x, over which absorption of  $\gamma$  radiation from an elementary radiating volume occurs

$$dv = r^2 \cos\theta d\theta d\varphi dr, \qquad (1.4)$$

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is determined by the relationship

 $\mathbf{x} = \mathbf{c} - \mathbf{d}.$ 



Figure 1.4. Calculation of Self-Absorption of  $\gamma$  Quanta of Fission in Charge Material

From the equation

$$d^2 - 2db\cos\alpha + b^2 - r_0^2 = 0 \tag{1.5}$$

considering the limits of change of the quantity

$$b - r_0 \leqslant d \leqslant b \cos \alpha \tag{1.6}$$

we find

$$d = b\cos\alpha - \sqrt{r_0^2 - b^2 \sin^2\alpha}.$$
 (1.7)

From the expression

$$r^2 = c^2 + b^2 - 2bc \cos \alpha \tag{1.8}$$

we determine

$$\cos \alpha = \frac{b - r \cos \theta \cos \varphi}{\sqrt{r^2 + b^2 - 2rb \cos \theta \cos \varphi}}.$$

Then

3

$$d = \frac{b^2 - rb\cos\theta\cos\varphi - \sqrt{r_0^2(r^2 + b^2 - 2rb\cos\theta\cos\varphi) - r^2b^2(1 - \cos^2\theta\cos^2\varphi)}}{\sqrt{r^2 + b^2 - 2rb\cos\theta\cos\varphi}};$$
(1.9)

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$$x = \frac{r^{2} - rb\cos\theta\cos\phi}{\sqrt{r^{2} + b^{2} - 2rb\cos\theta\cos\phi}} + \frac{\sqrt{r_{0}^{2}(r^{2} + b^{2} - 2rb\cos\theta\cos\phi) - r^{2}b^{2}(1 - \cos^{2}\theta\cos^{2}\phi)}}{\sqrt{r_{0}^{2}(r^{2} + b^{2} - 2rb\cos\theta\cos\phi) - r^{2}b^{2}(1 - \cos^{2}\theta\cos^{2}\phi)}}.$$
 (1.10)

Introducing the new variable  $m = r/r_0$  and  $p = b/r_0$ , we can write the expression for x in the form

 $\sqrt{r^2 + b^2 - 2rb\cos\theta\cos\phi}$ 

$$x = r_{0} \frac{m^{2} - mp \cos \theta \cos \varphi}{\sqrt{m^{2} + p^{2} - 2mp \cos \theta \cos \varphi}} + (1.11)$$
$$+ r_{0} \frac{\sqrt{m^{2} + p^{2} - 2mp \cos \theta \cos \varphi - m^{2}p^{2} (1 - \cos^{2} i\theta \cos^{2} \varphi)}}{\sqrt{m^{2} + p^{2} - 2mp \cos \theta \cos \varphi}}.$$

Then the desired attenuation of the dose or flux of instantaneous  $\gamma$  radiation due to self-absorption can be calculated as follows:

$$K(p, \mu r_0) = \int_0^1 \int_0^{\pi/2} \int_0^{\pi} \frac{e^{-\mu x} m^2 \cos \theta dm d\theta d\phi}{m^2 + p^2 - 2mp \cos \theta \cos \phi}.$$
 (1.12)

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The results of integration of this expression are shown in Table 1.2.

In this table, the dimensions of the charge of nuclear fuel are represented by the radius of the uranium sphere (in free path lengths -- $\mu r_0$  -- of fission  $\gamma$  quanta). The size of the nuclear charge can be judged on the basis of the critical mass necessary to initiate an explosive fission reaction. Data on the critical masses of fissionable elements have been published in the foreign literature. The most detailed data have been presented for charges consisting of  $U^{235}$ . For example, it is reported in [2] that the critical mass of a spherical charge (93.5%  $U^{235}$  and 6.5%  $U^{238}$ ) with a density of 18.8 g/cm<sup>3</sup> is 48 Kg. The radius of a sphere of this mass is about 8.5 cm. For  $\gamma$  quanta with an energy of 1 MeV, corresponding to the average of the spectrum of prompt  $\gamma$  radiation, expressed in free path lengths this is  $\mu r_0 \approx 12$ . Turning to Table 1.2, we note that with this size nuclear charge, the dose of prompt  $\gamma$  radiation at the surface of the weapon will be attenuated by approximately 100-1000 times due to

and

internal self-absorption<sup>1</sup>.

Therefore, although the number of prompt fission  $\gamma$  quanta and  $\gamma$  radiation of fission products are approximately equal, the contribution of prompt  $\gamma$  radiation during a nuclear explosion can be ignored.

The  $\gamma$  radiation of fission products is emitted mainly during the later stages of the explosion, after the material of the shell of the weapon has evaporated. Therefore, during the first minute following the explosion the quantity of  $\gamma$  radiation of fission products is approximately 100 times greater than the quantity of prompt  $\gamma$  radiation included in the composition of the initial ionizing radiation observed at equal distances from the center of the explosion. The production of fragments is a statistical process. The fission of  $U^{235}$  and  $Pu^{239}$  produces about 60 isotopes. With asymmetrical fission, the most probable reactions involve the formation of fragments with mass numbers 95 and 140. The yield of fragments as a function of mass number varies by a factor of  $10^5$ . It has been established that during fission of uranium and plutonium, fragments are produced with mass numbers from A = 72 to A = 161. The change in mass number of fissionable material from 233 to 238 is accompanied by a significant change in the production of fragments with mass numbers from A = 105 to A = 115. The group of heavy fragments experiences no significant change in yield.

Data on the yields of various fragments formed by fission of  $U^{235}$ ,  $U^{238}$  and  $Pu^{239}$  by fission neutrons, neutrons with energies of 14 MeV and neutrons from a thermonuclear explosion are presented in [12].

Each fragment formed undergoes an average of three  $\beta$  decay events with various periods. Thus, the fission fragments include about 200 radioactive nuclei with half lives from fractions of a second to decades. Many of the decay events are accompanied by  $\gamma$  radiation.

Since nuclei with various half lives participate in the radiation, the rate of decrease of  $\gamma$  activity of radiation with time does not follow a simple exponential rule. The effective half life increases as time passes, since the short-lived isotopes decay and only the longer-lived fragments remain.

<sup>1</sup>In this estimated calculation, the accummulation factor can be ignored, since its magnitude  $\simeq 5$ .

Table 1.2

# ATTENUATION OF PROMPT $\gamma$ RADIATION OF FISSION IN URANIUM, CALCULATED FROM FORMULA (1.12)

ŀ					LT DATU LEN	0TUC	
		RADIUS C	IF UKANIUM		EC LAIN LEN		
	0.4	0.8	2,0	4.0	8.0	10	15
	6,011.10-1 3,923.10-1 2,094.10-1 9,000.10-2 3,190.10-2 7,924.10-3	4,733.10-1 3,082.10-1 1,642.10-1 7,045.10-2 2,495.10-2 6,197.10-3	<b>2</b> ,718.10-1 1,763.10-1 9,366.10-2 4,000.10-2 1,414.10-2 3,511.10-3	1,528.10-1 9,869.10-2 5,255.10-2 2,233.10-2 7,857.10-3 1,947.10-3	8,526-10-2 5,446-10-2 2,922-10-2 1,230-10-2 4,306-10-3 1,054-10-3	7,244.10-2 4,598.10-2 2,474.10-2 1,036.10-2 3,625.10-3 8,807.10-4	5,725.10-2 3,580.10-2 1,930.10-2 7,992.10-3 2,803.10-3 6,695.10-3

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internal self-absorption<sup>1</sup>.

Therefore, although the number of prompt fission  $\gamma$  quanta and  $\gamma$  radiation of fission products are approximately equal, the contribution of prompt  $\gamma$  radiation during a nuclear explosion can be ignored.

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Data on the yields of various fragments formed by fission of  $U^{235}$ ,  $U^{238}$  and  $Pu^{239}$  by fission neutrons, neutrons with energies of 14 MeV and neutrons from a thermonuclear explosion are presented in [12].

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Since nuclei with various half lives participate in the radiation, the rate of decrease of  $\gamma$  activity of radiation with time does not follow a simple exponential rule. The effective half life increases as time passes, since the short-lived isotopes decay and only the longer-lived fragments remain.

<sup>1</sup>In this estimated calculation, the accummulation factor can be ignored, since its magnitude  $\simeq 5$ .

Table 1.2

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# ATTENUATION OF PROMPT $\gamma$ RADIATION OF FISSION IN URANIUM, CALCULATED FROM FORMULA (1.12)

		RADIUS C	<b>JF URANIUM S</b>	SPHERE IN FR	EE PATH LEN	GTHS µro	
<b>م</b> .	0.4	& `0	2,0	4,0	8.0	10	15
1,25 1,55 3,0 5,0	6,011.10 <sup>-1</sup> 3,923.10 <sup>-1</sup> 3,924.10 <sup>-1</sup> 9,000.10 <sup>-2</sup> 3,190.10 <sup>-2</sup> 7,924.10 <sup>-3</sup>	4,733.10-1 3,082.10-1 1,642.10-1 7,045.10-2 2,495.10-2 6,197.10-3	2,718.10-1 1,763.10-1 9,366.10-2 4,000.10-2 1,414.10-2 3,511.10-3	1,528-10-1 9,869.10-2 5,255.10-2 2,233.10-2 7,857.10-3 1,947.10-3	8,526-10-2 5,446-10-2 2,922-10-2 1,230-10-2 4,306-10-3 1,054-10-3	7,244.10-2 4,598.10-2 2,474.10-2 1,036.10-2 3,625.10-3 8,807.10-4	5,725.10-2 3,580.10-2 1,930.10-2 7,992.10-3 2,803.10-3 6,695.10-3

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The  $\gamma$  activity of fission products, i. e. the rate of emission of  $\gamma$  radiation energy u(t), can be calculated for U<sup>235</sup> and Pu<sup>239</sup> at various moments in time following the explosion using the following formula [13]:

$$u(t) = A_0 t^{-\gamma_0} \text{ MeV}/(\sec \cdot f).$$
 (1.13)

The values of  $A_0$  and  $\gamma_0$  for various time intervals are presented in Table 1.3.

Table 1.3

# VALUES OF COEFFICIENTS IN INTERPOLATION FORMULA (1.13) FOR CALCULATION OF $\gamma$ ACTIVITY OF FRAGMENTS AT VARIOUS TIMES FOLLOWING EXPLOSION

Time Sectors for Which Formula is Applicable	Unit of Measuremen of Time in Formula	t <sup>A</sup> 0	Υ <sub>0</sub> .	
0.05 sec< t < 1 sec	sec	0.4	0.27	
$1 \sec < t < 20 \sec$	11	0.4	0.72	
20 sec < t < 10 min	11	0.95	1.01	
10 min <t <10="" hr<="" td=""><td>11</td><td>7.1</td><td>1.32</td><td></td></t>	11	7.1	1.32	
10 hr <t <100="" hr<="" td=""><td>11</td><td>5.9</td><td>1.31</td><td></td></t>	11	5.9	1.31	
100 hr < t < 1000 hr	14	0.0276	0.89	
10 hr < t <100 hr	hr	1.3.10-4	1.31	
100 hr <t< 1000="" hr<="" td=""><td>11</td><td>1.9.10-5</td><td>0.89</td><td></td></t<>	11	1.9.10-5	0.89	
ì0 min <t 200="" <="" hr<="" td=""><td>sec</td><td>7.1</td><td>1.33</td><td></td></t>	sec	7.1	1.33	
20 sec <t<1000 hr<="" td=""><td>н</td><td>2</td><td>1.20</td><td></td></t<1000>	н	2	1.20	

The short-lived  $\gamma$  radiation of the products of fission of U<sup>235</sup> by thermal neutrons were studied in [14] from a few milliseconds to several seconds. Activities corresponding to decay times of 0.43 and 1.52 sec were detected.

Measurements performed by various authors with the fragments formed upon fission of uranium and plutonium by fast neutrons have shown that fragments with half-lives of less than one-tenth of one second are not present.

3

If we know u(t), we can calculate the energy radiated during the individual time intervals  $t_1 - t_2$ :

$$W_{t_1t_2} = \int_{t_1}^{t_2} u(t) dt$$
 (1.14)

or, expressing u(t) by an exponential rule:

$$W_{t_1 t_2} = \frac{A_0}{\gamma_0 - 1} \left( \frac{1}{t_1^{\gamma_0 - 1}} - \frac{1}{t_2^{\gamma_0 - 1}} \right).$$
(1.15)

The energy radiated upon full decomposition of radioactive fragments is

$$W = \int_{0}^{\infty} u(t) dt = 8.4 \frac{\text{MeV}}{\text{f}}$$
 (1.16)

The energy of  $\gamma$  radiation at various moments in time following fission is presented on Figure 1.5.



Time after fission, sec

Figure 1.5. Quantity of Energy of  $\gamma$  Radiation of Fission Fragments at Various Moments in Time Following Fission

The energy spectrum of the  $\gamma$  quanta of fission products consists of a large number of lines, the relationship between which changes

with time. The mean energy of the spectrum becomes lower, but significant softening occurs only during the first hour after fission. This is illustrated by Figure 1.6, based on data from [15-19]. Figure 1.7 shows the  $\gamma$  radiation spectra of fragments constructed on the basis of data of [20] for various moments in time following fission. Since the volatile fission products are not of direct danger from the standpoint of formation of local radioactive fallout, the  $\gamma$  radiation which they cause is excluded from the total spectrum.

In order to solve certain problems, it is necessary to know how the relative activity of  $\beta$  and  $\gamma$  radiation of fragments changes with time [17]. Figure 1.6 shows the change in relative  $\beta$  and  $\gamma$  activity with time, constructed according to the data of [16]. The spectral distribution of  $\beta$  radiation of the mixture of fission products of  $U^{235}$  at various moments in time following an explosion [20] is shown on Figure 1.8.

The  $\gamma$  radiation of an atomic explosion is not determined solely by the activity of the fission fragments.

The reason for the development of short-period  $\gamma$  radiation during a nuclear explosion is the  $\gamma$  radiation developing upon interaction of neutrons with the nuclei of the medium surrounding the explosion. One of the processes causing the appearance of  $\gamma$  radiation is inelastic scattering of fast neutrons in the materials of the bomb, in the air and in the earth. The duration of existence of this radiation in the shell of the bomb does not exceed the time required for the shell to burst, i. e. about  $10^{-5}$  sec. In addition to inelastic scattering, fast neutrons may enter into the reaction (n, p) or (n,  $\alpha$ ) with the emission of  $\gamma$  radiation, and the reaction (n,  $\gamma$ ). The duration of this radiation corresponds approximately to the duration of radiation by inelastic scattering.

The delayed neutrons are captured by nitrogen in the air or the earth, forming secondary  $\gamma$  radiation. The time of luminescence of this radiation is determined by the diffusion of neutrons in the air. Therefore, the  $\gamma$  radiation arising upon capture of moderated neutrons lasts for considerably longer than the radiation resulting from fast neutrons.

The life time of neutrons in air (relative to the reaction  $N^{14}$  (n, p)C<sup>14</sup> under normal atmospheric conditions) is about  $60 \cdot 10^{-3}$  sec. Thus, 0.3 sec following the explosion, practically all neutrons have been captured and the capture  $\gamma$  radiation ceases to be emitted.



Ε<sub>γ</sub>,

Figure 1.6. Change in Mean Energy of  $\gamma$  Quanta of Fragments of Fission of Pu<sup>239</sup> (curve 2) and Ratio of Summary  $\beta$  Activity to Summary  $\gamma$  Activity ( $\beta/\gamma$ ) of Fission Products (curve 1) as Functions of Time Following Fission

Although the life time of neutrons in the soil is brief, the air with its greater life time of neutrons may be a source of neutrons for capture in the soil. However, capture radiation in the soil is less significant in the formation of short-period radiation than capture radiation in the air. This results from the geometric factors determining the propagation of  $\gamma$  radiation in the soil, and the fact that the probability of capture of neutrons in the air is significantly higher than in the soil.

A significant portion of the neutrons formed in an explosion (90% for a shell permeability of 0.1) leaves the weapon with a mean energy of less than 1 KeV [4]. The mean distance traveled by neutrons in air to absorption is 210 m. The absorption of neutrons in air occurs primarily on nitrogen nuclei by two reactions:

N<sup>14</sup> (n, p) C<sup>14</sup> (
$$\sigma_{np} = 1,710^{-24} \ cm^2$$
);  
N<sup>14</sup> (n,  $\gamma$ ) N<sup>15</sup> ( $\sigma_{ny} = 0,1 \cdot 10^{-24} \ cm^2$ ).

In the reaction  $N^{14}(n, p)C^{14}$  with slow neutrons,  $\gamma$  radiation is not observed; consequently, the primary source of  $\gamma$  radiation upon neutron capture in air is the reaction  $N^{14}(n, \gamma)N^{15}$ .

-23-

The life time of neutrons in air at normal atmospheric pressure, considering the total capture cross section, is 60 msec. Therefore, the neutron cloud at the center of the explosion, with a radius of about 250-300 m, will be a source of  $\gamma$  radiation for approximately 60 msec. The neutrons leaving the cloud in the process of moderation are also sources of  $\gamma$  radiation, but since there are very few such neutrons, their contribution to the short-period radiation is slight.

Since at the same time the neutrons leave the bomb a shock wave is formed with high density at its leading edge, the life time of the neutrons is significantly changed. The solution of the problem of a strong shock wave gives us the following approximation for the density of air  $\rho_{\pm}$  at the leading edge of the shock wave [21]:

$$\frac{\rho_{\Phi}}{\rho_{0}} = \frac{(\overline{\gamma}+1)p + (\overline{\gamma}-1)}{(\overline{\gamma}-1)p + (\overline{\gamma}+1)}, \qquad (1.17)$$

where  $\rho_0$  is the density of air under normal conditions; p is the pressure at the leading edge of the shock wave;  $\overline{\gamma} = 1.4$  for air at normal temperature. For a very strong wave  $(p \rightarrow \infty)$  where  $\overline{\gamma} = 1.4$ , the maximum density of air in the shock front is 6. Since the air near the point of the blast is heated to very high temperatures, the value of  $\overline{\gamma}$  ceases to be constant under these conditions. At temperatures around 10<sup>4</sup> degrees, a significant portion of the energy is attributable to oscillating levels, dissociation and ionization of molecules occur; therefore, the heat capacity of the air changes. Under these conditions, the value of  $\overline{\gamma}$  may be 1.2. For  $\overline{\gamma} = 1.2$ , the maximum density of air in the shock front may be 11 times greater than the density of undisturbed air.

Since the life time of a neutron in a medium is inversely proportional to its density, the minimum life time of a neutron in the shock wave decreases from 60 to 5.5 msec. However, we must consider that the total life time of a neutron  $t_1$  is determined not only by the capture time  $t_c$  in the wave front, but also by the flight time  $t_{fl}$  in the cavity of low density:

$$t_{1} = \chi t_{3} + (1 - \chi) t_{f1}$$

where  $(1 - \chi)$  is the portion of neutrons reflected from the wave front into the cavity. Therefore, the minimum life time of a neutron in the leading edge of the shock wave exceeds 5.5 msec.



Figure 1.7. Spectral Distribution of  $\gamma$  Radiation of Fission Fragments at Various Moments in Time Following Fission





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Thus, if short-period  $\gamma$  radiation with periods of less than 5.5 msec exists, it cannot be explained as capture  $\gamma$  radiation. It results from the radiation of isomers formed upon interaction of fast neutrons with materials in the structure of the weapon. The short-period  $\gamma$  radiation from 5.5 to 20-30 msec may be partially caused by radiation of isomers and partially by capture of Maxwellian neutrons in the leading edge of the shock wave. The short-period  $\gamma$  radiation with a period of 30 msec or more results from capture of neutrons by nitrogen in the air.

The cross section of the reaction  $N^{14}(n, \gamma)N^{15}$  is 17 times less than the cross section of the reaction  $N^{14}(n, p)C^{14}$ ; therefore, only 6% of all neutrons will create short-period  $\gamma$  radiation.

According to the data of [22], each capture event liberates 10.82 MeV of  $\gamma$  quantum energy. Thus, for each neutron absorbed in the air, 0.64 MeV of  $\gamma$  radiation energy is radiated. Since each fission event emits 1.5 (for U<sup>235</sup>) to 2 (for U<sup>239</sup>) free neutrons, if all neutrons are absorbed by nitrogen in the air, the  $\gamma$  radiation liberated is 0.96-1.28 MeV/fission. This quantity is near the energy of  $\gamma$  radiation of fragments (1.8 MeV/fission) emitted in the first 10 sec (the mean time of application of  $\gamma$  radiation from the radioactive cloud to objects on the ground).

Since the time of capture radiation is less than the time of radiation of fragments, the dose rate of this short-period radiation will be significantly greater than the dose rate of the fragment  $\gamma$ radiation. At the initial moment, the intensity of capture  $\gamma$  radiation is about 21.4 MeV/(fission·sec), whereas the intensity of fragment  $\gamma$ radiation is 0.8 MeV/(fission·sec) [23]. Therefore, in the first moments following an atomic blast, the dose rate of capture  $\gamma$  radiation exceeds the dose rate of fragment  $\gamma$  radiation by over an order of magnitude. However, within a few tenths of a second, the primary contribution to the  $\gamma$  radiation is that of the fission fragments. Data on the spectrum of capture  $\gamma$  radiation are presented in [22].

Table 1.4 presents information on the energy and relative intensity of the capture radiation of nitrogen.

As we can see, the hardness of the capture  $\gamma$  radiation is considerably greater than the hardness of the fragment radiation. The mean energy of capture radiation is 6 MeV, of fragment radiation -- 2 MeV. As a result of capture of moderated neutrons, the upper layer of the earth (water) becomes itself a source of  $\gamma$  rays. Radiation from the earth occurs both at the moment of capture of neutrons, and during subsequent decay of artificially radioactive neuclei. The total yield of energy of capture  $\gamma$  radiation per absorbed neutron for all elements except hydrogen is about 8 MeV. During capture of a thermal neutron by a hydrogen atom,  $\gamma$  radiation is liberated with an energy of 2.23 MeV.

Energy MeV	of y	Quantum,	Relative Intensity, %	Energy of γ Quantum,MeV	Relative Intensity, %
		4,48 5,29 5,55 6,32 7,16	18,3 30,5 2,44 1,22 0,6	7,36 8,28 9,16 10,82	4,9 2,4 0,6 6,1

 $\gamma$  RADIATION SPECTRUM OF THE REACTION  $N^{14}(n,\ \gamma)N^{15}$ 

The radiation occurring upon capture of neutrons in soil is shortperiod radiation. The life time of neutrons in soil is about  $0.5 \cdot 10^{-3}$ sec, i. e. significantly less than in air ( $\simeq 60 \cdot 10^{-3}$  sec). However, the air is the source which delivers neutrons to the ground for capture. Therefore, the time during which capture of neutrons in the soil may occur, with emission of  $\gamma$  radiation, is equal to the time of existence of neutrons in the air, i. e. about 0.3 sec.

Let us study the main characteristics of fragment  $\gamma$  radiation resulting from activation of the soil by the neutrons of a nuclear explosion.

According to the data of [4], only a small number of elements included in the composition of the soil play a significant role in the formation of  $\gamma$ -active isotopes. During the first few minutes following bombardment, one such element is  $A1^{28}(T_{1/2} = 2.3 \text{ min})$ . Its activity is almost one hundred times greater than the activity of all other isotopes. After a few hours have passed, the most significant activity is that of Na<sup>24</sup> and Mn<sup>56</sup>. Between 20 and 200 hours, the activity of the bombarded soil is almost completely determined by the activity of Na<sup>24</sup>; therefore, during this interval, the effective period of decrease

Table 1.4

of soil activity corresponds to the half life of Na<sup>24</sup> (15 hours). After 200 hours have passed **fol**lowing the explosion, Fe<sup>56</sup> remains active ( $T_{1/2} = 47$  days).

Later (t > 200 hr), the soil activity drops more slowly than calculations would indicate. This fact is apparently related to the presence of small quantities of dispersed elements in the earth. For example, cobalt, traces of which are always present in the earth, due to its large capture cross section and long half life (5.3 years), may significantly change the form of the curve for the interval of time during which short-period elements decay.

The residual  $\gamma$  radiation in the region of a blast may also result from fission fragments which fall to the ground from the radioactive cloud. The quantity of fragments which fall out on the ground is determined to a great extent by the conditions of the explosion and may differ significantly for blasts of the same power performed under different conditions: on the surface or in the air. Therefore, the degree of participation of fission fragments and activated soil in the creation of the dose of residual  $\gamma$  radiation may change significantly. With blasts in the air, when the quantity of fission fragments falling out of the radioactive cloud is not great, the activated soil is the primary source of residual radiation. When a blast is performed at low altitude over the ground, the residual radiation is primarily created by fission fragments.

The energy of  $\gamma$  radiation of fission fragments radiated between 10 seconds (the approximate time of completion of  $\gamma$  radiation of the radioactive cloud to objects on the ground) up to infinity is 6.6 MeV/fission (see Figure 1.5). For activated soil, this quantity is 0.48 MeV per neutron captured by the soil. If we base our calculations on the greatest possible  $\gamma$  radiation by the activated soil, i. e. the radiation produced by absorption of one-half of all neutrons in the soil, the energy radiated by the activated soil, ignoring selfabsorption in the soil, is 0.48 MeV/fission when a plutonium bomb is exploded. Thus, even when a surface explosion occurs, the total quantity of energy radiated by the active soil is 14 times less than the energy of  $\gamma$  radiation of the fission fragments.

Direct comparison of the possible significance of fission fragments and activated soil in the creation of the dose of residual radiation is difficult due to the different spatial distribution of radiation from activated soil and fission fragments. Whereas the activity of the soil decreases with increasing distance from the point of the blast approximately according to the same rules as the decrease in neutron flux, the spatial distribution of fission fragments follows quite different regularities. First of all, the spatial distribution of fission fragments is very asymmetrical and is determined by the direction of the wind which carries away the fission products. If the particles which fall out to the ground from the radioactive cloud form a thin layer of radioactive material, the soil is activated to a great depth -- some tens of centimeters. Therefore, a great portion of the  $\gamma$  radiation of the soil is absorbed in the soil itself. This decreases the significance of activated soil in the creation of residual  $\gamma$  radiation. The differing role of fragments and activated soil in the formation of residual radiation is also determined by the nature of the changes in intensity with time.

It follows from [24] that during the first 10 hours the activity of fission fragments decreases more rapidly than the activity of the soil. After 50 hours, the induced activity decreases more rapidly than the fragment activity. From 10 to about 50 hours, the rule of change of fission fragment activity and soil activity is approximately the same.

As we know, a significant portion of the fission fragments, even with surface blasts, falls out over tremendous areas [25]. If 1% of all fission fragments formed fall out over the region of the epicenter the residual radiation will be determined not by activated soil, but rather by fission fragments.

Far from the epicenter with surface blasts the residual radiation is determined by induced activity. With a surface blast of a pure hydrogen bomb, an order of magnitude more neutrons may be liberated per unit energy than when a fission bomb is exploded. Then we can expect that the energy of radiation of the activated earth will be on the same order of magnitude as the energy of radiation of fragments upon explosion of a fission bomb of the same power.

We note in conclusion that the residual activity formed upon capture of neutrons by the elements which compose the air is low in comparison to the activity of the fragments. For this reason, the products of activation of atmospheric air can be ignored in calculations.

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## CHAPTER 2. THE FIELD OF NEUTRONS AND $\gamma$ RADIATION OF A NUCLEAR BLAST

§ 1. Propagation of  $\gamma$  Radiation of a Nuclear Blast

(the influence of the ground will not as yet be analyzed).

At the moment of a blast, a near-point source of  $\gamma$  radiation appears in the form of the fission fragments, the activity of which g(t)/(MeV/sec) and the spectrum  $\Phi(E_{\gamma})$  of which are known. The  $\gamma$  quanta emitted upon fission are ignored, since, as we have demonstrated, they are significantly attenuated by the material of the weapon. The  $\gamma$  radiation emitted by the fission fragments following evaporation of the shell of the weapon acts on objects on the ground for several tens of seconds, until the radioactive cloud rises to a considerable height. The propagation of  $\gamma$  radiation occurs through a practically infinite air medium

Let us analyze briefly the case of propagation of  $\gamma$  radiation from a point source through a homogeneous infinite medium. Since this problem has spherical symmetry, the distribution of intensities and doses of  $\gamma$  radiation in space depends only on one spatial coordinate -- the distance from the source R.

The intensity of radiation of a point isotropic source I(R) decreases with increasing R due to increasing distance from the source, and also due to interactions with the medium. In correspondence with this, for a monoenergetic source it will be:

$$I(R) = \frac{g(t)}{4\pi R^2} B(E_{0\gamma}, R) e^{-\mu_0 R} \text{ MeV}/(cm^2 \cdot sec), \qquad (2.1)$$

where  $\mu_0$  is the attenuation factor of  $\gamma$  radiation in the air;  $B(E_{0\gamma}, R)$ 

is the factor of accummulation of intensity, equal to the ratio of total radiation intensity to the portion related only to direct (non-scattered) radiation. Expression (2.1) was written for the intensity of  $\gamma$  radiation; however, it can be extended to cover all effects related to the action of quanta.

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The dose rate P for the case of monoenergetic radiation is related to the intensity I by the relationship

$$P = I \cdot 1,48 \cdot 10^{-5} \mu_e(E_{\rm Cy}) \ r/sec,$$
 (2.2)

where  $\mu_{e}(E_{0_{\gamma}})$  is the absorption factor of the energy of  $\gamma$  radiation in air.

In the case of non-monoenergetic radiation, expression (2.2) is written as

$$P = 1,48 \cdot 10^{-5} \int I(E_{0\gamma}) \mu_e(E_{0\gamma}) dE_{\gamma} \text{ r/sec.}$$
 (2.3)

As  $\gamma$  radiation propagates through matter, it becomes non-monoenergetic, even if a source of monoenergetic radiation with  $\gamma$  quantum energy  $E_{0\gamma}$  is acting on the matter. In this case,  $\mu_e(E_{\gamma})$  is taken away from the integral sign:

$$P = 1,48 \cdot 10^{-5} \,\mu_e(E_{\gamma}) \int \frac{\mu_e(E_{\gamma})}{\mu_e(E_{0\gamma})} I(E_{0\gamma}) \,dE_{\gamma} \quad r/sec. \qquad (2.4)$$

Since over a broad range of quantum energies  $\mu_e$  depends little on  $E_\gamma$ , the factor of accummulation of dose power practically corresponds with the factor of accummulation of intensity; therefore

$$P(R) = 1,48 \cdot 10^{-5} \frac{g(t) \mu_e(E_{0\gamma}) B(E_{0\gamma}, R)}{4\pi R^2} e^{-t_0 R} \text{ r/sec.} \qquad (2.5)$$

Determination of the coefficient  $B(E_{0_{\gamma}}, R)$  is a problem in the theory of transfer of quanta or modeling experiments.

As a result of model experiments, the authors of [1] suggested the following dependence for dose rate

$$P(R) = \frac{1,48 \cdot 10^{-5} \,\mu_e(E_{0\gamma}) \,g(l) \,\overline{\alpha}}{4\pi R^2} \,\mathrm{e}^{-R/l} \,\mathrm{e}^{ff} \,\mathrm{r/sec}, \qquad (2.6)$$

in which multiple scattering is considered by the dimensionless coefficient  $\alpha$  and the effective length of absorption of  $\gamma$  radiation energy under conditions of multiple scattering  $\lambda_{\text{eff}}$ .

Expression (2.6) is formally equivalent to expression (2.5), suggested in [2], but in place of the accummulation factor B, which changes greatly with distance, and the free path length  $\lambda_0 = 1/\mu_0$ , which is independent of distance, the quantities  $\overline{\alpha}$  and  $\lambda_{eff}$ , which change slightly with distance, are used. In many cases, it is preferable to use formula (2.6). Table 2.1 shows the values of  $\overline{\alpha}$  and  $\lambda_{eff}$  for air and two intervals of distances from the point isotropic source of  $\gamma$  radiation in the air [3].

In order to describe the dose rate of an atomic blast, we can express g(t) in formula (2.6) by the number of fission events and the activity of the fragments of one event u(t):

$$g(l) = 1,45 \cdot 10^{23} gu(l) \text{ MeV/sec.}$$
 (2.7)

The coefficients  $\overline{\alpha}$ ,  $\lambda_{eff}$ ,  $\mu_e(E_{\gamma})$  depend on the energy of the quanta emitted by the fragments. In correspondence with the data presented in § 3 of Chapter 1, we note that in the time interval of interest to us (several seconds following fission), the mean quantum energy is  $\simeq 2 \text{ MeV}$ . Substituting the numerical value of coefficient  $\alpha$  taken in Table 2.1 for this energy into formula (2.6), as well as the value of  $\lambda_{eff}$  (exp.) and expressing R in meters, we produce

$$P(t, R) = \frac{9.32 \cdot 10^{8} qu(t)}{R^{2}} e^{-R/255} \text{ r/sec.}$$
(2.8)

The source of fragment  $\gamma$  radiation following evaporation of the weapon is the radioactive cloud. Its diameter may reach some hundreds of meters. We must therefore determine the extent to which the nature of the propagation of  $\dot{\gamma}$  quanta from this volumetric source will correspond with the case just studied of a point source in an infinite air medium.

For further analysis, let us make the following assumptions: the fission fragments are evenly distributed throughout the entire volume of the cloud; attenuation of radiation within the volume of the radioactive cloud occurs just as it occurs outside its volume.

Table 2.1

e <sub>v</sub> , MeV		α	<sup>\lambda</sup> eff, m Interval of Dis- tances, m			
	Interval tances	of_Dis , m				
	400-1000	1000-2000	400-1000	1000-2000		
0,4	. 9,00	50,0	105	99		
0,6	4,80	18,4	128	108		
0,8	3,70	12,4	146	120		
1,0	2,90	8,90	164	134		
2,0	1,80	3,85	230	200		
4,0	1,30	1,90	333	295 🖌		
6,0	1,18	1,50	409	372		
8,0	1,14	1,30	450	425		
10,0	1,10	1,10	482	458		

# VALUES OF COEFFICIENTS $\overline{\alpha}$ AND $\lambda_{\mbox{eff}}$ for calculation of dose rate of point source of $\gamma$ radiation in Air

The power of the dose from a volumetric spherical  $\gamma$  source at any distance from the source can be calculated using the formula

$$P = \int \int \frac{\omega(r) f(R'\lambda)}{R'^2} dv \quad r/sec, \qquad (2.9)$$

where  $\Gamma$  is the  $\gamma$  constant of the source; w(r) is the specific activity of the  $\gamma$  source; f(R'/ $\lambda$ ) is the attenuation of the power in the medium as a result of absorption and scattering of the  $\gamma$  radiation. Integration is performed with respect to the volume of the source. Looking at Figure 2.1, we note:

$$dv = 2\pi y dr \frac{dx}{\sin\beta}; \quad \sin\beta = \frac{y}{r},$$

so that

 $dv := 2\pi r \, dr \, dx.$ 

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The power of the dose from a spherical layer of the source will

$$\Delta P(R) = 2\pi\Gamma \int_{-r}^{+r} \frac{\omega(r) rf(R'/\lambda)}{R'^2} dr dx, \qquad (2.10)$$

where

3

be

$$R' = \sqrt{y^2 + (R - x)^2} = \sqrt{r^2 + R^2 - 2Rx}.$$



Figure 2.1. Geometry for Calculation of Dose of  $\gamma$  Radiation from Radio-active Cloud of Nuclear Explosion

After substituting this dependence into formula (2.9), we produce

$$\Delta P(R) = 2\pi \Gamma w(r) r dr \int_{-r}^{+r} \frac{f\left(\frac{R'}{\lambda}\right) dx}{r^2 + R^2 - 2Rx};$$

the power of the dose from the entire volume of a spherical source of radius  $\mathbf{r}_0$  will be

$$P = 2\pi\Gamma \int_{0}^{r_{0}} w(r) r dr \int_{-r}^{+r} \frac{f(R'/\lambda) dx}{r^{2} + R^{2} - 2Rx}.$$
 (2.11)

Since we are interested not in the absolute value of dose power from a spherical  $\gamma$  source, but rather only its ratio to the power of the dose from a point source of the same activity, which is

$$P_{\rm T} = \frac{\Gamma Q f(R'/\lambda)}{R_0^2},$$

where

$$Q == 4\pi \int_{0}^{r_{0}} w(r) r^{2} dr,$$

it will be

$$K_{f} = \frac{P}{P_{T}} = 0.5 \frac{\int_{0}^{r_{0}} w(r) r dr \int_{-r}^{+r} \frac{f(R'/\lambda) dx}{r^{2} + R^{2} - 2Rx}}{\frac{f(R'/\lambda)}{R^{2}} \int_{0}^{r_{0}} w(r) r^{2} dr}; \qquad (2.12)$$

$$K_{f} = \frac{\int_{0}^{r_{0}} r dr \int_{-r}^{+r} \frac{f(R'/\lambda) dx}{r^{2} + R^{2} - 2Rx}}{\frac{r_{0}^{3} f(R'/\lambda)}{R^{2}}}.$$
 (2.13)

The results of calculations using this formula for various sizes of radioactive cloud and various distances from the point of measurement to the center of the source (in values of  $R/\lambda$ ) are presented in Table 2.2.

We see from Table 2.2 that the dose power of fragment  $\gamma$  radiation from a spherical source with dimensions  $r_0/\lambda \leq 2$  is equal to the dose power from a point source of the same activity, beginning at distances determined by the boundary of the source. We note that the

dimensions of the source of fragment  $\gamma$  radiation in all cases of practical importance are always less than  $2r_0/\lambda$ .

Table 2.2

## RELATIONSHIP OF DOSE POWERS FROM SPHERICAL RADIOACTIVE CLOUD AND POINT SOURCE OF SAME ACTIVITY

	$r_0/\lambda = 4$		$r_0/\lambda = 2$		$r_0/\lambda = 1$		$r_0^{/\lambda} = 0.5$		$r_0/\lambda = 0,25$	
	R/\	K <sub>f</sub>	R/X	ĸ	R/).	κ <sub>j</sub>	<b>R</b> /λ	ĸŗ	R/À	ĸ
-	6 4 3 2	1,8 2,6 1,3 0,4	3 2 1,5 1,0	1;0 1,2 0,8 0,4	2 1,0 0,75 0,5	1,0 1,0 0,83 0,43	1,0 0,5 0,37 0,25	1,0 1,0 0,82 0,46	0,5 0,25 0,188 0,125	1,0 1,0 0,8 ] 0,45

Dimensions of Radioactive Cloud

Thus, using expression (2.8), we can find the dose power of fragment  $\gamma$  radiation at any distance from the center of the explosion. However, this solution plan for the problem is insufficient. The authors of [4] turned their attention to the fact that the absorption of  $\gamma$  radiation propagating from the center of an explosion should be sensitive to redistribution of air masses resulting from the passage of the shock wave. The shock wave forms a cavity in the air, in which there is no absorption of  $\gamma$  radiation. The size of the cavity increases with increasing explosion energy. Therefore, the dose of  $\gamma$  radiation at a given distance from the center of the explosion does not increase in proportion to the power as it increases, but rather more rapidly. The calculation of the transmission of  $\gamma$  radiation should take into consideration the actual distribution of air density created by the shock wave during the emission of the  $\gamma$  radiation. A plan for considering the influence of the shock wave was presented in [4].

Let us study the distribution of the density of air in the shock wave, which influences the propagation of  $\gamma$  radiation.

In a strong shock wave propagating from the center of an explosion, all the air is concentrated practically in a narrow edge around the leading edge of the wave and moves away from the center. Thus, inside the spherical shock wave front an area is created with low air density, i. e. a cavity. After the shock wave has passed and the initial air pressure has been restored, the area with low air density remains. The entropy in the shock wave increases [5, 6], the air does not return to its initial state, and remains heated after passage of the shock wave and restoration of the initial pressure. In the area where the shock wave was strong (i. e. the pressure at the leading edge was significantly greater than atmospheric pressure), the residual heating is great and causes glowing of the air. The fireball formed after an atomic explosion is the "wake" of the passing shock wave. The pressure in the fireball is equal to atmospheric pressure, while the density is low as a result of the high temperature of the air. As an example, Figure 2.2 shows the distribution of air density across the leading edge of a shock wave 0.765 sec after the explosion of a nuclear weapon with a power of 300 Kt.

We can see from Figure 2.2 that 0.765 sec after the explosion at a distance of 600 m from the center  $\rho_{\phi}/\rho_{0} \ll 1$ . The absorption of  $\gamma$  radiation in this area is negligible and it can be considered that there is a cavity inside the strong shock wave. The radius of the fireball  $r_{fb}$  determines the maximum distance over which the shock wave which has passed was still strong. As time passes and the radius of the shock front increases, the pressure in the leading edge of the shock wave decreases, the shock wave becomes weak, the air partially returns to the central areas and atmospheric pressure is restored everywhere with the exception of a slight peak at the shock wave. However, the temperature in the central area remains high, the air density remains low.



Figure 2.2. Distribution of Air Density in a Strong Shock Wave

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Subsequently, the shock wave, propagating to an ever greater distance, is continually attenuated and degenerates to a sound wave. The shock wave phenomenon ends, but the air remains heated after it passes. Cooling of the air (and restoration of normal density) occurs by radiation and convection over an extended period of time, significantly greater than the time of emission of most of the  $\gamma$  radiation.

The  $\gamma$  radiation passes through the cavity practically without absorption, its flux attenuating within the cavity in inverse proportion to the square of the distance from the center, as is the case for propagation of  $\gamma$  radiation in a vacuum.

Consideration of the distribution of air density with a shock wave, which creates a heterogeneity in the medium surrounding the source, is performed by introducing the concept of the "optical thickness" for absorption of  $\gamma$  radiation. We know that if the absorption of radiation occurs exponentially, the intensity of the radiation which has traveled distance R in a homogeneous medium with length of absorption  $\lambda(R)$  is determined by the so-called optical thickness

$$x_0 = \int_0^R \frac{dr}{\lambda} \, .$$

Correspondingly, the attenuation of radiation over path R is  $e^{-x_0}$ .

The absorption of  $\gamma$  radiation differs somewhat from exponential absorption due to the presence of repeated scattering. However, within a certain interval of distances it can be represented in exponential form. This gives us reason to extend the concept of "optical thickness" to the case of absorption of  $\gamma$  radiation in a heterogeneous medium and to consider that the attenuation of intensity due to absorption is

$$\exp(-x_0) = \exp \left[-\frac{1}{\lambda_{eff}(\rho_0)\rho_0} \int_{0}^{R} \rho dr\right],$$
 (2.14)

where  $\rho_0$  and  $\rho$  are the air density before and after arrival of the shock wave.

In the undisturbed atmosphere, the optical thickness over path R is  $R/\lambda_{eff}$ . Suppose the radius of the leading edge of the strong shock

wave is  $R_{\phi}$ . All air in the strong shock wave is practically concentrated in a layer of thickness  $\Delta R$  at the leading edge of the shock wave, where for  $\Delta R \ll R_{\phi}$  (see Figure 2.2), the optical thickness is

$$x_0 = \frac{1}{\lambda_{eff}(\rho_0)\rho_0} \int_0^R \rho dr = \frac{1}{\lambda_{eff}} \frac{\rho \Delta R}{\rho_0}.$$
 (2.15)

For approximate estimation, it is assumed that within the interval of  $\Delta R$  the density  $\rho = const$ .

The value of the product  $\Delta R(\rho/\rho_0)$  can be found from the condition that the mass of air at the edge of the shock wave near the surface with radius  $R_{\phi}$  is equal to the mass of air which was formerly in the sphere of radius  $R_{\phi}$  before the shock wave developed:

$$\rho 4\pi R_{\Phi}^2 \Delta R = \rho_0 \frac{4}{3} \pi R_{\Phi}^3;$$
$$\Delta R \frac{\rho}{\rho_0} = \frac{R_{\Phi}}{3}.$$

Substituting this expression into formula (2.15), we find

$$x_0 = \frac{1}{3} \frac{R_{\phi}}{\lambda_{eff}}.$$

When there is no strong shock wave, the optical thickness is  $R/\lambda_{eff}$ . Thus, the redistribution of the mass of air in a strong shock wave causes the optical thickness along path R to decrease by two-thirds  $R_{\phi}/\lambda_{eff}$ . Absorption of  $\gamma$  radiation within the shock wave occurs as if it contained a cavity of radius L:

$$L = R - \frac{\int \rho dR}{\rho_0} \,. \tag{2.16}$$

In the case of a strong shock wave

$$L = \frac{2}{3} R_{\phi}.$$
 (2.17)

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Absorption of  $\gamma$  radiation occurs outside the cavity in layer  $\Delta R$  of dense air at the leading edge of the wave:

$$\Delta R = \frac{R_{\phi}}{3} \frac{\rho_0}{\rho} \,. \tag{2.18}$$

Thus, at distance R from the center of the explosion (R > L), the optical thickness is  $(R - L)/\lambda_{eff}$ . The decrease in flux will be proportional to

$$\exp\left[-(R-L)/\lambda_{\text{eff}}\right]$$
,

whereas without the shock wave the intensity would be proportional to exp (-  $R/\lambda_{eff}$ ).

