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away from our goal. A more compliant suspension for the apparatus was ruled out because the interferometer is too sensitive to low frequency rotational noise.

We completed our vibration isolation with the active position servo system. The three atom diffraction gratings are mounted on separate translation stages in the vacuum envelope. The servo's position sensor was an optical interferometer whose diffraction gratings were mounted on the same translation stages used for the atom gratings. By running the optical interference signal through a feedback network and applying the correction signal to a piezoelectric transducer on one of the translation stages, the relative motion of the atom gratings can be reduced to $\sim 40\text{nm rms}$.

Several other advantages were gained from using this optical interferometer to measure the grating position. This servo loop gives excellent stability over the time we take data, and even between adjacent data sets. Another advantage is that we can vary the reference voltage (and therefore the reference position) of the servo loop to scan a grating controllably over the interference pattern. This is how we have taken all of our data.

In addition to the required position stability in inertial space, the three gratings must be aligned so that their grating lines are parallel to $\sim 10^{-4}$ radians. To achieve this, we developed a system of alignment that relies on laser diffraction from the support structure of the gratings (the grating bars are too closely spaced to diffract a HeNe laser beam). This method also has the benefit of aligning the gratings in the earth's gravitational field, important because of the interferometer's sensitivity to gravitationally induced phase shifts.

Improved data analysis allowed us to demonstrate an excellent signal to noise ratio in our interferometer. We made improvements in our original data analysis techniques in both software and hardware. The hardware we are using is a Macintosh IIfx computer. The software consists of a set of macros and external subroutines used in a commercial data analysis package named Igor. The macros import the data, deconvolve the position from the lock loop error signal and provide the interface to our external subroutines that we have written in C. The external subroutines remove noise bursts from our data, bin the count data into positional bins, and extract the phase and amplitude of the interferometer signal.

Because each grating/slit assembly in the interferometer is neither in the near field nor the far field of the others, it is not possible to produce an analytic expression for the interference signal [CLA91]. We have therefore undertaken a program to model the interferometer numerically. In doing so we have advanced the state of the art for these calculations by devising a way to cast the multiple grating problem as a convolution problem, enabling us to use Fast Fourier Transforms [TUR91, TPK91]. A ten minute run on a CRAY can simulate the interferometer with an incoherent source possessing the actual velocity profile. The ability to model our experiment not only enables us to predict the size of our signals and to optimize design changes to the interferometer, but also to plan experiments to study other atom optical effects (eg. Talbot fringes [JAL79, SUT79] in the intermediate field of a single grating). Numerical simulations have so far allowed us to investigate several important issues in interferometer design. We have investigated the rate at which fringe contrast degrades with mis-spacing of the three gratings and due

to the spread of initial velocities (and corresponding change of deBroglie wavelengths) in the source beam. We have recently examined the possibilities of constructing interferometers with varying degrees of beam collimation, and we plan to study the effects of source coherence (the collimator does not really have a blackbody source behind it).

At MIT Submicron Structures Laboratory (SSL)

The key precursor to our successful program to build a three grating interferometer came from our demonstration [KSS88] that the 2000 Å period transmission gratings constructed for x-rays in the MIT Submicron Structures Laboratory could diffract atoms. We collaborated with M. Schattenberg in the SSL to produce gratings with higher transmission for atoms, but these gratings — although they gave excellent diffraction patterns — proved unsuitable for our interferometer due to previously undetected large scale irregularities. (Gratings for an interferometer must not deviate from a perfect grating by more than a small fraction of a period over the active area.) We then went to the National Nanofabrication Facility and learned to make gratings ourselves.

At National Nanofabrication Facility (NNF)

With the help of the staff (particularly M. Rooks), we used the facilities at NNF to develop procedures for making atom optics in thin (210 nm) silicon nitride membranes supported by conventional silicon wafers. We developed a new nanofabrication process for this — a special reactive ion etching gas mixture which etches Si₃N₄ faster than PMMA (plexiglass). This enables us to write a high resolution pattern in a thin overlayer of PMMA using an electron beam writer, chemically remove the exposed material, and then use the resulting pattern as a direct positive mask for etching the Si₃N₄ membrane, circumventing the need for making an intermediate metal mask as was previously necessary. This simplification increases the resolution and reliability of the fabrication process while shortening the cycle time, and may have other applications in nanotechnology. Although we made 2000 Å gratings which gave good diffraction patterns, the electron beam writer had some large scale pattern alignment problems which prevented their use in the interferometer — therefore we used our 4000 Å gratings to demonstrate the interferometer.

F. References

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