Thus, the strong shock wave causes the intensity of radiation to be determined by formula (2.10), in which the exponent contains R - L rather than R.

Coefficient  $\overline{\alpha}$  remains practically unchanged in this case:

$$P(t, R) = \frac{1,48 \cdot 10^{-5} \,\mu_c(E_{0\gamma}) \,g(t) \,\overline{\alpha}}{4\pi R^2} e^{L/\lambda_{eff}} e^{-R/\lambda_{eff}}.$$
 (2.19)

In the idealized case which we have studied, when the distribution of densities in the leading edge of the shock wave corresponds to that shown on Figure 2.2 by the dotted line, the radius of the cavity is equal to two-thirds the radius of the leading edge of the shock wave (2.17). With the actual distribution of densities in the shock wave, the effective radius of the cavity

$$L_{eff} < \frac{2}{3} R_{\phi}.$$

By the effective radius of the cavity we shall mean the radius of an idealized cavity formed as a result of expulsion of air from the space around the source.

The effective radius of the cavity can be determined on the basis of the assumption that the density of the air at the leading edge of the shock wave can be described by the dependence [3]

$$\rho(l) = \rho_{\phi} \left(\frac{l}{R_{\phi}}\right)^{n}, \qquad (2.20)$$

which is correct only for a strong shock wave. In this expression, 1 is the distance from the leading edge of the shock wave back toward the source.

The exponent n depends on the value of  $\rho_{\phi}$  and is determined by calculating the quantity of air in the shock wave.

Before formation of the shock wave, the quantity of air contained in a volume of radius  $R_{\phi}$  was  $4/3(\pi R_{\phi}^{3}\rho_{0})$ . Thus,

$$\frac{4}{3}\pi R_{\Phi}^{3}\rho_{0} = 4\pi \int_{0}^{R_{\Phi}} \rho(l) l^{2} dl.$$
(2.21)

Substituting expression (2.20) in the integral in place of  $\rho(1)$ , we find that

$$4\pi \frac{\rho_{\phi}}{R_{\phi}^{n}} \frac{R_{\phi}^{n+3}}{n+3} = \frac{4}{3} \pi \rho_{0} R_{\phi}^{3},$$

from which

$$n=3\left(\frac{\rho_{\phi}}{\rho_{0}}-1\right). \tag{2.22}$$

On the basis of these assumptions, we can calculate the actual optical thickness. Let us study two possible cases of placement of the source relative to the center of the explosion: 1) the detector is located within the shock wave (R <  $R_{\phi}$ ); 2) the detector is located outside the shock wave (R >  $R_{\phi}$ ).

In the first case

$$x_{0} = \mu_{eff} \int_{0}^{R} \frac{\rho}{\rho_{0}} dR = \mu_{eff} \left( -R + R + \frac{1}{\rho_{0}} \int_{0}^{R} \rho dR \right) =$$

$$= \mu_{eff} R - \mu_{eff} \left( R - \int \frac{\rho}{\rho_{0}} dR \right) = \mu_{eff} R - \mu_{eff} Leff,$$
(2.23)

where

$$L_{eff} = R - \int_{0}^{R} \frac{\rho}{\rho_{0}} dR. \qquad (2.24)$$

Substituting expression (2.20) for  $\rho$  in formula (2.24) and performing integration, we produce

. . . . . . . .

$$L_{eff} = \dot{R} \left[ 1 - \frac{\frac{\rho_{\phi}}{\rho_{0}}}{\frac{3\rho_{\phi}}{\rho_{0}} - 2} \left(\frac{R}{R_{\phi}}\right)^{3} \left(\frac{\rho_{\phi}}{\rho_{0}} - 1\right) \right].$$
(2.25)

Substituting formula (2.25) into expression (2.19), we produce

$$P(t, R) = \frac{1,48 \cdot 10^{-5} \,\mu_e(E_{0\gamma}) g(t) \,\overline{\alpha}}{4\pi R^2} \, e^{\frac{L_{eff}}{\hbar_{eff}}} \, e^{-R/\lambda_{eff}}. \qquad (2.26)$$

In the second case (R >  $R_{\phi}$ ), when where  $R_1 > R_{\phi}$ ,  $\rho/\rho_0 = 1$ ,

$$x_{0} = \mu_{eff} \int_{0}^{R} \frac{\rho}{\rho_{0}} dR = \mu_{eff} \int_{0}^{R_{\phi}} \frac{\rho}{\rho_{0}} dR + \mu_{eff} \int_{R_{\phi}}^{R} dR =$$

$$= -\mu_{eff} \int_{0}^{R_{\phi}} \frac{\rho}{\rho_{0}} dR + \mu_{eff} [R - R_{\phi}] = \mu_{eff} R -$$

$$(2.27)$$

$$- \mu_{eff} \left( R_{\phi} - \int_{0}^{R_{\phi}} \frac{\rho}{\rho_{0}} dR \right) = \mu_{eff} R - \mu_{eff} L_{eff},$$

where

$$L_{eff} = R_{\Phi} - \int_{0}^{R_{\Phi}} \frac{\rho}{\rho_{0}} dR. \qquad (2.27)$$

Substituting expression (2.20) into (2.27) and performing integration, we produce

$$L_{eff} = R_{\phi} \frac{2 \frac{\rho_{\phi}}{\rho_{0}} - 2}{3 \frac{\rho_{\phi}}{\rho_{0}} - 2}.$$
 (2.28)

Let us substitute formula (2.27) into expression (2.19); we produce a formula identical to expression (2.26), which covers both cases of placement of the detector relative to the shock wave. The difference lies in the expressions for  $L_{eff}$  of (2.24) and (2.27), which differ in that (2.24) includes R in place of  $R_{\phi}$ .

Thus, the distribution of air density caused by the shock wave increases the dose power by  $\exp(\mu_{eff}L_{eff})$  times. The factor showing the number of times by which the dose power when there is a shock wave exceeds the dose power at the same moment in time but without a shock wave is called the dose power cavity factor K. In formula (2.26), the value of the cavity factor is approximated by the expression

$$K_{p} = e^{\mu_{eff} L} e^{ff(t, q)}. \qquad (2.29)$$

Considering this, let us place  $K_{p}$  into the initial formula (2.8):

$$P(t, R) = \frac{9.32 \cdot 10^{8} qu(t) K_{p}(t)}{4\pi R^{2}} e^{-R/255} \text{ r/sec.} \qquad (2.30)$$

The dependence of  $L_{eff}$  on t and q can be expressed using  $R_{\phi}$  and  $\rho_{\phi}/\rho_{0}$ , the values of which depend on the energy of the explosion and the time which has passed since the explosion. With a very strong shock wave, the following limiting expressions are produced for  $L_{eff}$ : 1).  $L_{eff} = R$ , when  $R < R_{\phi}$  and, consequently, the effective radius of the cavity is independent of time; 2)  $L_{eff} = \frac{2}{3}R_{\phi}$ , when  $R > R_{\phi}$ .

The precise value of  $L_{eff}(t)$  can be calculated on the basis of the expression

$$L_{eff}(t) = a(t)k\sqrt[3]{q},$$
 (2.31)

where k = 100, and the dose power is expressed in kilotons. The dependence of a(t) for explosions of various powers is shown on Figure 2.3.

It should be noted that in determining the dimensions of the cavity we can use the rule of similarity of shock waves. If the values of  $R_{d1}$  and  $\rho_{d}$  are known with explosion of a weapon of power  $q_1$ , the same

values of  $\rho_{\phi}$  are produced when a weapon of power  $q_2$  is exploded, but at a different moment in time  $t_2$  and at a different distance, namely at moment

$$t_2 = t_1 \sqrt[3]{\frac{q_2}{q_1}}$$

and at distance

$$R_{\phi_2} = R_{\phi_1} \sqrt[3]{\frac{q_2}{q_1}}.$$



Figure 2.3. Size of Air Cavity as a Function of Time for Explosions of Various Powers: 20 Kt (curve 1); 200 Kt (curve 2); 2000 Kt (curve 3)

The values of  $R_{\phi}$  and consequently of  $L_{eff}$  depend also on the conditions of the explosion. With a surface explosion,  $R_{\phi}$  is greater than with an atmospheric explosion of the same power. This occurs due to the fact that with a surface explosion that portion of the explosive energy which would have been transmitted into the lower half space, i. e. half the explosive energy, remains in the upper half space, thus creating twice the energy density which would be created by an air blast. Therefore, the parameters of the shock wave with a surface blast are the same as would obtain with an air blast of twice the energy. In particular, in comparing air and surface explosions, the energy should be doubled for the surface explosion.

§ 2. Dose Power of  $\gamma$  Radiation With a Nuclear Explosion<sup>1</sup>

If there were no shock wave, the dose of fragment  $\gamma$  radiation from an atomic explosion would decrease monotonically with the passage of time, in proportion to the decreasing activity of fission fragments, and also as a result of the rise of the radioactive cloud. However, due to the formation and growth of the cavity, the absorption of  $\gamma$ radiation decreases with the passage of time, and after the shock wave passes the point of measurement, this point is inside the cavity, and the  $\gamma$  quanta are practically unattenuated. In other words, the dose rate of fragment radiation as a function of time is defined by the  $\mu_{eff}^{L} eff^{(t)}$ , one of which (the activity of the fragments) decreases, while the other (the dose power cavity factor) at first increases due to the growth of the cavity, then decreases to unity as the disturbed atmosphere returns to its normal

Due to the simultaneous influence of these two factors, the dose power of fragment  $\gamma$  radiation passes through a minimum and a maximum.

state.

Let us determine the dose power of capture  $\gamma$  radiation. The life time of neutrons in air with normal density, considering the total capture cross section, is  $\tau = 60$  msec. Therefore, the neutron cloud around the explosion (R = 250-300 m) will be a source of  $\gamma$  radiation, the intensity of which is determined by the value of  $\tau$ .

As concerns the dimensions of the source of capture  $\gamma$  radiation, calculations similar to those presented above for fragment  $\gamma$  radiation indicate that the factor, considering the influence of the finite dimensions of the cloud of Maxwellian neutrons on the formation of the dose of capture  $\gamma$  radiation during a nuclear explosion, can be written as follows:

<sup>&</sup>lt;sup>1</sup>This book will not analyze capture  $\gamma$  radiation formed by absorption of neutrons by the earth, since its contribution at great distances and high altitudes of explosion is relatively slight. At slight distances from the explosion (several times the neutron mean free path length), capture  $\gamma$  radiation from the earth should be considered. It can be calculated in the same manner as the radiation scattered from the earth, using the "albedo" of capture  $\gamma$  radiation (see Engineering Compendium on Radiation Shielding, edited by Jaeger R. G. et al., Vol. 1, Springer-Verlag, Berlin-Heidelberg-New York, 1968, p. 33; Goryachev, I. V. et al., <u>Report 56 at CEMA Symposium</u> "Problems of Shielding From Penetrating Radiation of Reactors," Melekess, 21-26 April, 1969).

$$K_{c} = 0.5 \cdot \frac{\int_{0}^{r} e^{-r/\lambda_{0}} \frac{dr}{r} \int_{-r}^{+r} \frac{f(R'/\lambda) dx}{r^{2} + R^{2} - 2Rx}}{\frac{f(R'/\lambda)}{R^{2}} \lambda_{0} (1 - e^{-r_{0}/\lambda_{0}})}.$$
(2.32)

Factor  $K_c$  is equal to the ratio of dose power at distance R from the center of a volumetric source of radius  $R_s$  to the dose power at the same distance from a point source of equivalent activity.

This expression was produced by substituting the function  $w(r) \approx e^{-r/\lambda_0}$ , describing the spatial distribution of sources of capture  $\gamma$  radiation ( $\lambda_0$  is the effective length of attenuation of neutrons in the cloud) into formula (2.12). Calculations of coefficient K<sub>c</sub> using this expression have shown that the source of capture  $\gamma$  radiation can be considered a point source where R'/ $\lambda \geq 0.1$ , i. e. at practically all distances from the center of the explosion.

This indicates that we can determine the dose power of capture  $\gamma$  radiation by using formula (2.19). For capture  $\gamma$  radiation it will be as follows:

$$P(R) = \frac{1,48 \cdot 10^{-5} Nq \varepsilon_{\gamma} \mu_{e}(E_{0\gamma}) \sigma_{n\gamma} e^{-\frac{t}{\tau}}}{\tau (\sigma_{n\gamma} + \sigma_{n\rho})} \times \frac{\overline{\alpha} (1-\alpha) \Delta (1-\zeta)}{4\pi R^{2}} e^{\frac{R-L e_{f\gamma}(t)}{\gamma_{eff}}} [r/sec], \qquad (2.33)$$

where N =  $3 \cdot 10^{23}$  is the number of neutrons formed upon explosion of a weapon with a power of 1 Kt; q is the power of the explosion, Kt;  $\varepsilon_{\gamma} = 10.75$  MeV is the energy liberated upon capture of one neutron by a nitrogen nucleus; a is the permeability of the shell;  $\Delta$  is the portion of capture  $\gamma$  radiation resulting from the reaction of capture of neutrons by atmospheric nitrogen (according to [7],  $\Delta = 0.9$ );  $\zeta$  is the share of neutrons from the cloud absorbed by the soil; with an air blast  $\zeta = 0$ , with a surface blast,  $\zeta = 0.5$ ;  $\sigma_{n\gamma} = 0.1 \cdot 10^{-24}$  cm<sup>2</sup> is the cross section of the reaction N<sup>14</sup> (n,  $\gamma$ )N<sup>15</sup>;  $\sigma_{np} = 1.7 \cdot 10^{-24}$  cm<sup>2</sup> is the cross section of the reaction N<sup>14</sup> (n, p)C<sup>14</sup>.

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At the initial moment of the explosion, the size of the cavity is not great. Since the dose power is defined at distances which are large in comparison with the dimensions of the area in which capture of Maxwellian neutrons occurs, we can assume without great error that

$$\cdot e^{\frac{R-L_{eff}(1)}{\hbar e^{ff}}} \approx e^{-\frac{R}{\hbar e^{ff}}}.$$

If we substitute into equation (2.33) the numerical values of the quantities included in it, for an air blast ( $\zeta = 0$ ) with a permeability of the shell of 0.1, the dose power will be

$$P(R, t) = \frac{5.4 \cdot 10^{6} q \,\overline{\alpha} \, e^{-R/2} \text{eff}}{R^{3}} \, e^{-\frac{t}{0.062}} \, r/\text{sec.}$$
(2.34)

If we consider the cavity effect, q in this equation must be replaced by  $q^{1.05}$ .

Figure 2.4 shows graphically the dependence of dose power on time which has passed since the moment of the explosion. The unit of power used is the dose power of fragment radiation at the first moment following the explosion.

The moment of the maximum of dose power depends on the distance and power of the explosion. At short distances, where the shock wave is strong, the maximum occurs earlier than at longer distances, namely occurring when the strong shock wave has passed the point of the detector.

At long distances, the power maximum may occur before passage of the shock wave through the point in question. As the power of the explosion increases, the maximum occurs later (ignoring short distances).

It should be noted that the ratio of ordinates of curves 2 and 3 on Figure 2.4 is equal to the dose power cavity factor  $K_{p}$ .

The dependences of dose power on time shown on Figure 2.4 were calculated without considering the rise of the cloud following the nuclear explosion, which causes an increase in distance from the source to the detector and, consequently, a decrease in  $L_{eff}$  which, in turn,

should lead to a decrease in the maximum in the area of greater time since the moment of the explosion and more rapid decrease in P(t) than that shown on Figure 2.4.



Considering the rise of the radioactive cloud, the total dose power of  $\gamma$  radiation can be calculated using the formula

$$P(t, R) = \frac{5.12 \cdot 10^{9} qu(t) \overline{\alpha}}{R^{2}} e^{-\frac{R - L_{eff}(t)}{\lambda_{eff}}} \left(\frac{R}{R_{t}}\right)^{2} \cdot e^{-\frac{\Delta R}{\lambda_{eff}}} + \frac{5.4 \cdot 10^{\theta} q^{1.05} \overline{\alpha} e^{-R/\lambda_{eff}}}{R^{2}} e^{-\frac{t}{0.062}}.$$
(2.35)

Here  $R_t$  is the distance from the center of the cloud at moment in time t;  $\Delta R$  is the increment in the distance to the center of the cloud following its rise:  $\Delta R = R_t - R$ .

The distance to the center of the cloud at moment in time t can be found on the basis of the rise rate of the radioactive cloud, which was determined in [7]:

$$R_t = \sqrt{(R^2 - H^2) + (H + 16t^2)^2},$$

where H is the height of the explosion over the surface of the earth; R is the distance to the center of the explosion. It should be noted that although the prompt  $\gamma$  radiation amounts to only about 1% of the total dose of  $\gamma$  radiation of an atomic blast, still due to the extremely brief time of its application it can provide a sharp burst of  $\gamma$  radiation. It was noted above that the time of radiation of  $\gamma$  quanta during fission is approximately  $(0.5-2.5)\cdot 10^{-9}$  sec, with 7.8 MeV liberated for each fission event. Considering the thousandfold attenuation in the materials of the weapon, the dose power of  $\gamma$  radiation still exceeds the dose power of fragment and capture  $\gamma$  radiation by several orders of magnitude.

### §. 3. Calculation of $\gamma$ Radiation Doses

If we know the dose rate of fragment  $\gamma$  radiation (2.19), we can calculate the integral dose at a given distance from the center of the blast:

$$D = \int_{0}^{t_{\text{eff}}} P dt = \frac{1,48 \cdot 10^{-5} \,\mu_c \,(E_{\gamma}) \,\overline{\alpha} \, N \mathrm{e}^{-R/\lambda} \mathrm{eff}}{4\pi R^2 \cdot 10^4} \times \\ \times \int_{0}^{t_{\text{eff}}} u(t) \,\mathrm{e}^{\frac{L_{\text{eff}} \,(q,\,t)}{\lambda \mathrm{eff}}} \,dt, \qquad (2.36)$$

where

as

$$Nu(t) = g(t) \text{ MeV/sec},$$

 $\mu_e$  is the electron conversion factor;  $t_{eff}$  is the effective time during which the effect of  $\gamma$  radiation must be considered (due to radioactive decay of fragments and the rise of the cloud, the intensity of  $\gamma$  radiation drops rapidly), which is taken as 10 sec.

The integral included in expression (2.36) can be represented

$$\int_{0}^{t} u(t) e^{\frac{1}{t} \frac{eff(q, t)}{t}} dt = e^{\frac{\overline{L}(q)}{\lambda eff}} W, \qquad (2.37)$$

where

$$W = \int_{0}^{10 \, \text{sec}} u(t) \, dt = 2,05 \text{ MeV/fission}$$

Substituting the numerical values of coefficients into (2.36) and (2.37), we produce

$$D = \frac{2 \cdot 10^{9} q e^{\frac{\bar{L}}{\lambda}} e^{-R/\lambda} e^{44}}{R^{2}}.$$
 (2.38)

An estimate of  $\overline{L}$ , the mean value of the cavity factor over the entire time of emission of quanta, can be made under the following assumptions: it is assumed that  $\overline{L}$  is similar to the final radius of the fireball  $(\overline{L} \simeq r_{fb})$ . This approximation is more precise, the less the time required for formation of the final dimensions of the cavity in comparison with the time of emission of  $\gamma$  quanta (since the expression with  $\overline{L} \simeq r_{fb}$ is eliminated from the integral with respect to time as a constant

quantity), i. e. it is more precise, the less the power of the explosion.

This approximation is clearly inaccurate for high power explosions, since a large portion of the  $\gamma$  radiation is liberated before the moment when the fireball reaches its final size, i. e. with a high power explosion  $\overline{L} < r_{fb}$ .

The final radius of the fireball can be estimated on the basis of the energy which it contains. It follows from the theory presented in [5] that the energy of the heated air is about 50% of the energy of the explosion. Then

$$\frac{4}{3}\pi r_{fg}^{3}\varepsilon = 0,5q;$$

$$\overline{L} \approx r_{fb} \frac{3}{4} \sqrt{\frac{0,5q}{4}}, \qquad (2.39)$$

where  $\epsilon$  is the energy per unit volume of the fireball.

If we ignore the increase in heat capacity with temperature and dissociation of air molecules, the energy of  $1 \ g$  of air will be

$$\frac{5R(T-T_{\rm H})}{28,8} = 0.35(T-T_{\rm H}) \quad {\rm cal/g.}$$

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ignoring the value of  $T_{\rm H}$  in comparison with T, we produce

$$\varepsilon = \rho_0 0.35T = \frac{\rho_0 T_0}{\rho_0} \frac{\rho_{\mu}}{T_{\rm H}} 0.35T = 1.23 \cdot 10^{-1} \rho_{\mu} \text{ cal/cm}^3, \qquad (2.40)$$

where  $p_{\rm H}$  is the pressure of the air before the explosion;  $T_{\rm H}$  and T are the temperatures of the air before and after the explosion;  $\rho_0$ ,  $T_0$  and  $p_0$  are the density, temperature and pressure of the air under normal conditions.

Substituting the value of  $\varepsilon$  into formula (2.39), we produce

$$\overline{L} = 100 \sqrt[3]{\frac{q}{p_{\rm H}}} \, \mathrm{M}. \tag{2.41}$$

Substituting expression (2.41) into formula (2.38), we can write

$$D = \frac{2 \cdot 10^{9} q e^{\frac{100}{\lambda} e f f}}{R^{3}} e^{-R/\lambda} e^{f f} \qquad (2.42)$$

For the case when  $p_H$  = 1 atm,  $T_H$  = 298° K,  $\lambda_{eff}$  is 255 m.

$$D = \frac{2 \cdot 10^{9} q \, \mathrm{e}^{0.4} \sqrt[3]{q} \, \mathrm{e}^{-R/255}}{R^{2}} = \frac{2 \cdot 10^{9} q \cdot 10^{0.174} \sqrt[3]{q}}{R^{2}} \, \mathrm{e}^{-R/255} \, p. \tag{2.43}$$

In expressions (2.38), (2.42) and (2.43), the factor

$$\exp\left(\overline{L}/\lambda_{eff}\right) = \exp\left(\frac{\frac{3}{100\sqrt{q/p_{H}}}}{\lambda_{eff}}\right),$$

which shows the number of times by which the dose is increased due to the effect of the shock wave, is specific. This factor is called the dose cavity factor  $K_d$ . In formula (2.43), the dose cavity factor is  $\frac{3}{d}$ 

$$K_{q} = e^{0.4 \sqrt[3]{q}}$$
 (2.44)

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Expression (2.43) is related to points lying outside the fireball, i. e. outside the cavity. Within the cavity, the absorption of  $\gamma$ radiation is zero and the dose changes in inverse proportion to the square of the distance from the center of the explosion.

The use of the quantity  $r_{fb}$  as the mean radius of the cavity  $\overline{L}$  in calculations produces approximate results, suitable only for the explosion of low-power weapons.

In the case of high explosion power, we cannot expect satisfactory accuracy of calculation results using formula (2.43), since  $\overline{L}$  is clearly less than  $r_{fb}$ . Furthermore, with increasing energy of the explosion,  $\overline{L}$  should increase more slowly than  $\approx \sqrt[3]{q}$ . With powerful explosions, the shock wave is strong throughout the time of emission of most of the  $\gamma$  radiation. According to the theory presented in [7], the radius of the leading edge of the strong shock wave at each moment in time is approximately proportional to  $\sqrt[5]{q}$ ; therefore, it must be considered that actually  $\overline{L}$  increases approximately as  $\sqrt[5]{q}$ .

Detailed analysis of the problem of doses leads to numerical values which can be described by various interpolation formulas. For example, the formula suggested in [3]:

$$D = 6.6 \cdot 10^{8} q \,\mathrm{e}^{0.925 \sqrt[5]{q} - 6.25 \cdot 10^{-5} q} \cdot \frac{\mathrm{e}^{-R_{\mathrm{Po}}/300}}{R^{2}} [p], \qquad (2.45)$$

is written in a form corresponding to the physical representation of the action of a shock wave.

We can also use a simpler expression:

$$D = 1, 4 \cdot 10^9 q \left(1 + 0, 2q^{0.65}\right) \frac{e^{-R_{\rm Po}/300}}{R^2} [p].$$
(2.46)

With a surface blast, the radius of the cavity is greater, corresponding to the radius of a cavity with an air blast of twice the power. Therefore, the dose with a surface blast is higher than with an air blast of a bomb of the same power:

$$D = 2.8 \cdot 10^9 q \left(1 + 0.3 q^{0.65}\right) \frac{e^{-R_{\rm Po}/300}}{R^2} [p].$$
 (2.47)

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The doubled volume of the cavity with a surface blast in comparison to an air blast is retained only during the time of emission of  $\gamma$  radiation, since the hemispherical area of heated air separates from the ground and becomes spherical as time passes.

However, we can calculate the dose with a surface blast using formula (2.47) with sufficient approximate accuracy.

Figure 2.5 shows the values of doses at various distances from the epicentr of the explosion at Hiroshima [8]. From the figure we can see that there is a hard component in the  $\gamma$  radiation spectrum: the slope of the curve log DR<sup>2</sup> = f(R) decreases with increasing distance. In the 600-1600 m distance interval,  $\lambda_{eff} = 282$  m, in the 2000-2600 m interval,  $\lambda_{eff} = 393$  m. This last value of  $\lambda_{eff}$  corresponds to a mean  $\gamma$  radiation energy  $\simeq$  6 MeV. The effect of capture  $\gamma$  radiation predominates over this sector.

The dose of the hard component, according to Figure 2.5, can be expressed by the empirical formula

$$D_{h} = \frac{5.5 \cdot 10^{8} q \, e^{-R/393}}{R^{2}} [p]. \qquad (2.48)$$

Let us estimate the dose of the hard component, assuming that it is formed only by capture radiation.

The dose of capture  $\gamma$  radiation can be found by integrating expression (2.33) or (2.34) with respect to time

$$D_{h} = \frac{5.4 \cdot 10^{9} q \overline{\alpha} e^{-R/\lambda} e^{tf}}{R^{2}} \int_{0}^{\infty} e^{-t/0.062} dt [p]. \qquad (2.49)$$

Integrating, we produce

$$D_{h} = \frac{1,48 \cdot 10^{-b} N_{\bar{e}\alpha} \mu_{e}(E_{0\gamma})}{4\pi R^{2} \cdot 10^{4}} e^{-R/\lambda} \epsilon^{4} f_{p}[p]. \qquad (2.50)$$

Since the main portion of capture radiation should be radiated within 0.1 sec, when the radius of the shock wave is still small, the cavity factor is not considered. Substituting the values of quantities in (2.50):

$$\begin{split} N &= 1.45 \cdot 1023 \text{ q (for } \varepsilon = 1 \text{ MeV/fission} \\ \hline \alpha &= 1.4 \\ \lambda_{eff} &= 375 \text{ m} \end{split} \right\} 1000 \text{ m} < R < 2000 \text{ m} \\ \mu_e(E_{0\gamma}) &= 1.9 \cdot 10^{-5} \text{ cm}^{-1} \text{ (for } E_{0\gamma} = 6 \text{ MeV}), \end{split}$$

we produce:

$$D_{h} = \frac{4,55 \cdot 10^{8} q \, \mathrm{e}^{-R/375}}{R^{2}} \, [p].$$

(2.51)



Figure 2.5. Spatial Distribution of Doses of  $\gamma$  Radiation During Explosion of Atomic Bomb at Hiroshima:

•, experimental data; 1,  $\lambda$  = 252 m; 2,  $\lambda$  = 282 m; 3,  $\lambda$  = 393 m

As we can see, the calculated estimate agrees satisfactorily with empirical formula (2.49). This confirms our assumption that the hard component of  $\gamma$  radiation observed in practice results from the capture of neutrons by nitrogen in the air.

The influence of the cavity on the dose of the hard component can be felt during strong explosions. It is always significantly less than the influence of the cavity on the dose of the fragment component, first of all because the hard component is radiated over a short period of time, when the cavity is small, and secondly because the influence of the cavity decreases with increasing hardness of the radiation, i. e. with increasing  $\lambda_{pff}$ .

In the case of hydrogen bombs, the number of fragments is determined not by the entire energy of the explosion, but only by the portion involving the fission reaction, but the size of the cavity is determined by the entire energy of the explosion. Therefore, in the case of a thermonuclear explosion, the total power of the explosion is included only in the expression for the cavity factor.

In certain cases, there may be interest in the dependence of the accummulated  $\gamma$  radiation dose on time. Figure 2.6, based on data from [3], shows the rate of accummulation of the dose upon explosion of nuclear weapons of various powers.



Figure 2.6. Rate of  $\gamma$  Radiation Dose Accummulation Upon Explosion of Nuclear Weapons With Powers of: a, q = 2 Kt; b, q = 200 Kt; c, q = 20,000 Kt.

We can see from the figure that the more powerful the explosion, the longer the dose continues to accummulate, since the cavity created by the shock wave continues to grow for a longer period of time.

The greater the distance, the more rapidly the dose accummulates, since the role of capture radiation increases.

### § 4. Spatial Distribution of Neutrons of Nuclear Explosion

The neutrons leaving the shell of the weapon interact with the nuclei of the surrounding medium: they are scattered, moderated and absorbed at various distances from the point of the explosion.

Let us first study the distribution of slow neutrons. The spatial distribution of slow neutrons is determined by the process of moderation and diffusion of neutrons in air. During the very first stage of the explosion, the temperature of the material of the bomb is about  $10^{7\circ}$  K, corresponding to an energy of about  $10^{3}$  eV, to which energy level the neutrons are absorbed as they achieve thermal equilibrium with the material. The neutrons subsequently cool together with the material of the shell. A few microseconds after the explosion, the temperature of the shell can be estimated as some hundreds of thousands or millions of degrees. Therefore, considering the lower temperatures in the outer layers of the shell, it is probably correct to consider its temperature as approximately  $5 \cdot 10^5 - 10^{6\circ}$  K as the neutrons fly outward, corresponding to a neutron energy of 50-100 eV. Thus, it can be considered that the energy of Maxwellian neutrons at the beginning of diffusion is some tens of electron volts.

One of the characteristics of air as a moderating medium is that the probability of absorption of low energy neutrons in air is high due to the high capture cross section of atmospheric nitrogen. Therefore, a neutron upon moderation has a high probability of being captured before it is moderated to thermal velocities. It can be considered approximately that a neutron in air is moderated on the average to an energy of about 0.2 eV and is captured at this energy by nitrogen [7].

Figure 2.7 shows the dependence of the flux of slow neutrons on the distance to the center of an explosion performed at high altitude, when the diffusion of neutrons occurs in air and the perturbing influence of the earth can be ignored [7]. As we can see from Figure 2.7, at distances less than 500 m, the flux of slow neutrons forming a cloud around the center of the explosion is determined primarily by neutrons leaving the shell as Maxwellian neutrons. At distances of greater than 500 m, the contribution of Maxwellian neutrons can be ignored. In this area the spatial distribution of slow neutrons is determined by neutrons leaving the shell during the process of moderation, while at great distances the primary contribution is that of neutrons having high initial energies. This flux of slow neutrons  $\Phi_{\rm T}({\rm R})$  at distances of greater than 500 m is described rather well by

the exponential formula

$$\Phi_{\mathbf{r}}(R) = C \, \mathrm{e}^{-R/\hbar} \,, \tag{2.52}$$

where C is a constant,  $\lambda = 135$  m is the relaxation length. This fact is determined by the fact that the value of  $\Phi_{T}(R)$ , as was shown in

[7], is the sum of Gaussian distributions created by various sectors of the initial neutron spectrum. At distances of less than 400-500 m, the neutron flux is determined by Maxwellian neutrons alone, since in this range of distances the distribution is near Gaussian.

Therefore, with any value of permeability of the shell of the weapon, the curve characterizing the distribution of slow neutron flux for distances greater than 500 m will be approximately described by formula (2.52), with the single difference that constant C changes for a weapon of the same power. At short distances, the more or less rapid rise in neutron flux will be determined by the permeability of the shell.

If an explosion is performed at the surface of the ground or at a slight altitude over the ground, the ground may have a significant perturbing influence on the distribution of neutrons near its surface.

Figure 2.7 shows the distribution of the flux of slow neutrons in air at various distances from the center of a surface blast with a power of 20 Kt, calculated [7] considering the influence of the ground; the permeability of the shell is assumed equal to 0.1. As we can see, consideration of the influence of the ground in the solution of the problem of diffusion of neutrons in the air did not cause significant distortion of the distribution function of the slow neutron flux at distances R > 100 m. Near the point of the explosion at R < 100 m, the influence of the ground causes a significant increase in the flux of slow neutrons in comparison with diffusion in air. This is explained by the fact that the presence of the denser medium increases the concentration of slow neutrons, whereas in the air these neutrons would propagate to greater distances.



 $\phi_{T}(R)$ , neutrons/cm<sup>2</sup>

Figure 2.7. Spatial Distribution of Slow Neutrons With a Nuclear Explosion: 1, total flux of slow neutrons, explosion on surface of the earth; 2, total flux of slow neutrons, explosion in the air; 3, flux of slow neutrons generated by neutrons with energies over Maxwellian; 4, flux of slow neutrons generated by fast neutrons in thermonuclear explosion

At great distances, the earth acts as a neutron sink, as a result of which the flux of neutrons changes in comparison with the flux in an infinite air medium. At distances exceeding 150-200 m, the disturbing influence of the earth on the propagation of slow neutrons is slight, and therefore without great error we can use the results of calculation for an infinite air medium.

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The numerical value of the preexponential term in formula (2.52) can be determined from Figure 2.7, which yields

$$\Phi_{r}(R) = 2,81 \cdot 10^{14} qa \, e^{-R/135} \text{ neutrons/cm}^2 (R > 500 \text{ m}), \qquad (2.53)$$

where a is the permeability factor of the shell.

The fact that formula (2.53) does not contain the term  $1/R^2$  has no physical sense. It would be equally correct to use the term 1/R or  $1/R^2$ . Then the value of the attenuation distance would not be 135 m.

The flux of slow neutrons for distances less than 500 m can be represented by the formula

$$\Phi_{\tau}(R) = 3,74 \cdot 10^{14} q (1-a) e^{-\frac{R}{3 \cdot 10^4}}.$$
 (2.54)

We should also discuss the problem of the spatial distribution of slow neutrons formed as a result of moderation and diffusion of delayed neutrons in the air. Figure 2.8 shows the distribution of slow (0.2 eV) neutrons, generated by delayed neutrons from a nuclear explosion with a maximum cavity radius of  $r_{\rm fb}$  and considering the rise of the fireball. The data were produced by calculation [7] for an explosion of 1 Kt. In order to determine the flux of slow neutrons at a given distance from the explosion of a weapon of q Kt power, the value of  $\Phi_{\rm T}(R)$ , determined from the graph of Figure 2.8 must be multiplied by the explosion power q, in correspondence with the calculated value of  $\overline{L} = 100 \sqrt[3]{q}$  for the weapon in question.

Let us see what fluxes of slow delayed neutrons are possible with explosions of various powers. For example, for q = 400 Kt ( $\overline{L} = 800$  m) with a shell permeability of 0.1, at a range of 1000 m the flux of slow neutrons generated by prompt fission neutrons, in correspondence with equation (2.54), is  $2.7 \cdot 10^{11}$  neutrons/cm<sup>2</sup>. With this same power level of the explosion, the flux of slow delayed neutrons is also about  $2.7 \cdot 10^{11}$  neutrons/cm<sup>2</sup>. In other words, at q = 400 Kt and R = 1000 m, the share of slow delayed neutrons is half the total quantity of slow neutrons at this range. With increasing explosion power, the percentage of delayed neutrons may be even higher. As a result of the effect of the shock wave, the relative quantity of delayed neutrons increases, but not in proportion to the explosion power.

With a thermonuclear blast, most of the neutrons leaving the weapon have high energy. Still, at distances of over 600-700 m, these neutrons form a field of slow neutrons. Figure 2.7 (curve 4) shows the spatial distribution of the flux of slow neutrons ( $E_n < 0.625$  eV) in an infinite air medium upon explosion of a thermonuclear weapon [9].





The spectrum of neutrons of the source was assumed to consist 17% of fission neutrons with energies over 0.4 MeV and 83% of neutrons, evenly distributed in the interval from 12 to 16 MeV. The authors of this work believe that this spectrum of the source is characteristic for thermonuclear explosions, although the exclusion of neutrons with energies of below 0.4 MeV from the calculation results in a reduction in the number of slow neutrons in the near zone. We must note that the

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spatial distribution of slow neutrons is insensitive to changes in the initial spectrum of neutrons over a rather broad range.

The neutrons leaving the shell of the weapon during the process of moderation have energies of from 100 eV to several MeV. Before the neutrons are moderated, they may travel a considerable distance in the air, the mean value of which is related to the energy of the neutrons leaving the shell. Thus, there are both slow and fast neutrons in the region of the blast.

Figure 2.9 shows the results of calculation of the spatial distribution of neutrons of point isotropic sources with various initial radiation spectra in an infinite air medium by the Monte Carlo method. The initial neutron energy used in the calculations correspond to the fission spectrum [10], to a source of monoenergetic neutrons with an energy of 14 MeV [10], and to the spectrum of neutrons from a thermonuclear blast [9]. In all cases, the results were normalized for a source with a power corresponding to one neutron/sec.



Figure 2.9. Spatial Distribution of Neutrons of Point Isotropic Source in Infinite Air Medium: 1, fission source. Flux and dose calculated for neutrons with  $E_n \ge 0.1$  MeV; 2, source of neutrons with initial energy 14 MeV. Flux and dose calculated for neutrons with  $E_n \ge 0.1$  MeV; 3, source of neutrons with spectrum modeling thermonuclear explosion. Flux and dose of neutrons calculated for neutrons with  $E_n \ge 0.5$  MeV.

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.

When there is a division boundary between two media -- earth and air -- its perturbing influence cannot be ignored. Detailed investigation of the influence of the earth on the spatial distribution of the dose or flux of neutrons during a nuclear blast was performed in [11]. The method of calculation suggested in this work consists of determining the influence of the earth on the distribution of the first collisions of neutrons with air nuclei near the source and last collisions near the detector. It is based on the assumption that the number of multipliscattered neutrons striking the detector from long distances is in the first approximation directly proportional to the number of first collisions of neutrons with nuclei of air near the source and last collisions near the detector.



Figure 2.10. Geometry for Calculation of Distribution of Radiation With Air-Ground Division Boundary

Let us analyze a point isotropic source of neutrons of unit power, located at height H over the surface of the earth (Figure 2.10). The number of first scattering events which occur per unit volume of air dv will be

$$dN = \frac{\sum_{s} e^{-\sum_{t} \sqrt{x^{s} + y^{s}}}}{4\pi (x^{2} + y^{2})} dv, \qquad (2.55)$$

where  $\Sigma_s$  is the macroscopic cross section of scattering of fast neutrons in air;  $\Sigma_t$  is the total cross section of the interaction of neutrons with air;  $dv = x dx dy d\phi$ .

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For fast neutrons, it is justified to assume that  $\Sigma_s = \Sigma_t$  [12]. Then the total number of first scattering collisions in the air near the source will be

$$N(H_s) = \frac{\sum_t \int_{-H_s}^{\infty} \int_{0}^{\infty} \frac{xe^{-\sum_t \sqrt{x^2 + y^2}}}{x^2 + y^2} dx dy.$$
(2.56)

For an infinite air medium  $N(H_s) = 1$ . Thus, the value of  $N(H_s)$ in the first approximation characterizes the portion of neutrons of the source which are scattered at least once in the air. Obviously, the best approximation is weighing of each first collision for the probability that the scattered neutron will have a later collision in the air.

If we assume that the first scattering event is isotropic in the laboratory system of coordinates, the probability that a neutron scattered once will not be absorbed by the earth is proportional to  $N(H_c)$ .

Thus, the number of first collisions of neutrons in the air, weighed on the basis of their probability of undergoing further scattering, is determined as follows:

$$N'(H_s) = \frac{\Sigma_t}{2} \int_{-H_s}^{\infty} \int_{0}^{\infty} \frac{f(y) \rho e^{-\Sigma_t \sqrt{x^2 + y^2}}}{x^2 + y^2} dx dy, \qquad (2.57)$$

where  $f(y) = N(H_s + y)$  is the weight function of centers of first collision.

In concluding formula (2.56), it was assumed that the earth is an absolutely black body. Actually, a portion of the neutrons which strike the earth without undergoing scattering in the air may be reflected back and, after further interactions, may strike the detector. The albedo of fast neutrons from the earth (see § 3 Chapter 4), depends on the incident angle, read from the normal as  $\cos^{2/3} \beta$ . If the integral albedo is equal to  $A_n^{\pi}$  for normal incidence, the total

number of neutrons scattered from the earth will be

$$N_{r}(H_{s}) = \frac{A_{n}^{n}}{2} \int_{0}^{\infty} \frac{xH_{s}^{*/s}e^{-\Sigma_{l}}\sqrt{x^{*}+H_{s}^{2}}}{(x^{2}+H_{s}^{2})^{*/s}} dx. \qquad (2.58)$$

Thus, the effective share of neutrons of the source experiencing single scattering near the source is expressed by the dependence

$$f(H_{2}) = N'(H_{2}) + N_{r}(H_{2}).$$
(2.59)

In order to estimate the influence of the earth near the detector, we assume that the centers of the last collisions are distributed evenly, and scattering is considered isotropic in the laboratory system of coordinates. The first assumption is rather obvious. Actually, the free path length of a neutron at distances of over 200 m from the source of fission, i. e. in the area where evenness of the spectrum of neutrons scattered in the air is observed, is about 72 m, while the effective length of attenuation of the flux is about 190 m. Thus, the density of neutrons does not change significantly in the area of the detector.

Although the assumption of isotropic scattering of neutrons is not usable for strict calculation of neutron transfer in an air medium, it should not have any significant influence on an estimation of the disturbing influence of the earth.

If the flux of primary neutrons near the source is considered the unit flux, the density of collisions at any point in the medium will be  $\Sigma_t$  Thus, following the geometry of Figure 2.10, the flux of neutrons scattered in the air with an absolutely black surface of the earth at the point of the detector (at height H<sub>d</sub> over the earth) will be

$$N(H_{d}) = \frac{\Sigma_{t}}{2} \int_{-H_{d}}^{\infty} \int_{0}^{\infty} \frac{x e^{-\Sigma_{t} \sqrt{x^{2} + y^{4}}}}{x^{2} + y^{2}} dx dy.$$
(2.60)

Since the reflection of neutrons from the earth has a significant influence on the effect at the detector only when it is located at a low altitude, let us assume that the neutrons scattered from the ground enter the detector without interacting with the nuclei of the air.

If  $A'_{H}$  is the share of neutrons reflected from the ground with isotropic distribution of incident neutrons (the flux of which in this case will be 1/2), with cosinusoidal distribution of neutrons reflected from the ground to the detector, they form the flux

$$N_{\rm r}(H_{\rm cl}) = A_{\rm fr}' \int_{0}^{\infty} \frac{xH_{\rm cl}}{(x^2 + H_{\rm cl}^2)^{s/2}} dx. \qquad (2.61)$$

The sum of the two quantities  $N(H_d)$  and  $N_g^N(H_d)$  can be looked upon as the fraction  $f(H_d)$  of the dose or flux observed in an infinite

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air medium, which will be recorded by a detector located at height  $H_d$  over the earth:

$$f(H_d) = N(H_d) + N_g(H_d). \tag{2.62}$$

If D(R) is the dose (flux) of neutrons from the explosion at distance R from the center of a nuclear explosion in an infinite air medium, the dose with the source at height  $H_s$  and the detector at

 ${\rm H}_{\rm d}$  over the surface of the earth can be expressed as

$$D(R, H_s, H) = D(R)f(H_s)f(H).$$
(2.63)

The factors  $f(H_s)$  and  $f(H_d)$  calculated on the basis of the discussions presented above are shown in Table 2.3. The calculation functions  $f(H_s)$  and  $f(H_d)$  presented in this table were checked experimentally with a suspended reactor. Good agreement was produced for source-detector distances of over 260 m [11].

Table 2.3

CORRECTION FACTORS  $f(H_s)$  AND  $f(H_d)$  FOR CONSIDERATION OF INFLUENCE OF THE EARTH ON SPATIAL DISTRIBUTION OF FAST NUCLEAR EXPLOSION NEUTRONS

	1	1 (H <sub>s</sub> )			$f(H_d)$		
		Expe	rimental		Exper	imental	
Н, м	Calcu- lated	R = 260 M	R = 350 м	Calcu- lated	R = 260 M	R = 350 м	
0 10 15 29 35 44 53	0,434 0,612 0,715 0,746 0,806	0,586 0,717 0,755 0,806	0,701 0,746 0,806	0,697 0,720 0,840 0,913 0,931 	0,714 0,841 0,884 0,931	0,828 0,858 0,932 0,931	

Figure 2.11 shows the results of measurements [13] of the spatial distribution of the flux and dose of fast neutrons from a nuclear explosion. The experimental explosion was performed at an altitude of 215 m over the surface of the earth, while detectors were placed on the ground. As we can see from Figure 2.11, the spatial distribution of fast neutrons can be approximated with sufficient accuracy of an exponential dependence. The authors of [13] suggested that the dose of

neutrons of the nuclear explosion be calculated using the empirical formula

$$D(R) = \frac{4.42q \cdot 10^{10}}{R^2} e^{-\frac{p}{p_0} \frac{R}{184}} \text{ rad.}$$
(2.64)

A similar expression can be used to calculate the flux of fast neutrons. For example, the neutron flux with  $E_n > 0.75$  MeV can be calculated using the formula

$$\Phi(R) = \frac{1,05 \cdot 10^{19} q}{R^2} e^{-\frac{P}{60} \frac{R}{190}} \text{ neutrons/cm}^2.$$
(2.65)

Although the number of delayed neutrons liberated upon fission is not great, the share of delayed neutrons acting at various distances from the center of the explosion may be comparable to the share of prompt fission neutrons. This results from the fact that as they are moderated in the shell, many of the fission neutrons reach the temperature of equilibrium with its atoms and into the air with an energy of about 1 KeV. As was shown above, these neutrons do not penetrate to distances of over 400 m. Therefore, at long distances, the neutron dose consists only of those fission neutrons which were not moderated to low energies upon passage through the shell.

In contrast to this, the delayed neutrons are liberated after the shell has flown apart. For example, for a shell with a permeability of 0.1, the share of delayed neutrons observed at distances of over 500 m should be greater than their initial share by approximately 10 times, i. e. should represent about 7% during fission of  $U^{235}$ .

The second factor facilitatinggthe transmission of delayed neutrons over great distances is the presence of the cavity with low air density around the explosion center, the radius of which, as was demonstrated above, increases with increasing explosion energy. As the delayed neutrons pass through this cavity, they are almost unattenuated by the air, while the prompt neutrons propagate before formation of the cavity. Therefore, the presence of the cavity increases the share of delayed neutrons at great distances. The number of prompt and delayed neutrons formed is proportional to the power of the explosion; however, due to the formation of the cavity, the share of delayed neutrons at great distances will increase with increasing power of the explosion. Thus, with explosions of sufficient power, it may be that, in spite of the small share of formation of delayed neutrons, a significant portion of the total neutron dose results from these delayed neutrons.



Figure 2.11. Fluxes of Neutrons Over Threshhold Value and Doses of Fast Neutrons as Functions of Distance to Center of Explosion: flux of neutrons  $E_n > 5$  eV: curve 1, calculated; •, experimental. Flux of neutrons  $E_n > 0.75$  MeV: curve 2, calculated;  $\Delta$ , experimental. Flux of neutrons  $E_n > 1.5$  MeV: curve 3, calculated;  $\square$ , experimental. Flux of neutrons  $E_n > 2.5$  MeV: curve 4, calculated;  $\nabla$ , experimental. Doses of fast neutrons: curve 5, calculated; o, experimental.

The total quantity of delayed neutrons recorded at a certain distance from the center of the explosion is determined by the number of neutrons formed, the effective size of the cavity and the rate of ascent of the cloud offfragments. Although the formation of the cavity increases the significance of delayed neutrons, the ascent of the cloud decreases their influence. Consequently, the number of delayed neutrons recorded for explosions of various powers will not be linearly related to the energy of the explosion, as is the case for fission neutrons. The number of delayed neutrons radiated per unit time at time t following the explosion is expressed by the formula

$$N(t) = N_0 \sum_i \lambda_i a_i e^{-\lambda_i t}, \qquad (2.66)$$

where  $N_0$  is the number of fission events;  $a_i$  is the total yield of delayed neutrons per fission event within the limits of the ith group;  $\lambda_i$  is the decay constant of neutrons in the ith group.

The values of the quantities included in this formula are presented in Table 1.1.

The total number of delayed neutrons liberated from the fragments formed in the explosion is

$$N_0 \sum_i a_i \int_0^\infty \lambda_i e^{-\lambda_i t} dt = N_0 \sum_i a_i.$$
(2.67)

This quantity of neutrons is liberated in a few tens of seconds following the explosion.

Since the delayed neutrons form groups with various half life periods, the effective period increases with passing time. The resulting decay rule can be written as the formula

$$S(t) = A + Be^{-Ct}$$
 neutrons/(sec fission), (2.68)

where t is the time following fission, sec.

The values of the coefficients included in this formula are:

	Uzaa	U238	Pu <sup>239</sup>
A	$8, 4 \cdot 10^{-6}$	$1,15\cdot10^{-5}$	3,87.10-6
B	7, 31 \cdot 10^{-3}	$3,08\cdot10^{-2}$	2,62.10-3
C	0, 542	0,663	0,544

The power of the dose of delayed neutrons at moment in time t at distance R from the center of the explosion, considering the effect of the cavity and the ascent of the cloud, can be written as:

$$P_{n,3}(R, t) = \frac{3 \cdot 10^{23} qS(t) r_1(E_n) e^{-\frac{R-L(t)}{\hbar} \epsilon_{ff}}}{4\pi R^2} \times (2.69)$$
$$\times \left(\frac{R}{R'}\right)^2 e^{-\frac{\Delta R}{\hbar} \epsilon_{ff}} rad/sec.$$

Here  $n(E_n)$  is the transition factor of the neutron flux to dose. It can be considered without great error that the energy of delayed neutrons, averaged over all groups, is 0.55 MeV. For neutrons of this energy, coefficient  $n(E_n)$  is  $2.15 \cdot 10^{-9}$  rad/(neutrons  $\cdot cm^2$ ) [14],  $\lambda_{eff}$ is the effective relaxation length of delayed neutrons in the air, which is about 157 m. The values of R' and  $\Delta R$  are calculated as for  $\gamma$  quanta (see § 3 of this chapter).

The dependence of the size of the cavity  $L_{eff}(t)$  on time for

explosions of various powers is shown on Figure 2.3. As we can see from Figure 2.3, for explosions with the power of < 2000 Kt at moment in time t = 2 sec the radius of the cavity is at least 80% of its final size. Since in this time only 20% of all delayed neutrons have been produced, for an explosion of between 1500 and 200 Kt yield, the radius of the cavity can be assumed constant over the entire time of For precise calculation action of the source of delayed neutrons. of the radius of the cavity, expression (2.31) should be used. However, it should be kept in mind that at high explosion powers, the value of k in this expression becomes less than 100, since a certain portion of the neutrons is emitted during the process of formation of the cavity. The total dose of delayed neutrons can be produced by integrating expression (2.69) with respect to time. Calculations show that with nuclear explosions with powers of up to 100 Kt, delayed neutrons can be ignored. For weapons with powers of 100 to 1000 Kt, the dose created by the delayed neutrons must be considered. With explosions of 1000 Kt and higher, at distances where the neutrons can be looked upon as a damaging factor, their dose is created primarily by the delayed neutrons.

§ 5. Spectral-Angular Distributions of Neutrons Scattered in the Air

The spectrum of the neutrons of a nuclear explosion in the air is determined by the form of the initial spectrum of the neutrons leaving the shell of the weapon, and the nuclear characteristics of the medium in which the propagation of the neutrons occurs. It was demonstrated above that the relative form of the spectrum of the neutrons leaving the shell of the weapon during a nuclear explosion based on the principle of fission is almost independent of the design of the weapon.

It follows from experiments and calculation that, beginning at a certain distance F, the spectrum of the neutrons in the air changes relatively little. Thus, in [13] the behavior of individual energy groups of neutrons was studied as a function of distance R during explosion of a nuclear weapon. The results of the investigations are shown on Figure 2.11.

In [10], the integral neutron spectra by angles at various distances from the center of a nuclear explosion in the air were calculated by the Monte Carlo method. The nuclear explosion in this case was modeled by a point isotropic source of fission neutrons in an infinite air medium. The results produced in this work are shown on Figure 2.12. The nature of the energy distributions shown on the figure indicates th slight transformation of the spectrum of the neutrons from the nuclear explosion with increasing distance from its center.

This same figures shows for comparison data on the spectral distribution of nuclear explosion neutrons produced by direct measurement at 1180 m from the center of an experimental blast [13, 15] performed at the Nevada Test Site (USA). The experimental and calculated spectra for the distance of 1180 m were normalized for an energy of 1 MeV. As we can see, the relative course of the experimental and calculation curves correspond satisfactorily. This indicates that the presence of the earth-air division boundary has little influence on deformation of the spectrum of the nuclear explosion neutrons in comparison to the spectrum in an infinite air medium. The data shown on Figure 2.12 characterize the spectral distribution of fast neutrons of a nuclear explosion, the energy of which exceeds 0.1 MeV. As concerns slow and intermediate neutrons, we can gain an idea of their energy distribution from Figure 2.13, showing the results produced by the method of moments [16] for a point fission source in an infinite air medium.

A comparison with the fission spectrum gives us an idea of the degree of deformation of the spectrum of nuclear explosion neutrons scattered in air in comparison with the spectrum of neutrons produced by fission.

During a thermonuclear explosion, in addition to the fission neutrons, neutrons with an energy of 14 MeV are generated. During the process of inelastic scattering of these neutrons in the shell of the weapon, neutrons may be produced with a most probable energy of about 4 MeV. Figure 2.14 shows the spectra of neutrons scattered in the air, calculated [10] for point isotropic sources with initial neutron



M (1580)

R=220 M N019

Medium at Various Distances From the Figure 2.12. Source Isotropic Fission Source in Infinite Air Spectra of Neutrons of Point

Figure 2.13. Nature of Spectral Distri-bution of Slow and Intermediate Neutrons Spectrum of Nuclear Explosion. ----- Fission

10

3

10-2 10-1 100 6 ğ

-74-

~2970M



energies of 4 and 14 MeV. In each case, the power of the source was assumed equal to one neutron/sec.

Figure 2.14. Spectra of Neutrons of Point Isotropic Sources With Initial Energies 4 MeV (- - - - -) and 14 MeV (------) in an Infinite Air Medium

In [9], the energy distributions of neutrons of a thermonuclear explosion with an initial spectrum consisting of fission neutrons (17%) and neutrons evenly distributed in the 12-16 MeV energy interval are calculated. The results produced in this work are shown on Figure 2.15. As we can see from this figure, with increasing distance to the center of a thermonuclear explosion, significant softening of the spectrum in the 10-16 MeV area occurs. However, below 8-9 MeV, the integral spectrum of fast neutrons is deformed slightly. This same figure shows data relating to the slow and intermediate neutrons. They allow us to conclude that in the range of energies from  $10^{-4}$  to 0.5 MeV, the spectral density of neutrons follows the rule  $1/E_m^m$ , where the

exponent m is either equal to or slightly greater than unity, while the fluctuztions of its value with increasing distance to the center of the nuclear blast are very slight.



Figure 2.15. Spectra of Neutrons of Thermonuclear Explosion in Infinite Air Medium at Various Distances From the Point of the Explosion: 1, 64 m; 2, 214 m; 3, 385 m; 4, 810 m; 5, 1150 m; 6, 1330 m; 7, 1500 m; 8, 1670 m; 9, 2000 m

In addition to the spectral characteristics of radiation and their spatial distribution, the design of shielding from penetrating radiation of a nuclear explosion also requires a knowledge of the angular distribution of radiation scattered in the air. In analyzing the problem of the angular distribution of penetrating radiation produced by a nuclear blast, it is a common practice to base calculations on the concepts of the "polar angle" of scattering  $\theta$  and the "angle of inclination"  $\xi$  of the scattered radiation. A diagram of the calculation of these angles is shown on Figure 2.10. In [17], it is demonstrated on the basis of analysis of calculation data produced by the Monte Carlo method that the angular distribution function of neutron doses from a point isotropic fission source in an infinite air medium at distances at least exceeding the relaxation length can be described by the equation

(2.70)

where  $F(\theta)$  is the field of the total dose of neutrons at a given point in space per unit solid angle from a direction characterized by the scattering angle  $\theta$ . Thus, the following condition is observed:

(2.71)

This dependence agrees well, as we can see on Figure 2.16, with the experimental data of [13], in which the angular distributions of doses of neutrons at the surface of the earth were studied during experimental nuclear explosions in the air. The satisfactory agreement of the results of these works indicates the identical form of the angular distribution of neutrons of a nuclear explosion in an infinite air medium and near the earth-air division boundary.



Figure 2.16. Angular Distribution Functions of Doses of Nuclear Explosion Neutrons Scattered in Air: \_\_\_\_\_, calculated [17]; experimental [13]: 0, R = 1240 m;  $\triangle$ , R = 1450 m;  $\square$ , R = 1660 m

It has been established in a number of works that the form of the angular distribution function of neutrons scattered in air does not change over a broad range of distances from the center of the explosion. This conclusion is illustrated on Figure 2.17, where the angular distributions of flux and dose of neutrons are shown at various distances from the center of a nuclear explosion [10]. As we can see, beginning at a distance of 165 m, the field of neutrons scattered in the air can be considered stable. Experimental confirmation of the equilibrium phenomena of the angular distributions of neutrons is presented by the data of Figure 2.16 [13], relating to various distances from the center of the explosion. It should also be noted that during measurement of angular distributions of neutrons scattered in the air [13], no dependence of the dose on angle of inclination & has been discovered for radiation striking a detector located on the surface of the earth and measuring radiation from the upper half space.





This indicates that the earth has no influence on the nature of the angular distribution of neutrons scattered in the air. Although function  $F(\theta)$  in an infinite air medium corresponds in form with the angular distribution function of neutrons near the earthair division boundary when the detector records neutrons arriving only from the upper half space, the numerical values of the free term and the preexponential factor in equation (2.70) will be different). Obviously, when the source and detector are located on the surface of theearth, the values of these coefficients will be double the values produced according to equation (2.70) for an infinite air medium, if we consider the earth to be an absolutely black body. With other positions of the source relative to the detector and surface of the earth, characterized by the elevation angle of the explosion  $\alpha$  between the direction of the source-detector line and the surface of the earth, these coefficients will have intermediate values. Therefore, we can write

$$F(\alpha, \gamma) = K_{\alpha}(0,033 + 0,4045 e^{-0.033450}), \qquad (2.72)$$

where  $K_{\alpha}$  is a coefficient which is dependent on angle  $\alpha$ .

Let us first study the hypothetical case of an absolutely black surface of the earth. Integration of equation (2.72) with respect to the upper half space can be performed as follows:

$$2K_{\alpha}\left[\pi\int_{0}^{\pi}\sin\theta \left(0,033+0,4045\,\mathrm{e}^{-0.03345\theta}\right)d\theta + \left(2.73\right)\right.$$

$$\left.+\int_{0}^{\pi-\alpha}\sin\theta \left(\pi-\arccos\frac{\mathrm{tg}\,\alpha}{\mathrm{tg}\,\theta}\right)\left(0,033+0,4045\,\mathrm{e}^{-0.03345\theta}\right)d\theta\right].$$

The angular distribution function of this integral should be equal to unity by definition. From this, it is easy to find the value of coefficients  $K_{\alpha}$  for any angles  $\alpha$ .

The influence of reflection of neutrons from the surface of the earth can be considered by using the results of [18, 19], which show (see § 3, Chapter 4) that the integral dose albedo in concepts of the flux of neutrons from the source of fission in the air from the surface of the **ea**rth is described by the expression

$$A_{n, \mathbf{s}}^{\pi}(\beta) = 0.435 \frac{\Sigma_{\mathbf{g}} - \Sigma_{\mathbf{H}}}{\Sigma_{\mathbf{g}}} \cos^{3/s} \beta, \qquad (2.74)$$

where  $\Sigma_g$  and  $\Sigma_H$  are the total macroscopic neutron cross sections of the ground and hydrogen contained in the ground;  $\beta$  is the angle of

incidence of neutrons on the surface of the earth, measured from the normal (see Figure 2.10).

Since a portion of the integral dose in the detector located near the surface of the earth is created by neutrons reflected from the earth, the values of coefficients  $K_{\alpha}$  in equation (2.72) will be

different than with an absolutely black medium. They can be found by solving the equation

$$\int_{2\pi} \left[ 1 + A_{n,d}^{\pi}(\beta) \right] K_{\alpha}(0,033 + 0,4045 \,\mathrm{e}^{-0,033450}) \,d\Omega = 1, \tag{2.75}$$

the integration of which is performed with respect to the upper half space.

If we know the integral value of  $A_{nd}^n$  ( $\alpha$ ), the albedo of neutrons from the nuclear explosion with a fixed elevation of the explosion  $\alpha$  (§ 6, Chapter 2), coefficient  $K_{\alpha}$  can be calculated using the formula

$$K_{\alpha}^{-1} = 2 \left[ 1 + A_{n, \pi}^{n}(\alpha) \right] \int_{0}^{\pi/2} \int_{0}^{\pi} \{0, 033 + 0, 4045 \times 0\}$$

 $\times \exp\left[-0.03345 \arccos\left(\sin\beta\cos\varphi\cos\alpha + \cos\beta\sin\alpha\right)\right] \sin\beta d\beta d\varphi.$ (2.76)

The values of coefficients  $K_{\alpha}$  for soils of varying moisture contents, characterizing their contents of hydrogen, and for water calcualted according to this equation are presented on Figure 2.18 [20].

Thus, the dose of neutrons arriving per unit solid angle from the upper half space from directions with low angle  $\theta$  can be calculated using formula (2.72) considering coefficient  $K_{\alpha}$ .

The form of the angular distribution of neutrons from a thermonuclear explosion is similar to that of a fission weapon. This is easy to see by comparing the data of Table 2.4, relating to a thermonuclear explosion [9], to the angular distribution functions of neutrons from a nuclear explosion shown on Figure 2.17. In both cases, the angular distributions of neutron flux are normalized per unit value of cos  $\theta$  (divided by  $\Delta \cos \theta$ ), Thus, by dividing the quantities in Table 2.4 and Figure 2.17 by  $2\pi$ , we produce the value of the angular distributions per unit solid angle. The results shown on Figure 2.17 relate to a flux of neutrons with energies  $E \ge 0.2$  MeV, in Table 2.4 --  $E \ge n = 0.5$  MeV.



 $\alpha$ , degrees

Figure 2.18. Coefficient  $\overline{K}_{\alpha}$  of

Neutrons for an Absolutely Black Surface of the Earth (curve 4) and  $K_{\alpha}$  for Scattering Media as Functions of  $\alpha$ : 1, dry soil, composition:

hydrogen 1.24%, oxygen 52.38%, aluminum 7.85%, silicon 38.53% by weight; 2, moist soil, composition: hydrogen 2.54%, oxygen 57.36%, aluminum 7.08%, silicon 38.32%; 3, water

As concerns slow neutrons, their angular distribution is near isotropic.

80 ° 1 8 00

In [10], the energy distribution of scattered neutrons arriving at the detector from various directions was also studied. Table 2.5 shows the spectral-angular distributions of neutrons from a nuclear explosion scattered in the air, calculated for a range of 595 m [10]. These data indicate that the energy distribution of neutrons scattered in the air is independent of angle  $\theta$ . The angular spectra correspond in form to the integral spectrum of them**neutron**s of a nuclear explosion (see Figure 2.12). Experimental studies during nuclear tests [13] have confirmed this conclusion.

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Table 2.5

9 SPECTRAL-ANGULAR DISTRIBUTION OF NEUTRONS SCATTERED IN AIR AT 595 m FROM CENTER NUCLEAR EXPLOSION  $\Phi(\Delta E_n, \Delta 0) 4\pi R^2$  neutrons/(cm<sup>2·sec</sup>)

0,1299(-5)0,2598(-5)0,4454(-6)0.2767(--7 0.3358(--6 0.3247(--6 0,2419(--6 0,4170(--6 0;1962(--6 0,2128(-6 0,4334(-6 0,7243(-7 0.1503(-6)0.2491(-6)0.9854(-7)0,1460(-5 0,3572(-7 0,7314(-7 0,3973(-6)0,1691(-7)2900 0,5253(--4 0,3067(--4 0,1127(--3 0,9168(-5 0,1289(-4 0,8637(-5 0,2398(--4 0,2615(--4 0,2235(--4 0,1436(--4 0,2561(--4 0,8686(--5 0,1330(--4 0,2510(--5 0,2723(--4 0,6524(--5 0,9053(--5 0,8145(--5 0,2068(--4 0,8351(--4 Ε 2200 Distance From Center of Blast,  $\begin{array}{c} 0.3612(-2\\ 0.2148(-2\\ 0.1584(-2\\ 0,1584(-2\\ \end{array}$ 0,3046(-2)0,1398(-2)0,1282(-2)0,8358(--3 0,1019(--2 0,1084(--2 0,2754(-2 0,8158(-3 0,8150(-3 0,6357(---3 0,4344(---3 0,3832(---3 0,8716(--3 0,1781(--2 0,7969(--3 0,6638(--3 0,6638(--3 1510 0,1717 0,1036 0,8110(—1 0,7549(-1 0,8293(-1 0,5071(-1 0,3183(-1 0,3157(-1 0,4031(-1 0,3333(—1 0,2908(—1 0,4266(—1 0,1924(-1 0,2573(-1 0,3080(-1 0,2283(-1)0,2440(-1)0,2465(-1)0,2093(-10,2437(-1 820 0.1579(+1)0.79390.67310,5184 0,4234 0,2667 0,2799 0,3364 0,1639 0,2090 0,2073 0,3018 0.2018 0.2583 0.1757  $\begin{array}{c}
0,2325\\
0,1849\\
0,1557
\end{array}$ 0,1338 0,1148 0.001781 300 R 0,3642(+1)0,1202(+1)0,99980.1781 (-2 216 0,6072 0,5692 0,4297 0,4983 0,4584 0,4558 0,4425 0,3289 0,3181 0,2267 0,2659 0,2417 0,3491 0,2912 ,3520 ,2564 ,2317 ō ంం Angular Interval Read -0.3 + -0.2-0.4 + -0.3-0.5 + -0.4-0,6+-0,5-0,7+-0,6-0,8+-0,7 $-0.9 \div -0.8$  $-1.0 \div -0.9$ 0,0+0,1-0,1+0,0-0,2+-0,10,9+1,00,8+0,90,7+0,8 $\begin{array}{c} 0.6 \pm 0.7 \\ 0.5 \pm 0.6 \\ 0.4 \pm 0.5 \end{array}$ 0,3+0,40,2+0,30,1+0,2Note.

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Table 2.4

sec)

ANGULAR DISTRIBUTION OF FLUX OF FAST NEUTRONS  $4_{\pi}R^2_{\Phi}(R)F\{\odot)$  AT VARIOUS DISTANCES FROM THE CENTER OF A THERMONUCLEAR EXPLOSION IN AN INFINITE AIR MEDIUM [2 $_{\pi}$  NEUTRONS·cm<sup>2</sup>/(cm<sup>2</sup>·ster)]

			nergy Int	erval, Me	V (ΔE <sub>n</sub> )				(vv)	c
Δ <sup>θ</sup> . degrees	0.20.4	0,40,6	0,6-0,75	0,75-1,5	1,5-2,5	2,5-4.0	4,06,0	6,0-10,0	neutrons/(	Cm
010 1020 3040	0,5367 (3 0,1874 (2 0,3308 (2 0,3037 (2	0,3560 (3 0,1726 (2 0,1334 (2 0,3604 (2	$\begin{array}{c} 0,3476 \left(-3\\0,1969 \left(-2\\0,1250 \left(-2\\0,2405 \left(-2\right)\right)\end{array}\right)$	$\begin{array}{c} 0,1670\ (-2\\ 0,3400\ (-2\\ 0,2801\ (-2\\ 0,4013\ (-2\\ -2\\ \end{array})$	$\begin{array}{c} 1,1698 (-2 \\ 0,2720 (-2 \\ 0,2028 (-2 \\ 0,1615 (-2 \\ -2 \\ 0,1615 (-2 \\ -2 \\ -2 \\ 0,1615 (-2 \\ -2 \\ -2 \\ -2 \\ -2 \\ -2 \\ -2 \\ -2 $	$\begin{array}{c} 0,1034 \left(-2 \\ 0,1302 \left(-2 \\ 0,1304 \left(-2 \\ -3 \\ 0,9497 \left(-3 \right) \end{array}\right) \end{array}$	0,8729 (3 0,6097 (3 0,2365 (3 0,5298 (3	0,3587(3 0,1596(3 0,1613(3 0,1457(3	0,6874 (-2 0,1376 (-1 0,1242 (-1 0,1630 (-1	
40-50 50-60 60-70 70-80	0,4235 (2 0,4373 (2 0,4758 (2 0,3201 (2	0,2370 (-2 0,2046 (-2 0,1350 (-2 0,1586 (-2	0,1655 (-2 0,1202 (-2 0,9089 (-3 0,8902 (-3	$\begin{array}{c} 0,5123 \left(-2 \\ 0,4189 \left(-2 \\ 0,2249 \left(-2 \\ 0,1942 \left(-2 \right)\right) \end{array}\right)$	$\begin{array}{c} 0,2119 \left(-2\\0,1590 \left(-2\\0,2798 \left(-2\\0,2085 \left(-2\right)\right)\end{array}\right)$	$\begin{array}{c} 0,1210 \left(-2 \\ 0,5791 \left(-3 \\ 0,7583 \left(-3 \\ -3 \\ 0,7110 \left(-3 \right) \right) \end{array}$	0,3704 (3 0,1127 (3 0,2563 (3 0,8636 (4	0,2390(4 0,1811(6 0,3613(5 0,7166(6	0,1711 (1 0,1409 (1 0,1308 (1 0,1050 (1	
8090 90100 100-110 110-120	0,2446 (-2 0,1287 (-2 0,2909 (-2 0,4504 (-2	0,1391 (-2 0,1759 (-2 0,3459 (-2 0,1235 (-2	0,5031 (-2 0,6692 (-3 0,7072 (-3 0,1019 (-2	0,2255 (-2 0,2355 (-3 0,1576 (-2 0,2061 (-2	0,1166 (-2 0,5286 (-3 0,7057 (-3 0,7133 (-3	0,3474 (-5 0,1883 (-4 0,1113 (-2 0,7306 (-4	$\begin{array}{c} 0,1380 \left(-\frac{4}{-4} \\ 0,6281 \left(-\frac{4}{-4} \\ 0,1050 \left(-\frac{3}{-4} \\ 0,4394 \left(-4\right)\right) \end{array}$	0,9743(5 0,4281(4 0,1187(4 0,1431(4	$\begin{array}{c} 0,1212 \ (-1)\\ 0,4604 \ (-2)\\ 0,1059 \ (-1)\\ 0,9762 \ (-2)\end{array}$	
120—130 130—140 140—150 150—160	0,1731 (2 0,25507 (2 0,2633 (2 0,8194 (3	0,1507 (-2 0,6497 (-3 0,6037 (-3 0,4932 (-3	0,2960 (3 0,4714 (3 0,3513 (3 0,3027 (3	0,8492 (3 0,1403 (2 0,6394 (3 0,1108 (2	0,8902 (3 0,6619 (3 0,3045 (3	0,9344 (4 0,8435 (4 0,4525 (4 0,1405 (3	0,2557 (-6 0,3786 (-5 0,1466 (-4 0,3192 (-12		0,5367 (2 0,5781 (2 0,4727 (2 0,3168 )2	
170-170	0,5457 (3 0,1018 (2	0,4603 (—3 0,1096 (—3	0,4318 (3 0,9362 (4	0,9272 (—3 0,5999 (—3	0,1474 (—3 0,3533 (—4	0,1007 (7 0,5295 (5	0,5437 (—17		0,2512 (-2 0,1862 (-2	
$\Phi(\Delta E_n)$ , reutrons (cm <sup>2</sup> · sec)	0,4562 (—1	0,2604 (—1	0,2000 (1	0,3704 (—1	0,2224·(1	0,9425 (—3	0,3319 (2	0,9325(3	0,1646 (0	

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 $\S$  6. Reflection of Neutrons From the Earth During a Nuclear Blast in the Air

In addition to the neutrons scattered in the air and acting on surface structures from the upper half space, during a nuclear blast the phenomenon of reflection of neutrons from the surface of the earth is observed. These neutrons also make a contribution to the dose of penetrating radiation within structures which rise over the surface of the earth. We analyze below the characteristics of neutrons scattered from the earth during a nuclear blast in the air.

Based on known dependence (2.74), the values of the integral dose albedo of neutrons from a nuclear blast

$$A_{n,J}^{n}(\alpha) = 2 \int_{0}^{\pi/2} \int_{0}^{\pi} 0.87 \frac{\Sigma_{g} - \tilde{\Sigma}_{g}}{\Sigma_{g}} \cos^{4/3} \beta \sin \beta \overline{K}_{\alpha} \{0,033 + 0.4045 \exp \left[-0.03345 \arccos \left(\sin \beta \cos \varphi \cos \alpha + + \cos \beta \sin \alpha\right)\right]\} d\beta d\varphi.$$

$$(2.77)$$

In this expression, we use the coefficient  $\overline{K}_{\alpha}$ , calculated using formula (2.73) for an absolutely black earth (Figure 2.18).

Similar calculations can be made for the case of reflection of neutrons of the atomic blast from water, if we use the expression describing the albedo for water [19, 21]:

$$A_{n, e}^{\pi}(\alpha) = 2 \int_{0}^{\pi/2} \int_{0}^{\pi} \sin\beta \left[ 0,22\cos\beta + 0,30\left(1 - \cos\beta\right) \right] \overline{K}_{\alpha} \times \left\{ 0,033 + 0,4045\exp\left[ -0,03345\arccos\left(\sin\beta\cos\varphi\cos\alpha + \cos\beta\sin\alpha\right) \right] \right\} d\beta d\varphi.$$
(2.78)

The dependence of the total albedo of nuclear blast neutrons from soil and water for various angles  $\alpha$  is illustrated on Figure 2.19.

In § 5 of Chapter 2, we have studied the problem of the angular distributions of nuclear blast neutrons arriving at the detector from the upper half space.

As concerns the angular distribution of neutrons arriving at the detector from the lower half space, in the case when the detector is located directly at the surface of the earth, this distribution is determined entirely by neutrons reflected from the ground, the distribution of the flux of which, as has been established [18, 22], is isotropic. If the detector is located at some height over the ground, the neutrons reflected from the surface of the earth in the direction of the detector will be attenuated in the layer of air next to the earth. On the other hand, neutrons entering the layer of air near the ground may be scattered in the direction of the detector and recorded by it as coming from the earth.

The contribution of these two phenomena to the distribution of neutrons propagating from the direction of the earth can be calculated if we know the integral albedo  $A_{n,d}^{\pi}$  ( $\alpha$ ) and the spatial distribution of neutrons of the atomic blast.



Figure 2.19. Integral Albedo of Nuclear Blast Neutrons for Water (Curve 3) and Soil. Composition of soil and symbols correspond to Figure 2.18.

Let us analyze the solution of this problem in the geometry shown on Figure 2.20. The source of fission neutrons, with power Q neutrons/ sec, is located at height H<sub>s</sub> and at distance R, which is greater than the relaxation length of neutrons in air, from detector d, located at height H<sub>d</sub> over the surface of the earth. The position of the source relative to the detector is fixed by angle  $\alpha$ 

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The total flux of neutrons at point B is

$$\Phi_B = f_s(H_s) f_d(H_B) \frac{Q}{4\pi l^2} e^{-\frac{l}{\lambda}}, \qquad (2.79)$$

where

$$H_B = H_d - x \cos \beta; \quad l^2 = R^2 + x^2 - Rx \cos \theta.$$

The number of fission events per element of volume

$$dN = \Phi_B \Sigma_S \, dv = \Phi_B \Sigma_S \, x^2 dx d\omega, \qquad (2.80)$$

where  $\Sigma_s$  is the macroscopic cross section of scattering of neutrons in air; dw is the solid angle subtended at point d by an elementary area ds.

If we assume that the scattering of fission neutrons in the air is isotropic in the laboratory system of coordinates, the neutrons scattered in an element of volume dv around point B will create at the detector d the flux:

$$d\Phi'_{d} = \frac{\Phi_B \Sigma_S x^2 d\omega dx}{4\pi x^2} e^{-\Sigma_t x}, \qquad (2.81)$$

where  $\Sigma_t$  is the total macroscopic cross section of the interaction of neutrons with atoms in the air.



Figure 2.20. Analysis of Angular Distributions of Neutrons Propagating From the Direction of the Earth

The flux per unit solid angle is

$$\frac{d\Phi'_{d}}{d\omega} = \frac{\Phi_B \Sigma_S}{4\pi} e^{-\Sigma_f x} dx.$$
 (2.82)

Substituting (2.79) into this expression and integrating along beam dm, we produce

$$\frac{d\Phi'_{a}}{d\omega} = Q \int_{0}^{H/\cos\beta} \frac{\Sigma_{s} e^{-l/\lambda} f_{s}(H_{s}) f_{d}(H_{B})}{16\pi^{2}l^{2}} e^{-\Sigma_{t} x} dx.$$
(2.83)

The total flux of neutrons at point d will be

$$\Phi_{d} = f_{S}(H_{S}) f_{d}(H_{d}) \frac{Q}{4\pi R^{2}} e^{-\frac{R}{\Lambda}}.$$
(2.84)

Therefore, the flux of neutrons arriving at point d per unit body angle from the direction characterized by angle  $\beta$ , and created by neutrons scattered in the layer of air near the ground in relationship to the total flux of neutrons at point d can be written as:

$$\frac{d\Phi'_{d}}{d\omega} = \frac{R^2 \Sigma_S}{4\pi} \int_0^{H_{d}/\cos\beta} \frac{e^{-l/\lambda}}{l^2} \frac{e^{-\Sigma_f x}}{e^{-R/\lambda}} \frac{f_d(H_d - x\cos\beta)}{f_d(H_d)}.$$
 (2.85)

Let us now study the reflection of neutrons from the earth at point M. The total flux at point M is

$$\Phi_{M} = f_{S}(H_{S}) f_{d}(0) \frac{S}{4\pi m^{2}} e^{-\frac{m}{h}}.$$
 (2.86)

If the integral albedo of neutrons at point M is  $A_{nh}^{\pi}$  (a'), the flux of neutrons reflected at point M will be

$$\Phi_{M}^{\prime\prime} = \frac{A_{nd}^{n}(\alpha^{\prime}) \Phi_{M}}{1 + A_{n\pi_{\gamma}}^{n}}, \qquad (2.87)$$

$$m^{2} = R^{2} + \left(\frac{H_{d}}{\cos\beta}\right)^{2} + \frac{RH_{d}}{\cos\beta}\cos\theta.$$

With isotropic distribution of neutrons reflected from the ground in any direction per unit solid angle the reflection is

$$\frac{d\Phi_{M}}{d\omega} = \frac{\Phi_{M}}{2\pi} \,. \tag{2.88}$$

Considering the attenuation of reflected neutrons as they propagate from point M in the direction toward detector d and relating the flux of reflected neutrons arriving from this direction to the total flux at point d, we can write

$$\frac{d\Phi_{d}}{d\omega} = \frac{A_{nh}^{\Pi}(\alpha') R^2}{2\pi \left[1 + A_{nh}^{\Pi}(\alpha')\right]} \frac{e^{-\frac{m}{\lambda}}}{m^2 t_{d}} (H_{d})} \frac{e^{-\Sigma_{t} \frac{H_{d}}{\cos\beta}}}{-\frac{R}{\lambda}}.$$
(2.89)

The total flux of neutrons at point d arriving from the direction of the earth per unit body angle from the direction with polar angle  $\beta$ and azimuth angle  $\phi$  is determined by the sum of expressions (2.85) and (2.89). In this case,  $\cos \theta = \sin \beta \cos \phi \cos \gamma + \cos \beta \sin \gamma$ .

Since the spectrum of neutrons of the **source** of fission is very slightly transformed as the neutrons propagate through the air and are reflected from the ground, all of the discussions presented above remain correct for the dose of neutrons as well. Thus, we can calculate the angular distribution function of neutron doses arriving at a detector placed at a certain height from the direction of the earth.

Figure 2.21 shows the results of calculation of the angular distributions of neutron doses arriving from the lower half space at a detector placed at various heights over the surface of the earth at a distance from 800 m from the center of the explosion with an elevation angle of the explosion  $\alpha = 20^{\circ}$ . As we can see, the angular intensity of neutrons arriving at the detector from the direction of the earth depends little on angle  $\beta$ , with the exception of directions near horizontal, where it increases significantly. The angular distributions of these neutrons show rather precise axial symmetry, which is rather clearly demonstrated by Figure 2.22. The maximum divergence in the values of angular intensities at angles  $\beta$  near horizontal is 20° for angles  $\phi = 0$  and  $\phi = 180^{\circ}$  and with the height of the detector over the surface of the earth H<sub>d</sub> = 2 m. This divergence increases to 35% for H<sub>d</sub> = 20 m. With lower values of angle  $\beta$ , it amounts to a few percent.

The data on Figure 2.21 also indicate that the relative form of angular distributions of neutrons depend little on the height of the detector over the surface of the earth. Only for heights beginning at 20 m does the course of the dependence of angular intensity begin to change for directions horizontal. However, this deformation is slight and can be ignored in practical calculations of the shielding of surface structures.

It is interesting to analyze the composition of neutron radiation arriving at the detector from the direction of the earth by components.



Figure 2.21. Angular Distributions of Neutrons Arriving From the Direction of the Earth at a Detector Located at Height H<sub>d</sub>, equal to: 1, 0.5 m; 2, 1.0 m; 3, 2.0 m, 4, 3.0 m; 5, 5.0 m; 6, 10 m; 7, 20 m. Data relate to azimuthal direction determined by angle  $\phi = 0$ , read from projection of source-detector line on surface of the earth.





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Figure 2.23 shows the contribution of neutrons reflected from the ground and scattered in the air, arriving at the detector from various directions depending on the position of the detector over the surface of the earth. As we can see from this figure, for low detector altitudes (up to 0.5 m) over the surface of the ground, the flux of neutrons arriving at the detector from directions near horizontal is approximately evenly divided between neutrons reflected from the earth and scattered in the near-earth layer of air. With increasing height  $H_d$ , the component

of neutrons scattered in the air begins to predominate. Nevertheless, as was shown on Figure 2.22, the form of the angular distributions of the summary dose of neutrons resulting from superposition of the dependences of the distribution of reflected and scattered neutrons remains practically unchanged for all heights  $H_d$  studied.



Figure 2.23. Contribution of Neutrons Reflected From the Ground (solid lines) and Scattered in the Near-Surface Layer of air (dotted lines) to Total Dose of Neutrons Arriving at Detector From Direction of Ground With Azimuthal Angle  $\phi = 0^{\circ}$ . Data relate to various positions of detector over the surface: 1, H<sub>d</sub> = 0.5 m; 2, H<sub>d</sub> = 1.0 m; 3, H<sub>d</sub> = 2.0 m; 4, H<sub>d</sub> = 10 m; 5, H<sub>d</sub> = 20 m. The functions of differential distributions of neutron doses produced allowed us to calculate the dependence of the total dose of neutrons arriving at a detector from the direction of the ground on the height of the detector over the surface of the ground. This dependence, produced by integrating the angular distribution functions with respect to the lower half space, is shown on Figure 2.24. As we can see, the share of neutrons striking the detector from the direction of the ground increases with increasing height of the detector over the surface of the ground.



Figure 2.24. Total Dose of Neutrons Arriving at Detector From the Direction of the Ground as a Function of Height of Detector Over Surface of Ground

It should be noted that the differential and integral characteristics of neutrons propagating from the direction of the earth depend very little on the distance to the center of the explosion. The maximum divergence was 5-10%. Therefore, the divergence can be ignored in practical calculations.

The results produced allow us to consider the contribution of neutrons propagating from the direction of the earth correctly in calculating the shielding of surface structures from the radiation of nuclear explosions.

§ 7. Spectral-Angular Distribution of Initial  $\gamma$  Radiation in a Nuclear Blast

As is the case for neutrons, when a certain disance from the source is reached (1-2 times the free path length of unscattered radiation), quasi-equilibrium is observed in the angular and energy distributions of  $\gamma$  quanta scattered in the air. This conclusion is confirmed by the results of [10], in which the energy and angular characteristics of  $\gamma$  radiation of an atomic explosion in the air were calculated by the Monte Carlo method.

Figure 2.25 shows the spectra of individual components of the  $\gamma$  radiation of a nuclear explosion, calculated for a range of 1240 m from the center of the blast. It should be kept in mind that in the calculations performed in [10], attenuation of the prompt  $\gamma$  radiation in the shell of the weapon was not considered, and the influence of the shock wave on the propagation of fragment  $\gamma$  radiation was not analyzed. Therefore, the spectra of the individual components shown on Figure 2.25 should be considered independently. Their share in the integral flux of  $\gamma$  quanta may be different than that indicated by Figure 2.25.





Figure 2.25. Spectra of  $\gamma$  Radiation of Point Isotropic Fission Source in Infinite Air Medium: 1, capture  $\gamma$  radiation; 2,  $\gamma$  quanta of fission fragments; 3, prompt fission  $\gamma$ radiation; 4, total flux of  $\gamma$  radiation



The fact of establishment of equilibrium of the angular distribution of  $\gamma$  radiation during nuclear blasts in the air has been confirmed by experimental data produced in [13]. These measurements also establish that the nature of the angular distribution of  $\gamma$  radiation is almost independent of the type of nuclear weapon and its power.

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Figure 2.26 shows the dependences of the angular distributions of prompt (attenuation in the shell of the weapon ignored), fragment and capture  $\gamma$  radiation produced in [10], as well as the total dose of  $\gamma$  radiation during a nuclear blast. Here also we see the experimental points borrowed from [13].

As we can see from Figure 2.26, the coincidence of calculated and experimental data on angular distribution of  $\gamma$  radiation dose is rather good for scattering angles  $\theta < 60^{\circ}$ . The divergence at larger angles, which is generally insignificant in relation to the integral dose, can be explained by failure to consider the scattering of  $\gamma$ radiation from the walls of the collimator and by the presence of capture  $\gamma$  radiation from the structural materials of the experimental installation. The divergence resulting from scattering on the walls of the collimator should increase with increasing collimator aperture, which was experimentally observed.

The angular intensity function of fragment  $\gamma$  quanta scattered in the air produced in [10] can be described by the dependence

$$F_f(\theta) = K_a^f[0,0092 + 3,22 \exp(-0,08\theta)].$$
 (2.90)

The angular distributions of capture  $\boldsymbol{\gamma}$  radiation dose follow the equation

$$F_{c}(\theta) = K_{a}^{c}[0,008 + 2,11 \exp(-0,065\theta)].$$
(2.91)

In these expressions, function  $F(\theta)$  characterizes the share of the total dose of scattered radiation of a given component arriving per unit body angle from the direction determined by scattering angle  $\theta$ , counted from the line connecting the source and detector.

K is a normalizing factor which depends on angle  $\alpha$  between the surface of the earth and the direction toward the center of the blast, as well as the reflecting properties of the soil. It can be calculated:

$$K_{\alpha} = \frac{\overline{K}_{\alpha}}{1 + A_{\nu,d}^{\mathrm{n}}(\alpha)}, \qquad (2.92)$$

where  $\overline{K}_{\alpha}$  is the normalizing factor calculated for a hypothetical absolutely black surface of the earth, not reflecting the  $\gamma$  radiation

which strikes it;  $A_{\gamma d}^{\pi}$  ( $\alpha$ ) is the integral albedo of the  $\gamma$  rays from the soil for a given elevation angle of the explosion  $\alpha$ .

With an absolutely black surface of the earth, coefficient  $\overline{K}_{\!\!\!\!\alpha}$  can be calculated by the expression

$$(\overline{K}_{a}^{t})^{-1} = 2 \left\{ \pi \int_{0}^{\pi} \sin \theta \left[ 0,0092 + 3,22 \exp \left( -0,080 \right) \right] d\theta + \int_{0}^{\pi - \alpha} \sin \theta \left( \pi - \arccos \frac{\operatorname{tg} \alpha}{\operatorname{tg} \theta} \right) \left[ 0,0092 + 3,22 \exp \left( -0,080 \right) \right] d\theta \right\}.$$
(2.93)

for fragment  $\gamma$  radiation and

$$(\overline{K}_{\alpha}^{c})^{-1} = 2 \left\{ \pi \int_{0}^{\pi} \sin \theta \left[ 0,008 + 2,11 \exp \left( -0,0650 \right) \right] d\theta + \int_{\pi}^{\pi-\alpha} \sin \theta \left( \pi - \arccos \frac{t_{\beta} \alpha}{t_{g} \theta} \right) \left[ 0,008 + 2,11 \exp \left( -0,0650 \right) \right] d\theta \right\}$$

$$(2.94)$$

for capture  $\gamma$  radiation.

The results of calculation of coefficients  $\overline{K}_{\alpha}$  using these expressions are presented on Figure 2.27.

The main portion of the capture  $\gamma$  radiation is radiated in 0.1 sec, and the influence of the capture  $\gamma$  radiation is practically ended by 0.3 sec after the explosion.

During this time, the cloud practically does not rise. Therefore, the source of capture  $\gamma$  radiation, like the source of neutrons, can be considered nonmoving.

The change in dose power of fragment  $\gamma$  radiation with time is determined by the decrease in activity of the fission fragments, as well as the influence of the ascent of the cloud and the effect of formation of the cavity as the shock wave passes through the atmosphere:

$$P_{t}(t) \sim u(t) K_{p}(t) K_{\mu}(t).$$
(2.95)

The cavity factor for the dose power is

$$K_{p}(t) = e^{\mu} \epsilon f^{L} \epsilon f^{(t,q)}. \qquad (2.96)$$

where  $L_{eff}(t, q)$  is the effective radius of the cavity at a given moment

in time after explosion of a weapon with a power of q Kt.  $K_{\rm H}(t)$  is a coefficient considering the influence of ascent of the cloud on the dose power at the given moment in time:

$$K_{\rm H}(t) = e^{-\mu} eff^{\Delta R} \left(\frac{R}{R_t}\right)^{\rm s}, \qquad (2.97)$$

where R is the distance to the point of the explosion;  $R_t$  is the distance to the center of the cloud at moment in time t;  $\Delta R$  is the change in the distance to the center of the cloud with time:  $\Delta R = R_t - R$ .

Naturally, when the source of fragment  $\gamma$  radiation is considered to move, the value of coefficient  $K^{f}_{\alpha}$  will not remain constant during the time of accummulation of the dose of fragment  $\gamma$  radiation.

If the calculation of the scattering angles  $\theta$  is performed for a moving polar axis corresponding with the direction to the center of the rising cloud of the atomic blast, the effective value of coefficient  $K^{f}_{\alpha}$  over the entire time of application of fragment  $\gamma$  radiation in the region of the atomic blast is determined from the expression

 $K_{a}^{i} = \frac{\int_{0}^{t \, eff} K_{a}^{i}(t) \, u(t) \, K_{p}(t) \, K_{H}(t) \, dt}{\int_{0}^{t \, eff} u(t) \, K_{p}(t) \, K_{H}(t) \, dt} \, (2.98)$ 

Here  $K_{\alpha}^{f}(t)$  is the value of coefficient  $K_{\alpha}^{f}$  at each given moment in time, determined by the change in the elevation angle of the center of the cloud with time  $\alpha(t)$ .

Integration is performed up to that moment in time  $t_{eff}$  when, as a result of radioactive decay of fragments and ascent of the cloud, the intensity of  $\gamma$  radiation drops to a value which, from the standpoint of accummulation of total dose, can be ignored. Usually  $t_{eff}$  can be considered equal to 10-15 sec.

However, it is inconvenient to read angles relative to a moving axis, since usually the initial conditions in the calculations are limited by values of angles  $\alpha$  and  $\theta$  read from the vector direction to the center of the explosion, and in order to determine the dose of fragment  $\gamma$ radiation arriving from any given direction fixed by the value of the initial scattering angle  $\theta$ , it would be necessary to trace the change of scattering angle with time  $\theta(t)$  relative to the moving center of the cloud by changing the elevation angle of the center of the cloud  $\alpha(t)$  and the related coefficient  $K_{\alpha}^{f}(t)$ , as well as the change in the quantities u(t),  $K_{p}(t)$  and  $K_{H}(t)$ . In this case, the dose of  $\gamma$  radiation per unit solid angle can be determined in relative units using the following formula

$$D(0, \varphi) = \frac{\int_{0}^{t} K_{x}^{f}(t) [0,0092 + 3,22 e^{-0,080(t)}] u(t) K_{p}(t) K_{H}(t) dt}{\int_{0}^{t} u(t) K_{p}(t) K_{H}(t) dt}, \qquad (2.99)$$

where

$$\theta(t) \quad \operatorname{arc} \cos \left[ \sin \left( \operatorname{arc} \operatorname{tg} \frac{H + 16t^3}{\sqrt{R^2 - H^2}} - \alpha \right) \sin \theta \cos \varphi + \\ + \cos \left( \operatorname{arc} \operatorname{tg} \frac{H + 16t^2}{\sqrt{R^2 - H^2}} - \alpha \right) \cos \theta \right].$$
(2.100)

Obviously, the use of this expression involves extremely cumbersome calculations for practical purposes. Still, when precise calculation is required, this formula should be used.

For practical purposes in calculations related to the angular distributions of fragment  $\gamma$  radiation in air, it is recommended to consider that the source of fragment  $\gamma$  radiation remains nonmoving throughout its entire time of action and is located at a certain effective height corresponding to the height where the radioactive cloud is located at the moment of radiation of half of the total dose of  $\gamma$  radiation of the explosion.

In contrast to neutrons, the energy distribution of  $\gamma$  radiation of a nuclear explosion is significantly transformed with changing angle  $\theta$ . As the scattering angle  $\theta$  increases, the spectrum of  $\gamma$ quanta softens. This follows from the results of calculation of the distribution of  $\gamma$  radiation of an atomic explosion in the atmosphere, performed by the Monte Carlo method [12]. Tables 2.6 and 2.7, composed on the basis of the results of this work, show the spectral-angular distributions of scattered  $\gamma$  quanta. The quantities presented in the tables characterized the distribution of  $\gamma$  ray doses scattered per unit body angle within the limits of a given interval of angles  $\theta$  by energy groups.

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The angular quantities are the ratio

$$\Xi(\theta, E_{\gamma}) = \frac{\int_{\Delta\theta} \int_{\Delta E_{\gamma}} F(\theta, E_{\gamma}) \sin \theta d\theta dE_{\gamma}}{\int_{\Delta\theta} F(\theta) \sin \theta d\theta}, \qquad (2.101)$$

where  $F(\theta, E_{\gamma})$  is the spectral-angular distribution function of the  $\gamma$  quanta scattered in the air, characterizing the dose of scattered  $\gamma$  radiation in the interval of energies from  $E_{\gamma}$  to  $E_{\gamma} + dE_{\gamma}$ , arriving in a unit body angle from the direction defined by the scattering angle  $\theta$ .

As we can see from the tables, significant softening of the scattered  $\gamma$  quanta is observed in the range of angles  $\theta$  from 0 to 90°. With a further increase in angle  $\theta$ , deformation of the spectrum is practically not observed.

The value of the spectral-angular characteristics of the  $\gamma$  radiation scattered in air allows us to estimate the influence of the phenomenon of reflection from the earth on the value of coefficient  $\overline{K}_{\alpha}$ . Considering the reflection of  $\gamma$  radiation from the earth, it can be calculated on the basis of equation (2.92).

The integral albedo of  $\boldsymbol{\gamma}$  radiation with a nuclear explosion can be calculated using the expression

$$A_{\gamma E}^{n}(\alpha) = \overline{K}_{\alpha} \int_{E_{\gamma}} \int_{0}^{\pi/2} \int_{0}^{\pi} \sin \beta F(\theta) \Xi(\theta, E_{\gamma}) A_{\gamma E}^{n} \cdot (\beta, E_{\gamma}) dE_{\gamma} d\beta d\varphi, \qquad (2.102)$$

where  $A_{\gamma E}^{\pi}$  ( $\beta$ ,  $E_{\gamma}$ ) is the albedo for the energy of  $\gamma$  radiation with initial energy  $E_{\gamma}$ , striking the surface of the earth at angle  $\beta$ . According to [23], it is

$$A_{\gamma E}^{\pi}(\beta, E_{\gamma}) = 3.2 \frac{1}{E_{\gamma}} \frac{1}{\cos \beta} \frac{\rho_{\alpha}}{Z^{3}}.$$
 (2.103)

Here  $\rho_g$  is the density of the ground ( $\simeq 1.7 \text{ g/cm}^3$ ), while its effective atomic number Z  $\neq 13$ .

Table 2.6

ENERGY COMPOSITION OF DOSE OF SCATTERED FRAGMENT  $\gamma$  RADIATION  $\Xi_f(\theta, \epsilon_{\gamma})$ 

(% DOSE IN GIVEN ANGULAR INTERVAL)

			Interva	l of Ener	aies <u>A</u> E	MeV				
Δθ. degree	S 0.1-0.2	c.2-0.3	0.3-0.5	0.5-0.75	0.1-27.0	1-1.5	1.5-2	23	3-4	<b>4</b> -6
0000 0000 00000 00000	92.0 14.1	0,53 0,69 1,19	0,23 4,66	0,67 2,85 4,28	2,23 3,92 6,24	4.7 8,55 28,84	12,50 16,52 21,0	12,93 40,25 28,55	37,45 17,3 3,83	28,0 7,42
	3,25 9,55 9,55	2,24 8,11 9,33	13,43 9,78 14,31	15,66 13,5 22,91	11,11 18,6 36,27	38,65 36,0 7,63	13,93 0,56	1,73		
6070 7080 8090	17,52 23,80 46,65	14,0 38,89 18,37	17,03 21,62 31,69	47,67 15,69 3,28	1,53	2,25		•	,	
90—100 100—110 110—120	59,75 43,36 51,13	5,15 14,67 23,83	.35,1 42,97 25,04	r	,					
$120 - 130 \\ 130 - 140 \\ 140 - 150$	45,14 48,9 56,86	34,64 40,78 37,49	20,22 5,65			,				
150—160 160—170 170—180	50,87 46,06 55,3	43,5 52,84 44,7	5,63							

Table 2.7

ENERGY COMPOSITION OF DOSE OF SCATTERED CAPTURE  $_{\gamma}$  RADIATION  $\Xi_{c}(\theta,~E_{\gamma})$ 

(% DOSE IN GIVEN ANGULAR INTERVAL)

	8—10	8,65 2,65		•				
	68	3,13 8,36						
	4—¢	68,6 22,46					•	
	34	13,53 25,96 13,62						
MeV	2-3	4,27 27,33 39,31	8,49		<u></u>			
es ∆Eγ,	1,5-2	0,98 6,83 22,75	25,93 7,63					
Energi	1—1,5	0,31 3,05 17,61	33, <b>2</b> 9 46,73 23,21	22,73				
rval of	0,75-1,0	0,29 1,58 2,34	12,95 17,63 46,47	9,34		•		
Intei	0.5-0,75	0,16 0,79 1,72	6,63 9,18 11,06	38,42 48,52 21,64				
	0,3-0,5	0,03 0,37 0,89	2,88 6,23 6,94	11,83 27,74 16,58	62,65 38,94 39,4	10,07 31,97 0,86	1,18 0,46	
	0,2-0,3	0,25 0,82	5,58 5,96 4,94	6,60 9,68 30,99	4,15 31,19 15,55	31,73 38,39 51,54	55,07 55,36 37,82	
	0.1-0.2	0,05 0,33 0,94	4,25 6,34 7,38	11,08 14,06 30,79	33,20 29,87 45,05	58,20 29,64 47,60	43,75 44,18 62,18	
	Δθ, degs	0—10 10—20 20—30	3040 4050 5060	60—70 70—80 80—90	90100 100110 110120	120—130 130—140 140—150	150—160 160—170 170—180	



 $\theta = \arccos(\sin\beta\cos\varphi\cos\alpha + \cos\beta\sin\alpha).$ 



α, degrees

Figure 2.27.

Q







The results of calculation of the integral albedo of  $\gamma$  radiation of a nuclear explosion are presented on Figure 2.28, while the corresponding values of coefficients K are shown by the dotted line on Figure 2.27.

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CHAPTER 3. RADIATION OF RADIOACTIVE FALLOUT

§ 1. Types of Radioactive Fallout

The residual radiation following a nuclear explosion results from three factors: fission fragments, induced activity and the remains of the nuclear charge.

As concerns the relationship between induced and fragment activity, this problem was discussed in considerable detail in Chapter 1. We will note only that with high altitude air bursts, the yield of activation products is low, since only a small fraction of the neutrons reaches the surface of the earth below the epicenter of the blast. The radioactive contamination resulting from matter with induced activity will in this case cover only a small region, where other damaging factors of the blast are more important -- the shock wave and the light radiation. However, even with a surface blast, for which a comparatively high yield of activation products is observed, fragments are predominant in the radioactive contamination of the surface of the earth by products of the blast.

In addition to this, particles of the portion of the charge which is not yet undergone fission: uranium or plutonium, settle to the earth. The isotopes of these elements which undergo fission emit  $\alpha$  particles, as well as a certain number of low energy  $\gamma$  quanta. The half life of U<sup>235</sup> is 7.10<sup>8</sup> years, of Pu<sup>239</sup> -- 24.3.10<sup>3</sup> years. Due to the slow decay of these isotopes, their radioactivity is negligible in comparison with the activity of fission fragments.

Due to the low penetrating capacity of  $\alpha$  particles (they are almost completely absorbed in a layer of air 2.5-5 cm thick), plutonium is dangerous only if it enters the organism.

Thus, of all products of a nuclear explosion, the fission fragments are the main products determining the radiation danger in the region of fallout of these products from the cloud.
The danger for radioactive contamination of the terrain is represented by a surface blast of a nuclear weapon. During air blasts, the radioactive isotopes condense primarily on particles of the material of the weapon and represent no significant danger, since due to their small size the particles precipitate slowly and are scattered over a great area. With surface explosions, large quantities of earth are caught up in the fireball and form particles, onto which the radioactive isotopes precipitate.

The particles of earth captured during the nuclear explosion and entering the central portion of the fireball at temperatures amounting to several millions of degrees, evaporate; the particles entering the area with lower temperature melt. A large quantity of crushed earth is caught up by the ascending air stream, drawn up into the cloud and mixed in it. As the particles of evaporated and melted earth cool, they condense and are formed into drop-shaped or spherical particles, which frequently adhere to the unmelted particles of earth.

The formation of the radioactive wake begins immediately after the explosion of a nuclear weapon, when the earth drawn up into the atomic cloud begins to fall back onto the surface. The rate of fallout and the location of the point where particles return to the earth in relation to the point of the explosion depend on the size of the particles and the wind speed. The atomic cloud is moved by the wind, and the precipitating particles form a trail behind it.

All radioactive fallout can be divided into three main types [1, 2].

1. Local -- with large particles (up to  $50 \mu$ ). These particles fall out during the first 10-20 hours following the explosion over a territory up to 400-500 km from the center of the explosion. It has been established [3] that about 80-90% of the radioactive material formed during a surface blast falls out in the form of local fallout.

2. Continental fallout consists of fine particles which enter the upper layers of the troposphere. The precipitation of these particles is slower due to their small diameter (5  $\mu$  and less). They continue to fall out onto the earth for 2-3 weeks following a blast. During this time, the cloud covers a great distance over the earth's surface. Most of the short-lived radioactive isotopes in continental fallout decay before reaching the ground.

3. Global fallout consists of fine particles (less than 1  $\mu)$  carried into the stratosphere. These particles fall out very slowly.

They continue to precipitate onto the surface of the earth over a period of time which has been estimated variously as from one half year to 5-7 years.

Special cases arise when radioactive clouds mix with rain or snow clouds. In these situations, the rain drops or snow flakes capture the radioactive particles and accelerate fallout. It is assumed [4] that the effectiveness of capture of radioactive particles by raindrops and snowflakes is the same.

When rain acts on an explosion cloud, local areas of increased contamination may be produced. For example, as calculations and observations performed in the USA in 1955 showed, about 29% of all particles with diameters of 10  $\mu$  are washeddouttby rain with an intensity of 1 mm/ hr in 15 minutes.

In the following, we will analyze the phenomena accompanying local fallout.

§ 2. Formation of the Field of  $\gamma$  Radiation Over a Sector of Radioactive Contamination

The radioactive particles precipitating from a nuclear cloud settle onto the surface of the earth. The surface radioactive contamination of the terrain thus formed exists for a relatively short time. Gradually, over the course of the next few weeks, various processes cause the radioactive products to begin to penetrate into the depth of the soil either together with their carrier particles or as a result of being washed from the surface of these particles, or as a result of breakdown of the particles. Volumetric contamination of the upper layer of soil then develops to a depth of several centimeters. According to the data of [5], 80-95% of all activity is contained in the top layer (5-6 cm).

Objects located on the contaminated surface of the earth are acted upon by  $\beta$  and  $\gamma$  radiation from the fission fragments. The main problem of protection from radioactive fallout consists of shielding from  $\gamma$ radiation, which is the most penetrating and biologically dangerous type produced.

Let us study first a simplified model of an area of radioactive contamination, evenly covered with radioactive particles. Any surface distribution of  $\gamma$  activity can be considered to consist of individual point isotropic sources and the dose power over the ground can be calculated as the sum of the doses from these sources.

For a point isotropic source of  $\gamma$  radiation in an infinite air medium, the dose power at a point separated from it by a distance of R(m) can be found using the expression

$$P_{i}(R) = 5.09 \frac{\mu_{ai}u_{i}}{4\pi R^{2}} B_{i}(R) e^{-\nu_{i}R} [p/u], \quad [\mathbf{r/hr}], \quad (3.1)$$

where  $\mu_i$  is the linear coefficient of absorption of the dose of  $\gamma$ radiation of the sourse with initial energy of  $E_i$ ;  $\mu_{ai}$  is the coefficient of absorption of the energy of  $\gamma$  radiation in the air, cm<sup>-1</sup>;  $u_i$  is the activity of the source, MeV/sec;  $B_i$  is the dose accummulation factor in the air, which, according to the results of [6], can be calculated using the equation

$$B_{i}(R) = [1 + 1,09e^{-0.155E_{i}}(\mu_{i}R) + 1,73e^{-2.3+1F_{i}}(\mu_{i}R)^{2}] \times \leq \exp \{(-0,085)\exp\{-3,0E_{i}(\mu_{i}R)\}\}.$$
(3.2)

Another expression which approximates the dose accummulation factor B<sub>i</sub> (R) with good accuracy in the 0.5-2.0 MeV energy interval is the following formula, suggested in [7]:

$$B_i(R) = 1 + \mu_i R + \frac{(\mu_i R)^2}{7E_i^{2.4}}.$$
 (3.3)

The presence of the division boundary prevents the accummulation factors calculated for homogeneous infinite media from being applied directly; corrections must be introduced to the calculation formulas, considering the influence of the heterogeneity of the medium. The influence of the underlying surface is generally considered by the correction factor  $\Delta_{\pi}$ , defined as the ratio of power of the dose from

a source located at the earth-air boundary to the power of the dose from a source located in a homogeneous air medium. The value of this factor depends on the atomic number z and density of the underlying medium, as well as the energy of the  $\gamma$  radiation from the source, the height of the detector over the surface of the ground and its distance from the source.

It was demonstrated in [8] that if heterogeneous media are similar in atomic number, as can be considered in the case of air and soil, the correction factor can be calculated using the expression

$$\dot{\Delta}_{n} = \frac{1 + \frac{1}{2} \left[ B_{i}^{\prime}(\mu_{i}^{\prime}R) - 1 \right]}{B_{i}^{\prime}(\mu_{i}^{\prime}R)} + \frac{e^{-\mu_{i}^{\prime}R} \left[ B_{i}^{\prime}(\mu_{i}^{\prime}R) - 1 \right]}{2e^{-\mu_{i}^{\prime}R} B_{i}^{\prime}(\mu_{i}^{\prime}R)}, \qquad (3.4)$$

where  $B'_i$ ,  $B''_i$  are the accummulation factors for infinite air and soil homogeneous media respectively.

Let us analyze some particultar cases. Near the source, when R is  $-\mu_i^R$  low, it is  $\mu_i^R \approx 0$ , e = 1 and  $B_i^L(\mu_i^R) = 1$ ,

$$\Delta_{\mu} = 1 + \frac{1}{2} e^{-\mu_i^* R} [B_i(\mu_i R) - 1]. \qquad (3.5)$$

As you can see, at very close distances to the source  $\Delta_H > 1$ , i. e. the power in the detector is greater than for a homogeneous medium, due to the contribution of back scattering from the underlying surface.

If R is great, i. e. 
$$e^{-\mu_{i}^{\prime\prime}R} \rightarrow 0$$
,  

$$\Delta_{\mu} = \frac{1 + \frac{1}{2} \left[ B_{i}^{\prime}(\mu_{i}^{\prime}R) - 1 \right]}{B_{i}^{\prime}(\mu_{i}^{\prime}R)}.$$
(3.6)

This expression demonstrates that for great distances the dose power at the detector is less than for the corresponding distances in an infinite medium. We note that expression (3.6) is applicable for source-detector distances greater than 0.5 m.

The power of the dose at the point of observation over a flat isotropic source is determined by integrating the contributions from elementary sources dP located on the plane over the entire surface of the plane:

$$dP_i(R) = 5,09 \frac{\mu_{ai} a_i}{4\pi R^2} B_i(R) e^{-\mu_i R} ds.$$
(3.7)

If we use expression (3.3) for the dose accummulation factor, assuming ds =  $2\pi RdR$  (Figure 3.1) and performing integration from H to  $R_0$ , we produce an expression for the dose power P(H,  $R_0$ ) from an area with radius  $r_0 = \sqrt{R_0^2} = H^2$  at the point of observation, located over the center of this area at height H:

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$$P(H, R_{0}) = 5,09 \frac{\mu_{ai} u_{i}}{2} [E_{1}(\mu_{i} H) - E_{1}(\mu_{i} \sqrt{r_{0}^{2} + H^{2}}) + \frac{e^{-\mu_{i} H}}{7E_{i}^{2,4}} (1 + 7E_{i}^{2,4} + \mu_{i} H) - \frac{e^{-\mu_{i} \sqrt{r_{0}^{2} + H^{2}}}}{7E_{i}^{2,4}} (1 + 7E_{i}^{2,4} + \mu_{i}) / (r_{0}^{2} + H^{2})].$$

$$(3.8)$$



Figure 3.1. Plan of Calculation of Dose Power Over Radioactively Contaminated Sector

For an area of infinite radius  $(r_0 = \infty)$ 

- ---

$$P(H) = 5,09 \frac{\mu_{ai} u_{l}}{2} \left[ E_{1}(\mu_{l} H) + \frac{e^{-\mu_{i} H}}{7E_{i}^{2,4}} (1 + 7E_{l}^{2,4} + \mu_{i} H) \right].$$
(3.9)

Here  $E_1(x) = E_1(-x)$  is the integral exponential function. When it is necessary to consider the influence of the underlying surface, expression (3.4) must be multiplied by correcting factor  $\Delta_H$ .

Formulas (3.8) and (3.9) allow us to determine the share contributed by an area with any radius r to the total dose power from an area with infinite radius. Table 3.1 shows the results of calculation of the radius of sectors determining 50, 75 and 90% of the total dose at the point of measurement over an infinite contaminated sector for three different  $\gamma$  quantum energies. The calculations were performed considering the earth-air division boundary.

### Table 3.1

1

# RADII OF SECTORS (IN METERS) DETERMINING 50, 75 AND 90% OF THE TOTAL DOSE OF RADIATION AT A POINT OVER A CONTAMINATED SECTOR OF INFINITE EXTENT

	Ei	= 0,41 Me	V	E	l = 1,25	MeV	E	1 = 2,81	ηc¥
Н, М	50%	75%	90%	50%	75%	90%	50%	75%	90%
1	6	40	122	. 6	42	130	7	50	140
10	8	86	170	38	92	180	45	115	200
50	12	60	258	98	173	280	112	205	300

As we can see from the table, at low height over the earth, the dose of  $\gamma$  radiation is primarily determined by sources located not over 100-150 m from the point of measurement. The  $\gamma$  activity of more remote sources creates only a low percentage of the total dose.

We note that the radii of areas making an identical contribution to the total dose power at the point of observation depends only slightly on the primary energy of the  $\gamma$  radiation.

As we have noted, radioactive contamination of an area causes contamination of the upperrsoil layer. The individual point sources of  $\gamma$  radiation in this case are distributed through this layer of soil. In order to calculate the dose power over the surface of a  $\gamma$ -active soil layer, we must consider the absorption of radiation in this layer.

Let us analyze the case of volumetric contamination with an activity decreasing with depth exponentially

$$u_{V}(x) = u_{V}(0) e^{-mx}, \qquad (3.10)$$

where  $u_V(0)$  is the volumetric activity of the surface layer; m is a constant quantity characterizing the rate of decrease of concentration with depth.

As was demonstrated in [9], this type of source if formed several weeks following fallout of the radioactive products onto terrain as a result of their penetration into the soil. The value of m may reach  $0.3-6.0 \text{ cm}^2/\text{g}$  for various soils and times following fallout of the radioactive precipitate.

The same type of volumetric contamination is observed upon activation of the earth by the neutrons of a nuclear explosion.

The dose power in the air over an infinitely large sector of contamination with a distribution of  $\gamma$  activity through the upper layer is

$$P_{V}(H) = 5,09 \frac{u_{V}(0) \mu_{al}}{2m} \left\{ E_{1}(\mu_{l}H) - e^{(B-1)\mu_{l}H} E_{l}(B\mu_{l}H) + \frac{B-1}{B} e^{-\mu_{l}H} \left[ 1 + \frac{1}{7E_{l}^{2,4}} \left( 1 + \mu_{l}H + \frac{1}{B} \right) \right] \right\},$$
(3.11)

where  $B = 1 + m/\mu_i$ .

It is similarly easy to produce a formula for a sector of limited size. It should be noted that the effective radius of action of a volumetric source is considerably less than the effective radius of action of a surface source. For example, at a height of 1 m over the surface of the earth, 90% of the total dose of  $\gamma$  radiation is collected from an area limited by a radius of 15-20 m, while at a height of 25 m it is collected from an area with a radius of 100 m.

The form of the contaminated territory (radioactive wake) varies widely and depends primarily on the weather conditions -- speed and direction of the wind at various altitudes reached by the cloud. The actual gradience of the density of surface contamination are such [10] that within the areas providing practically complete collection of the dose at the point of observation over the contaminated surface, the density of contamination can be considered constant with high reliability.

§ 3. Angular and Energy Distributions of  $\gamma$  Radiation Over a Contaminated Surface

The structure of the  $\gamma$  field over a radioactively contaminated surface is determined first of all by the radioactive decay of the

explosion products and the related change in the energy composition of the  $\gamma$  quanta emitted and, secondly, by the nature of interaction of the  $\gamma$  radiation of radioactive decay with the air and the soil.

Works [11, 12] present the results of calculation of the differential spectra of  $\tilde{\gamma}$  radiation over an infinite plane, evenly covered with fission products of various ages.

Table 3.2 shows the results of calculation of the spectral and angular distributions of  $\gamma$  radiation of fission products with ages from 1.12 hours to 9.82 days over an ideally flat plane of infinite size. These same tables contain information on the angular distributions of the total flux and dose of  $\gamma$  quanta, as well as the integral spectra of  $\gamma$  radiation [12].

The data of Table 3.2, which relate to fluxes of  $\gamma$  quanta, are expressed in units of quanta/(cm<sup>2</sup> sec), while the dose is expressed in rad/hr. All results are normalized for the surface activity corresponding to one U<sup>235</sup> fission event per 1 cm<sup>2</sup> of surface of the earth. It must be noted that the fluxes of quanta shown in Table 3.2 are integral for the energy and angular intervals indicated. In order to make the transition to differential characteristics of the angular [quanta/(cm<sup>2</sup> sec ster)] and energy [quanta/(cm<sup>2</sup> sec MeV)] distributions, the quantities presented must be divided by the solid angles corresponding to the angular intervals and the energy intervals.

Analysis of the angular distributions of the dose of  $\gamma$  quanta corresponding to fission products of various ages indicates that the angular dose distribution function of  $\gamma$  quanta from fission products is insensitive to transformations of the  $\gamma$  radiation spectra of the fission products within limits of 1 hour to 10 days after the explosion. This is understandable if we consider that below the horizon the primary contribution to the dose over the contaminated area is that of unscattered radiation of sources located on the surface of the earth. The attenuation of unscattered radiation in the air is insensitive to changes of energy within the limits of the transformation of the  $\gamma$ radiation spectra of fission products. The dose power of  $\gamma$  quanta arriving from the upper half space over the surface of the earth, i. e. scattered in the air, is only 7-8% of the total dose power. The dose power of scattered radiation (in the air and in the ground) amounts to a total of about 15% of the complete dose power of  $\gamma$  quanta of fission products over the contaminated sector.

Data on the angular distributions of dose power of  $\gamma$  radiation at various heights over a flat isotropic source with an age of fission

	D (AB).	rad/hr		1,276       -12         3,928       -12         7,026       -12         1,418       -11         1,418       -11         1,418       -11         2,5518       -11         3,186       -11         5,518       -11         5,518       -11         5,518       -11         5,518       -11         7,360       -12         7,360       -12         7,360       -12         2,038       -12         2,038       -12	3,859 (10		1,256 (15 3,857 (15 7,082 (15	9,911 (15 1,413 (14 1,979 (14	3,116 (11 5,370 (14 2,017 (13	1,192 (11 6,089 (15 3,819 (15
	Φ (13)	(cw:.sec)		1,090 6,280 6,280 6,280 6,280 6,280 1,166 1,55 1,166 1,315 1,315 1,566 1,55 1,55 1,556 1,315 1,3	3,070 (4		1,421 (-9 4,586 (-9 8,306 (-9	$\begin{array}{c} 1,122 (-8\\ 1,517 (-8\\ 2,066 (-8\end{array})$	3,196 (8 5,101 (8 1,611 (7	2,329 (8 1,593 (8 1,334 (8
		3,500		9,886 (-8 9,886 (-8 1,720 (-7 2,583 (-7 7,5279 (-7 7,957 (-7 7,957 (-7 1,523 (-6 6,930 (-6 6,930 (-6	1,075 (5		$ \begin{bmatrix} 7,897(-12)\\ 2,416(-11)\\ 2,204(-11) \end{bmatrix} $	6,314(-11) 9,022(-11) 1,290(-10)	$\begin{array}{c}1,915(-10\\3,450(-10\\1,694(-9\end{array})\end{array}$	3,722(11
NFINITE /cm <sup>2</sup> )	-	2,500		1.294         3.958         6,888         6,888         6,888         1.477         6,888         7,921         6,888         7,147         1.477         6,586         7,147         6,596         6,501         6,501         6,501         6,501         6,501         7,801         7,801         8,801	4,027 (—5		$\begin{bmatrix} 1,721(-10)\\5,264(-10)\\9,161(-10)\end{bmatrix}$	1,375(-9) 1,965(-9) 2,810(-9)	1,231(-9) 7,487(-9) 3,285(8)	4,210(-11
OVER l issions	-	1,500		1,558 (	3,521 (5	N	$\begin{bmatrix} 2,281(-11) \\ 6,981(-11) \\ 1,214(-10) \end{bmatrix}$	$\begin{array}{c} 1,823(-10\\2,604(-10\\3,761(-10\end{array})\end{array}$	7,961(-10 $1,441(-9$ $4,952(9$	9,673(-10 1,193(10
TITUDE		1,000	12 hr	2,5266 (	2,575 (5	9.8 day:	$\begin{bmatrix} 1,426(-10) \\ 4,361(-10) \\ 7,588(-10) \end{bmatrix}$	1,139(9 1,686(9 2,399(9	3,643(-9) 6,403(-9) 2,448(-8)	1,078(9 3,691(11 2,004(10
AT 1 M Al (cm <sup>2</sup> ·see		0.750	losion 1.	<b>1, 292</b> <b>1, 292</b> <b>1, 1, 292</b> <b>1, 1074</b> <b>1, 688</b> <b>1, 1688</b> <b>1, 688</b> <b>1, 191</b> <b>1, 466</b> <b>1, 799</b> <b>1, 191</b> <b>1, 466</b> <b>1, 799</b> <b>1, 191</b> <b>1, 19</b> <b>1, 19</b> <b>1, 19</b> <b>1, 19</b> <b>1, 19</b> <b>1, 19</b>	4,210 (5	plosion ?	1,285 (10 3,960 (10 8,536 (10	1,131 (9 -1,641 (9 2,966 (9	3,750 (9 6,755 (9 2,271 (8	1,705 (9 1,132 (9 1,597 (10
Υ QUANTA φ(Δβ, ΔΕ <sub>γ</sub>	Ŋ	0,500	After Exp	5,617 (-7 1,179 (-7 1,179 (-7 1,179 (-7 1,297 (-7 1,933 (-6 2,649 (-7 2,649 (-7 2,649 (-7 1,550 (-6 8,139 (-7 1,550 (-6 1,550 (-7 1,550 (-7 1,513 (-8 1,413 (-8) (-8 1,413 (-8) (-8) (-7 1,513 (-8) (-8) (-7 1,513 (-8) (-7 1,513 (-8) (-7 1,513 (-8) (-7 1,513 (-7)1)))))))))))))))))))))))))))))))))))	4,863 (5	AFter Ex	$\begin{array}{c}1,687 \ (-10\\5,062 \ (-10\\1,255 \ (-9\end{array})\end{array}$	$\begin{array}{c} 1.161 (-9) \\ 1.700 (-9) \\ 2.452 (-9) \end{array}$	4,282 (-9 7,061(-9 2,103(-8	2,936(9 1,920(9 9,550(10
TIONS OF SECTOR,	inergy, Me	0,300	Time /	$\begin{array}{c} 2,001 \\ 5,958 \\ -2,958 \\ -1,736 \\ -7 \\ -1,736 \\ -6 \\ -1,736 \\ -6 \\ -1,736 \\ -6 \\ -1,736 \\ -6 \\ -3,583 \\ -6 \\ -3,583 \\ -6 \\ -1,736 \\ -6 \\ -1,736 \\ -6 \\ -1,736 \\ -6 \\ -1,736 \\ -6 \\ -1,736 \\ -6 \\ -1,736 \\ -6 \\ -1,736 \\ -7 \\ -2,829 \\ -7 \\ -2,829 \\ -1 \\ -1 \\ -2,829 \\ -1 \\ -1 \\ -2 \\ -2 \\ -2 \\ -2 \\ -2 \\ -2$	3,655 (5	Time	3,040 (-10 9,013 (-10 1,545 (-9	2,136 (9 3,049 (9 3,841 (9	5,990 (9 8,994 (9 2,196 (8	4,493 (9 4,053 (9 2,035 (9
DISTRIBU AMINATION	Quantum E	0,180		$\begin{array}{c} 1, 4.35 \\ 5, 847 \\ 5, 597 \\ 7, 706 \\ 1, 581 \\ 1, 581 \\ 1, 581 \\ 1, 581 \\ 1, 581 \\ 1, 581 \\ 1, 582 \\ 1, 582 \\ 1, 582 \\ 1, 582 \\ 1, 582 \\ 1, 582 \\ 1, 532 \\ 1, 5$	3,655 (5		3,096 (10 1,116 (9 1,810 (9	2,634 (9 3,413 (9 4,364 (9	$\begin{array}{c} 6,494 (9\\9,108 (9\\2,485 (8\end{array})\end{array}$	5,871 (9 4,015 (9 4,564 (9
AR ENERGY FLAT CONT	<b>*</b> ** 1	0,100		$\begin{array}{c} 6,957 (-8) \\ 2,666 (-7) \\ 5,783 (-7) \\ 5,783 (-7) \\ 5,783 (-7) \\ 5,783 (-7) \\ 1,190 (-6) \\ 1,237 (-6) \\ 1,237 (-6) \\ 1,237 (-6) \\ 1,237 (-6) \\ 1,237 (-6) \\ 1,237 (-6) \\ 1,237 (-6) \\ 1,237 (-6) \\ 1,237 (-6) \\ 1,236 (-7) \\ 2,028 (-7) $	2,321 (5		1,489 (-10 5,456 (-10 8,992 (-10	$1,264 (-9) \\ 1,249 (-9) \\ 1,802 (-9)$	<b>2</b> ,334 (9 3,129 (9 5,701 (9	5,199 (9 3,122 (9 3,853 (9
ANGUL		0.060		1,157      8         5,084      8         6,084      8         7,707      8         7,707      8         7,707      8         1,440      7         6,084      7         6,084      7         6,084      7         6,084      7         6,086      7         6,556      7         8,910      7         7,181      7         7,181      7         7,181      7         7,4815      7         1,455      7         1,455       -7	7,967 (—6		1,620 (—11 6,460 (—11 1,015 (—10	$1,263 (-10)\\1,099 (-10)\\2,234 (-10)$	2,388 (-10 3,171 (-10 8,265 (10	9,653 (10 1,523 (9 1,571 (9
		B, degrees	-	015334556653888888888888888888888888888888	$\Phi (\Delta E_{\gamma}),$ <b>quanta</b> ( $cu^2$ • <b>See</b>		170 150	140 130 120	001 001 001	80 90 90 90

Table 3.2

(Table 3.2)

	rad/h1	2,611 (	3,721 (-1
	$\phi(\Delta\beta)$ , quanta $cm^2 \cdot sec$	$\begin{array}{c} 1,091 \left(8\\1.740 \left(9\\6.757 \left(9\\2.990 \left(9\\1,188 \left(9\right)\right)\right)\end{array}$	3,978 (7
	3,500		2,627 (9
	2,500	×	5,241 (8
	1,500	· · ·	
	1,100	-	4,242 (—8
WoW 100	0,750	7,124 (-11 6,377 (-12 5,161 (-12 1,555 (-12 7,026 (-13 9,661 (-13	4,273 (—8
	0,500	8,788 (10 1,207 (10 2,270 (10 5,369 (11 2,275 (11	4,681 (8
Ċ	005.0	$\begin{array}{c} 1,118 (-9) \\ 1,540 (-9) \\ 6,997 (-10) \\ 3,333 (-10) \\ 3,333 (-11) \\ 2,876 (-11) \end{array}$	6,672 (—8
	0,180	$\begin{array}{c} 3,161 (9 \\ 1,966 (-9 \\ 1,129 (-9 \\ 1,817 (-9 \\ 1,810 (-10 \\ 8,102 (-10 \\ 1,521 (-10 \end{array}) \end{array}$	7,759 (8
•	0,100	$\begin{array}{c} 4,112 (-9) \\ 3,615 (-9) \\ 1,818 (-9) \\ 2,762 (-9) \\ 8,895 (-10) \\ 8,243 (-10) \end{array}$	4,327 (8
	S 0.060	$\begin{array}{c} 1,566(-9\\ 2,901(-9\\ 8,611(-10\\ 1,367(-9\\ 1,367(-10\\ 1,566(-10\\ 1)\end{array})\end{array}$	1,390 (8
	<sup>3</sup> . degree	01033455	$\Phi (\Delta E_{\gamma}),$ quanta $c M^2 \cdot \text{sec}$

1. Energy of  $\gamma$  quanta correspond to upper boundary of energy interval. For example, quantities in column under figure 0.500 are values of  $\gamma$  quanta fluxes in the 0.300-0.500 MeV interval.

2. First column corresponds to energy interval 0.04-0.060 MeV.

For 3. Values of angle ß indicated correspond to lower boundary of angular interval. example, quantities in row opposite 170 correspond to  $\gamma$  quanta in angular interval ß = 170-180°.

4. Read 1.197 (-2 = 0.01197.

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products of 1.12 hours are presented in [13]. Figure 3.2 shows the angular distribution functions of dose  $F(\beta)$  for an ideal flat fallout source, normalized to the dose at an altitude of 1 m over the ground, so that

$$2\pi \int_{0}^{\pi} F(\beta) \sin \beta d\beta = \frac{D(H)}{D(H=1)}.$$
 (3.12)

The angular distribution functions corresponding to a height of 1 m over the surface of the ground can be described with sufficient accuracy by the empirical expression

$$F(\beta) = \begin{cases} 0.0122 + 0.0484e^{-9\cos^2\beta} & 0^{\circ} < \beta < 90^{\circ} \\ 0.04 + 1.3e^{10\cos\beta} & 90^{\circ} < \beta < 180^{\circ}. \end{cases}$$
(3.13)

As a result of the influence of irregularities of the underlying surface, the angular distributions of dose power over an actual sector of contamination should differ from the angular distributions of  $\gamma$ radiation over an ideal flat source. It should be expected here that the form of the spectrum of scattered radiation will change comparatively slightly due to the significant contribution of multiply scattered quanta. The intensity of unscattered quanta changes strongly. For this reason we can expect that the angular distributions of dose power of  $\gamma$  radiation might undergo significant change in the area of angles  $\beta$  near 90°.

Figure 3.3 shows experimental data on the angular distributions of of dose power at 1 m over a contaminated area with various types of relief: over the flat bottom of a dry lake (curve 3) and an area of dry steppe (curve 4) [14], as well as over a very level virgin grassy area (curve 2) [15]. The experimental distributions produced in these cases correspond in the first case to the "effective" height of measurements (with an ideally flat surface) of 6 m, in the second case -- 12 m, in the third case -- 3 m. This same figure shows for comparison the angular dependence of dose power at a height of 1 m with no microrelief (curve 1). As was assumed, the angular distribution is distorted with values of angle  $\beta$  near 90°.

The angular spectra of  $\gamma$  quanta over a radioactively charged area with an age of fission products of 9 days were measured in [16]. The results of these measurements are shown on Figure 3.4.



Figure 3.2. Angular Distributions of Dose Power of  $\gamma$  Quanta for Ideal Flat Source of Radioactive Fallout Various Heights Over the Ground



Figure 3.3. Angular Distribution of Dose Power at 1 m Over Contaminated Surface.

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Figure 3.4. Spectral-Angular Distributions of  $\gamma$  Quanta at 1 m Over Terrain Contaminated With 9 day old Nuclear Fission Products

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CHAPTER 4. PENETRATION OF NEUTRONS THROUGH SHIELDING

§1. Attenuation of Unscattered Neutrons

As we know, the flux of unscattered neutrons beyond shielding resulting from a flat monoenergetic source with energy  $E_n$  is determined by the expression

$$\Phi_{n,\mu}(E_n, \delta) = \Phi_{-\Sigma_t}(E_n)^{\delta'}, \qquad (4.1)$$

where  $\Phi(E_n)$  is the flux of neutrons without the shielding;  $\delta'$  is the thickness of the material of the shielding, measured in the direction of movement of the neutrons of the initial beam;  $\Sigma_t(E_n)$  is the total macroscopic cross section of interaction of neutrons for energy  $E_n$ .

For a continuous spectrum, expression (4.1) becomes

$$\Phi_{n, u}(E, \delta) = \int_{E_n} \Phi(E_n) e^{-\Sigma_l(E_n)\delta^*} dE_n.$$
(4.2)

The penetration of the dose of unscattered neutrons through the barrier is described similarly:

$$D_{nu}(E_n, \delta) = \int_{E_n} \Phi(E_n) \eta(E_n) e^{-\Sigma_t(E_n)\delta'} dE_n.$$
(4.3)

Here  $\eta(E_n)$  is the transfer factor from the flux of neutrons to the dose.

Thus, the dose of unscattered neutrons beyond the shielding barrier is determined by the product of the dose of neutrons without the shield times a certain factor  $K_{n, u}(E_n, \delta)$ , which will be referred

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to in the following as the dose penetration coefficient (or factor) of unscattered neutrons:

$$K_{nu}^{*}(E_{n},\delta) = \frac{\int_{\mathbb{R}^{n}} \Phi(E_{n}) \eta(E_{n}) e^{-\Sigma_{I}(E_{n})\delta^{*}} dE_{n}}{\int_{\mathbb{R}^{n}} \Phi(E_{n}) \eta(E_{n}) dE_{n}} . \qquad (4.4)$$

However, expressions (4.3) and (4.4) are correct only if  $\Sigma_t(E_n)$  is a smooth function or if all resonances in the cross section are defined. For the dependence  $\Sigma_t(E_n)$  with a strongly expressed resonant structure located beyond the limits of resolution of the research apparatus, infiltration of neutrons along the minima in the cross section is observed. As a result of this, the effective value of the cross section decreases, i. e. neutrons are attenuated less than is indicated by the total cross section [1].

Measurements performed in [2, 3] have shown that infiltrtion does not appear significantly with all resonant structures. It is significant only when there are frequent and deep resonances. This sort of course of the total cross section of neutrons with energies corresponds to elements with ordinal numbers Z > 15-20 and in the energy area  $E_n < 1.5-2$  MeV. In a mixture of elements, the minima and maxima in the cross sections of individual components will partially overlap, as a result of which the effect of infiltration decreases.



Figure 4.1. Penetration of Total Dose of Unscattered Neutrons in Reactor Spectrum as a Function of Thickness of Iron Shield: 1, 0, experimental data; 2, calculated data

One of the most characteristic elements in which the effect of infiltration appears clearly is iron. Figure 4.1 shows the experimentally measured dependence of the transmission factor of the dose of unscattered neutrons on thickness of iron shield [4, 5]. Measurements were performed with a reactor, the spectrum of which was quite similar in form to the energy distribution of neutrons from a nuclear explosion, scattered in air. For comparison, this same figure shows the calculated attenuation based on total cross sections, borrowed from [6]. For the spectrum of the nuclear explosion, the effective unscattered neutron dose transmission factor for iron can be found from the empirical expression

$$K_{\pi,\mu}(\delta) = e^{-0.247\delta'} [1 + 3 \cdot 10^{-3} (\delta')^{3,24}].$$
(4.5)

It should be kept in mind that the "slant" thickness  $\delta$ ' of the screen is  $\delta/\cos \gamma$ , where  $\gamma$  is the angle of incidence of the initial beam at the surface of the shield, read from the normal;  $\delta$  is the thickness of the shield measured in the normal cross section.

Table 4.1

		١r	on Thio	ckness,	cm		
r <sub>n</sub> , nev	5	7	10	15	22	31	38
0,1 0,2 0,3 0,4 0,5 0,6 0,7 0,8 0,9 1,0 1,2 1,4 1,6 1,8 2,0	0,240 0,296 0,30 0,18 0,40 0,37 0,253 0,316 0,38 0,38 0,38 0,34 0,31 0,30 0,29 0,288	0,155 0,200 0,20 0,107 0,30 0,263 0,263 0,263 0,214 0,275 0,214 0,275 0,27 0,234 0,2 0,195 0,19 0,178	0,091 0,126 0,127 0,050 0,195 0,167 0,165 0,126 0,17 0,165 0,132 0,11 0,103 0,094 0,086	0,045 0,066 0,067 0,018 0,102 0,087 0,043 0,060 0,083 0,079 0,059 0,059 0,059 0,059 0,044 0,039 0,0324 0,0262	0,0224 0,0324 0,0338 0,0050 0,045 0,0406 0,017 0,0252 0,035 0,0315 0,022 0,0138 0,0117 0,0087 0,0063	$\begin{array}{c} 0,0115\\ 0,0167\\ 0,0173\\ 0,00152\\ 0,0181\\ 0,02\\ 0,0068\\ 0,0098\\ 0,014\\ 0,012\\ 0,0079\\ 0,0039\\ 0,00296\\ 0,002\\ 0,0012\\ \end{array}$	0,0076 0,0117 0,0115 0,00071 0,0103 0,0115 0,0036 0,0053 0,0076 0,0063 0,00165 0,00115 0,000115

TRANSMISSION FACTORS OF UNSCATTERED MONOENERGETIC NEUTRONS  $K_n$ ,  $(E_n, \delta)$  FOR IRON

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If it is required to calculate the transmission factor for the dose of unscattered neutrons through an iron shield for any other spectrum, the data of Table 4.1 can be used, showing the results of measurements of the transmission factors of monoenergetic neutrons for iron.

§ 2. Transmission of Scattered Neutrons Through Shield Barriers

The field of neutrons of a nuclear explosion in air can be represented as a set of flat, singly directed beams, oriented in space in correspondence with the angular distribution function. Therefore, in order to design a shield for the neutrons of a nuclear explosion, we must first of all obtain information on the penetration of neutrons of a flat, singly directed beam with various incidence angles at the surface of the screen.

A number of works have been dedicated to investigation of the problem of the attenuation of neutrons by flat shields of various materials. In [7, 8], the attenuation of the total dose of a broad, singly directed beam of monoenergetic neutrons with normal and inclined incidence at barriers of various materials was studied. According to the data of [7], the transmission factor for the total dose of monoenergetic neutrons can be represented by the relationship

$$K_{n, t}(E_n, \gamma, \delta) = \zeta(\gamma, E_n) e^{-\frac{\delta}{\lambda(\gamma, E_n)}}.$$
(4.6)

The values of the preexponential factor  $\zeta(\gamma, E_n)$  and the relaxation lengths  $\lambda(\gamma, E_n)$  of the total dose of neutrons, taken from [7], are presented in Tables 4.2 and 4.3.

As analysis of the data of these tables indicates, the angular dependence of the preexponential factors  $\zeta(\gamma, E_n)$  for iron and polyethylene can be represented with an error of not over 30% by the simple relationship

$$\zeta(\gamma, E_n) = \zeta(\gamma = 0, E_n) \cos \gamma, \qquad (4.7)$$

which is correct for measured thickness of layers throughout the entire investigated interval of energies.

Table 4.2

••••						
	۲		λ, см	۲		λ, см
ν. Degrees	Experi-	Calcu- lated	Experi- mental	Experi- mental	Calcu- lated	Experi- mental
	$E_n =$	= 3 MeV		En	= 0,83 Me	eV
0 26 38 47 56 62	$1,00\pm0,031,00\pm0,030,93\pm0,04$	1,0 0,90 0,79 0,68 0,56 0,47	$\begin{vmatrix} 14,2\mp0,3\\8,4\mp0,2\\7,4\mp0,3\\-\\.\\.\\.\\.\\.\\.\\.\\.\\.\\.\\.\\.\\.\\.\\.\\.\\.\\.\\$	$1,00\pm0,050,88\pm0,060,85\pm0,060,73\pm0,04$	1,0 0,90 0,79 0,68 0,56 0,47	$ \begin{array}{r} 19,0\pm0,4\\18,0\pm0,3\\9,3\pm0,5\\7,3\pm0,4\\-\end{array} $
	$E_n =$	0,22MeV	- 	En En	= 0,024	MeV
0 20 38 47	$1,00\pm0,05$ $0,88\pm0,05$ $0,80\pm0,04$		$\begin{vmatrix} 19,0\pm0,3\\18,0\pm0,3\\-\\9,3\pm0,5\\7,3\pm0,4 \end{vmatrix}$	$\begin{array}{c} 0,83\pm0,03\\ -\\ 0,75\pm0,05\\ \end{array}$	0,83 0,75 0,66 0,56 0,47	$27 \pm 1$  11,0±0,5

RELAXATION LENGTH  $\lambda(\gamma, E_n)$  AND FACTOR  $\zeta(\gamma, E_n)$  FOR CALCULATION OF NEUTRON DOSE TOTAL TRANSMISSION FACTOR FOR IRON

The angular dependence of the relaxation length of the dose of neutrons with initial energies of 3 and 0.83 MeV for polyethylene can be expressed with an error of not over 8% using a similar relationship:

 $\lambda(\gamma, E_n) = \lambda(\gamma = 0, E_n) \cos \gamma, \qquad (4.8)$ 

0,39

correct for approximately  $2 \le \delta \le 10$  cm.

In the area of polyethylene thicknesses  $\simeq 10 \le \delta \le 30-40$  cm for monoenergetic neutrons with energies of 0.8-3 MeV, the relaxation length should depend on angle  $\gamma$  significantly more weakly than following the rule of the cone. This is confirmed by data on the penetration of neutrons with energies of 0.22 MeV and 24 KeV through layers of polyethylene. With thickness of  $1-2 \le \delta \le 6-10$  cm for  $E_n = 0.22$  MeV, the

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Table 4.3

RELAXATION LENGTH  $\lambda(\gamma, E_n)$  and coefficient  $\zeta(\gamma, E_n)$  for calculation of neutron dose TOTAL TRANSMISSION FACTOR FOR POLYETHYLENE

Calcu-lated 3,06 2,3 ł 3,4 2,7 I 1 6.1 . 1,6 1 1 Š Experimen-3,15±0,07 ډ 2,0±0,15  $2,3\pm 0,07$ 2,0±0,1 2,0±0,1  $3,4\pm0,1$ I I ł I Ţ I  $E_n = 0,024 \, \text{MeV}$  $E_n = 0,83$  MeV Experimen- Calcu-tal 1,10 0,90 0,87 0,75 0,62 0,52 1,33 1,05 0,74 0,62 16'0 5  $1,33\pm0,06$ 1,10±0,04 0,89±0,05  $1,10\pm 0,06$  $1,00\pm 0,05$ 0,83±0,05 ١ I I 1 ł ł  $2,3\pm 0,05$  $2,2\pm 0,12$ 1,9±0,1 1,8±0,1 Experimen-| Calcu-tal 3,6 3,0 5,8 4,4 . 6,4 5,1 ł λ., CM  $2,9\pm 0,15$  $6,3\pm 0,2$  $5,2\pm0,3$  $6,4\pm 0,2$ 1.1 1 ł I I I ł  $E_n = 0, 22$  MeV  $E_n = 3$  MeV Calcu-lated **0**,..5 0,99 0,87 1,30 1,03 0,88 0,73 0,61 1,17 1,1 0,62 0,52 Experimen-1,17±0,05  $0,85\pm0,04$  $1, 1 \pm 0, 04$ 1,1<sub>±``</sub>.1  $1, 30 \pm 0, 04$  $1,25\pm0,04$ **0,78±0,08**  $0,49\pm0.05$ I ł I 1 tal negrees 56 62 0 26 38 47 0 25 38 47 56 62

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relaxation length depends very little on angle  $\gamma$ . For  $E_n = 24$  KeV with thicknesses over 4 cm the relaxation length of the dose is independent of angle  $\gamma$ . With polyethylene thicknesses over 30-40 cm, the relaxation length of the dose apparently should not depend on angle  $\gamma$  throughout the entire range of energies studied.

Table 4.4

	the second se			
Energy of	Polyet	hylene	Wate	r
Neutrons, MeV	λ	ζ	λ	<b>c</b>
15 3 0,83 0,22 0,24	$18,4\pm0,46,4\pm0,23,4\pm0,12,30\pm0,052,00\pm0,10$	${}^{1}_{1,18\pm0,02}_{1,30\pm0,04}_{1,10\pm0,04}_{1,10\pm0,04}_{1,10\pm0,04}_{1,33\pm0,06}$	$18,4\pm0,47,6\pm0,153,40\pm0,072,61\pm0,07$	$1,10\pm0,031,20\pm0,041,00\pm0,051,00\pm0,05$

EXPERIMENTAL VALUES OF  $\lambda$  (cm) AND  $\zeta$  FOR WATER AND POLYETHYLENE  $\gamma = 0^{\circ}$ 

As neutrons propagate through iron, the dependence of  $\lambda(\gamma, E_n)$  on angle  $\gamma$  is expressed more sharply than the law of the cone.

Table 4.4 presents for comparison data relating to the transmission of neutrons through water and polyethylene ( $\rho = 0.93 \text{ g/cm}^3$ ) with normal incidence of neutrons with initial energies from 0.024 to 15 MeV on the surface of the shield [8].

Particularly valuable are the results of measurement of the transmission factors of the total dose of neutrons of a nuclear reactor through protective shielding, since the spectrum of these neutrons is quite similar to the spectrum of nuclear explosion neutrons scattered in air. Table 4.5 shows the transmission factors for various thicknesses of iron, concrete and polyethylene, measured using the neutrons of a reactor with normal and inclined incidence of the beam on the surface of the shield.

The important characteristics of the field of neutrons passing through the shielding are their angular and energy distributions. A large number of works published up to the present time has been

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dedicated to the study of the angular distributions of scattered neutrons with normal incidence to the barrier. In [9], a report is presented on a study of the angular distributions of the dose of reactor neutrons beyond layers of polyethylene and lead. The angular distributions of doses of fast neutrons beyond shielding of polyethylene of various thicknesses with normal incidence of the beam of neutrons with reactor spectrum are presented. In [10], a study is described of the spectral and angular distribution of fast reactor neutrons beyond a shield of **poly**ethylene 23 cm thick under conditions of normal incidence of the beam to the screen.

The angular characteristics of neutrons scattered by shielding were studied in the greatest detail in [26]. The energy and angular characteristics of reactor neutrons beyond shielding made of iron, concrete and polyethylene were studied experimentally. The experimental results of this work confirm the calculations performed for the same conditions by the Monte Carlo method and in the  $2P_7$  approximation using the R.O.Z. program [11].

As was shown in [12, 13], the angular distributions of the dose of fast neutrons in relative units with normal incidence to the barrier with a thickness of at least 1-2 free path lengths can be described by the dependence

$$f(\psi) := \frac{1 + \frac{1}{\psi_0^2}}{2\pi \left(1 - \frac{e^{-\frac{\pi}{2}\psi_0}}{\psi_0}\right)} e^{-\frac{\psi}{\psi_0}}.$$
 (4.9)

This function characterizes the share of the dose of scattered neutrons leaving the protective barrier per unit solid angle in the direction characterized by polar angle  $\psi$ , read from the normal to the surface of the shield. Thus, the form of the field of scattered neutrons is determined only by empirical constant  $\psi_0$ , called the char-

acteristic angle. Table 4.6 shows values of characteristic angles for various materials. In multilayered shields, the angular distribution of neutrons is determined by the last layer (next to the detector), if its thickness is over  $1-2 \lambda$ .

In order to find the dose of neutrons propagating beyond a barrier in the direction of angle  $\psi$ , expression (4.9) must be multiplied by the total dose of scattered neutrons on the surface of the barrier, which is equal to

$$D_{n,s}(\delta) = D_{n,0}[K_{n,t}(\delta) - K_{n,u}(\delta)] = D_{n,0}K_{n,u}[B_n(\delta) - 1], \qquad (4.10)$$

Table 4.5

# TOTAL NEUTRON DOSE PENETRATION FACTOR $\kappa_{n_{t}}(\lambda,\,\delta)$ for reactor spectrum AND VARIOUS SHIELDING MATERIALS

	Thick	cness of C	concrete,	Th <b>i</b> ckn	iess of	lron,	E,	Thick	ness of P	olyethyle	ne, cm
Jourses	J	E				-			0	16	. 29
いいい いいい	v	5	25	4	8	14	20	*	0	2	
82002000 82000 82000 8000000	0,771 0,770 0,750 0,736 0,736 0,736 0,736 0,736 0,736 0,736 0,246 0,246 0,246	0,382 0,364 0,364 0,366 0,366 0,150 0,150 0,110 0,0509 0,151 0,151 0,151 0,151 0,151 0,151 0,151 0,151 0,151 0,151 0,151 0,151 0,151 0,152 0,152 0,152 0,152 0,152 0,152 0,152 0,152 0,152 0,152 0,152 0,152 0,152 0,152 0,152 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,155 0,150 0,1100000000	0,133 0,128 0,117 0,101 0,0792 0,0792 0,0320 0,0148 0,440 (2	0,822 0,810 0,740 0,740 0,573 0,588 0,588 0,588 0,236 0,0930	0,539 0,512 0,512 0,473 0,420 0,340 0,237 0,131 0,0504	0,341 0,335 0,320 0,283 0,283 0,283 0,283 0,283 0,283 0,283 0,283 0,283 0,283 0,0723	0,216 0,195 0,195 0,174 0,174 0,076 0,0452 0,0194	0,567 0,567 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,542 0,556	0,2240 0,227 0,212 0,191 0,152 0,152 0,152 0,158 0,0638 0,920 (2	0,0600 0,0547 0,0547 0,0364 0,0364 0,0158 0,761 (-2 0,290 (-2 0,683 (-2	0,890 (-2 0,781 (-2 0,781 (-2 0,457 (-2 0,298 (-2 0,162 (-3 0,162 (-3 0,683 (-3 0,781 (-3)) (-3)
Note	Rea	-)044.0 be	-2 = 0.004	40.)	- , -	-		:			

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where  $K_{n, t}$  and  $K_{n, u}$  are the total dose transmission factors for neutrons for the barrier and the doses of neutrons in the unscattered component respectively;  $D_{n, 0}$  is the dose of neutrons without the barrier;  $B_{n}(\delta)$  is the neutron accummulation factor beyond a barrier of thickness  $\delta$ .

Table 4.6

# VALUES OF CHARACTERISTIC ANGLE $\psi_0$ for Normal incidence of reactor neutrons to protective barriers

Shield Material	Density, g/cm <sup>3</sup>	Thicknesses Studied, cm	Threshold of Record- ing of Neu- trons, MeV	Ψ <mark>0</mark> , Degs	Reference
Water Polyethyle Concrete Iron	1,0 ne 0,89 0,93 2,16 7,8	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	07 Total Dose 0,2 Total Dose	30 30 30 45 50	[12] [14] [4] [4] [4] [4]

Table 4.7 shows several values of neutron dose accummulation factors with normal incidence beyond various shields.

lable 4.7

	Param	- T	hickr	ness d	of Prot	ective	Shieldi	ng, cm
Materia	l eter	5	10	15	20	30	40	50
Iron	$B_n(\delta)$ $K_{n,u}(\delta)$	2,6 0,31	4,6 0,13	7,7 0,065	11 0,033	17,2 0,13	22 0,0057	26 0,0026
Concret	e B <sub>n</sub> (δ) K <sub>n, u</sub> (δ)	1,5 0,48	2,3 0,24	3,2 0,12	4 0,06	6,5 0,0146	10,8 0,004	17 0,0012
Polyeth ylene (p = 0.93) $g/cm^3$	B <sub>n</sub> (δ) K <sub>n, u</sub> (δ)	3 0,12	5,8 0,017	9,7 0,004	16 0,0012	<b>34,5</b> <sup>•</sup> 0,000135	70 0,000021	130 0,0000040

VALUE OF DOSE TRANSMISSION FACTOR FOR UNSCATTERED REACTOR NEUTRONS  $K_{n,\,u}$  AND ACCUMMULATION FACTOR  $B_n(\delta)$  WITH NORMAL INCIDENCE

Table 4.8

ANGULAR DEPENDENCE OF SCATTERED NEUTRON DOSE POWER BEYOND SHIELDS OF IRON AND POLYETHYLANE --  $f(\gamma, \psi, \tau)$ , ster<sup>-1</sup> (rel. units)

23				ψ <b>. d</b>	egrees		
degrees	degrees	5 0	10	20	30	50	70.
,			Iron	— 5 см			·
0 30 50 70	0 0 180 0 180 0 180	0,102 0,069 0,052	0,183		0,110 0,077 0,084 0,060 0,075 0,049	0,088 0,140 0,07 0,056 0,101 0,039	0,079 0,108 0,062 0,097 0,052 0,043
	•	-	Iron	-14 см			
0	0	0.069	0,148		0,089	0,057 0,069	0,057 0,060
30	180 0	0.051			0,052 0,051	0,044	0,039 0,041
50	180 0	0,030			0,038 0,024	0,034 0,025	0,031
70	180				0,024	0,021	0,019
		Polye	thylene	— 9,5 см	:	,	
0 30	0	0,018	0,032	0.008	0,020	0,010	0,004
50	180 0 180	0,008		0,000	0,012 0,003	0,003	0,003

Very few data on the angular distributions of neutrons with inclined incidence of the initial beam to the barrier can be found in the literature.

In Table 4.8, based on [26], we present the angular distributions of the doses of neutrons beyond iron and polyethylene shields with various angles of incidence  $\gamma$  of a flat, unidirectional beam of reactor neutrons.

The energy distributions of neutrons as functions of angles of incidence  $\gamma$  and scattering angle  $\psi$  have certain common regularities for

barriers of various materials. With increasing angles  $\gamma$  and  $\psi$ , the neutron spectra soften. The greatest softening in the low energy area of the spectrum is characteristic for hydrogen-containing media. With large angles  $\gamma$  and  $\psi$  (over 50°) practically 90-95% of all neutrons fall in the area of energies below 0.5 MeV (contribution to dose at least 70-80%). For shields of concrete and iron, these estimates are true for the energy area  $E_n < 1$  MeV.

The spectra of neutrons with energies below 0.2 MeV in these media can be described by a dependence of the form  $\Phi(E_n) \approx E_n^{-1}$ , where 1 varies from 0.7 to 1.1. The low energy portion of the spectrum is important for calculation of the dose of secondary  $\gamma$  radiation produced by neutron capture.

# § 3. Reflection of Neutrons From Barriers

No unified terminology has been established for description of the characteristics of neutrons reflected from barriers; therefore, it is expedient first to introduce definitions and symbols. The differential albedo (angular distribution of back scattering of radiation) is a function of the initial radiation energy E, angle of incidence  $\gamma$ , polar angle  $\psi$  and azimuthal angle  $\tau$  of reflection. Angles  $\gamma$  and  $\psi$  are counted from the normal to the surface. The differential albedo will be represented by the letter a. The first left subscript indicates the type of radiation (n = neutrons,  $\gamma = \gamma$  radiation), while the second subscript indicates the albedo of the number of particles (little h) or dose (d). The superscripts determine the quantities in which the incident (first) and reflected (second) radiation are expressed. The current density (or simply current) will be represented by T, the flux density (or flux) by  $\pi$ . For example, the differential dose albedo for neutrons for an incident current and reflected flux is written as  $a_{\pi,\pi}^{\tau,\pi}(E, \gamma, \psi, \tau)$ .

If the incident and reflected radiation (both) are expressed in current or flux terms, the second (identical) superscript is omitted. The differential current and flux albedos are related by the formula

$$a^{n}(E, \gamma, \psi, \tau) = a^{\tau}(E, \gamma, \psi, \tau) \frac{\cos \gamma}{\cos \psi}; \qquad (4.11)$$

integration of the differential albedo with respect to angles produces the integral albedo

$$A(E, \gamma) = \int_{0}^{2\pi} d\tau \int_{0}^{\pi/2} a(E, \gamma, \psi, \tau) \sin \psi d\psi. \qquad (4.12)$$

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The integral albedo is given the same indexes as the differential albedo. We present below a brief review of works dedicated to study of the reflection of neutrons, and certain results which may be useful in calculations.

Among the first works on the determination of albedo is experimental work [15]. It presents data on the measurement of the integral numerical albedo  $A_{n, h}$  of monoenergetic neutrons from paraffin and water with normal and inclined incidence of neutrons to the surface of the reflector. The range of neutron energies studied in the initial beam was from the energy of thermal neutrons to 5 MeV.

It has been found that the reflection factor of neutrons from hydrogen-containing material decreases with increasing energy. The range of incidence angles of the neutron beam studied (counted from the normal) was from 0 to 75°. The angular dependence of the albedo on incidence angle  $\gamma$  for all energy values studied can be expressed by the relationship

$$1 - A_n^{\rm H} \, (\gamma) = \left[ 1 - A_n^{\rm H} \, b \, (\gamma = 0) \right] \cos \gamma \,. \tag{4.13}$$

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Work [16] was dedicated to the determination of the empirical regularities for description of the characteristics of the field of neutrons reflected from iron, concrete and earth on the basis of Monte Carlo calculations [17].

It was found in this work that the differential "dose" albedo of neutrons from concrete, iron and earth of various compositions (various water contents) can be described by the empirical expression

$$a_{n,\sigma}^{n,\tau}(E, \gamma, \psi) = a_{n,\sigma}^{n}(E, \gamma = 0, \psi = 0) \cos^{2/3} \gamma \cos \psi \text{ ster}^{-1},$$
 (4.14)

where  $a_{n, d}^{T}$  is the differential dose albedo with normal incidence and reflection of the flat beam of neutrons (the tabular value, depending on initial energy E and type of scattering material).

It is demonstrated in this work that for neutrons with the fission spectrum and for the spectrum leaving the shield, the angular albedo for all materials studied can be expressed by the formula

$$\mathbf{a}_{n,d}^{\text{IT}}\left(E, \gamma, \psi\right) = \frac{1}{\pi} \frac{0.435 \left(\Sigma_{t} - \Sigma_{H}\right)}{\Sigma_{T}} \cos^{2/2} \gamma \cos \psi, \qquad (4.15)$$

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where  $\Sigma_t$  is the total macroscopic cross section of all elements included in the composition of the scattering material, weighted with respect to to the neutron spectrum;  $\Sigma_H$  is the macroscopic cross section of hydrogen contained in the material, weighted for the neutron spectrum.

Expression (4.15) describes the differential albedo of neutrons for all materials studied except water. For the integral dose albedo of fast neutrons from water with normal incidence, [16] suggests the dependence

$$A_{n,a}^{\Pi}(E, \gamma = 0) = \frac{0.435 (\Sigma_t - \Sigma_H)}{\Sigma_t}$$
 (4.16)

In [18], the Monte Carlo method was used to calculate the integral values of numerical (in terms of "current") and dose (in terms of "flux") albedos of monoenergetic neutrons and neutrons with the fission spectrum with inclined incidence to reflectors of various thicknesses, and the energetic distributions of the reflected neutrons are presented.

In [19], a broad range of information is presented on determination of the characteristics of the neutron field reflected from concrete. These characteristics were calculated by the Monte Carlo method and measured experimentally (in terms of "current"). The calculations were performed for initial energies of neutrons E from 0.2 to 8 MeV and angles of incidence of the unidirectional beam  $\gamma = 0$ , 45, 60, 75 and 85°. The experimental data agreed well with the calculated data.

In [20], an attempt was made to produce a semi-empirical expression, using a number of assumptions concerning the nature of the interaction of neutrons with the scattering nuclei. A four-parameter semi-empirical formula is presented in this work for the differential dose albedo for iron, aluminum, dry earth, earth with 50% moisture content, earth with 100% moisture content and concrete. This formula is written in terms of neutron "current" and covers the area of energies from 0.1 to 14 MeV.

The formula produced has the following form:

$$\mathbf{a}_{n,n}^{\tau}(E, \gamma, \psi, \tau) = \mathbf{E}_{1}(E)\cos\psi + \frac{\mathbf{E}_{2}(E) + \mathbf{E}_{3}(E)\cos\psi_{s}}{1 + \mathbf{E}_{4}(E)\frac{\cos\gamma}{\cos\psi}} \quad [\text{ster}^{-1}], \quad (4.17)$$

where  $\psi_{\rm S}$  is the total scattering angle, determined from the expression

$$\cos \psi_s = \sin \gamma \sin \psi \cos \tau - \cos \gamma \cos \psi, \qquad (4.18)$$

Table 4.9

ENERGY DEPENDENCE OF PARAMETERS FOR FOUR-PARAMETER FORMULA OF DIFFERENTIAL DOSE ALBEDO FOR FAST NEUTRONS

	Ref lector				En.	MeV			,	
Param- eters	Material Index	0.1	0,25	0.5	1.0	2.0	3,0	5,0	14,0	
E (F)		0 1401	0.1385	0,1817	2620,0	0,0904	0,1147	0,0756	0,0657	
(	• 04 0	0,1528	0,1491	0,1523 0,0834	0,0940	0,1287 0,0691	0,1020	0,0442	0,0288	
	<b>بر</b> د	0,0648	0,0687	0,0813	0,0583	0,0547	0,0617	0,0431	0,0269	
	יט ע	0,0746	0,0820	0,0693	0,0477	0,0488	0,0407	0,0218	0,0287	
F. (E)	~	0.0776	0,0672	0,0796	0,0643	0,0863	0,0858	6,0710	0.0575	
	.010	0,0488	0,0874	0,0649	0,1008	0,0921	0,0991	0,1214	0,0697	
	4 CI	0.0711	0,0727	0,0793	0,0740	0,0971	0,0836	0,0787	0,0812	
	5	0,0561	0,0514	0,0982	0,0838	0,0972 0,1259	0,1010	0,1042	0,0567	
í 1	•	0,0003	0.005	0 0684	-0 0067	-0.0118 ·	0,0140	0,0293	0,0195	
E <sub>3</sub> ( <i>E</i> )	- 27	-0,0014	0,0329	0,0112	-0,0381	-0,0101	-0,0122	0,0172	0,0147	
	ω,	0,0406	0,0278	0,0585	-0,0228	0,0189	0.0154	0,0177	0,0444	
	4 4	0.0517	0,0294	0,0671	-0,0295	0,0149	0,0133	0,0289	0,0240	
•	000	0,0555	0,0345	0,0964	-0,0118	0,0393	0,029/	0,0212	0,400	
E. (E)	/	1,1654	1,0115	2,5708	0,9052	0,8566	1,2569	0,9325	9877 9,6300	
	2	0,4868	2,9778	1,2733	1,4044 0,6403	1,2034	2,3264	2,6435	1,6450	
	- 00	1 166	0,0/00	1.2315	0,5448	1,1314	1,4523	1,3720	1,8531	
	H 10	1,0278	1,2818	1,3826	0,7913	1,4744	1,5615	1,8382	1,4914	
	•	1,41/0	C177,1	***					tivo moietur	0
Note:	l, iron;	2, alumi	num; 3, c	oncrete;	4, dry ea	rth; 5, e	arth with	bug rela		<u>د</u>

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in which  $\tau$  is the azimuthal angle of reflected radiation, counted from the projection of the direction of the incident beam on the surface of the reflector.

The values of parameters  $E_1(E)$ ,  $E_2(E)$ ,  $E_3(E)$  and  $E_4(E)$  are presented in Table 4.9.

It is indicated in [20] that the results calculated by the fourparameter formula agree well with the data produced by the Monte Carlo method [17].

Works [21-24, 26] present the results of studies performed for determination of the characteristics of the field of reflected neutrons, performed experimentally for the reactor spectrum (water, earth, iron) by a calculation method in the  $2P_7$  approximation (for the same conditions and media) and using the Monte Carlo method [21] for water and various groups of incident neutrons  $\Delta E$ . Figure 4.2 shows the dependence  $a_{n,d}^{\pi}(\gamma, \psi, \tau)$  and  $a_{n,h}^{\pi}(\gamma, \psi, \tau)$  for water, earth and iron for the spectrum of neutrons leaving a reactor [23].

We can see from Table 4.2 that the flux albedo is always greater than the corresponding value of dose power and that this difference increases with improvement of the moderating properties of the medium. The angular distributions of flux density of reflected neutrons are similar to the isotropic distributions for materials with high atomic number and sharply anisotropic for light shielding materials such as water and polyethylene. As we can see from the figure, even for an iron shield, total isotropy is not observed, and therefore the empirical dependence (4.14) suggested by the authors of [16], in which the angular distribution of dose power of reflected neutrons is assumed proportional to cos  $\psi$  must be looked upon as approximate.

In the case of a light reflector (water, polyethylene), the angular distributions of flux and dose power of reflected neutrons depend significantly on the azimuthal angle  $\tau$ , whereas for earth and iron this dependence is slight.

The difference in the behavior of the angular distributions of flux and dose power for various materials can be easily explained, keeping in mind the angular distribution of neutrons uponscattering (in an elementary event) and the absorption cross section.

The dependence of the integral dose albedo  $A_{n,d}^{\pi}$  ( $\gamma$ ) on incidence angle of a flat, infinite beam of neutrons with the reactor spectrum to

barriers of iron, earth and water is shown on Figure 4.3 [22]. We can see from Figure 4.3 that the albedo for iron and earth acts similarly, while other cases since as angle  $\gamma$  increases in the case of water the energy of the neutrons scattered at slight angles  $\psi$  increases significantly. The effect of the influence of hydrogen at high values of  $\gamma$  is noted in the case of earth as well. As we can see from Figure 4.3, the agreement of experimental and calculated (Monte Carlo method) data is quite satisfactory.



Figure 4.2. Angular Distributions of Dose and Flux of Neutrons, Reflected From Water, Earth and Iron With Various Incidence Angles of a Beam With Reactor Spectrum to the Surface: a, 10 cm water; b, 30 cm earth; c, 19 cm iron.

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Analysis of the results illustrated by Figure 4.3 shows that the integral dose albedo for iron and earth is described well by an empirical dependence [see formula (4.14)].

The dependence  $A_{n,d}^{\pi}(\gamma)$  for water and polyethylene [26] can be described by the following empirical relationship:

$$A_{n,a}^{\mathrm{n}}(\gamma) = A_{n,a}^{\mathrm{n}}(\gamma = 0) \cos \gamma + 0,3 (1 - \cos \gamma). \qquad (4.19)$$

This expression is correct with  $\gamma < 70^{\circ}$ , the accuracy of the empirical expressions presented being no worse than about 10%, which is quite satisfactory for practical purposes.

Analysis of the experimental and calculation data (Monte Carlo method) indicates that for a medium such as water, the value of  $A_{n,d}^{\pi}$  depends little on the form of the incident neutron spectrum. Thus, for the neutron spectrum of a Po-Be source ( $E_{av} \approx 4 \text{ MeV}$ ), as for the spectrum of reactor neutrons ( $E_{av} \approx 0.65 \text{ MeV}$ ) and for the softened neutron spectrum from a Po-Be source ( $E_{av} \approx 2 \text{ MeV}$ ),  $A_{n,d}^{\pi}$  ( $\gamma = 0$ ) = 0.18 (with a threshold for recording of reflected neutrons  $\approx 1 \text{ KeV}$ ).



Figure 4.3. Dependence of Integral Dose Albedo on Angle of Incidence of Flat, Unidirectional Beam of Neutrons With Reactor Spectrum to Barriers of Iron ( $\Delta$ ), Earth ( $\square$ ), and Water (0). ———, Data produced by Monte Carlo calculation.

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Table 4.10

DIFFERENTIAL DOSE ALBEDO  $a_n^{\pi}$ ,  $(\Delta E, \pi, \psi) \cdot 10^{-3}$ , remester  $n^{-11}$ 

	\$= 60°	-= 180	0 2 2 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3 3	5,55 33,55 44,55	- 55 58 58 58 58 58 59 59 59 59 59 50 50 50 50 50 50 50 50 50 50 50 50 50	27.9 32.6 41,1	24,6 26,4 35,4 33,2
		t = 90°	16 20,6 29,0	21,1 23,8 31,0 30,7	18,0 22,7 33,8 33,8	26,4 30,0 40,4	28,8 35,0 45,0 46,1
		.0 = 1	28,8 60,7 125,7 193,0	28,2 60,0 128,0 209,0	29,1 66,0 158,0 246,0	36,2 85,0 174,0 251,0	40,1 83,0 167,0 247,0
	ψ = 45°	τ = 180°	13,2 15,8 17,7 20,0	23,9 36,4 42,4 38,4 38,9 42,4 38,4 38,4 38,4 39,9 39,9 30,9 30,9 30,9 30,9 30,9 30,9	18,1 26,7 31,6 31,5	27,3 37,0 41,4 45,5 43,6	26,2 34,1 39,8 39,8 8 6 8 7 8 7 8 7 8 7 8 7 8 7 8 7 8 7 8
		<del>ء</del> = 90°	20,5 24,4 30,55	28,4 33,1 32,4	24,1 28,7 34,4 36,8	35,2 46,5 43,8	37,4 43,7 50,3
		۲ == 0°	27,8 42,3 93,0 142,6	29,7 42,0 88,0 136,0	30,6 47,3 104,0 186,0	40,3 59,5 201,0	44,1 66,7 132,0 187,0
	ψ= 30°	-= 180°	14,8 18,9 20,4 21,9 21,9 21,9	31,1 40,0 39,3 37,7	22,0 33,0 33,0 33,0 5,0 5,0 5,0 5,0 5,0 5,0 5,0 5,0 5,0 5	33,6 48,0 38,4 83,4 84,0 8,0 8,4 8,0 8,0 9,0 9,0 9,0 9,0 9,0 9,0 9,0 9,0 9,0 9	29,6 40,2 47,4 44,0
		1 = 90°	21.9 25.5 29,8	33,5 32,8 34,8 32,2	28,1 31,8 37,1 37,0	41,4 45,5 49,5 44,4	42,3 48,3 55,2 51,3
		<b>4</b> = 0°	25,9 34,0 51,6 79,9	32,0 35,8 79,0	31,7 39,5 59,2 96,8	42,1 51,5 74,3 113,0	46,3 58,6 79,4 124,0
	¢ = 0°	l	16,0 25,3 29,9 29,9	37,9 37,4 35,3 31,3	25,1 31,5 34,0 39,2 37,5	39,5 44,2 51,57 44,57 44,45	33,3 45,4 56,1 56,1 51,5
	v. degrees		83 7 66 <del>5</del> 0 83 7 90 0	8750550 8750 8750 8750 8700 8700 8700 87	600 80 80 80 80 80 80 80 80 80 80 80 80 8	45 0 85 50 55 0	0 7 60 5 0 8 7 5 0 8 5
		$\Delta E_n$ . MeV	0,40,8	0,8—1,4	1,42,5	2,54	45

The value of  $A_{n,d}^{\pi}$  with a recording threshold  $E \approx 0.2$  MeV, although it differs slightly for spectra of neutrons of various "hardnesses," is still very low: from 0.14 for the spectrum of the Po-Be source to 0.125 for the reactor spectrum. For a water (polyethylene) reflector, the contribution of neutrons with energies E < 0.2 MeV to the dose is quite significant (about 30%), and cannot be ignored. For such media as iron, concrete, earth, the contribution of neutrons with energies E < 0.2 MeV to the total dose is slight (not over 10%).

Since the differential albedo for water is not described by simple approximation expressions Table 4.10 presents numerical data from [21].

Investigations of the dependence of integral albedo  $A_{n,d}^{\pi}(\gamma)$  on thickness of barriers of iron, aluminum, concrete, graphite, etc. performed on the basis of the results of [18, 22], have shown that it can be expressed by the following empirical relationship:

$$A_{n,d}^{\pi}(\gamma, \delta) = A_{n,d}^{\pi}(\gamma=0)\cos^{2/3}\gamma\left(1-e^{-\frac{0.14\lambda}{\cos\gamma}}\right). \tag{4.20}$$

In conclusion we note that the form of the **en**ergy distributions of neutrons reflected from barriers (for example, see [25]) differs little from the distributions beyond the protective shields with comparable scattering angles.

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CHAPTER 5. PENETRATION OF  $\gamma$  AND  $\beta$  RADIATION THROUGH SHIELDING

§ ]. Attenuation of Unscattered  $\gamma$  Radiation

The transmission factor of unscattered monochromatic  $\gamma$  radiation is calculated using expressions similar to formula (4.1), in which the total macroscopic cross section of the interaction of  $\gamma$  radiation with the medium is usually referred to as the linear attenuation factor and represented by  $\mu(E_{\gamma})$ .

The thickness of the absorber can be conveniently expressed in units of mass  $(g/cm^2)$ . Correspondingly,  $\mu$  is frequently expressed in square centimeters per gram  $(cm^2/g)$  and referred to as the mass absorption factor, represented as  $\mu/\rho$ . The mass absorption factor of chemically complex materials or mixtures consists of the sum of the mass absorption factors of the component elements, taken with their statistical weight, proportional to the content of each element.

The total attenuation factors for  $\gamma$  quanta in materials are complex functions of energy  $E_{\gamma}$ . At the present time, they have been determined with a mean accuracy of about 1% and presented in the form of tables or graphs in many works (for example, in [1]. However, in certain cases, particularly when calculations are performed by computer, expressions are required for the attenuation factors, written in the analytic form. A parabolic approximation for the total mass attenuation factor was suggested in [2]:

$$\frac{\mu(E_{\gamma})}{\rho} = c_0 + c_1 \varepsilon + c_2 \varepsilon^2, \qquad (5.1)$$

where  $\varepsilon$  is the wave length of a  $\gamma$  quantum in Comptom units;  $\varepsilon = E_{\gamma} (MeV) / 0.511$ .

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Formula (5.1) can be written in a form more convenient for practical calculations

$$\frac{\mu(E_{\gamma})}{\rho} = b_0 + \frac{b_1}{E_{\gamma}} + \frac{b_3}{E_{\gamma}^2}, \qquad (5.2)$$

the coefficients b and c being related by the formulas:  $b_0 = c_0; b_1 = 0.511 c_1; b_2 = 0.261 c_2.$ 

Table 5.1 presents the coefficients of the parabolic approximation for various materials. The maximum relative error of approximation using these coefficients, according to estimates of the authors of [2], will not exceed 4% in the 0.03-0.08 MeV energy range, 3% in the 0.08-0.03 MeV range and 5% in the 0.3-10 MeV range.

Approximate expressions with this level of accuracy should be used carefully, since when calculations with low error are required they are suitable only for low values of attenuation. For example, with an accuracy of 5% and an attenuation of the primary beam by a factor of 10, the error in calculating the attenuation factor of the unscattered radiation is 10%, while when the attenuation is by a factor of 1000 the error may reach 60%.

For complex chemical compounds not included in Table 5.1, the mass attenuation factors can be calculated based on the known dependences  $\mu_i(E_v)/\rho_i$  of the individual elements using the expression

$$\frac{\mu(E_{\gamma})}{\rho} = \sum \frac{\mu_i}{\rho_i} a_i, \qquad (5.3)$$

where  $a_i$  is the weight share of the ith element in the mixture or chemical compound.

If the shielding surface is struck by nonmonochromatic  $\gamma$  radiation, the attenuation of the unscattered component does not follow the simple exponential rule due to the filtration of hard  $\gamma$  quanta. In this case the attenuation factor of the unscattered component can be found by integration with respect to the entire spectrum  $\Phi(E_{\gamma})$  of the initial  $\gamma$  radiation similarly to expression (4.2).

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Table 5.1

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COEFFICIENT OF PARABOLIC APPROXIMATION, × 10<sup>-3</sup>

	$E_{\gamma} =$	0,03 + 0,08	MeV	E y =	0,08 + 0.3	MeV	Ey	, = 0,3 + 10	4eV
Element, Material	<i>b</i> <sub>0</sub>	Iq	<i>b</i> 3	•q	p,	p.	þe	Iq	P2
Carbon	135	2,07	0,201	71,5	23,3	1,56	14,9	113	
Oxygen	174	-8,65	1,04	75,4	21,1	-1,20	16,3	108	34,5
Sodium	246	-33,0	3,24	70,4	20,2	-0,808	17,2	99 <b>,</b> 3	30,7
Maqnesium	319	-53,5	4,95	76,1	18,4	0,381	18,3	100	30,4
Aluminum	353 ~	-65,3	6,10	69,8	20,3	-0,430	18,4	96'9	29,1
Silicon	411	-83,2	7,92	73,9	19,5	0,0400	19,7	98,2	28,9
Potass ium	866	-227	21,4	86,4	4,75	4,12	22,0	86,9	-22,8
Calcium	666		25,6	90,3	2,14	5,34	23,1	87,8	22,6
l ron	1730	-512	51,2	103	-21,7	13,8	24,8	72,1	-14,4
	2110	646	66,6	118	-40,4	20,1	25,9	65,1	9,85
Water	182 .	6,27	0,891	75,4	28,5	-2,0	0'11	124	40,5
Sand	289	-44,6	4,32	67,2	26,6	1,34	17,8	104	-32,1
Concrete	365	68,2	6,59	61,9	24,2	-0,739	18,2	103	31,6
					•				

#### § 2. Penetration of Scattered $\gamma$ Radiation Through Shielding

The concept of the transmission factor  $K_{\gamma s}$  for the dose of scattered radiation of a flat unidirectional beam, the ratio of the integral dose of scattered radiation reaching the internal surface of the barrier to the total dose of  $\gamma$  quanta striking its outer surface, is introduced in order to characterize the attenuation of scattered  $\gamma$  radiation in the barrier.

In [3], the Monte Carlo method was used to calculate the transmission factors for scattered radiation for barriers of concrete and iron (3.7-60.0 g/cm<sup>2</sup> thickness) upon incidence of a flat, unidirectional beam to the surface of the barrier at various angles. Table 5.2 shows the results relating to transmission of the energy of scattered  $\gamma$  radiation by concrete shields with various incident  $\gamma$  radiation energies. It is remarkable that for low barrier thicknesses, the maximum intensity of transmitted scattered radiation is observed at high angles of incidence of the primary  $\gamma$  radiation. The data of these same tables can be used to calculate the energy factor of accummulation with inclined incidence of  $\gamma$  radiation:

$$B_E(E_{\gamma}, \delta, \gamma) = e^{-\mu(E_{\gamma})^{\delta} \left(\frac{1}{\cos \gamma} - 1\right)} + K_{\gamma s}(E_{\gamma}, \delta, \gamma) e^{\mu(E_{\gamma})^{\delta}}.$$
 (5.4)

In [4], the dose factors of accummulation of  $\gamma$  radiation with initial energies from 0.2 to 10 MeV are calculated for inclined incidence of a broad unidirectional beam to concrete barriers (Table 5.3).

If the accummulation factor  $B(E_{\gamma}, \delta, \gamma)$  is known for a given barrier, the transmission factor for scattered  $\gamma$  radiation can be calculated using the expression

$$K_{\gamma s}(\dot{E}_{\gamma}, \, \delta, \, \gamma) = [B(E_{\gamma}, \, \delta, \, \gamma) - 1] e^{-\mu(E_{\gamma}) \frac{\delta}{\cos \gamma}}.$$
 (5.5)

For a complex spectrum of initial y radiation

$$K_{\gamma s}(\delta, \gamma) = \int_{E_{\gamma}} \Phi(E_{\gamma}) \left[ B(E_{\gamma}, \delta, \gamma) - 1 \right] e^{-\mu(E_{\gamma})} \frac{i}{\cos \gamma} dE_{\gamma}.$$
(5.6)

Broad information on the transmission of scattered  $\gamma$  radiation by screens of iron was produced in [5], in which calculations were performed by the Monte Carlonmethod. The energy groups of initial  $\gamma$ 

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Table 5.2

#### Barrier Thickness, g/cm<sup>2</sup> cos y 60 30 3,75 7.5 15 Initial Energy 0.35 MeV 0,0617 0,0004 0,0126 0,1250 0,1960 0 0,0254 0,0011 0,2480 0,1190 0,2210 0,25 0,0028 0,2290 0,0471 0,50 0,75 0,1660 0,1990 0,1980 0,0073 0,0786 0,2170 0,1640 0,0162 0,1980 0,2040 0,2020 1,00 0,1410 Initial Energy 0,66 MeV 0,0006 0,1150 0,1990 0,0170 0,0631 0,1725 0 0,1260 0,0034 0,25 0,50 0,75 0,0391 0,2050 0,1780 0,0728 0,0086 0,1590 0,2040 0,1130 0,0161 0,2030 0,1340 0,1900 0,0325 0,1490 0,1730 0,2030 1,00 0,1150 Initial Energy 1.25 MeV 0,0177 0,0011 0,0534 0,1460 0,1080 0 0,0039 0,25 0,1870 0,1690 0,1280 0,0450 0,1570 0,0159 0,0895 0,1700 0,50 0,75 0,1150 0,1880 0,1300 0,0341 0,1510 0,0891 0,0604 0,1570 0,0716 0,1820 0,1370 1,00 Initial Energy Corresponding to Spectrum of y Radiation of U<sup>235</sup> Fission Products 1.12 Hours After Explosion 0,0505 0.0018 0,0145 0,1520 0,1000 0 0,0049 0,1580 0,1170 0,0925 0,1770 0,1200 0,0460 0,25 0,1630 0,0905 0,0157 0,50 0,75 0,1640 0,1680 0,0386 0,1270 0,1520 0,0667 0,1360 0,1690 0,1430 0,0809 1,00 Initial Energy Corresponding to Spectrum of ¥ Radiation of U<sup>235</sup> Fission Products 23.8 hours After Explosion 0,0152 0.0015 0,0555 0,1140 0,1660 0 0,1240 0,1680 0,0042 0,0365 0,25 0,1970 0,1870 0,0108 0,0246 0,0417 0,0765 0,50 0,1870 0,1450 0,1850 0,1110 0,1150 0,1800 0,75 0,1880 0,1440 0,1640 1,00 0,0991 .

#### TRANSMISSION FACTORS OF SCATTERED Y RADIATION ENERGY FOR CONCRETE BARRIERS

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radiation were selected considering the convenience of application of these data for calculation of the attenuation of  $\gamma$  radiation of a nuclear explosion. The spectral intensity of  $\gamma$  quanta was considered constant within the limits of each energy group. The results of calculation of the transmission factors for scattered  $\gamma$  radiation are presented in Table 5.4.

In [6, 7], an empirical formula is presented for calculation of the accummulation factors from a flat isotropic source for a heterogeneous shield with any number of layers of various materials in the 1-7 MeV  $\gamma$  radiation energy range:

$${}^{B}_{het} = \sum_{n=1}^{N} B_n \left( \sum_{i=1}^{n} \mu_i \delta_i \right) - \sum_{n=2}^{N} B_n \left( \sum_{i=1}^{n-1} \mu_i \delta_i \right), \qquad (5.7)$$

where N is the total number of shielding layers;  $B_n$  is the accummulation factor of the homogeneous material of the i-th layer (the count of layers is performed from the source inward), taken with the corresponding thickness  $\sum_{n=1}^{n} \mu_n \delta_n$  in free path lengths.

Formula (5.7) was checked experimentally with a large number of combinations in the barrier geometry for a point isotropic and flat isotropic source of  $\gamma$  radiation. In both cases, the formula gave good correspondence. This allows us to assume that it will be correct for a flat unidirectional beam of initial  $\gamma$  radiation as well.

[Pages 126-129 of source missing]

Table 5.3

				-					
		For	'nð			-	For	μδ	
cos y	0,5	1,0	2,0	4,0	<b>c</b> os γ	0,5 .	1,0	2,0	4,0
	.	$E_{\gamma} = 0$	,2 Me	v I			$E_{\tau} = 4$	,0 MeV	
<b>.</b> 1.0	1,59	2,15	3,54	6,95	1,0	1,28	1,54	2,12	3,25
0.75	1,67	2,34	4,18	11,1	0,75	1,34	1,70	2,48	4,13
• 0,50	1,72	2,79	7,43	66,3	0,50	1,46	1,99	3,52	0,78
0.25	2,16	6,37	106		0,25	1,74	3,29	17,0	
0,10	8,76	_	-		0,10	4,92	-		-
,	1 '	$E_{\tau} =$	1 :0,5 .M	eV		•	$E_{\gamma} =$	6,0 Me	V
. 1.0	1.62	2,23	3,73	8,10	1,0	1,08	1,46	1,89	2,63
0.75	1.68	2,39	4,51	13,4	0,75	1,31	1,59	2,17	3,22
0.50	1.80	2,78	8,00	64,6	0,50	1,41	1,84	2,66	4,93
0.25	2,16	6,50	104	-	0,25	1,64	2,70	12,2	-
1.0	9,06		-		0,10	3,99		-	-
·	1	$E_{\gamma} =$	=1,0 <sup>°</sup> M	ĨeV .	1		$B_{\gamma} = 1$	0,0 Me	∍V́
1.0	1,45	1,88	2,90	5,78	1,0	1,16	1,32	1,61	2,07
0,75	1,49	2,04	3,49	9 8,94	0,75	1,20	)   1,41	1,82	2,81
0,50	1,65	5 2,42	5,70	0 38,5	0,50	1,28	3 1,56	2,18	5,25
0,25	1,88	3 4,96	62,2	-	0,25	1,48	3 2,28	9,03	-
0,10	6,98	3 -	-		0,10	3,20	v  −	-	-
	•	<i>Έ</i> γ ==	= 2,0 ີ	MeV					
1,0	1,38	8   1,73	3 2,5	2 4,29	Ð				
0,75	1,4	7 1,9	1 2,9	6,3	8				
0,50	1,5	5 2,2	2 4,6	5 17,1			、		
0,25	1,8	2 4,2	4 39,0	)					
0,10	6,0	2 –	-	-					

1

### DOSE ACCUMMULATION FACTORS FOR CONCRETE

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Table 5.4

# DOSE TRANSMISSION FACTORS OF SCATTERED $\gamma$ RADIATION FOR IRON SHIELDS



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We can see from Table 5.5 that characteristic angle  $\psi_0$  has a tendency to decrease with increasing energy of primary  $\gamma$  quanta and increasing atomic number Z of the scattering medium.

Information on the angular characteristics of scattered  $\gamma$  radiation with inclined incidence of the initial beam to the surface of the barrier is quite sparse. As a rule, the periodical literature presents only individual examples of results. In [15], the angular distributions of doses of scattered Co<sup>60</sup>  $\gamma$  radiation were investigated with inclined incidence to shields of iron and aluminum. Figure 5.1 shows the nature of the azimuthal dependence of the angular distribution function of  $\gamma$ quanta scattered on iron.





In [10, 12], it is demonstrated on the basis of experimental studies that the angular distributions of the dose of  $\gamma$  quanta scattered on a shield, as well as the differential and integral spectra are determined by the last layer of the shield if its thickness is over 1-2 free path lengths. The relative values of  $\gamma$  radiation doses depend on the order and are determined by the accummulation factors of the heterogeneous mixtures.

#### § 3. Reflection of $\gamma$ Quanta From Protective Shielding

The integral energy albedo of  $\gamma$  quanta in the range of energies of primary radiation 0.3 MeV <  $E_{\gamma}$  < 2.0 MeV is inversely proportional to the energy  $E_{\gamma}$  and the square of the atomic number of the scattering material Z. With increasing density of the material  $\rho$  and angle of incidence of the beam for a given Z, the energy albedo increases. The empirical dependence of the energy albedo on  $E_{\gamma}$ , Z,  $\rho$  and  $\gamma$ , according to [16], is

$$A_{\gamma E}^{n}(E_{\gamma}, \gamma) = 3.2 \frac{1}{E_{\gamma}} \frac{1}{\cos \gamma} \frac{\rho}{Z^{2}}$$
(5.12)

where 6 < Z < 50 and  $\gamma < 80^{\circ}$ ; and

$$A_{\gamma E}^{\pi}(E_{\gamma}, \gamma) = 3.2 \frac{1}{E_{\gamma}} \frac{1}{\cos^{2} \gamma} \frac{\rho}{Z^{2}}$$
(5.13)

where Z > 50 and  $\gamma$  < 80°.

In the area of energies of  $\gamma$  quanta of primary radiation 2.0 <  $E_{\gamma}$  < 10 MeV, according to [17, 18], the dependence of the energy albedo on  $E_{\gamma}$  is retained, while the dependence on atomic number Z of the scattering material is weaker due to the increasing influence of secondary annihilation radiation. With energy of primary  $\gamma$  quanta  $E_{\gamma}$  < 0.3 MeV, due to the strong photoelectric absorption, the energy albedo of  $\gamma$  quanta decreases with decreasing  $E_{\gamma}$  and increasing Z of the scattering material (Figure 5.2)[19].

The dose albedo of  $\gamma$  quanta differs somewhat from the energy albedo. Table 5.6 presents several values of dose and energy albedo for concrete.

The dependences (5.12) and (5.13) presented characterize the albedo of  $\gamma$  radiation for scattering materials of semi-infinite-thickness. With a thin reflector layer, the albedo does not reach its maximum value. Experiments [20] have shown that the dependence of  $\gamma$  quantum albedo on thickness of the scattering material  $\delta$  with normal incidence of the primary radiation is described for materials with Z  $\leq$  26 by the empirical expression

 $A_{\gamma q}^{\pi}(\delta) = A_{\gamma q}^{\pi}(\infty) \left[1 - e^{-2\gamma \{E_{\gamma}\}^{\delta}}\right], \qquad (5.14)$ 

where  $A_{\gamma d}^{\pi}(\infty)$  is the albedo for a semi-infinite scattering material. For heavy materials (lead) this dependence is

$$A_{y = 1}^{n}(\delta) = A_{y = 1}^{n}(\infty) \left[1 - e^{-8\nu (E_{y})^{2}}\right].$$
(5.15)

#### Table 5.6

	· · · · · · · · · · · · · · · · · · ·		Energy of	γ Quanta	E <sub>y</sub> , MeV
Degrees	Albedo	0,2	0,5	1,0	2,0
	$A^{\mathrm{n}}_{\gamma, E}(\gamma, E_{\gamma})$	0,154	0,085	0,041	0,017
0	$A^{\Pi}_{\gamma, \beta}(\gamma, E_{\gamma})$	0,138	0,074	0,040	0,020
	$A^{\mathfrak{n}}_{\gamma, E}(\gamma, E_{\gamma})$	0,170	0,097	0,051	0,021
30	$A^{n}_{\gamma, A}(\gamma, E_{\gamma})$	0,148	0,083	0,045	0,023
	$A^{\mathfrak{n}}_{\gamma, E}(\gamma, E_{\gamma})$	0,242	0,160	0,098	0,048
60	$\cdot A^{\mathbf{n}}_{\gamma, \mathbf{A}}(\gamma, E_{\gamma})$	0,220	0,146	0,099	0,055
	$A^{\pi}_{\gamma, E}(\gamma, E_{\gamma})$	0,500	0,413	0,348	0,266
~90	$A^{n}_{\gamma, A}(\gamma, E_{\gamma})$	0,470	0,395	0,355	0,303

DOSE AND ENERGY ALBEDO OF Y QUANTA FOR CONCRETE

With thicknesses equal to approximately two mean free path lengths, the scattering material can be considered semi-infinite from the standpoint of back scattering of  $\gamma$  radiation.

The spectra of the reflected radiation are significantly softer than the primary  $\gamma$  quanta. As an example, Figure 5.3 shows the distribution of Co<sup>60</sup>  $\gamma$  quanta reflected from concrete by energies for cases of normal and inclined incidence of primary  $\gamma$  radiation to the concrete. As we can see, the spectrum of the reflected radiation in all cases has a maximum at 0.16-0.30 MeV, regardless of the angle of incidence of the primary radiation. With inclined incidence of the  $\gamma$  radiation to the barrier surface, the spectrum of the reflected radiation becomes harder. This is natural, since in this case the  $\gamma$  quanta may return through the surface of the absorber after being scattered at lower angles and, consequently, changing their energy less than in reflection of the normal incident radiation.



Figure 5.2.

The nature of the angular dependence of back scattered radiation is determined by the competition of two processes: on the one hand, the probability of scattering and hardness of  $\gamma$  quanta increase with decreasing scattering angles; on the other hand, the path traveled by the scattered radiation in the material increases, consequently increasing the probability of absorption, which is greater, the greater the Z of the scattering material. Therefore, the angular distribution of intensities of scattered  $\gamma$  radiation is anisotropic.





Figure 5.3. Distribution of Co $^{60}$   $_{\gamma}$  Quanta Reflected From Concrete by Energies for Cases of Normal and Inclined Incidence of Primary Radiation to Concrete

The differential albedo of the dose of  $\gamma$  radiation, describing the relative value of the dose of  $\gamma$  quanta scattered per unit solid angle in the direction characterized by reflection angle  $\psi$  is determined by the following semi-semirical formula, suggested in [21]:

$$a_{\gamma d}^{\Pi \tau}(E_{\gamma}, \gamma, \psi, \tau) = \frac{C \Gamma(\psi_s) 10^{26} + C'}{1 + \cos \gamma \sec \psi}, \qquad (5.16)$$

where  $\Gamma(\psi_s)$  is the differential cross section of Comptom scattering per electron; C and C' are empirical constants, depending only on the energy of the incident radiation  $E_{\gamma}$  and the composition of the reflector. The scattering angle  $\psi_s$  is determined from expression (4.18). Expression (5.16) describes the differential albedo in concepts of current. For transition to flux quantities, it must be divided by cos  $\psi$ .

Based on expression (5.16), we can also calculate the integral albedo:

$$A_{\gamma d}^{\mathrm{hr}}(E_{\gamma}, \gamma) = C \left\{ \int_{0}^{1} \frac{d(\cos \gamma)}{1 + \cos \gamma \sec \psi} \int_{0}^{2\pi} \Upsilon(\psi_{s}) 10^{26} d\tau \right\} + C' \left\{ \left[ 1 - \cos \gamma \ln \frac{1 + \cos \gamma}{\cos \gamma} \right] \right\}.$$
(5.17)

The values of empirical constants C and C' produced for various materials in [22] are presented on Table 5.7.

#### Table 5.7

Scattering Material	E <sub>y</sub> ,MeV	с	C*
Water	0,2 0,662 1,00 2,50 6,13	$\begin{array}{c} -0,0187\pm0,0027\\ 0,0309\pm0,0047\\ 0,0470\pm0,0053\\ 0,0995\pm0,0068\\ 0,1861\pm0,0107\end{array}$	$\begin{array}{c} 0,1327\pm 0,0054\\ 0,0253\pm 0,0034\\ 0,0151\pm 0,0025\\ 0,0058\pm 0,0010\\ 0,0035\pm 0,0005\end{array}$
Concrete	0,2 0,662 1,00 2,50 6,13	$\begin{array}{c} 0,0023 \pm 0,0038 \\ 0,0347 \pm 0,0050 \\ 0,0503 \pm 0,0056 \\ 0,0999 \pm 0,0078 \\ 0,1717 \pm 0,0103 \end{array}$	$\begin{array}{c} 0,0737 \pm 0,0065 \\ 0,0197 \pm 0,0035 \\ 0,0118 \pm 0,0025 \\ 0,0051 \pm 0,0011 \\ 0,0048 \pm 0,0005 \end{array}$
Iron	0,2 0,662 1,00 2,50 6,13	$ \begin{vmatrix} 0,0272 \pm 0,0033 \\ 0,0430 \pm 0,0045 \\ 0,0555 \pm 0,0049 \\ 0,1009 \pm 0,0073 \\ 0,1447 \pm 0,0101 \end{vmatrix} $	$\begin{array}{c} 0,0100\pm 0,0062\\ 0,0063\pm 0,0030\\ 0,0045\pm 0,0021\\ 0,0044\pm 0,0010\\ 0,0077\pm 0,0006\end{array}$
Lead	0,2 0,662 1,00 2,50 6,13	$\begin{array}{c} 0,0044 \pm 0,0002 \\ 0,0308 \pm 0,0015 \\ 0,0452 \pm 0,0013 \\ 0,0882 \pm 0,0014 \\ 0,1126 \pm 0,0048 \end{array}$	$\begin{array}{c} 0,0050 \pm 0,0004 \\ 0,0100 \pm 0,0007 \\ 0,0083 \pm 0,0004 \\ 0,0001 \pm 0,0002 \\ 0,0063 \pm 0,0002 \end{array}$

## PARAMETERS C AND C' FOR CALCULATION OF DIFFERENTIAL DOSE ALBEDO USING SEMI-EMPIRICAL FORMULA (5.16)

In [22], the analytic dependences of constants C and C' on incident radiation energy are suggested for calculation of the differential albedo of concrete:

$$C = \exp \left[-2,92 + 0,681 \ln E_{\gamma} + 0,0111 (\ln E_{\gamma})^2 - 0,0413 (\ln E_{\gamma})^3\right];$$

. .

$$C' = \exp\left[-5,89 + 0,275 (E_{\gamma} - 3,25)^{2}\right] \text{ where } 0,1 \le E_{\gamma} \le 1,0 \text{ MeV};$$
(5.18)  
$$C' = \exp\left[-4,86 + 0,36 (E_{\gamma} - 2,0)^{2}\right] \text{ where } 1,0 < E_{\gamma} \le 2 \text{ MeV};$$

$$C' = \exp[-4,83 - 0.013E_{y}]$$
 where  $2.0 < E_{y} < 10$  MeV. (5.19)

# § 4. Attenuation of $\beta$ Radiation in Protective Shielding

For thick shielding with mass thickness  $\delta > 1-2 \text{ g/cm}^2$ , in which an electron is fully decelerated, the total energy of the Bremstrahlung radiation generated is proportional to the first power of the atomic number of the absorbing material Z [23]

$$\hat{C}(E_e, Z) = 5,77 \cdot 10^{-4} Z E_e^2 \text{ MeV/electron},$$
 (5.20)

where  $E_{e}$  is the energy of monoenergetic electrons.

The dose of Bremstrahlung at the internal surface of the shield can be calculated using the equation

$$D_{\rm rp} = \int_{E_{\gamma}} K_{\rm rp} \left( E_{\gamma} \right) \eta_{\rm rp} \left( E_{\gamma} \right) \Delta N_{\rm rp} \left( E_{\gamma} \right) E_{\gamma} dE_{\gamma}, \qquad (5.21)$$

where  $\Delta N_{tp}(E_{\gamma})$  is the number of Bremstrahlung quanta with energies in the interval from  $E_{\gamma}$  to  $E_{\gamma} + dE_{\gamma}$ , emitted by the source per unit time per unit area;  $n_{tp}(E_{\gamma})$  is the transfer factor from flux to dose power; it is equal to  $1.6 \cdot 10^{-8} \mu_a(E_{\gamma})$ , where  $\mu_a(E_{\gamma})$  is the electron conversion factor for biological tissue (cm<sup>2</sup>/g);  $K_{tp}(E_{\gamma})$  is the Bremstrahlung radiation attenuation factor of the shield material.

For electrons with energy  $E_{e}$ , the Bremstrahlung radiation spectrum

$$N_{\rm rp}(E_{\gamma}) = I(E_e) \Phi_{\rm rp}(E_{\gamma}) \widehat{C}(E_e, Z), \qquad (5.22)$$

(8 250

where  $I(E_e)$  is the current of electrons with energy  $E_e$  through the surface of the screen, related to each unit energy interval;  $F_{tp}(E_{\gamma})$  is the relative contribution of Bremstrahlung radiation of the various spectral components.

The spectral distribution of Bremstrahlung radiation with complete deceleration of the electrons, roughtly speaking, is independent of the atomic number of the absorber and the electron energy. It is demonstrated in [20], that the energy radiated per unit interval of energies is expressed approximately by the formula

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$$\Phi_{\rm rp}(E_{\rm y}) = {\rm const} \left[ 4 \left( 1 - \frac{E_{\rm y}}{E_{\rm e}} \right) - 3 \frac{E_{\rm y}}{E_{\rm e}} \ln \left( \frac{E_{\rm e}}{E_{\rm y}} \right) \right]. \tag{5.23}$$

Table 5.8 presents the relative contributions of various spectral components of Bremstrahlung radiation calculated according to this formula.

#### Table 5.8

E <sub>y</sub> /E <sub>e</sub>	$\Phi_{\tau p}(E_{\gamma})$ . %	E <sub>y</sub> /E <sub>e</sub>	Φ <sub>τρ</sub> ( <i>Ε</i> <sub>γ</sub> ), %
00, f	26,9	0,5-0,6	6,5
0,10,2	20,5	0,6-0,7	4,5
0,20,3	15,8	0,7-0,8	2,8
0,30,4	12,1	0,8-0,9	1,5
0,40,5	9,0	0,9-1,0	0,4

#### RELATIVE CONTRIBUTIONS OF VARIOUS SPECTRAL COMPONENTS OF BREMSTRAHLUNG

The final expression for Bremstrahlung radiation dose power considering attenuation in the shield will be:

$$D_{rp} = 1.8 \cdot 10^{-2} \int_{E_e} I(E_e) \widehat{C}(E_e, Z) dE_e \int_{E_y} K_{rp}(E_y) \times \\ \times \mu_a(E_y) \Phi_{rp}(E_y) dE_y. \quad , \qquad (5.24)$$

The attenuatio of Bremstrahlung radiation by the shield material can be considered if we assume that the radiation formed in an elementary deceleration event is sufficiently isotropic. With a shield thickness of  $\delta$ , the attenuation factor for Bremstrahlung radiation can be calculated using the expression

$$K_{\tau p}(E_{\gamma}) = B_{\tau p}(E_{\gamma}, \delta) Ei[-\mu(E_{\gamma})\delta], \qquad (5.25)$$

where the function Ei is the integral exponent

$$Ei(x) = \int_{x}^{\infty} \frac{e^{-y} \, dy}{y} \, .$$

The Bremstrahlung  $\gamma$  radiation dose accummulation factors  $B_{tp}(E_{\gamma})\delta$ ) can be produced on the basis of the data of [24, 25].

In [26], based on analysis of Monte Carlo method calculations of the transmission of electrons with energies of 0.4-10 MeV through various materials, apportential is suggested for calculation of the attenuation factor of electrons with normal incidence to the screen:

$$K_{e}(E_{e^{3}}\overline{\delta}_{s} Z) = \exp\left[-\left(\frac{0.634E_{e} Z^{-0.23}}{\overline{\delta}^{0.848}}\right)^{-7(Z-3.25)^{-0.24}}\right].$$
 (5.26)

where  $E_e$  is the energy of the incident electrons, MeV;  $\overline{\delta}$  is the mass thickness of the shield,  $g/cm^2$ .

The thickness of the absorbed dose of penetrating electrons with initial spectrum  $\Phi(E_{e})$  can be calculated from the expression

$$D_{e} = 0.8 \cdot 10^{-8} \int_{E_{e}} \Phi(E_{e}) \frac{E_{e}}{\lambda(E_{e})} K_{e}(E_{e}, \delta, Z) \times \\ \times dE_{e} \frac{\mathrm{mrad}}{\mathrm{hr}} / \frac{\mathrm{electrons}}{\mathrm{cM}^{2} \cdot \mathrm{sec}}$$
(5.27)

where  $\lambda(E_e)$  is the path of an electron with energy  $E_e$  in biological tissue, g/cm<sup>2</sup>. The value of the paths of electrons can be produced using formula [25]

 $\lambda(E_e) = (0.542E_e - 0.133) \text{ g/cm}^2.$  (5.28)

It should be noted on the whole that with a shielding thickness  $\delta < 1.0 \text{ g/cm}^2$ , the majority of the total dose of  $\beta$  radiation is created by electrons penetrating through the shield. With a shield thickness  $\overline{\delta} > 1-2 \text{ g/cm}^2$ , the dose is determined entirely by Bremstrahlung.

#### § 5. Yield of Secondary $\gamma$ Radiation From the Shield

As they pass through the shield and interact with the nuclei of elements in the composition of the shield, neutrons generate secondary  $\gamma$  radiation. The formation of secondary  $\gamma$  radiation primarily results from the reaction of radiation capture of neutrons. Secondary  $\gamma$  radiation in some cases makes a significant contribution to the total dose beyond the shielding. In order to calculate the dose of secondary  $\gamma$  radiation beyond the shielding, it is necessary to know the spectral composition of the capture  $\gamma$  radiation, its total production and its angular distribution.

The spectral composition of prompt  $\gamma$  quanta with radiation capture of thermal neutrons has been well studied, and the primary results were collected in [27]. The yield of capture  $\gamma$  radiation from the protective barrier is determined primarily by the distribution of sources of capture  $\gamma$  radiation through the volume of the barrier. The distribution of sources of capture  $\gamma$  radiation in the absorbing layer corresponds to the spatial distribution of thermal neutrons in the shield and neutrons moderated to thermal velocities.

The rate of absorption of neutrons per unit volume can be written as

$$\sum_{a} \Phi_{th}$$
(5.29)

where  $\Sigma_a$  is the macroscopic cross section of absorption of thermal neutrons;  $\Phi_{th}$  is the flux of thermal neutrons.

The distribution of the flux of thermal neutrons in the barrier can be found by solving the diffusion equation

$$\frac{d^2 \Phi_{th}(x)}{dx^2} - \frac{\Phi_{th}(x)}{L^2} + \frac{S(x)}{D} = 0, \qquad (5.30)$$

where S(x) is the density of generation of thermal neutrons in the shield barrier itself; L is the diffusion length of thermal neutrons in the shielding medium; D is the coefficient of diffusion of thermal neutrons.

If generation of thermal neutrons does not occur in the barrier and the only source of these neutrons is the external field of thermal neutrons diffusing inward, the solution of the diffusion equation is as follows:

$$\Phi_{th}(x) = I_{th}(0) \frac{L}{D} \frac{\sinh \frac{\delta + 2D - x}{L}}{\cosh \frac{\delta + 2D}{L}},$$
(5.31)

where  $I_{th}(0)$  is the current of thermal neutrons striking the surface of the shield. Actually, the formation of thermal neutrons in the protective shield always occurs. A portion of the fast neutrons, being moderated, goes over into the group of thermal neutrons, which may be captured by nuclei of the media and thus become sources of capture  $\gamma$  radiation.

In order to solve the diffusion equation in this case it is convenient to use the two-group theory in the form in which it is used for calculation of the fluxes of thermal and fastnneutrons in reactor design. In the case of neutrons of moderate energies, the theory is quite sufficient, but for neutrons having high energy and penetrating to greater distances, the theory produces significant error.

The solution in the two-group approximation is

$$\Phi_{th}(0) = \frac{I_f(0)}{\left(\frac{L_s}{L}\right)^3 - 1} \left( \frac{L_s \operatorname{sh} \frac{\delta + 2D - x}{L_s}}{D \operatorname{ch} \frac{\delta + 2D}{L_s}} - \frac{L}{D} \frac{\operatorname{sh} \frac{\delta + 2D - x}{L}}{\operatorname{ch} \frac{\delta + 2D}{L}} \right) + I_{th}(0) \frac{L}{D} \frac{\operatorname{sh} \frac{\delta + 2D - x}{L}}{\operatorname{ch} \frac{\delta + 2D}{L}}.$$
(5.32)

where  $L_s$  is the length of moderation.

The nature of the distribution of sources of capture  $\gamma$  quanta in the shield medium corresponds with the distribution of thermal neutrons in the medium. The number of  $\gamma$  quanta produced in a unit volume about point N is proportional to  $\Phi(x)\Sigma_{a}$ .

Thus, in order to determine the yield of capture  $\gamma$  quanta, we must find the flux of capture  $\gamma$  quanta at the surface of the shield generated at various depths in the barrier.

Let us study an infinitely thin flat isotropic source of  $\gamma$  quanta located at depth 8 in a plate with a thickness of  $\delta$  (Figure 5.4).

Let us determine the flux  $\Phi_\gamma$  -- the number of  $\gamma$  quanta which are not scattered -- at point A.



Figure 5.4. Calculation of Yield of Capture  $\gamma$  Radiation From Shielding

We represent by r the distance from a certain point in the plane of the source to the x-axis, passing through point A perpendicular to the plane of the source. The number of photons leaving each unit area of the source per unit time is equal to  $I_x$ . If these photons are radiated isotropically, their flux at point A will be

$$d\Phi_{\gamma} = I_x e^{-\mu (E_{\gamma})^{\natural} \sec \psi} \frac{dr}{r \csc^2 \psi}.$$
 (5.33)

The flux of photons  $\Phi_{\gamma}$  at point A from the entire plane of the source will be:

$$\Phi_{\gamma} = I_{x} \int_{\mu b}^{\infty} e^{-y} \frac{dy}{y} = -I_{x} Ei(-\mu \delta), \qquad (5.34)$$

where -Ei (-z) =  $\int_{z}^{\infty} e^{-y} \frac{dy}{y}$  is the integral exponent.

In order to consider scattered photons, we must introduce the accummulation factor B, i. e.

$$\Phi_{\gamma} = -I_{x}BEi (-\mu\delta). \tag{5.35}$$

Let us now determine the value of current dI of  $\gamma$  quanta at point B from point sources located in the circle r, r + dr:

$$dI = I_x = e^{-\mu\delta \sec\psi} \frac{dr}{r \csc^2\psi} \cos\psi. \qquad (5.36)$$

The total current of  $\gamma$  quanta at point A from the entire flat source will be

$$I = I_{x} [e^{-\mu \delta} + \mu \delta Ei (-\mu \delta)].$$
 (5.37)

If the thermal neutrons are captured during the process of diffusion through the plate shown on Figure 5.4, we can use equation (5.31), where there are no fast neutrons, to determine the rate of capture of thermal neutrons per unit area in a layer of thickness dx:

$$\Sigma_a I_{th}(0) \frac{L}{D} \frac{\operatorname{sh} \frac{\delta + 2D - x}{L}}{\operatorname{ch} \frac{\delta + 2D}{L}} dx.$$
(5.38)

With sufficiently large values of  $\frac{\delta+2D}{L}$ , this expression is simplified:

$$\eta_{\gamma} \frac{I_{th}(0)}{L} e^{-x/L} dx, \qquad (5.39)$$

where  $\eta_{\gamma}$  is the yield of  $\gamma$  quanta per captured neutron.

This formula gives us the number of capture  $\gamma$  quanta arising in a layer of thickness dx per unit surface area.

In layer dx, the currents of  $\gamma$  quanta in direction x and in the opposite direction are equal; therefore,

$$I_x = \frac{1}{2} \frac{I_{th}(0)}{L} e^{-x/L} dx.$$
 (5.40)

Let us now determine the total current of all capture  $\gamma$  quanta generated in a flat screen of thickness  $\delta$ . For this, in expression (5.37) we substitute the value of I<sub>x</sub> from formula (5.40). A convenient approximation is produced if coefficient  $\mu$  is replaced by the linear  $\gamma$  quantum energy absorption factor  $\mu_{\rm E}$  and at the same time B is assumed equal to 1. Then we produce the following expression for the resulting currente of  $\gamma$  quanta formed in layer dx:

$$\frac{1}{2} \varepsilon_{\gamma} \frac{I_{th}(0)}{L} e^{-x/L} \left[ e^{-\mu_E x} + \mu_E x Ei \left( -\mu_E x \right) \right] dx,$$

and from all capture  $\gamma$  quanta generated in the barrier and leaving its surface in the direction of the neutron source,

$$I_{c} = I_{th}(0) \frac{\varepsilon_{\gamma}}{2L} \int_{0}^{\delta} e^{-x/L} \left[ e^{-\mu_{E}x} + \mu_{E}xEi(-\mu_{E}x) \right] dx =$$

$$= I_{th}(0) \frac{\varepsilon_{\gamma}}{2L} \left\{ -\left[ \frac{e^{-(1/L+\mu_{E})x}}{\frac{1}{L}+\mu_{E}} \right]_{0}^{\delta} + \mu_{E} \int_{0}^{\delta} e^{-x/L} xEi(-\mu_{E}x) dx \right\},$$
(5.41)

where  $\varepsilon_{\gamma}$  is the total energy of  $\gamma$  radiation generated upon capture of one neutron, equal to the product of the yield  $\eta_{\gamma}$  times the mean energy of capture  $\gamma$  quanta.

Integrating this expression [28], we produce

$$I_{c} = I_{th}(0) \frac{E_{\gamma}}{2} \mu_{E} L \left\{ \frac{1}{\mu_{E} L} \left[ 1 - e^{-(1 + \mu_{E} L)^{\delta/L}} \right] - \left( \frac{\delta}{L} + 1 \right) \times e^{-\delta/L} Ei \left( -\mu_{E} L \frac{\delta}{L} \right) + Ei \left[ -(1 + \mu_{E} L) \frac{\delta}{L} \right] - \ln \frac{1 + \mu_{E} L}{\mu_{E} L} \right\} = I_{th}(0) K_{c}.$$
(5.42)

Calculation of the flux through the surface opposite to the surface struck by the neutrons is performed similarly:

$$I_{c} = I_{ih} (0) \frac{\epsilon_{v}}{2L} \int_{0}^{b} e^{-x/L} \left\{ e^{-\mu_{E}(\delta - x)} + \mu_{E}(\delta - x) Ei \left[ -\mu_{E}(\delta - x) \right] \right\} dx = I_{ih} (0) \frac{\epsilon_{v}}{2L} \left\{ -\frac{e^{-\mu_{E}\delta}}{\frac{1}{L - \mu_{E}}} \left[ e^{-(1/L - \mu_{E})x} \right]_{0}^{\delta} + \mu_{E} e^{-\delta/L} \int_{0}^{b} e^{\delta - x/L} (\delta - x) \times X_{Ei} \left[ -\mu_{E}(\delta - x) \right] \right\} dx = L_{ih} (0) \frac{\epsilon_{v}}{2L} \left\{ -\frac{e^{-\mu_{E}\delta}}{\frac{1}{L - \mu_{E}}} \left[ e^{-(1/L - \mu_{E})x} \right]_{0}^{\delta} + \mu_{E} e^{-\delta/L} \int_{0}^{b} e^{\delta - x/L} (\delta - x) \times X_{Ei} \left[ -\mu_{E}(\delta - x) \right] \right\} dx = L_{ih} (0) \frac{\epsilon_{v}}{2L} \left\{ -\frac{e^{-\mu_{E}\delta}}{\frac{1}{L - \mu_{E}}} \left[ e^{-(1/L - \mu_{E})x} \right]_{0}^{\delta} + \mu_{E} e^{-\delta/L} \int_{0}^{b} e^{\delta - x/L} (\delta - x) \times X_{Ei} \left[ -\mu_{E}(\delta - x) \right] \right\} dx = L_{ih} (0) \frac{\epsilon_{v}}{2L} \left\{ -\frac{e^{-\mu_{E}\delta}}{\frac{1}{L - \mu_{E}}} \left[ e^{-(1/L - \mu_{E})x} \right]_{0}^{\delta} + \mu_{E} e^{-\delta/L} \int_{0}^{b} e^{\delta - x/L} (\delta - x) \times X_{Ei} \left[ -\mu_{E}(\delta - x) \right] \right\} dx = L_{ih} (0) \frac{\epsilon_{v}}{2L} \left\{ -\frac{e^{-\mu_{E}\delta}}{\frac{1}{L - \mu_{E}}} \left[ e^{-(1/L - \mu_{E})x} \right]_{0}^{\delta} + \mu_{E} e^{-\delta/L} \int_{0}^{b} e^{\delta - x/L} (\delta - x) \times X_{Ei} \left[ -\mu_{E}(\delta - x) \right] \right\} dx = L_{ih} (0) \frac{\epsilon_{v}}{2L} \left\{ -\frac{e^{-\mu_{E}\delta}}{\frac{1}{L - \mu_{E}}} \left[ e^{-(1/L - \mu_{E})x} \right]_{0}^{\delta} + \mu_{E} e^{-\delta/L} \int_{0}^{b} e^{\delta/L} \left[ e^{-(1/L - \mu_{E})x} \right]_{0}^{\delta} + \mu_{E} e^{-\delta/L} \left[ e^{\delta/L} \right]_{0}^{\delta} + \mu_{E} e^{-\delta/L} \left[ e^{-(1/L - \mu_{E})x} \right]_{0}^{\delta} + \mu_{E} e^{-\delta/L} \left$$

After integration [28] we produce the function K', characterizing the yield of energy of capture  $\gamma$  quanta generated per incident thermal neutron on the opposite surface of the shield

$$I_{c}' = I_{th}(0) \varepsilon_{\gamma} \frac{\mu_{E} L}{2} e^{-\delta/L} \left\{ \frac{1}{\mu_{E} L} \left[ e^{(1-\mu_{E} L) \delta/L} - 1 \right] + \overline{E}i \left[ (1-\mu_{E} L) \frac{\delta}{L} \right] - \ln \frac{1-\mu_{E} L}{\mu_{E} L} \right\} = I_{th}(0) K_{c}', \qquad (5.44)$$

where

e.

$$\overline{E}i(z) = \int_{-\infty}^{z} \frac{e^{y}}{y} dy.$$

If (1 -  $\boldsymbol{\mu}_E^{}L)$  is negative, we must replace the expression

$$\overline{E}i\,\mu_E\left[\left(1-\mu_E\,L\right)\frac{\delta}{L}\right] - \ln\frac{1-\mu_E\,L}{\mu_E\,L}$$

with another expression

$$Ei\left[-\left(-\mu_E L-1\right)\frac{\delta}{L}\right] - \ln\frac{\mu_E L-1}{\mu_E L}.$$
(5.45)

In designing shielding, we must consider also the capture  $\gamma$  radiation formed upon capture of moderated fast neutrons. To do this, in place of expression (5.39) we use the first portion of formula (5.32) which, with sufficiently high values of  $\delta$  + 2D/L and  $\delta$  + 2D/L, will be:

$$\frac{\frac{I_f(0)}{L_s} \left(\frac{L_s}{L}\right)^{\mathbf{a}}}{\left(\frac{L_s}{L}\right)^{\mathbf{a}} - 1} \cdot e^{-x/L_s} - \frac{\frac{I_f(0)}{L}}{\left(\frac{L_s}{L}\right)^{\mathbf{a}} - 1} e^{-x/L} \cdot (5.46)$$

This expression must be introduced to formula (5.44).

In order to simplify calculation, we can replace L with  $L_s$  in (5.44) [28]:

$$\overline{I}_{cf}' = I_f(0) \varepsilon_{\gamma} \frac{\mu_E L_s}{2} e^{-ifL_s} \left\{ \frac{1}{\mu_E L_s} \left[ e^{(1-i_E L_s)} - 1 \right] + \left( \frac{\delta}{L} - 1 \right) e^{i/L_s} Ei \left( -\mu_E L_s \frac{\delta}{L_s} \right) + \overline{Ei} \left[ (1-\mu_E L_s) \frac{\delta}{L_s} \right] - \frac{1-\mu_E L_s}{\mu_E L_s} \right\} = I_f(0) \overline{K}_{cf}'.$$
(5.47)

According to expression (5.46), the value of  $\overline{K}_{cf}$ ' produced must be multiplied by

$$\frac{\left(\frac{L_s}{L}\right)^{\mathbf{s}}}{\left(\frac{L_s}{L}\right)^{\mathbf{s}}-1},$$

and the value of  $K_{\mbox{\scriptsize C}}^{\,\prime},$  corresponding to the same conditions, by

$$\frac{1}{\left(\frac{L_s}{L}\right)^2 - 1}$$

then we subtract the second product from the first product:

$$K'_{ef} = \overline{K}'_{ef} \frac{\left(\frac{L_s}{L}\right)^3}{\left(\frac{L_s}{L}\right)^2 - 1} - K'_e \frac{1}{\left(\frac{L_s}{L}\right)^3 - 1}.$$
(5.48)

Similarly, we can find the yield of capture  $\gamma$  radiation energy on the side of the source:

$$K_{cf} = \overline{K}_{cf} \frac{\left(\frac{L_s}{L}\right)^2}{\left(\frac{L_s}{L}\right)^2 - 1} - K_c \frac{1}{\left(\frac{L_s}{L}\right)^2 - 1}$$
(5.49)

These functions characterize the yield of capture  $\gamma$  radiation on the surface of the shield resulting from moderated neutrons.

Table 5.9 shows constants for calculation of the yield of capture  $\gamma$  radiation from certain materials.

Table 5.9

Material	L, см	D, сж	Σ <sub>а</sub> , см	L <sub>s</sub> , см	$n_{\gamma}, Quanta Neutrons$	$n_{\overline{\gamma}}$ , MeV	μ <sub>E,</sub> cm <sup>-1</sup>
Water tron Concrete	2,9 1,5 4,85	0,14 0,461 0,342	0,017 0,205 0,0146	5,6 16,0 8,65	1,0 1,1 1,0	2,23 6,4 6,0	9,0241 0,177 0,048

CONSTANTS FOR CALCULATION OF YIELD OF CAPTURE  $\gamma$  RADIATION

The results of calculation of the yield of capture  $\gamma$  radiation from shield barriers of various compositions and thickness are presented on Figure 5.5.

One important characteristic of secondary  $\gamma$  radiation is its angular distribution. If the protective shield is represented as a barrier of infinite area and thickness  $\delta$ , the angular distributions of the capture  $\gamma$  radiation leaving the barrier will be independent of the azimuthal angle **and** the points near which they are studied.

Since the distribution function of neutron capture density can be expressed by an exponential dependence, simple calculations allow us to produce the angular distribution function of capture  $\gamma$  radiation energy flux generated by thermal neutrons diffusing from without:

$$f'_{c}(\psi) = \frac{I_{th}(0)}{I'_{c}(\delta)} \varepsilon_{\gamma} \frac{\cos \psi}{4\pi} \frac{e^{-\frac{\mu_{E}\delta}{\cos \psi}} - e^{-\frac{\delta}{L}}}{\cos \psi - \mu_{E}L}; \qquad (5.50)$$

$$f_{c}(\psi) = \frac{I_{th}(0)}{I_{c}(\delta)} \varepsilon_{\gamma} \frac{\cos \psi}{4\pi} \frac{1 - e^{-\frac{\delta}{\cos \psi}} \left(\mu_{E} + \frac{\cos \psi}{L}\right)}{L\mu_{E} + \cos \psi}. \qquad (5.51)$$

 $f_{ef}(\psi) = \frac{I_f(0) \epsilon_{\psi}}{I_{cf}'\left[\left(\frac{L_s}{L}\right)^2 - 1\right]} \frac{\cos\psi}{4\pi} \left[\left(\frac{L_s}{L}\right)^2 \frac{e^{-\frac{ik_E^2}{\cos\psi}} - \frac{i}{L_s}}{\cos\psi - \mu_E L_s} - \frac{i}{L_s}\right]$  $-\frac{e^{-\frac{\mu E^{i}}{\cos\psi}}-e^{-\frac{i}{L}}}{\cos\psi-\mu_{E}L};$ (5.52) $f_{cf}(\psi) = \frac{I_f(0) \varepsilon_{\gamma}}{I_{cf} \left[ \left( \frac{L_s}{L} \right)^2 - 1 \right]} \frac{\cos \psi}{4\pi} \left[ \left( \frac{L_s}{L} \right)^2 \times \right]$  $\frac{1-e^{-\frac{\lambda}{\cos\psi}}\left(\mu_E-\frac{\cos\psi}{L_s}\right)}{\frac{L_s\,\mu_E-\cos\psi}{L_s\,\psi}} \frac{1-e^{-\frac{\lambda}{\cos\psi}}\left(\mu_E-\frac{\cos\psi}{L}\right)}{\frac{L_{\mu_E}+\cos\psi}{L_{\mu_E}+\cos\psi}}\right].$ (5.53). 06-10 Concrete 0,5.10 Concrete rad/(neutrons/cm<sup>2</sup>) Q.4·10 H20 H2 0,3 10 Concrete 0210 H\_O ₽С Н20 0,140 Concrete Fe Fe 18 20 14 16 12 4 6 8 ð си

Figure 5.5. Functions of Yield of Secondary  $\gamma$  Radiation From Protective Barrier: A, on side of source (K); B, on opposite side (K'); -----, due to capture of fast neutrons moderated in barrier (K<sub>cf</sub>); -----, due to capture of thermal neutrons striking barrier from without (K<sub>c</sub>)

The angular distribution functions of capture  $\gamma$  radiation energy flux resulting from moderated fast neutrons can be represented as

These functions characterize the portion of the total energy (dose) of capture  $\gamma$  radiation on the surface of the barrier leaving per unit solid angle in the direction defined by angle  $\gamma$ , which is counted from the normal to the surface of the barrier.

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CHAPTER 6. DESIGN OF SHIELDING OF SURFACE STRUCTURES FROM THE RADIATION OF A NUCLEAR EXPLOSION

\$ l. General Form of Angular Distribution Functions of Doses of Neutrons and  $\gamma$  Radiation During a Nuclear Explosion

The angular energy distribution functions of neutrons and  $\gamma$  radiation from a nuclear explosion presented in Chapter 2 describe the field of radiation near the surface of the earth when the position of the center is known. However, the specific position of the center of an explosion is always unknown.

If the distribution of probability densities of appearance of centers of explosions in the azimuthal direction around a position to be designed is considered even, corresponding to the rule of circular scattering of shells relative to an aiming point, the characteristics of the shielding properties of a structure can be determined as the mean result produced for various azimuthal positions of the epicenter of the explosion around the object being designed. This approach greatly increases thevvolume of calculation in determining the protective properties of a structure. A significant reduction of the required volume of calculation can be achieved if we do not use the expressions relating to a fixed explosion center as the angular distribution functions, but rather use certain general functions considering the equal probability of appearnace of an explosion at various azimuthal positions [1].

For neutrons, this function can be calculated as follows:

 $F_{n}(\Omega, \beta) = \frac{1}{\pi} \int_{0}^{\pi} \{0,033 + 0,4045 \exp\left[-0,03345 \arccos \times (6.1)\right] \times (\sin\beta\cos\varphi\cos\alpha + \cos\beta\cos\alpha)\} d\varphi \quad \text{ster}^{-1}.$ 

the results of calculation of this generalized angular distribution function for neutrons of a nuclear explosion, based on expression (6.1) are shown on Figure 6.1. This function describes the share of the total dose of neutrons arriving at a detector per unit solid angle from the direction determined by angle  $\beta$ , which is read from the normal to the surface of the earth.

The angular distribution functions of doses of fragment  $\gamma$  radiation are calculated similarly:

$$F_{f}(\Omega, \beta) = \frac{1}{\pi} \int_{0}^{\infty} \{0,0092 + 3,22 \exp \left[-0,08 \arccos \left(\sin \beta \times \cos \varphi \cos \alpha + \cos \beta \sin \alpha\right)\right] \} d\varphi$$
(6.2)

as are the functions for capture  $\boldsymbol{\gamma}$  radiation

$$F_{c}(\Omega, \beta) = \frac{1}{\pi} \int_{0}^{\pi} \{0,008 + 2,11 \exp[-0,065 \arccos(\sin\beta) \times (6.3)]\}$$

 $\times \cos \varphi \cos \alpha + \cos \beta \sin \alpha$ ] dq.



Figure 6.1. Generalized Angular Distribution Functions of Doses of Nuclear Explosion Neutrons  $F_n(\Omega, \beta)$ 

The results of calculations of the generalized angular distribution functions of the  $\gamma$  radiation of a nuclear explosion scattered in the air are shown on Figure 6.2. When the data of Figure 6.2 are used in practical calculations for the design of shielding, the quantities F( $\Omega$ ,  $\beta$ ) must be multiplied by the corresponding values of coefficients K<sub> $\alpha$ </sub>, cal-

culated in correspondence with the design elevation angle of the explosion  $\alpha$  (symbols for angles shown on Figure 2.10).

The generalized angular distribution functions of doses of  $\gamma$  radiation from a nuclear explosion can be conveniently used to estimate the penetration of radiation through spaces in structures, when the  $\gamma$  quanta scattered in the air can act directly on the people in the structure without undergoing the attenuating effects of the walls and roof. These data are insufficient to design closed shielding or sectors covered with various shields, since the spectral characteristics of the radiation change significantly as a function of angle  $\beta$ .

The data presented in Chapter 2 on the spectral and angular distributions of  $\gamma$  quanta scattered in the air following annuclear explosion allow us to calculate the generalized characteristics of the spectral-angular distributions of the  $\gamma$  radiation. These calculations were performed using the equation

$$F_{f}(\Omega, \beta, E_{\gamma}) = \frac{1}{\pi} \int_{0}^{\pi} \Xi_{f}'(0, E_{\gamma}) \left[0,0092 + 3,22 \exp\left(-0,080\right)\right] d\phi \quad (6.4)$$

for fragment  $\gamma$  radiation and

$$F_{c}(\Omega, \beta, E_{\gamma}) = \frac{1}{\pi} \int_{0}^{\pi} \Xi_{c}'(\theta, E_{\gamma}) [0,008 + 2,11 \exp(-0,065\theta)] d\phi \qquad (6.5)$$

for capture  $\gamma$  radiation.

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In these expressions  $\theta$  = arccos (sin  $\beta$  cos  $\phi$  cos  $\alpha$  + cos  $\beta$  sin  $\alpha$ ). The value of these functions calculated for various energy intervals  $\Delta E_{\gamma}$ , angles  $\beta$  and  $\alpha$  are presented in Tables 6.1 and 6.2. When performing calculations using equations (6.4) and (6.5), the functions  $\Xi(\theta, E_{\gamma})$ were selected from Tables 2.6 and 2.7 in correspondence with the expression

$$\Xi'(0, E_{\gamma}) = \frac{\Xi(0, E_{\gamma})}{100}, \qquad (6.6)$$

keeping in mind the assumption that the spectral distribution of the doses

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shown in Tables 2.6 and 2.7 for angular intervals  $\Delta \theta$  remain unchanged with all angles  $\theta$  within the limits of the interval  $\Delta \theta$  in question.

Thus, the functions  $F(\Omega, \beta, E_{\gamma})$  produced characterize the portion of the total dose of scattered radiation of this component in the energy interval  $\Delta E_{\gamma}$  arriving per unit solid angle from the direction determined by angle  $\beta$ , under the condition that the distribution function of probability densities of explosion centers in the azimuthal direction relative to the design position is constant. These results represent exhaustive information on radiation acting on structures from the upper half space over the surface of the earth following a nuclear explosion.

As concerns radiation scattered from the ground, no additional calculations are necessary with respect for neutrons, since the angular distribution function of the dose of neutrons reflected from the ground is isotropic. The angular distribution of neutrons incident to the surface of structures from the earth from any direction depends only on angle  $\beta$ , read from the normal to the surface of the earth.

The angular distribution of the dose of initial  $\gamma$  radiation reflected from the earth does not have azimuthal symmetry. However, the use of the assumption of evenness of distribution of the probability of appearance of explosion centers in the azimuthal direction around the calculated position makes it possible to produce a generalized distribution function of  $\gamma$  quantum reflected from the earth, in which the angular dose power depends only on one parameter -- angle  $\beta$ .

The differential albedo of  $\gamma$  quanta of a nuclear explosion can be calculated using the formula

$$\mathbf{a}_{\boldsymbol{\gamma},\boldsymbol{d}}^{\boldsymbol{\pi}} = 2 \sum_{E_{\boldsymbol{\gamma}}} \int_{0}^{\pi/2} \int_{0}^{\pi} \overline{K}_{\alpha} F(\Omega, \beta, E_{\boldsymbol{\gamma}}) \mathbf{a}_{\boldsymbol{\gamma},\boldsymbol{d}}^{\boldsymbol{\pi}\boldsymbol{\pi}}(E_{\boldsymbol{\gamma}}, \psi_{s}) \sin\beta \, d\beta \, d\varphi \,, \qquad (6.7)$$

where  $a_{\gamma, d}^{\pi t}(E_{\gamma}, \psi_s)$  is the differential albedo for a broad parallel beam of  $\gamma$  quanta for earth described by formula (5.16).

The total scattering angle  $\psi_{\rm S}$  of reflected radiation is calculated using the equation

 $\cos \psi_s = \sin \beta \sin \psi \cos \varphi - \cos \beta \cos \psi, \qquad (6.8)$ 

where  $\psi$  is the exit angle of the reflected radiation, read from the normal to the surface of the earth. The remaining angles have values corresponding to Figure 2.10.

Table 6.1

SPECTRAL-ANGULAR DISTRIBUTION OF FRAGMENT  $\gamma$  RADIATION

 $F_f(Q, \beta, E_V) = \frac{rad}{ster}$  (rel. units)

•

000000000000000000000000000000000000000	3.         0.1-0.2           70         0.00693           10         0.00633           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00554           10         0.00556           10         0.00556           10         0.00556           10         0.00556           10         0.00556           10         0.00556           10         0.00556           10         0.00556           10         0.00556           10         0.00556           10         0.00556           10         0.00556	Energy Interval ∆E <sub>Y</sub> , MeV	0.2-0.3 0.3-0.5 0.5-0.75 0.75-1 1-1.5 1.5-2 2-3 3-4 4-6	$\alpha = 0^{\circ}$	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	α==10°	$ \begin{bmatrix} 0,00267 & 0,00461 & 0,477 (-3) \\ 0,00460 & 0,00401 & 0,00158 \\ 0,00224 & 0,00427 & 0,00508 \\ 0,00521 & 0,00505 & 0,00508 \\ 0,00505 & 0,00503 & 0,229 (-3) \\ 0,00408 & 0,00503 & 0,00602 \\ 0,00408 & 0,00808 & 0,00040 \\ 0,00821 & 0,00821 & 0,00822 \\ 0,00406 & 0,00808 & 0,00835 \\ 0,00821 & 0,00835 & 0,02813 \\ 0,00432 & 0,00745 & 0,00835 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00745 & 0,00917 \\ 0,00432 & 0,00536 & 0,00917 \\ 0,00432 & 0,00536 & 0,00917 \\ 0,00432 & 0,00536 & 0,00917 \\ 0,00432 & 0,00536 & 0,00917 \\ 0,00432 & 0,00536 & 0,00917 \\ 0,00432 & 0,00536 & 0,00917 \\ 0,00432 & 0,00536 & 0,00917 \\ 0,00431 & 0,00536 & 0,00917 \\ 0,00431 & 0,00536 & 0,00536 \\ 0,00536 & 0,00536 & 0,00981 \\ 0,00431 & 0,00536 & 0,00981 \\ 0,00440 & 0,00536 & 0,00917 \\ 0,00471 & 0,00536 & 0,00536 \\ 0,00536 & 0,00536 & 0,00981 \\ 0,00471 & 0,00536 & 0,00536 \\ 0,00536 & 0,00536 & 0,00561 \\ 0,00746 & 0,00536 & 0,00561 \\ 0,00746 & 0,00531 & 0,00516 \\ 0,00746 & 0,00531 & 0,00536 \\ 0,00536 & 0,00536 & 0,00561 \\ 0,00746 & 0,00531 & 0,00536 \\ 0,00536 & 0,00550 & 0,00561 & 0,00561 \\ 0,00745 & 0,00578 & 0,005519 & 0,00516 \\ 0,00745 & 0,00551 & 0,00551 & 0,00551 \\ 0,00745 & 0,00551 & 0,00551 & 0,00551 \\ 0,00745 & 0,00551 & 0,00551 & 0,00551 \\ 0,00745 & 0,00551 & 0,00551 & 0,00551 \\ 0,00745 & 0,00551 & 0,00551 & 0,00551 \\ 0,00745 & 0,00551 & 0,00551 & 0,00551 \\ 0,00745 & 0,00551 & 0,00551 & 0,00551 \\ 0,00745 & 0,00551 & 0,00551 & 0,00551 \\ 0,00745 & 0,00551 & 0,00551 & 0,00551 \\ 0,00745 & 0,00550 & 0,00550 & 0,00550 & 0,00550 \\ 0,00745 & 0,00745 & 0,00550 & 0,00550 & 0,00550 \\ 0,00745 & 0,00745 & 0,00550 & 0,00550 & 0,00550 & 0,00550 & 0,00550 \\ 0,00745 & 0,00745 & 0,00057 & 0,00550 & 0,00550 & 0,00550 & 0,00550 & 0,00550 & 0,00550 & 0,00550 & 0,00$																																																																									
	Teex         0.1-0.2           30         0.00693         0           40         0.00554         0.00554           70         0.00554         0.00554           70         0.00554         0.00554           70         0.00554         0.00554           70         0.00554         0.00554           70         0.00554         0.00554           70         0.00554         0.00554           70         0.00554         0.00556           80         0.00553         0.00554           70         0.00554         0.00556           80         0.00556         0.00556           70         0.00557         0.00556           80         0.00557         0.00556           80         0.00557         0.00556           80         0.00557         0.00556           80         0.00557         0.00557           80         0.00557         0.00557           80         0.00557         0.00557           80         0.00557         0.00557	-	0.2-0.3 0.3-																																																																												• 597 (3)         • 0,00407           • 00155         • 0,00395           • 00155         • 0,00395           • 003383         • 0,00387           • 003383         • 0,00387           • 003383         • 0,00387           • 003383         • 0,00387           • 003383         • 0,00387           • 003383         • 0,00387           • 00463         • 0,00787           • 00463         • 0,00787           • 00463         • 0,00787           • 00463         • 0,00787           • 00463         • 0,00787           • 00463         • 0,00787           • 004643         • 0,00787           • 000267         • 0,0046           • 000294         • 0,0040           • 00221         • 0,0040	0,00267 0,00460 0,00460 0,00460 0,00321 0,00408 0,00408 0,00408 0,00465 0,00412 0,00412 0,00412 0,00412

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	0,111(-3 0,00893 0,00837 0,111(-3	_	0,01153 0,01153 0,0099	0,90367 0,40767 0,04891
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•	0,112 (		0,112(4 0,00126 0,00126 0,02957 0,02957 0,0299 0,02957 0,0299 0,02142 0,112(3	0,4173 0,18825 0,26534 0,0860 0,002429
·	0,18 (-5 0,23 (-3 0,00626 0,01747 0,02781 0,02781 0,02781 0,01582		0,38 (3 0,377 (3 0,00961 0,02479 0,03606 0,04171 0,03122 0,03122 0,00556	0,40342 0,10890 0,01956 0,38 (3
20°	0,32(3 0,00150 0,00150 0,00972 0,02073 0,3071 0,02685 0,02685 0,02497 0,02679	40°	0,02454 0,02050 0,02257 0,03257 0,033258 0,03156 0,03156 0,03156 0,01821 0,0078	00° 0, 15169 0, 06843 0, 06843 0, 05636 0, 05427 0, 803 (3 0, 803 (3
α Ξ	0,272 (	a=r	0,01336 0,01890 0,01890 0,01371 0,01371 0,01371 0,01371 0,01130 0,0056 0,0056 0,0056	a=5 0,07197 0,03247 0,01880 0,01880 0,01560 0,01366 0,01366 0,01366 0,01366 0,01366
	0,00331 0,00802 0,00663 0,00667 0,00667 0,00677 0,00883 0,00815 0,00815 0,00878 0,00878		0,0092 0,01257 0,01891 0,01290 0,01290 0,01033 0,00888 0,00888 0,00908	0,02162 0,00975 0,01879 0,01289 0,01289 0,01701 0,1771 (-3
	0,00456 0,00432 0,00527 0,00527 0,00501 0,00941 0,00710 0,00710 0,00710		0,00666 0,00856 0,01336 0,01145 0,00922 0,00915 0,00915 0,00916 0,00484	0,00742 0,00742 0,01015 0,01404 0,01886 0,006666 0,00608 0,00456 0,00456 0,00456
	0,00821 0,00537 0,00537 0,00439 0,00439 0,00465 0,00465 0,00408		0,00553 0,00553 0,00512 0,00556 0,00491 0,00454 0,00391 0,00453 0,00453	0,01710 0,00772 0,00455 0,00358 0,00553 0,00553 0,00553 0,00553 0,00553 0,00267 0,00267
	0,005022 0,00460 0,00525 0,00525 0,00615 0,00674 0,00568 0,00569 0,00554		0,00848 0,00870 0,00671 0,00512 0,00594 0,00534 0,00634 0,00614 0,00614	0,02453 0,001106 0,00125 0,00425 0,00425 0,00425 0,00625 0,00679 0,00679
	00000000000000000000000000000000000000		98096004400 9806004400 980000	88398299899900

.

Table 6.2

SPECTRAL-ANGULAR DISTRIBUTION OF CAPTURE  $\gamma$  RADIATION F<sub>c</sub>( $\Omega$ , B, E<sub> $\gamma$ </sub>) rad ster (rel. units)

	8-10		0,726(4 0,00225 0,00878		0,72(4 0,00392 0,00241																											
	68		0,113(3 0,00698 0,00633		0,11(3 0,00702																											
	4-6	į	0,151(3 0,0700		0.151(3 0.07128																											
	3-4		0,97 ( <u>-4</u> 0,0833 0,02634		0,97 (4 0,00866 0,0268																											
	23	•	0,0264 0,0264 0,0253		0,647(4 0,0355 0,00355 0,02597 0,02671 0,03482																											
4eV	1,5-2		0, 33(-4 0, 00205 0, 01205 0, 01201 0, 01195		0,33 (4 0,00232 0,0115 0,01911 0,01253 0,01566																											
ΔΕν, Ν	1-1.5	α=0°	0,00251 0,00486 0,004881 0,01881 0,01881 0,0163	a=10°	0,00319 0,00559 0,02077 0,02295 0,01621 0,01786																											
Energy Interval	0,75-1									0,00620 0,00787 0,00787 0,00897 0,006802 0,006802 0,00620		0,00139 0,00887 0,00887 0,00844 0,00844 0,00649 0,00649																				
	0.5-0.75																0,00196 0,00196 0,00488 0,00469 0,00463 0,00538 0,00473 0,00473		0,00425 0,00778 0,00762 0,00567 0,00567 0,00552 0,00552 0,00454													
	0,3-0.5																													0,00382 0,00502 0,00502 0,00398 0,00418 0,00389 0,00389 0,00324 0,00327 0,00302 0,00302		0,00325 0,00492 0,00377 0,00471 0,00471 0,00471 0,00413 0,00413 0,00292
	0.2-0.3																															
	0.1-0.2	•	0,00285 0,00285 0,00285 0,00385 0,00425 0,00482 0,00482 0,00429 0,00429		0,00605 0,00429 0,00385 0,00385 0,00471 0,00452 0,00452 0,00427																											
	degrees	-	000000000000000000000000000000000000000		000000000000000000000000000000000000000																											
					Fnerav	Interva	1 DE	MeV																								
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degrees	0,1-0,2	0,2-0,3	0.3-0.5	0.5-0.75	0,75-1	1-1,5	1,5-2	4 2-3	34	46	6-8	8-10																				
		-		-	-	a=26°	-																									
C	0 00426 1	0.00293	0.00840	0,01469	1		1	1	1		1	1																				
01	0,0040	0,00257	0,00576	0,01358	0,00196	0,00457	1	1	1	<u> </u>	1	<b>\</b>																				
20	0.00470	0,00368	0,00457	0,00827	0,01123	0,00682	0,33(-4	1		1		11																				
30	0,00411	0,00357	0,00633	0,00738	0,01029	0,02429	0,00255	0,65(-4]	.	 	 																					
40	0,00509	0,00525	0,00554	0,0077	0,01129	0,02570	0,01303	0,00407	0,9/ (4	15/_3	0 113/3	.73(4																				
50	0,00457	0,00402	0,00460	0,00618	0,00855	0,02601	0,02108	0,0264	0,0000	0 00100	0 00770	0 0026																				
8	0,00423	0,00363	0,00420	0,00541	0,00789	0,01967	0,01746	0,0384	0,0007	0,02120	0 00703	0,00929																				
70	0,00426	0,00348	0,00367	0,00494	0,00697	0,01698	0,01349	0,0283	0,0201	0 01996	0 00727	0.00243																				
80	0,00438	0,00367	0,00359	0,00501	0,00747	0,01842	0,01624	0,0036	0,00833	0 15 (-3	0.113(-3	0.726(4																				
06	0,00435	0,00407	0,00323	0,00493	0,00682	0,02257	0,01823	0,0240	,																							
	-					a==40°																										
0	1 0,0057	0,00535	0,00559	0,00824	0,01583	0,0604	0,00712	1	1	1		1 1																				
01	0,00645	0,00542	0,00623	0,00947	0,02666	0,0503	0,00513	0,65 (4		1		ł																				
20	0,00672	0,00664	0,00585	0,01299	0,02010	0,0431	0,0803	0,00059	0,01315	0_151 (-3	0.113 (3	0,73(-4																				
30	0,00504	0,00449	0,00546	0,01076	0,01181	0,03/8	0,0293	0.04954	0.03654	0,02738	0,00995	0,00343																				
40	0,00504	0,00467	0,00463	0,00891	0,01030	0,020,0	· · ·																									
	,							_	-	-	-																					

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				Energ	y Inte	rval <u>A</u> E	v, MeV					
<b>в.</b>	0,1-0.2	0,2-0,3	0,3-0.5	0,5-0,75	0,75—1	1-1.5	1.5-2	2—3	3-4	<b>4</b> 6	68	- 8-10
8	0,00363	0,00343	0,00530	0,00690	0,00851	0,0223	0,0176	0,03511	0,03528	0,09102	0,00861	0,01131
80	0,00404	0,00377	0,00497	0,0061	0,00821	0,0224	0,0197	0,04255	0,03184	0,02333	0,00893	0,00283
20	0,00458	0,00404	0,00461	0,00620	0,00857	0,0261	0,0212	0,02846	0,00931	l	1	١
ଛ	0,00542	0,00494	0,00471	0,00644	62600,0	0,0229	0,0115	0,00355	0,97 (4	I.	!	1
8	0,00482	0,00381	0,00440	0,00538	0,0076	0,0188	0,00205	0,65 (4	ļ	1	1	1
	-	-	-			a=90°						-
0	0.00106		0,63(-3	0,00339	0,00614	0,00656	0,0207	0,09038	0,2864	1,45203	0,6625	0,18309
10	0,555(-3	ļ	0,33(3	0,00177	0,00322	0,00344	0,01087	0,04737	0,15011	0,7611	0,03473	0,09597
20	0,001924	0,00146	0,00216	0,00460	0,00921	0,01778	0,03982	0,15933	0,15134	0,1309	0,04874	0,01568
30	0,002896	0,00253	0,00274	0,0053	0,00721	0,00543	0,0701	0,12113	0,04197	I	1	1
40	0,00699	0,00919	0,00476	0,01092	0,0213	0,0548	0,0427	0,01398	ł	1	!	1
50	0,00569	0,00935	0,00559	0,00824	0,0158	0,0604	0,00712	1	1	1	1	1
. 60	0,00562	0,00335	0,005997	0,01948	0,00473	0,0115	1	1	ļ	1		1
70	0,00426	0,00293	0,00840	0,01470	1	ļ	1	1	.1	1	1	1
8	0,00604	0,00608	0,00325	0,00425	I		1	٢	I	1	1	1
8	0,52(4	0,58(3	0,00882	۱	ļ	[	1	Ι.	I	1	1	1
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The results of calculation of the generalized characteristics of the angular distributions of  $\gamma$  radiation scattered from the earth are shown on Figure 6.3.

## § 2. Method of Design of Shielding for Surface Structures

The radiation of a nuclear explosion and radioactive fallout, scattered in the air, strikes the walls of structures rising over the surface of the ground or buried below the surface, penetrates through them and acts on the people inside the structures.

Let us study the general nature of this effect in greater detail using the example of a simple protective structure shown on Figure 6.4. If the integral dose in the open terrain of any type of nuclear explosion radiation is  $D_0$ , in correspondence with the form of the function

and spectral-angular distributions of the radiation in the air, the dose generated at the surface of the wall by the radiation arriving from direction  $\overline{\Omega}$  with energy E will be  $D_0F(\overline{\Omega}, E)$ . With anaangle of incidence

to the surface of the wall of  $\gamma$ , the total dose of this type of radiation at point A on the internal surface of the wall will be

$$D_{At} = D_0 \int_{2\pi} \int_E F(\Omega, E) K_t(\delta, \gamma, E) d\Omega dE, \qquad (6.9)$$

where  $K_t(\delta, \gamma, E)$  is the transmission factor for the total dose of the parallel beam of radiation striking the surface of the wall at angle  $\gamma$  with initial energy E, and integration is performed with respect to the entire spectrum of bombarding particles and all directions in the half space over the outer surface of the wall.

If we ignore boundary effects at points of contact between the wall and other elements of the structure due to the slight contribution of these effects to the total illuminated surface, we can consider that this value of total dose  $D_{At}$  will be characteristic for all points on the internal surface of the flat wall.

Dose  $\mathrm{D}_{\mathrm{At}}$  is formed by the scattered and unscattered radiation leaving the wall, so that

$$D_{At} = D_{As} + D_{Au}, \qquad (6.10)$$

where  $D_{As}$  and  $D_{Au}$  are the doses of scattered and unscattered components of radiation at point A.

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Figure 6.4. Analysis of Method of Calculating Shielding for Surface Structures.

The dose of the scattered component at point A on the internal surface of the wall can be calculated from the equation

$$D_{As} = D_0 \int_{2\pi} \int_E F(\Omega, E) K_s(\delta, \gamma, E) d\Omega dE, \qquad (6.11)$$

where  $K_{s}(\delta, \gamma, E)$  is the transmission factor of the dose of the scattered component of radiation, striking the surface of the wall at angle  $\gamma$  with initial energy E.

The dose of unscattered radiation at a detector located on the inner surface of the wall will be

$$D_{Au} = D_0 \int_{2\pi} \int_E F(\Omega, E) K_u(\delta, \gamma, E) d\Omega dE, \qquad (6.12)$$

where  $K_u(\delta, \gamma, E)$  is the dose transmission factor for unscattered radiation in a parallel beam, if

$$D_{As} = D_0 \Gamma(\delta); \tag{6.13}$$

$$D_{Au} = D_0 \Lambda (\delta), \qquad (6.14)$$

where  $\Gamma(\delta)$  and  $\Lambda(\delta)$  are the generalized transmission factors, called the barrier factors.

If the detector is moved from point A to point D, located at a certain distance from the inner surface of the wall, the dose at the detector is changed [2, 3]. The attenuation resulting from movement of the detector away from the wall will be referred to as geometric attenuation. Geometric attenuation can be looked upon as the effect of the finite extent of the source. Actually, the wall acts as a secondary source, and if this source were infinite in size, there would be no difference between positions of the detector A and D, since in both cases the detector would collect radiation from a  $2\pi$  space. Geometric attenuation occurs because the wall is finite in size and as the detector is moved from A to D, the solid angle in which radiation is collected is reduced.

Thus, the change in the dose at point D from the scattered component can be calcul**ate**d by introducing the so-called geometric radiation attenuation factor G

$$D_{J_s} = D_{As} G_{J_s}.$$

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If the angular distribution function of the dose of scattered radiation leaving the barrier is known, the geometric factor is determined by integrating the angular distribution function  $f(\omega)$  within limits of solid angle  $\omega_d$ , subtended by the wall in question or an

element of the wall at point D:

$$G_{ds} = \int_{\omega_d} f(\omega) \, d\omega \,. \tag{6.16}$$

In the simplest case, when the scattered radiation leaving the barrier is distributed isotropically, expression (6.16) is simplified:

$$G_{ds} = \frac{\omega_d}{2\pi} , \qquad (6.17)$$

and the dose of scattered radiation at point D can be calculated as

$$D_{ds} = D_{As} \frac{\omega_d}{2\pi} \,. \tag{6.18}$$

This expression shows that the dose of scattered radiation at the detector at some distance from the wall decreases in comparison with the dose of scattered radiation at the surface of the wall, in inverse proportion to the square of the distance to the wall.

The assumption of isotropy of radiation scattered in shields of a structure and radiating within it is basic to the method of calculation used. Although consideration of the angular distribution of radiation by the scattered component represents no difficulty in principle and equation (6.16) can be rather simply realized when the problem is solved by electronic computer, the lack in many cases of exhaustive information on angular distributions of radiation, particularly with inclined incidence to the surface of the barrier, hinders consideration of the angular characteristics. However, the assumption of isotropy of the scattered radiation should not lead to any significant errors.

As concerns the contribution of the unscattered component, its dose at point D is determined in correspondence with the expression

$$D_{du} = D_0 \int_{\omega_d} \int_E F(\Omega, E) K_u(\delta, \gamma, E) d\Omega dE, \qquad (6.19)$$

Integration in which is performed within limits of the solidaangle subtended by the wall or wall element being studied at the point of placement of the detector. Physically, this corresponds to the fact that the particles of the unscattered component do not change the direction of their initial movement upon penetration through the barrier and therefore that portion of the radiation of the nuclear explosion which strikes the surface of the wall, moving in directions passing by point d, cannot act on the detector located at this point. However, since the particles of the unscattered component do not interact with the materials of the wall, the wall cannot be looked upon as a secondary source. Therefore, in contrast to the scattered radiation, the unscattered radiation does not undergo geometric attenuation upon movement from the surface of the wall to the detector. Actually, the sources of those sparticles which strike the detector without interacting with the material of the wall are the centers of the last scattering events in the atmosphere. The degree of geometric attenuation of the dose of unscattered radiation propagating from the wall to the detector depends on the relationship between distances from the scattering center to the wall and from the wall to the detector.

If a particle has undergone scattering in the direction of the detector at distance R from the wall and, without interacting with the material of the wall, strikes the detector, located at distance L from it, the factor of geometric attenuation is  $R^2/(R + L)^2$ . Let us estimate the value which may be reached by this factor under actual conditions. Let us assume for simplicity that the density of scattering centers in the air surrounding the structure is identical throughout and **that** the

number of particles scatteredpper unit volume about point M is N. In an element volume dv, this will amount to

$$Ndv = NR^2 d\omega dR.$$

Their flux at point B per unit solid angle is

$$\frac{NR^2 d\omega dR}{4\pi R^2 d\omega} e^{-R/\lambda} = \frac{N e^{-R/\lambda} dR}{4\pi}.$$

Integrating with respect to ray BM, we produce

$$\Phi_B = \frac{N}{4\pi} \int_0^\infty e^{-\frac{R}{\lambda} dR} = \frac{N\lambda}{4\pi}.$$
 (6.20)

At point D this amounts to

$$\Phi_{\mathbf{d}} = \frac{N}{4\pi} \int_{0}^{\infty} \frac{R^2 e^{-R/\lambda}}{(R+L)^2} dR.$$
(6.21)

The geometric factor determining the attenuation of the unscattered component upon propagation from the wall to a detector located at distance L from it is

$$G_{\mathbf{ch}} = \frac{\Phi_{\mathbf{d}}}{\Phi_{B}} = \frac{1}{\lambda} \int_{0}^{\infty} \frac{R^{2} \mathrm{e}^{-R/\lambda}}{(R+L)^{2}} dR.$$
(6.22)

Calculations performed using this equation show that the geometric factor is 0.994 for a distance L = 1 m from the wall; 0.986 for L = 2 m; 0.958 for L = 5 m; 0.922 for L = 10 m.

Thus, for practical purposes in the design of shielding of surface structures, geometric attenuation of the dose of the unscattered component can be ignored. The total dose at point D within the structure is determined by adding the contribution from the individual walls and the structural elements

$$D_{cl} = \Sigma \left( D_{ds} + D_{du} \right). \tag{6.23}$$

In some cases in order to estimate the protective properties of structures, we can use a simplified calculation method. According to this method, the contribution of any given structural element to the total dose at point D is calculated using the formula

$$\hat{D}_{d} = D_{A}G_{d,s} = (D_{As} + D_{Au})G_{d,s}.$$
(6.24)

It is not difficult to see that in expression (6.24), the unscattered radiation component reaching the internal surface of the wall is given the same geometric attenuation as the scattered component. This causes a reduction in the dose at point D in comparison to the actual dose by

$$\Delta D_{d} = D_{du} - D_{Au} G_{ds} = D_{0} \left[ \int_{\omega_{d}E} F(\overline{\Omega}, E) K_{u}(\delta, \gamma, E) d\Omega dE - G_{ds} \int_{2\pi} \int_{E} F(\overline{\Omega}, E) K_{u}(\delta, \gamma, E) d\Omega dE. \right]$$

$$(6.25)$$

Thus, the results of calculations performed using the two methods outlined above are interrelated by the relationship

$$D_{d} = \widetilde{D}_{d} - \Delta D_{d}. \tag{6.26}$$

In many cases, however, the value of correction  $\Delta D_d$  is insignificant and expression (6.24) can be used for approximateldesign of shielding.

§ 3. Calculation of Barrier and Geometric Factors in Attenuation of Radiation

The value of the barrier factor depends on the material and thickness of the shield  $\delta$ , its position in space and the spectral angular distribution of the radiation scattered in the air.

If the position of the wall of a structure is fixed by angle  $\theta$  (see Figure 6.5), calculated between the normal to its outer surface and the normal to the earth (top downward), the barrier factor of transmission of the total neutron dose is calculated using the expression

$$\Lambda_{n, t}(\delta, \vartheta, \alpha) = 2 \int_{0}^{\pi/2} \int_{0}^{\pi} \sin \gamma K_{nt}(\delta, \gamma) F_n(\Omega, \beta) \, d\gamma d\tau, \qquad (6.27)$$

where  $\beta = \arccos(\sin \gamma \cos \tau \sin \vartheta + \cos \gamma \cos \vartheta)$ .

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The barrier transmission factor for the unscattered component of neutron dose is defined as

$$\Lambda_{nu}(\delta, \vartheta, \alpha) = 2 \int_{0}^{\pi/2\pi} \int_{0}^{\pi} \sin \gamma K_{nu}(\delta, \gamma) F_n(\Omega, \beta) d\gamma d\tau. \qquad (6.28)$$

Then the barrier transmission factor for the scattered component of neutrons is

$$\Lambda_{ns}(\delta, \theta, \alpha) = \Lambda_{nt}(\delta, \theta, \alpha) - \Lambda_{nu}(\delta, \theta).$$
(6.29)

The results of calculation of the barrier transmission factors for the total neutron dose and the scattered component of neutrons are presented in Appendix 1 (Table a. 1 and Table a. 2.).



Figure 6.5. Geometry of General Design of Shielding of a Structure

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The barrier transmission factor for the total dose of fragment  $\boldsymbol{\gamma}$  radiation is calculated using the equation

$$\Lambda_{\gamma t}^{f}(\delta, \vartheta) = 2 \sum_{E_{\gamma}} \int_{0}^{\pi/2} \int_{0}^{\pi} \sin \gamma F_{f}(\Omega, \beta, E_{\gamma}) K_{\gamma t}^{f}(\delta, \gamma, E_{\gamma}) d\gamma d\tau.$$
(6.30)

Addition is performed over the entire spectrum of  $\boldsymbol{\gamma}$  quanta scattered in the air.

Similarly, the barrier transmission factor of the unscattered component of  $\gamma$  quanta is calculated by

$$\Lambda_{\gamma\mu}^{f}(\delta, \vartheta) = 2 \sum_{E_{\gamma}} \int_{0}^{\pi/2} \int_{0}^{\pi} \sin \gamma F_{c}(\Omega, \beta, E_{\gamma}) e^{-\mu(E_{\gamma})} \frac{\delta}{\cos \gamma} d\gamma d\tau.$$
(6.31)

The barrier transmission factor for the scattered component is

$$\Lambda_{\gamma s}^{f}(t, \vartheta) = \Lambda_{\gamma t}^{f}(\delta, \vartheta) - \Lambda_{\gamma t}^{f}(\delta, \vartheta).$$
(6.32)

The results of calculation of the barrier transmission factors for the total dose of  $\gamma$  quanta and the dose of scattered fragment  $\gamma$  radiation are presented in Appendix 2 (Tables a. 3 and a. 4).

Using expressions similar to (6.30) and (6.31), the barrier transmission factors for the total dose of capture  $\gamma$  radiation in the air  $\Lambda_{\gamma t}^{c}(\delta, \theta, \alpha) \gamma$  radiation from radioactive fallout  $\Lambda_{\gamma t}(\delta, \theta)$  and the corresponding factors for the unscattered component  $\Lambda_{\gamma u}^{c}(\delta, \theta, \alpha)$  and  $\Lambda_{\gamma u}^{r}(\delta, \theta, t)$  were calculated. The factors  $\Lambda_{\gamma t}^{c}(\delta, \theta, \alpha)$  and  $\Lambda_{\gamma s}^{c}(\delta, \theta, \alpha)$ are presented in Appendix 3 (Tables a. 5 and a. 6), while  $\Lambda_{\gamma t}^{r}(\delta, \theta, t)$ and  $\Lambda_{\gamma s}^{r}(\delta, \theta)$  are presented in Appendix 4 (Tables a. 7. and a. 8).

The barrier transmission factors for capture  $\gamma$  radiation formed in the materials of the shielding structure were determined using the expression

$$\Lambda_{\gamma}^{s}(\delta, \vartheta, \alpha) = 2K_{cf}'(\delta) \int_{0}^{\pi/2} \int_{0}^{\pi} \sin \gamma \cos \gamma F(\beta) \, d\gamma d\tau =$$
  
=  $\Lambda_{\gamma i}^{s}(\delta) \Lambda_{\gamma i}^{s}(\vartheta, \alpha).$  (6.33)

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Factors  $\Lambda_{\gamma i}^{3}(\delta)$  and  $\Lambda_{\gamma j}^{3}(\theta, \alpha)$  are presented in Table 5 (Tables a. 9 and a. 10).

As concerns calculation of the geometric attenution factors of radiation within the shielding structure, with the assumption made of isotropy of the distribution of the scattered radiation leaving the barrier it is reduced, as follows from expression (6.17), to calculation of the solid angles subtended by the structural elements in question at the point of observation.





In many cases of practical importance, the elements of the structure are rectangular (walls, ceilings) or can be represented as a sum of rectangular elements. Calculation of the solid angles in this case is rather simple [4].

If we use the geometry of Figure 6.6, on which the coordinate origin is selected in the plane of the rectangular element in question beneath the location of the detector, the distance from the detector to the plane of the rectangle is represented by  $z_c$ , and the coordinates of the four angles of the rectangle are represented by  $x_i$ ,  $y_i$  (i = 1, 2, 3, 4), then if the coordinate system is selected so that its axes are parallel to the sides of the rectangle and the numeration of the angles is performed clockwise, the solid angle subtended by the rectangle in question at point D will be

 $\omega_{d} = |q_{1}'\tau_{0}(\varepsilon_{1}, \eta_{1}) - q_{2}'\tau_{0}(\varepsilon_{2}, \eta_{2}) + q_{3}'\tau_{0}(\varepsilon_{3}, \eta_{3}) - - q_{4}'\tau_{0}(\varepsilon_{4}, \eta_{4})| \text{ ster,}$ (6.34)

$$\tau_0(\varepsilon_i, \eta_i) = \operatorname{arctg} \frac{\varepsilon_i}{\tau_{il} \sqrt{\varepsilon_i^2 + \tau_i^2 + 1}}; \qquad (6.35)$$

$$\varepsilon_i = \left| \frac{y_i}{x_i} \right|; \quad \eta_i = \left| \frac{z_c}{x_i} \right|; \quad q'_i = \frac{y_i}{\varepsilon_i x_i}.$$
 (6.36)

In more complex cases, when the form of the structure cannot be represented by flat rectangular elements, the method described below must be used to calculate the solid angles.

If the protective shield has a complex shape, it can be represented as a set of triangular elements of arbitrary form. The solid angle subtended at any point by a triangular element oriented arbitrarily in space is relatively easy to calculate.



Figure 6.7. Calculation of Solid Angle of Element of Triangular Shape

Suppose O M<sub>1</sub> M<sub>2</sub> M<sub>3</sub> is a triangular solid angle fixed in coordinate system Oxyz by points M<sub>i</sub> (x<sub>i</sub>, y<sub>i</sub>, z<sub>i</sub>) of each side and point O (x<sub>0</sub>, y<sub>0</sub>, z<sub>0</sub>) (Figure 6.7).

If we construct a sphere of unit radius with its center at point 0, it is not difficult to see that the value of quantity  $\omega$ , the solid angle, is exactly equal to the value of the area of the surface of the spherical triangle formed by the trace of each edge on the surface of the sphere:

$$S_{\Delta M_1'M_2'M_3'} = \omega = (\alpha' + \beta' + \gamma' - \pi), \qquad (6.37)$$

where

where  $\alpha'$ ,  $\beta'$  and  $\gamma'$  are the angles at point  $\Delta M'_1 M'_2 M'_3$ , equal to:

$$\alpha' = 2 \arcsin \sqrt{\frac{\sin (n-b) \sin (h-c)}{\sin b \sin c}}; \qquad (6.38)$$

$$\beta' = 2 \arcsin \sqrt{\frac{\sin (n-a)\sin (h-c)}{\sin a \sin c}}; \qquad (6.39)$$

$$\gamma' = 2 \arcsin \sqrt{\frac{\sin(n-a)\sin(h-b)}{\sin a \sin b}}; \qquad (6.40)$$

The values of a, b and c determine the lengths of the sides of this triangle and are calculated according to the relationships of the scalar product of the vectors  $M_1$ ,  $M_2$  and  $M_3$ , constructed on the faces of the solid angle:

$$a = \arccos \frac{(M_2 M_3)}{\|M_2\| \|M_3\|}; \qquad (6.41)$$

$$b = \arccos \frac{(M_1 M_3)}{\|M_1\| \|M_2\|}; \qquad (6.42)$$

$$c = \arccos \frac{(M_1 M_2)}{\|M_1\| \|M_2\|}.$$
 (6.43)

$$2h = a + b + c.$$
 (6.44)

§ 4. Sequence of Design of Shielding From the Radiation of Nuclear Explosions

The calculation of the shielding ability of a surface structure<sup>1</sup> against the radiation of nuclear explosions is reduced to arbitrary subdivision of the surfaces of the structure into individual elements, determination of the position of these elements in space, determination of the barrier transmission factors for these elements and the geometric attenuation factors relative to a selected design position within the structure, with subsequent addition of radiation doses created at the detector by the individual structural elements.

<sup>1</sup>The calculation plan for subterrainean structures differs only in that there is only one surface in this case.

Let us first study a simple method of estimating the shielding of a surface structure, the use of which does not require computer equipment. This method is recommended to be used when the surface of the structure to be calculated can be represented as a set of a small number of flat rectangular shields. The edges of these elements are selected so that the thickness of the shield of homogeneous material within each is identical.

The position of each of the rectangular shields in this type of structure is determined by the value of angle  $\theta$  between the normal to the surface of the earth and the external normal of the element (see Figure 6.5).

In correspondence with this angle and the known shield thickness  $\delta$ , the appendices of this book can be used to find the barrier transmission factors for the total dose of neutrons and  $\gamma$  radiation from a nuclear explosion.

After this, the geometric attenuation factor is calculated for each element of the structure and the selected design points within the structure. The geometric factor is calculated using expression (6.17). The solid angle subtended by the rectangular element at the detection point is determined using equation (6.34).

The contribution of each element to the total transmission factor for the structure at the design point is calculated using the expression

$$\Pi_t = \Lambda_t G_{d, s}. \tag{6.45}$$

The total transmission factor of the structure for the dose of each type of radiation is determined by adding the values for each structural element. In order to produce the transmission factor for all types of radiation existing in a nuclear explosion, the total transmission factors for the individual types of radiation must be added in correspondence to their relative share of the total dose of nuclear radiation on the open terrain.

If a shielding structure has a complex shape and complex shield structure, analysis of the shielding properties may require dividing the entire surface into a large number of elements. In particular, this necessity arises when the form of the structure is curved. In this case, the dimensions of the elements must be selected so that the thickness of the shield does not change significantly within the limits of an element, and the surface can be considered flat. The mathematical operations involved in calculating this type of shielding are simple to perform if the elements are triangles of arbitrary shape. The position of each triangular element can be fixed by the coordinates of its points on the external surface  $M_1(x_1, y_1, z_1)$ ,  $M_2(x_2, y_2, z_2)$ ,  $M_3(x_3, y_3, z_3)$ , calculated in the main system of coordinates XYZ, coupled to the structure being calculated. This main system of coordinates will be selected in the remainder of this book such that the positive direction of the Z axis forms a perpendicular to the surface of the earth. The thickness of the shielding can be fixed by the mean thickness  $\delta$  of a shield within the limits of each of the geometric elements selected, measured in the normal cross section.

Thus, the <u>first stage</u> in calculation consists of preparation of a system of initial data concerning the geometry of the shielding structure, which includes the set of coordinates of points selected on the external surface of the structure and the thickness of the shielding materials at these points. Preparation of this type of initial information for a structure of any configuration represents no great difficulties.

The second stage in calculation consists of determination for each element of the angle  $\theta$  readabetween the normal to the element and the normal to the surface of the earth.

The normal to the element will be considered the unit vector n, perpendicular to its external surface (Figure 6.7). We note also that in the system of coordinates which we are using, the normal to the surface of the earth corresponds in its direction to the axis OZ. Thus, angle  $\theta$  between the normal to the element and the normal to the surface of the earth is the angle between the normal to the element and the element and the OZ axis.

The vector of the normal to the element is found as the vector product of two noncolinear vectors, lying in the plane of the surface of the element. These vectors can be the vectors  $M_1 M_2$  and  $M_1 M_3$ , constructededn the sides of the triangle defining the element (see Figure 6.5).

The vectors  $M_1M_2$  and  $M_1M_3$  can be defined as follows:

$$\mathbf{M}_{1}\mathbf{M}_{2} = \{ (x_{2} - x_{1}); \quad (y_{2} - y_{1}); \quad (z_{2} - z_{1}) \}; \\ \mathbf{M}_{1}\mathbf{M}_{3} = \{ (x_{3} - x_{1}); \quad (y_{3} - y_{1}); \quad (z_{3} - z_{1}) \}.$$
(6.46)

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The resulting vector product will be the vector N:

$$\mathbf{N} = [\mathbf{M}_{1}\mathbf{M}_{2}; \ \mathbf{M}_{1}\mathbf{M}_{3}] = \begin{vmatrix} \mathbf{i} & \mathbf{j} & \mathbf{k} \\ (x_{2} - x_{1}) & (y_{2} - y_{1}) & (z_{2} - z_{1}) \\ (x_{3} - x_{1}) & (y_{3} - y_{1}) & (z_{3} - z_{1}) \end{vmatrix} = \\ = \begin{vmatrix} (y_{2} - y_{1}) (z_{3} - z_{1}) - (y_{3} - y_{1}) (z_{2} - z_{1}) \\ (x_{2} - x_{1}) (z_{3} - z_{1}) - (x_{3} - x_{1}) (z_{2} - z_{1}) \\ (x_{2} - x_{1}) (y_{3} - y_{1}) - (y_{2} - y_{1}) (x_{3} - x_{1}) \end{vmatrix} = \begin{vmatrix} x_{N} \\ y_{N} \\ z_{N} \end{vmatrix}$$
(6.47)

In order that the direction of vector N might correspond to the direction of the vector of the normal to the element n, the numeration of points  $M_1$ ,  $M_2$  and  $M_3$  should be selected counterclockwise as seen from the end of vector n.

Thus,

$$\vartheta = \arccos \frac{z_N}{\sqrt{z_N^2 + y_N^2 + z_N^2}} \,. \tag{6.48}$$

The <u>third stage</u> in the calculation consists of calculating the doses of the scattered radiation component of each type from the structural element. In correspondence with the known thickness of the shielding material  $\delta$  of the element and the angle  $\theta$  calculated for it, the barrier transmission factors of the doses of the scattered component of neutrons and  $\gamma$  radiation from the nuclear explosion are determined from the appendices.

Then, calculation of the geometric attenuation factor is performed for each structural element and the selected design points within the structure, fixed by the coordinates D,  $x_d$ ,  $y_d$ ,  $z_d$ ). The geometric factor is calculated using expression (6.17). The solid angle subtended by the triangular element  $M_1M_2M_3$  at the point of detection D is determined in correspondence with equation (6.37). We should note only that the vectors  $M_1$ ,  $M_2$  and  $M_3$  used in calculation of the solid angle are the vectors, the origins of which are located at point D ( $x_d$ ,  $y_d$ ,  $z_d$ ), the ends at points  $M_1(x_1, y_1, z_1)$ ,  $M_2(x_2, y_2, z_2)$  and  $M_3(x_3, y_3, z_3)$ respectively, i. e.

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$$\mathbf{M}_{1} = \{ (x_{1} - x_{c}); \quad (y_{1} - y_{c}); \quad (z_{1} - z_{c}) \}; 
\mathbf{M}_{2} = \{ (x_{2} - x_{c}); \quad (y_{2} - y_{c}); \quad (z_{2} - z_{c}) \}; 
\mathbf{M}_{3} = \{ (x_{3} - x_{c}); \quad (y_{3} - y_{c}); \quad (z_{3} - z_{c}) \}.$$
(6.49)

The contribution of each element to the total transmission factor of the scattered component at the point in question is calculated using the expression

$$\Pi_{s} = \Lambda_{s} G_{ds}.$$
 (5.50)

The total transmission factor of the structure for the scattered component of the dose of each type of radiation is determined by adding the values of  $\Pi_{c}$  for all structural elements.

The fourth stage is the calculation of angle  $\beta$  between the lines connecting the point to the center of the element and the normal to the surface of the earth.

The coordinates  $x_B^{}$ ,  $y_B^{}$ ,  $z_B^{}$  of the center of gravity of the triangular element, fixed by the coordinates of points  $M_1^{}(x_1^{}, y_1^{}, z_1^{})$ ,  $M_2^{}(x_2^{}, y_2^{}, z_2^{})$  and  $M_3^{}(x_3^{}, y_3^{}, z_3^{})$  are calculated using the equations

$$x_{B} = \frac{x_{23} - x_{1}}{y_{23} - y_{1}} y_{B} + \frac{x_{1}y_{23} - y_{1}x_{23}}{y_{23} - y_{1}};$$
  

$$y_{B} = \frac{(y_{12} - y_{3})(x_{1}y_{23} - y_{1}x_{23}) + (y_{23} - y_{1})(y_{3}x_{12} - x_{1}y_{12})}{(x_{12} - x_{3})(y_{23} - y_{1}) - (y_{12} - y_{3})(x_{23} - x_{1})};$$
  

$$z_{B} = z_{1} + \frac{x_{B} - x_{1}}{x_{23} - x_{1}}(z_{23} - z_{1}),$$
  
(6.51)

where

$$\begin{aligned} x_{12} &= \frac{x_2 - x_1}{2} + x_1; \quad y_{12} &= \frac{y_2 - y_1}{2} + y_1; \quad z_{12} &= \frac{z_2 - z_1}{2} + z_1; \\ x_{23} &= \frac{x_3 - x_2}{2} + x_2; \quad y_{23} &= \frac{y_3 - y_2}{2} + y_2; \quad z_{23} &= \frac{z_3 - z_2}{2} + z_2. \end{aligned}$$

The unit vector corresponding in direction to the direction from point  $D(x_d, y_d, z_d)$  to the center of element  $B(x_B, y_B, z_B)$  is the vector

$$\mathbf{B} = \{ (x_B - x_d); (y_B - y_d); (z_B - z_d).$$
(6.52)

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Therefore, the desired angle  $\beta$  is

$$\beta = \arccos \frac{z_B - z_1}{\sqrt{(x_B - x_1)^2 + (y_B - y_1)^2 + (z_B - z_A)^2}}$$
(6.53)

The <u>fifth stage</u> in the calculation consists of calculation of angle  $\psi$  between the line connecting the point to the center of the element and the normal to the element.

Angle  $\psi$  is defined as the angle between vectors B and n and is calculated using the formula

$$\psi = \arccos \frac{(nB)}{\|n\| \| B\|} = \arccos \frac{(nB)}{\|N\| \| B\|}.$$
 (6.54)

The coordinates of vectors B and n are calculated using formulas (6.52) and (6.47) respectively.

Thus,

$$\dot{\Psi} = \frac{x_N (x_B - x_A) + y_N (y_B - y_A) + z_N (z_B - z_A)}{\sqrt{(x_B - x_A)^2 + (y_B - y_A)^2 + (z_B - z_A)^2} \sqrt{z_N^2 + y_N^2 + z_N^2}}$$
(6.55)

During the <u>sixth stage</u>, the dose of the unscattered component of neutrons at the point of detection is calculated. In correspondence with the angle  $\beta$  calculated for the element, using the data of § 1 of Chapter 6, the values of the generalized angular distribution functions of radiation scattered in the air or reflected from the earth  $F(\Omega, \beta)$  are found. The contribution of each element to the total transmission factor of the unscattered component of each type of radiation at the point in question is

 $\Pi_{\mu}(\delta, \beta, \psi) = F(\Omega, \beta) \omega_{\mu} K_{\mu}(\delta, \psi), \qquad (6.56)$ 

where  $K_u(\delta, \psi)$  is the transmission factor for the dose of the unscattered component of radiation by a shield of thickness  $\delta$  with a beam incident angle to its surface  $\gamma = \psi$ .

In the case of  $\gamma$  radiation, addition is performed for the entire spectrum of energies:

$$\Pi_{\gamma, u}(\delta, \beta, \psi) = \sum_{E_{\gamma}} F(\Omega, \beta, E_{\gamma}) \omega K_{\gamma u}(\delta, \psi, E_{\gamma}).$$
(6.57)

The total transmission factor of the structure for the unscattered component of each type of radiation is determined by addition of the values of  $\Pi_{u}^{\cdot}$  for all structural elements. The transmission factor of the structure for the total dose of radiation of each type is equal to the sum of the transmission factors for the scattered and unscattered components:

$$\Pi_t = \Pi_s + \Pi_u. \tag{6.58}$$

In order to produce the transmission factor for all types of radiation acting during a nuclear explosion, the total transmission factors for the individual types of radiation must be added in correspondence to their relative shares in the total dose of nuclear radiation on the open terrain.

This method of calculation considers the presence in the structure of windows and open doorways. For these elements with 0 shielding thickness, the transmission factor of the scattered component of radiation is zero, the transmission factor for the unscattered component is one. Therefore, using the method which we have outlined it is possible to calculate the radiation situation at windows and openings.

§ 5. Estimate of Radiation Dose Repeatedly Scattered Within a Structure

The radiation from a nuclear explosion which passes through the shielding may undergo repeated reflection from the walls of the structure, thus increasing the dose recorded at any point within the structure. An estimation of the degree of increase of the dose within a structure due to reflection can be performed using the example of a room of spherical shape.

If for simplicity we assume that the angular distribution of radiation leaving the walls of the spherical room is isotropic, and the total transmission factor for the walls of the structure is identical everywhere and equal to  $K_t$ , the transmission factor for radiation for the entire structure for any point within the room is  $\Pi_+ = 2K_+$ .

Due to the symmetry of the problem, the angular distribution of radiation striking the surface of the wall at any point on the wall is isotropic. If the albedo of radiation from the material of the walls of

the room is  $A_d^{\pi}(\gamma)$ , where  $\gamma$  is the angle of incidence of the radiation to the surface of the wall, the dose of singly reflected radiation at the point of observation is

$$D' = K_t \int_0^{\pi/2} \sin \gamma A_d^{\pi}(\gamma) d\gamma. \qquad (6.59)$$

For neutrons in the case of concrete or iron walls

$$D'_{n} = K_{t} A^{\pi}_{n\alpha} (\gamma = 0) \int_{0}^{\pi/2} \sin \gamma \cos^{2/3} \gamma d\gamma =$$
  
= 0.6K\_{t} A^{\pi}\_{n\alpha} (\gamma = 0), (6.60)

where  $A_{nd}^{\pi}(\gamma = 0)$  is the albedo with normal incidence on the surface of the wall, about 0.9 for iron and 0.6 for concrete.

Since as been shown in Chapter 4 the distribution of reflected neutrons is isotropic, and their spectrum is slightly deformed in comparison to the spectrum of the incident neutrons to the barrier, the geometric conditions of repeated reflection will correspond precisely with the conditions of first reflection. Therefore, the dose recorded due to multiple reflection at any point M within the structure will be

$$D_n^{M} = K_t \frac{0.6A_{nd}^{n}(\gamma = 0)}{1 - 0.6A_{nd}^{n}(\gamma = 0)}$$
 (6.61)

In the case of a concrete structure,  $D_n^m = 0.563 K_t$ , for iron walls  $D_n^m = 1.175 \cdot K_t$ . Recalling that the transmission factor for radiation of the structure for any point within the structure is  $\Pi = 2 K_t$ , we can conclude that in order to consider the contribution of multiply reflected neutrons within a concrete structure the total transmission factor for neutrons  $\Pi_t$  should be multiplied by 1.28. For structures with iron shielding, this correcting factor is about 1.585. In other words, the increase in the dose due to multiple reflection within a concrete structure about 58.5%. For polyethylene (water), this contribution is about 8-10%.

In the case of  $\gamma$  radiation, the yield of singly reflected radiation D' is D = A<sup> $\pi$ </sup> K, i. e. 0.073 K for water, 0.069 K for concrete

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and 0.066  $K_{\rm t}$  for iron based on the mean  $\gamma$  radiation energy [3].

It was established in Chapter 5, with the exception of cases of incidence of  $\gamma$  radiation at very high inclination, the cosinusoidal dependence is very satisfactory as an approximation of the angular distribution of radiation scattered back from a barrier. As concerns the spectral distribution of reflected  $\gamma$  radiation, most  $\gamma$  quanta undergoing single or multiple reflection have energies in the intervals 0.1-0.5 MeV (see Figure 5.3). We can see from Figure 5.2 that in this energy interval the albedo of the  $\gamma$  radiation does not change significantly at all and can be assumed equal to 0.065 for iron, 0.160 for concrete and 0.25 for water with normal incidence.

The contribution of the second reflection of  $\boldsymbol{\gamma}$  quanta from the walls is

$$D'_{\gamma} = 2\pi D'_{\gamma} \int_{0}^{\pi/2} A^{\pi}_{\gamma d} \sin \gamma \cos \gamma d\gamma = D'_{\gamma} A^{\pi}_{\gamma d}, \qquad (6.62)$$

 $A_{\gamma d}^{\pi}$  is the albedo of  $\gamma$  radiation from the wall with an energy of 0.1-0.5 MeV and normal incidence.

The dose recorded due to multiple reflection at any point within the structure will be

$$D_{\gamma}^{M} = D_{\gamma}' + D_{\gamma}' \frac{A_{\gamma d}^{n}}{1 - A_{\gamma d}^{n}}.$$
 (6.63)

For concrete, this expression is 0.082, for iron 0.0707 and for water 0.0985. Thus, in order to consider the contribution of multiply reflected  $\gamma$  quanta from the walls within a concrete structure, the total transmission factor for  $\gamma$  quanta should be multiplied by about 1.041, within an iron structure -- by about 1.0304 and within a structure having a shield of water or polyethylene -- about 1.050. In other words, the contribution of multiply reflected  $\gamma$  quanta does not exceed 5% and can generally be ignored.

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APPENDIX 1

Table a. 1

## BARRIER TRANSMISSION FACTORS OF TOTAL DOSE OF NEUTRONS FROM NUCLEAR EXPLOSION $\Lambda_{nt}^{}(\delta,\,\theta,\,\alpha)$ for various materials

	180		0253 0235 0235 0278 0278 0278 0313 0341 0363 0363		0173 0178 0183 0194 0194 0211 0225 0245 0245 0261 0261
	160		0322 0360 0378 0378 0378 0319 0318 0318 0314 0314 0374		0183 0183 0186 0196 0204 0232 0232 0252 02596
	140		0432 0448 0448 0403 0312 0335 03352 03352 03352 03352 03352		0241 0233 0233 0223 0223 0227 0227 0227 0227
ies -	120	-4)	0666 0627 0597 0456 0411 0411 0411 0441 0441 0441	-4	0320 0320 0283 0271 0272 0279 0279 0279
₀, degre	100	16 см (×10 <sup>-</sup>	0781 0800 0688 0619 0455 0455 0457 0447	-20 cm (×10 <sup>-</sup>	0413 0400 0363 0363 0315 0318 0315 0313 0313 0313
	80	Lon .	0989 0911 0857 0677 0577 0587 0587 0547 0547 0547 0538	Iron , 🚈	0529 0529 0454 0454 0354 0378 0378 0356 0356
	60		1084 0958 0958 0792 0792 0701 0678 0678 0659 0653 0653 0623	-	0553 0557 0555 0555 0555 0555 0555 0555
	40	-	1080 0956 0904 0831 0809 0811 0804 0811 0804 0738 0738	-	0599 0599 0599 0571 0571 0571 0554 0554 0551 0514 0514
	20	-	1038 0925 0876 0876 0876 0876 0880 0880 0880 0868 0845 0845 0830	_	0617 0615 0615 0615 0613 0613 0613 0612 0612 0586 0586
	0		0941 0941 0883 0879 0885 0897 0917 0931 0873 0873	-	0624 0624 0622 0620 0616 0630 0646 0657 0657 0657 0619
	Degrees	-	00000000000000000000000000000000000000		88282220 88282220

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Table a. l. continued

	. 180		0292 0296 0298	0310 0318 0345	0367	0473		0321 0336 0349	0363	0403	0429	0498 0553	_		0761 0793 0825 0859			
	160		0272 0287 0303	0316 0330 0351	0373	0434 0479		0337 0345 0352	0367	0407	0434 0473	0504 0558		,	0793 0824 0855 0889			
4	140		0322 0334 0341	0345 0350	0387	0410 0443 0486		0350 0367 0380	0330	0423	0448 0486	0517 0570	_		1150 1125 1091 1062			
s	120	(1-0)	0471 0447 0428	0416 0406	0419	0440 0458 0498	()()	0499 0481 0479	0463	0468	0482 0512	0536 0586		_	1543 1468 1396 1327			
9. degree	100	3=16 см (X1	0724 0671 0625	0579 0540	0501	0503 0525	=29 см (×1(	0854 0789 0796	0668	0591	0577 0579	0583 0613	=5 cw (×104)		1981 1864 1759	1		
	. 08	hylene, õ	0978 0910 0838	0772	0601 0601	0547 0542 0542	lene, ô	0283	2160	0747	0694	0618 0618		ורו ברביי	2439 2285 2150 2020	2424		
	. 09	Polyet	2131 0142 0985	0928 0876	1220	0682 0682 0652	olyethy	1352 1275 1190	1116	1049	0907	0787 0748	<u>·</u> (	5	2651 2541 2430 2398	0707		
	40				1027 1020 1012	1008 0991 2025	c/60	0900 0900 0862	·č	1200	1202	1168	1161	1080	<b></b> ,	•	2749 2698 2644 7507	7007
	20		0921 0954 0983	1010	1123	11126 11126 1118	-	1019 1080 1138	1182	1309	1367	1391			2870 2831 2799 9767	2101		
	0		0904 0935 0967	0998	1114	1113 1113	-	0984 1042	1161	1242 1341	1459	1476			2960 2897 2845	2133		
 	degrees		2000	893	201	288	-	000	28	5 <del>0</del>	88	288			3000 2000	2		

Table a. l. continued

									•	
						9, degre	ses			
degrees	0	20	40	60	80	100	120 .	140	160	180
40	2767	2731	2524	2233	1890	1564	1259	1049	0924	0877
00	2762	2702	2467	2143	1786	1501	1233	1068	0988	0955
60	2783	2682	2440	2080	1723	1456	1211	1011	1052	1010
20	2791	2628	2377	2033	1678	1436	1230	1109	1903	110/
88	2618 2451	2542 2486	2264	1949	1612	1422 1463	1209	1324	1342	1309
-		2 2 2	- - -	Conc	rete, ð=	-15 cm (X10-	(•			
			, oc. o		1000	6010	1166	3221	1 2447	94M
0	8551	8404 8404	9529 8508	8998 8619	520 <del>1</del>	5831	4074	3274	2551	2500
	2000	1000	0000	2100	7084	5441	3986	3224	2666	2599
0,5	0400		8479	0010	6527	5078	3854	3187	2773	2704
89	8853	8823	8316	7369	6024	4779	3721	3179	2893	2783
50	9168	2006	8162	1969	5575	4587	3695	3283	3082	3006
, 09	9657	0110	8088	6612	5298	4470	3724	3414	3200	248.1
70	9666	9196	7889	6319	2038	1044	3084	2849	3813	3717
20	5497 8899	9054 8962	7182	6766 6767	4864	4510	4295	4182	4197	4121
			_	Conc	rete, 🕯 =	:25 см (×10—	•			
-		1 200	0000	1010	0100	01 80	1554	1110	0871	0836
00	3001	3014	3021	3057	2704	2067	1473	1110	8060	0877
20	3047	3061	3056	2908	2507	1940	1408	1106	0940	0918
300	3090	3101	3023	2756	2327	1807	1355	1103	2260	0900 0700
40	3182	3152	2962	2617	2139	1692	1306	6011	0201	0/60
50	3303	3222	2906	2474	1980	1615	1298	7911	1000	1003
60	3456	3283	2881	2353	1877	1570	1310	1204	9611	1931
70	3567	3284	2817	2252	1799	1561	1361	2621	1200	1071
88	3372	3225	2680 9576	2128	172/	1518	1517	1490	1481	1457
	2100	cole	2010	1 0003	1110	~ ~ ~ ~ ~ ~	• • > -			

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Table a. 2.

BARRIER TRANSMISSION FACTORS FOR SCATTERED NEUTRONS FROM A NUCLEAR EXPLOSION  $\Lambda_{\rm n,s}$ , (8, 0,  $\alpha$ ) FOR VARIOUS MATERIALS

1	· .		~				
		180		0228 0211 0210 0247 0253 0254	0283 0309 0329 0351		0156 0157 0167 0177 0177 0177 0223 0223 0223 0266
		160		0297 0336 0354 0294 0252 0290	0314 0342 0365 0383		0166 0172 0172 0179 0179 0198 0230 02312 0230 0272
		140	, ,	0403 0415 0372 0372 0386 0305 0305	0320 0340 0358 0375		0222 0217 0211 0208 0209 0213 0221 0223 02246 02246 02267
-	es	120	4)	0607 0569 0541 0415 0342 0342	0375 0384 0398 0423	•	0285 0274 0261 0258 0258 0241 0241 0240 0248 0251 0269
	<sub>e</sub> , degre	100	16 см (×10—	0685 0719 0611 0553 0419 0411	0409 0406 0398 0398	20 см (×10-	0363 0355 0343 0343 0317 0303 0303 0292 0289 0289 0289
	•	80	l ron, à=	0861 0796 0750 0595 0518 0518	0513 0490 0483	lron, <sup>č</sup> =	0477 0467 0450 0430 0375 0375 0375 0375 0375 0321 0321 0322
		60		0937 0836 0784 0618 0618	0596 0582 0560 0556		1521 0509 0499 0482 0482 0439 0439 0439 0439 0394 0383
		40		0952 0843 0796 0704 0708	0701 0685 0655 0645	•	0542 0545 0534 0509 0509 0491 0468 0468 0453
		20		0934 0822 0775 0770 0763	0758 0744 0722 0708	·	05667 05667 05667 0560 0560 0550 0550 05
		Q		0840 0813 0796 0776 0773	0787 0794 0743 0698		0576 0577 0571 0571 0566 0572 0583 0591 0583 0556 0556
		degrees		055855	88288		582855692920 2829256

Table a. 2. continued

ะ						<sup>8</sup> . degre	es.			
degrees	0	20	40	60	.08	100	120	140	160	180
				Polyet	hy lene,	ð=16 см (×	10-5)			
0	10864	1880	2260	1260	0918	1 0684	0441	0312	0262	0282
10	0895	0914	0260	0937	0860	0631	0417	0324	0267	0286
88	0927	0943	0972	0935	0788	0585	0398	0331	0293	0288
2	0000	0/60	0060	0000	25/0	00500	0390	0330	0000	0.000
₽ ₽ ₽	1000	1037	1020	0530	00/2	0200	0300	0358	0320	0335
30	1131	1073	0928	0741	0571	0481	0399	0367	0363	0357
20	0611	1090	2060	0698	0541	0481	0420	0398	0387	0380
80	1130	1076	0860	0652	0527 0529	0483	0438	0423	0414	0407
2					7700					0010
				Polyet	hylene,	д=29 см (×1	(90)			<b>`</b>
0	1 0964	6660	L 1180	1322	0660	0834	0489	0341	0329	0313
10	1022	1060	1180	1245	0988	0769	0471	0359	0337	0328
20	1087	1318	1181	0211	0973	0106	0462	0372	0344	0341
ନ୍ନ	1141	1162	1182	1096	0897	0659	0453	0381	0359	0355
40	1222	1219	1166	1029	0816	0611	0366	0390	0375	0363
50	1311	1279	1148	6690 7990	0/3/	0567	0428	0413	0393	0394
38	1519	1366	1118	0827	0644	0569	0502	0476	0463	0457
80	1446	1361	1060	0777	0614	0573	0526	0507	0494	0488
8	1354	1357	1012	0738	0608	0603	0576	0560	0548	0543
	•.	-		Con	crete, ô	=5 см (×10-		_		
001	1291	1017	0868 1026	0801 0881 0881	0691	0611 0599	0315 0312	0315 0267	0162 0182	0289 0356
88	1295	1272	1059	0874 0936	0646 0836	0523 0537	0291 0490	0244 0292	0191 0267	0391 0343

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Table a. 2. continued

	180	0346 0398	0434 0461	0494 0581		0963	1222 1272	1324	1413	1638 1748 1937		_	0546	0666	0678	0788	0858	0916
	160	0398 0379	0391 0427	0452		1007 1180	1288 1331	1382	1577	1717 1833 9091			0580 0631 0631	0686 0686	0715	0763	0885	0944 1051
	140	0458 0438	0440	0485 0547		1705 1460	1408 1435	1670	1605	1681 1737 1937		-	0798 0765	0764 0764	0804	0812	8680	0942 1041
es	120	0563	0440	0438 0494		1236 1008	0985 1730	2071	1604	1611 1584 1796	1120		1082 0863	0803 0947	0860	0886	0800	0910 0993
e, degre	100	0721	0623	0598 0640	15 см (×10 <sup>-1</sup>	1906 2009	1700	2609	2417 2253	2216 1171 9306	-010		1311 1288	1172	1262	1181	1107	1095 1143
	80	0857 0730	0694	0640	oncrete³=	2527 - 2341	2160 2952	3060	2565 2457	2401 2265	6077	ncrete, 🖳	1724 1616	1489	1540	1373	1257	1198 1198
	60	1008	0898	0836 0829	່ ບັ	2817 3428	3384 3701	3556	3298 3179	3105 2962 2962	1067	00	1931 2007	1940	1834	1733	1616	1536 1501
	40	1137	1052	0966		3358 3971	<b>4</b> 119 4063	3970	3833 3833 3833	3705 3505 3705	0/00		1969 2092	2164 9195	2073	2029	1961	1869
	20	1233	1198	1124		4250 4368	4520	4281	4185	4040 3922	3049		2125 2178	2238 9936	2231	2235	2242	2151 2110
	0 -	1254	1239	1152	-	4400	4676	4275	4196	4526 4060	3801		2111 2200	2255	2262	2292	2339	2226
	degrees	40	888	288	•	00	28	99	88	2283	 F		001	50	2 €	22	35	886

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Table a. 3.

APPENDIX 2

BARRIER TRANSMISSION FACTORS FOR TOTAL DOSE OF FRAGMENT  $\gamma$  RADIATION OF .

NUCLEAR EXPLOSION  $\Lambda^f_{\gamma t}(\delta,\,\theta,\,\alpha)$  FOR IRON

		180	-	0175 0175 0160 0167 0223 0223 0223 0223 0229	1610	0160 0158 0153 0153 0153 0188 0188 0188 0204 0214 0214 0214 0175
		160	_	0193 0183 0183 0183 0183 0281 0219 0245 0245 0244	2020	0175 0171 0173 0164 0204 0208 0208 0208 0208 0225 0185
;		140		0313 0313 0280 0280 0223 0228 0228 0264 0264 0264		0347 0352 0352 0279 0217 0234 0234 0233 0235 0245 0245 0202
ı	es	120		0519 0488 0434 0434 0337 0337 0238 0291 0288 0288 0288 0288	) ) 	0759 0724 0572 0347 0347 0347 0275 0275 0272 0272 0272
•	<sup>9</sup> .degre	100	см (×10-4)	0771 0768 0512 0512 0332 0332 0259 0259 0204	сж (×10-5)	1333 1418 1418 0570 0570 0365 0365 0365 0365 03229 0186
		80	õ=5 .	0699 0756 0827 0827 0827 0827 0826 0826 0326 0376 0328	6 0 10	1053 1223 1223 1258 0803 0803 0409 0409 0409 0385 0409 0385
		60		0653 0787 0787 0780 0780 0780 0875 0763 0611 0611	_	1065 1492 1344 1381 1674 1600 1600 0887 0916 0934 0934
		40		0396 0528 0714 0714 0865 0952 1077 1387 1387 1384 1375	_	0564 0892 1325 1710 1954 2716 2716 2716 2716 2716 2716 2717 2717
		20		0254 0322 0694 0694 0989 1274 1274 1984 2300 2552	-	0301 0438 0438 0438 0456 1149 1149 2394 2394 2394 3392 2394 3392 2394 2595 5959
		0		0150 0244 0244 0578 0578 0906 1289 1888 1888 2652 2770 2134		0140 0577 0577 0577 0577 0577 0577 1505 1505
	<b>.</b>	degrees		88388498999 <b>0</b>		0.000000000000000000000000000000000000

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Table a. 3. continued

4. Table a. 180

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BARRIER TRANSMISSION FACTORS FOR SCATTERED DOSE OF FRAGMENT  $\gamma$  RADIATION OF NUCLEAR EXPLOSION  $\Lambda_{\gamma}^{f}$  s(s,  $\theta$ ,  $\alpha$ ) FOR IRON

180		0155 0152 0150 0147 0181 0181 0181
160		0170 0166 0163 0160 0193 0200
140		0236 0201 0200 0192 0192 0220 0211 0220
120		0340 0313 0313 0299 0263 0278 0278 0278 0278
100	=5 см	0473 0439 0390 0327 0304 0304 0255 0216
80	.0	0452 0452 0486 0486 0482 0454 0367 0367
60		0403 0424 0431 0449 0449 0469 0469
40	-	0243 0290 0379 0458 0526 0587 0587 0707
20	-1	0160 0186 0271 0378 0530 0674 0879
0		- 0113 - 0155 0240 0328 0497 0696 0991
dedrees		884985°o

Table a. 4. continued

180	0205 0202 0168		0150 0147 0150 0176 0176 0196 0196 0196 0196		0017 0017 0016 0020 0020 0021 0021 0021 0023 0023 00218
160	0216 0215 0178		0164 0160 0154 0154 0197 0197 0197 0197 0193 0173		0017 0017 0017 0017 0021 0021 0023 0023 0019
140	0231 0231 0191		0319 0320 0259 0218 0218 0217 0227 0188 0188		0029 0023 0025 0024 0024 0022 0022 0022 0022 0022
120	0253 0252 0209		0560 0537 0465 0317 0262 0254 0252 0252 0252		0155 0145 0091 - 0037 0031 0032 0032 0032 0032
100	0219 0211 0173	н (×10 <sup>—3</sup> )	0818 0827 0712 0557 0465 0322 0322 0270 0270 0214 0174	ж (×10—с)	0376 0403 0145 0146 0084 0040 0027 0027 0023 0023 0023
80	0286 0276 0224	δ=10 cı	0735 0818 0818 0856 0912 0350 0350 0350 0339 0339 0339	д=20 с	0244 0342 0406 0406 0304 0304 0105 0016 0016 0046 0046
é0	0378 0354 0300		0688 0855 0855 0851 0851 0851 0853 0648 0648 0675 0683		0283 0428 0378 0378 0378 0378 0378 0304 0175 0177 0177
40	0708 0751 0676		0406 0594 0795 0795 0955 0955 1064 1225 1585 1585 1584 1584	,	0111 0216 0356 0461 0463 0528 0777 0777 0707 0703
-20	1024 1190 1229	•	0269 0361 0361 0559 0798 1135 1442 1944 1944 2611 2250 3003	-	0022 0050 0144 0144 0253 0435 0659 0988 0988 1920 1928
0	1286 1255 1133	-	0132 0263 0488 0488 0656 1024 1481 1481 1481 1481 2155 2942 2616 2540	-	0005 0015 00131 0131 0135 0131 0135 1135 1
dearees	0.080	-	9829898990 88298898990	_	00000000000000000000000000000000000000

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Table a. 5

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APPENDIX 3

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BARRIER TRANSMISSION FACTORS FOR TOTAL DOSE OF CAPTURE  $\gamma$  RADIATION PRODUCED IN AIR FOR NUCLEAR EXPLOSION  $\Lambda_{\gamma t}^c(\delta,\,\theta,\,\alpha)$  FOR IRON

	180		0132 0114 0125 0120	0150 0145 0170 0170	0175		0157 0134 0146 0140	0176 0183 0201	0198
	160		0150 0126 0139 0132	0162 0159 0180 0180	0184		0168 0145 0158 0150	0187 0184 0196	0211
	140		0322 0288 0254 0185	0190 0175 0179 0186	0188	÷	0357 0323 0286 0213	0226 0211 0221	0234
es	120		0488 0440 0439 0377	0296 0214 0206	0221		0810 0720 0622 0464	0365 0291 0270	0270 0270 0299
<sub>9</sub> , degre	100	4 (×10-4)	0662 0633 0633 0685	0474 0348 0299 0560	0242	см (×10 <sup>-6</sup> )	1399 1375 1232 1089	0653 0429 0343	0263 0210
	80	ð=5 c	0661 0659 0727 1216	0757 0630 0501 0507	0480	§=10	1188 1329 1516 2886	1391 1007 0705	0560
	60		0643 0712 0779 0959	0920 1002 0861 0819	0810		1236 1516 1640 2125	1882 1993 1546	1301 1228 1021
	40		0472 0581 0739 0979	1178 1178 1422	1285		0736 1047 1406 2180	2006 2387 3007	3102 2727 2590
	20		0400 0468 0643 1265	1112 1386 1650 1815	2095 2142		0409 0557 0915	2095 3015 3793	3910 4925 5784
	0	-	0274 0399 0809 0928	1067 1423 1792 2264	2180		0235 0402 0798	1924 3025 4243	5829 5067 5330
	degrees		0000 0000	29000 29000 20000	88	I	30 <u>0</u> 0	9999 9999 9999 9999 9999 9999 9999 9999 9999	908 90 90 90

Table a. 5. continued

													ì
	180		0016	0014	0015	0015	6100	0018	6100	0021	0021	0025	
	160		2100	6100	0016	0015	. 0020	6100	0021	0022	0022	. 0226	,
	140		0034	0032	0028	0020	0024	0023	0024	0026	0026	0025	
es	120		0242	0216	0133	0064	0042	0032	0031	0032	0032	0031	
8. degree	100	см (×10—6)	0736	0765	0539	0335	0136	0000	0037	0045	0026	0026	
	80	5=20 c	0459	0636	0764	1688	0554	0296	0151	0166	6900	0068	
	60		0577	0838	0851	1157	0963	0380	0690	0549	0344	0371	-
	40		0214	0392	0617	1200	0960	1150	1579	1389	1941	1480	-
	20		0041	0088	0225	0777	0877	1561	9046	0120	2033	2723	
	0		0034	0035	0111	0283	0660	1350	0001	1020	-0986	3377	
	Degrees	•       	0	10	۶ ۲	18	ŝ		ŝ	300		6	

Table a. 6.

BARRIER TRANSMISSION FACTORS OF SCATTERED DOSE OF CAPTURE  $\gamma$  RADIATION FROM AIR FOR NUCLEAR EXPLOSION  $\Lambda^{C}_{\gamma s}(\delta, \; \theta, \; \alpha)$  FOR IRON

degrees	2065 2643 4001	200 2660 3745 3745	40 2774 3061 3761 4640	60 3441 3320 3779 4409	80 8=5 ( 3745 3367 3691 5111	*. degr           100           200           c.k (×10 <sup>-5</sup> )           3336           3335           3335	<b>ees</b> 120 25949 2594 25740 25740	140 2307 1956 1874	160 1286 1083 1136	180 1127 0968 1063 1025
9	4555	5492	6494	4440	4040	2995	2228	1582	1387	1280
50	6639	6219	5607	4744	3644	2447	1872	1467	1357	1238

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Table a. 6. continued

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						θ, degre	es			
degrees	0	20	40	60	80	100	120	140	160	180
8688	7498 8414 8843 8881	7133 7543 7644 7832	6297 7081 6420 6481	4591 4664 4637 4612 4612	3277 3397 3268 3215	2227 2112 1952 1968	1762 1713 1722 1722	1500 1550 1558 1574	1440 1535 1540 1562	1324 1440 1443 1468
				-	$\delta = 10 c$	ж (×10-г)				5,10
822622433220 82260	0217 0347 0645 0938 1244 1775 2178 2687 2687	0350 0441 0665 1412 1275 1635 2016 2264 2465	0506 0506 0662 0839 0839 1159 1159 1617 1617 1608	0738 0197 0908 0908 1136 1118 1118 0969 0981 0981 0981	0776 0776 0876 0876 0880 0880 0712 0548 0524	0783 0723 0749 0749 0508 0366 0366 0262 0244	0569 0503 0399 0328 0328 0258 0250 0250 0250	0352 0286 0260 0211 0211 0215 0215 0215 0215 0215	0157 0135 0148 0175 0175 0197 0197 0198	0146 0125 0136 0131 0159 0174 0187 0187 0187
88	2590	2702	1592	0139	0497	7070	0170			
					$\delta = 20$	см (×10 <sup>-6</sup> )			1 0017	0016
00 <u>0</u> 00	0031 0033 0099 0227	0038 0075 0175 0566	0165 0279 0386 0558	0341 0439 0488 0654	0266 0370 0431 0918	0405 0400 0342 0248 0111	0183 0162 0058 0040	0033 0031 0027 0019 0023	001500000000000000000000000000000000000	1000 1000 1000
50 0	0492	0561	0565 0685	0567	0220	0053	0031	0022	0020	0019
9689 9689 9689	1315 1696 1683 1854	11575 1575 1300	0915 0794 0801	0345 0265 0282	0157 0065 0066	0044 0025 0240	0031 0031 0030	0025 0025 0023	0027 0025	0020

APPENDIX 4

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Table a. 7

BARRIER TRANSMISSION FACTORS FOR TOTAL DOSE OF  $\gamma$  RADIATION OF RADIOACTIVE FALLOUT FROM NUCLEAR EXPLOSION  $\Delta_{\gamma}^{r}(\delta,\,\theta,\,\alpha)$  FOR IRON

					¢	degrees				
Time Following Expfosion	0	20	40	60	80	001	120	140	160	180
				=5 см (×1	(1-0)					
1,12 hr 23,8 hr 4,57 days 9,8 days	0012 0011 0011 0012	0103 0099 0100	0190 0172 0165 0188	0427 0240 0250 0385	0613 0404 0300 0587	0667 0531 0343 0634	0550 0604 0345 0500	0522 0563 0274 0512	0792 0604 0893	0484 0491 0361 0612
			ð=	=10 см (Х	10-5)					
1,12 hr 23,8 hr 4,57 days 9,8 days	0015 0015 0015 0014 0015	0146 0130 0125 0143	0354 0189 0205 0327	0655 0255 0313 0604	1050 0476 0405 0405	1153 0672 0452 1113	0849 0763 0424 0764	0889 0727 0345 0934	1675 0807 0374 1962	0982 0567 0514 1322
	ł		0= -	=20 см (Х	(2-01)					
1,12 hr 23,8 hr 9,8 days	6000 6000	0167 0087 0141 0167	0413 0144 0223 0425	1013 0254 0264 0994	2917 0771 0574 2715	3211 1017 0684 2931	1617 1624 0492 1537	1035 1014 0413 1093	2610 2630 0517 3230	1531 1524 0703 2617
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Table a. 8.

BARRIER TRANSMISSION FACTORS OF DOSE OF SCATTERED  $\gamma$  RADIATION FROM RADIOACTIVE FALLOUT FROM NUCLEAR EXPLOSION  $\Lambda^r_{\gamma s}(\delta, \theta, \alpha)$  FOR IRON

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	180		0431 0420 0345 0524		0941 0542 0500 1191		1520 1509 0690 2114	
	160		0733 0523 0233 0815		1622 0774 0351 1880		2519 2521 0504 2924	
<ul><li>degrees</li></ul>	140		0430     0461     0384     0433       0321     0425     0384     0433       0321     0425     0482     0477       0234     0262     0260     031       0450     0471     0383     0413       010-1     0383     0383     0443		0850 0688 0312 0901	-	1007 1005 0392 1060	
	120				0701 0697 0372 0692	-	1410 1420 0457 1425	
	100				0912 0596 0404 0940	_	2704 0921 0602 2724	
	4) 89			0871 0422 0361 0901	(1-0)	2509 0725 0521 2503		
	60	5 си (×1(	0315 0205 0215 0323	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	0572 0239 0301 0591	 20 см (X1	0901 0245 0263 0981	
	40	=2	0164 0141 0142 0162		0349 0186 0205 0326	-	0409 0137 0222 0411	
• •	20		010 8600 8600 8600 8600		0143 0128 0125 0140		0165 0087 0139 0165	
	0		0012 0012 0012		0015 0014 0014 0015		6000 6000	
time Following Explosion			1,12 hr 23,8 hr 4,57, days 9,8	<b>.</b>	1,12 hr 23,8 hr 4,57 days		1,12 hr 23,8 hr 4,57 days 9,8 *	

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APPENDIX 5

Table a. 9.

YIELD OF CAPTURE  $\gamma$  BADIATION OF VARIOUS MATERIALS ON SIDE OPPOSITE SIDE BOMBARDED BY NEUTRONS  $\Lambda^3_{\gamma i}(\delta)$  (rad of  $\gamma$  QUANTA PER 1 rem NEUTRONS FROM NUCLEAR

EXPLOSION WITH NORMAL INCIDENCE TO SURFACE OF SHIELD)

Thicknes	Conc	crete		uo	Wat	er
of Bar-	Thermal	Fast	Thermal	Fast	Thermal	Fast
rier, cn	Neutrons	Neutrons	Neutrons	Neutrons	Neutrons	Neutrons
2	0,0754	0,0850	0,132	0,0370	0,0630	0,0535
ę	0,0954	0,1135	0,122	0,0465	0,0727	0,0690
4	0,1081	0,1352	0,102	0,0526	0,0860	0,0850
9	0,1194	0,1644	0,065	0,0580	0,091	0,1081
8	0,1178	0,1770	0,040	0,058	0,0885	0,1192
10	0,1110	0,1830	0,0248	0,0554	0,0826	0,1276
12	0,1043	0,1796	0,0153	0,0515	0,0768	0,1250
14	0160'0	0,1532	0,0097	0,0471	0,0701	·0,1228
16	0,0810	0,1620	0,0061	0,0427	0,0643	0,1178
18	0,0710	0,1503	0,0038	0,0384	0,0593	0,1110
20	0,10622	0,1393	0,0024	0,0343	0,0543	0,1044
24	0,0476	0,1152	0,0010	0,0272	0,0459	1660'0

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Table a. 10.

YIELD OF CAPTURE  $\gamma$  RADIATION ON SIDE OPPOSITE INCIDENT NEUTRONS  $\Delta^3_{\gamma j}(\theta, \ \alpha) \ (rel. units) \ \times 10^{-4}$ 

	160	1364	1575	1671	1345	1121	1228	1420	15-13	1655	1661
	140	17.12	1876	1889	1727	1331	1381	1481	1567	1646	1605
S	120	2497	2483	2447	1938	1661	1680	1811	1848	1910	1801
<sub>9,</sub> degre	100	2631	2682	2747	2587	2078	2053	2093	2075	2079	2240
i i	80	3286	3343	3353	2820	2611	2596	2715	2696	2694	2727
	60	3492	3514	3540	3347	3728	3156	3239	3244	3224	3291
	40	3606	3577	3661	3487	3587	3684	3761	3858	3825	3812
	20	3549	3491	3579	3751	3923	4051	4169	4202	4232	4247.
	0	3169	3458	3647	3790	3970	4133	4322	4-162	4357	4188
	degrees	0	10	20	30	01	50	60	20	S	06
	s, degrees	degrees 0 20 40 60 80 100 120 140 160	degrees s, degrees   0 20 40 60 80 100 120 140 160   0 3169 3549 3606 3492 3286 2631 2497 1742 1364	degrees s, degrees   0 20 40 60 80 100 120 140 160   0 3169 3549 3606 3492 3286 2631 2497 1742 1364   10 3458 3491 3577 3514 3343 2682 2183 1876 1575	degrees 0 20 40 60 80 100 120 140 160   0 3169 3549 3606 3492 3286 2631 2497 1742 1364   10 3458 3491 3577 3514 3343 2682 2497 1772 1364   20 3647 3579 3661 3533 2747 2447 1889 1671	3, degrees     degrees   0   20   40   60   80   100   120   140   160     0   3169   3549   3606   3492   3286   2631   2497   1742   1364     10   3458   3491   3577   3514   3343   2682   2447   1876   1575     20   3647   3579   3661   3540   3353   2747   2447   1889   1671     30   3790   3751   3487   3347   2820   2587   1938   1575	s, degrees     degrees   0   20   40   60   80   100   120   140   160     0   3169   3549   3606   3492   3236   2631   2497   1742   1364     10   3458   3491   3577   3514   3343   2682   2497   1742   1364     20   3458   3491   3577   3514   3343   2682   2447   1876   1575     20   3647   3579   3661   3540   3353   2747   2447   1889   1671     30   3790   3751   3747   2820   2587   1938   1727   1345     40   3970   3923   3587   3728   2611   2078   1661   1331   1121	4 EGFees   , degrees     0   2169   3549   3606   3492   32386   2631   2497   1712   1361     0   3169   3549   3606   3492   32386   2631   2497   1712   1361     10   3458   3491   3577   3514   3343   2682   2183   1876   1575     20   3647   3573   3514   3343   2682   2147   1876   1575     20   3647   3540   3353   2747   2447   1889   1671     30   3790   3751   3487   3347   2820   2587   1938   1727   1345     40   3570   3681   3728   2611   2078   1661   1331   1121     50   4133   4051   3684   3156   2596   2053   1680   1331   1121     50   4133   4051   3641   2796   2053   1680   1331   1121	4 egrees       0     20     40     60     80     100     120     140     160       0     3169     3549     3606     3492     3286     2631     2497     1742     1364       10     3169     3549     3606     3492     3286     2631     2497     1742     1364       20     3647     3577     3514     3343     2682     2483     1876     1575       20     3647     3577     3514     3343     2682     2487     1876     1575       20     3647     3540     3353     2747     2447     1876     1575       20     3790     3751     3487     3347     2820     2587     1938     1671       40     3970     3923     3728     2611     2078     1661     1331     1121       40     3923     3761     3239     2747     2038     1727     1345       50     4133     405	4 6 9 $\frac{1}{2}$ , degrees       degrees     0     20     40     60     80     100     120     140     160       10     3169     3549     3606     3492     3236     2631     2497     1742     1364       10     3458     3491     3577     3514     3343     2682     2483     1876     1575       20     3451     3579     3661     3353     2747     2447     1889     1671       30     3790     3751     3343     2682     2631     2447     1889     1671       30     3790     3751     3347     2820     2587     1938     1727     1345       40     3970     3923     3587     3728     2611     2078     1661     1331     1121       40     3970     3584     3156     2736     2053     1660     1331     1121       50     4133     4051     3284     2736     2053	4 GGT cest       0     20     40     60     80     100     120     140     160       0     3169     3549     3606     3492     3236     2631     2497     1742     1364       10     3458     3491     3577     3514     3343     26682     2483     1876     1575       20     3451     3579     3661     3540     3353     2747     2417     1889     1671       20     3750     3751     3487     3343     26682     2483     1876     1575       30     3750     3571     3587     3343     26682     2483     1876     1575       30     3750     3751     2487     1336     1727     1345       40     3970     3751     3681     3156     2505     1680     1331     1121       70     4162     4051     3634     2696     2053     1680     1381     1420       70 <t< th=""></t<>

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