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CERENKOV RADIATION AND ELECTROMAGNETIC PULSE PRODUCED BY ELECTRON BEAMS TRAVERSING A FINITE PATH IN AIR

by

F. R. Buskirk and J. R. Neighbours

March 1986

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# CERENKOV RADIATION AND ELECTROMAGNETIC PULSE PRODUCED BY ELECTRON BEAMS TRAVERSING A FINITE PATH IN AIR

by

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# ABSTRACT

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Coherent Cerenkov radiation has been investigated previously in the time domain for an infinite path. The present calculations for a finite path length show an effect analogous to diffraction (in the frequency domain) in which radiation fields appear both at Cerenkov angles and at other angles. The latter have previously been named electromagnetic pulse fields.

# I. INTRODUCTION

Cerenkov radiation occurs when charge moves faster than radiation in a medium. Most work  $1^{-4}$  is concerned with optical radiation produced by a point charge, and involves the Fourier spectra of these fields. Here, in contrast, we explore the time dependence of the radiation fields, using our earlier<sup>5</sup> formulation to describe the radiation from a bunch of electrons passing through an infinite medium; however, now the medium has finite length, which causes diffraction so that the radiation is produced at other angles than the usual Cerenkov angle. This spreading by diffraction was investigated earlier in the Fourier expansion approach,  $6^{-8}$  but using the present time dependent fields, new insights are developed, and for shorter paths, the Cerenkov fields are related to other forms of radiation, namely, what is referred to as electromagnetic pulse  $(EMP)^9$ , transition and ordinary dipole radiation. It is much easier to understand Cerenkov radiation from a finite-size charge than from a point charge; in the former case, the Cerenkov fields remain finite but become singular for a point charge. $^{10}$ 

#### II. TIME DEPENDENCE OF FIELDS

Let all charges within a bunch move along the z-axis with the same velocity v, which is larger than the velocity c of radiation in the medium. Let  $c_0$  be the velocity of radiation in

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vacuum, let  $s^2 = x^2 + y^2$ , and assume the volume charge density  $\rho_{V}$  has the form

$$\rho_{\mathbf{v}}(\vec{\mathbf{r}}, \mathbf{t}) = \rho_{\mathbf{0}}(z - v\mathbf{t})\delta(\mathbf{x})\delta(\mathbf{y})$$
(1)

where  $\rho_0$  represents an arbitrary line density. Following our earlier work^5, the vector potential is

$$\vec{A}(\vec{r},t) = \frac{\vec{v}}{c_0} \int R^{-1} \rho_0(u) dz^2$$
(2)

where the variable  $u = z^{-vt}$  becomes

$$u = z^{2} - vt + \frac{v}{c} [s^{2} + (z - z^{2})^{2}]^{1/2}$$
(3)

Retardation is included by the form of u in Eq. 3. The magnetic field B has radiation terms resulting from taking the derivatives of  $\rho_0$  in Eq. 2. This leads to B in the 0 direction of cylindrical coordinates, with a magnitude

$$B = \frac{v^2}{cc_0} \int \frac{s}{R^2} \rho_0(u) dz^* .$$
 (4)

An approximate evaluation of Eq. (4) proceeds as follows: u(z') is plotted as a function of z'. For further calculation assume  $\rho_0'$  has linear rising and falling ramps, of width a, separated by a distance b. Then  $\rho_0'$  consists of two opposite polarity square

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pulses. For early times, the u curve is high (see Reference 5, Fig. 1) and the integrand is zero for all values of z<sup>2</sup>. As time increases, the u curve drops and the minimum intersects the  $\rho_0^2$ pulse, and contributes to the integral in Eq. (4). The field  $\vec{E}$ may also be calculated; in Ref. 5 it was shown that  $\vec{E}$  is perpendicular to  $\vec{B}$  and to  $\vec{R}_m$ , and  $\vec{R}_m$ , the vector from the particle (at the retarded time) to the observer, is at an angle of -c to the z-axis. Then E/B =  $c/c_0$ ; both the fields fall off as  $R^{-1/2}$ , appropriate for radiation from a cylindrical source, and the total energy radiated agrees with the Fourier approach. The two opposite pulses of the radiation field have the same separation as the front and rear slopes of the current pulse.

We now calculate the radiation fields for the case of a finite path. The physical situation shown in Fig. 1 is one in which a beam emerges at z' = 0 from an accelerator, passes through a dielectric medium and at z' = Z, stops in an absorber. The path (usually air) from z' = 0 to Z is the radiator. The beam bunch has a linear rising ramp of length a, is then constant and has a linear decrease of length a, with an effective length b, from the midpoint of the rise to the mid point of the fall.

The calculation of the fields is based on Eq. (4) for B; corresponding results will hold for E. We again assume that  $s/R^2$ is about constant in the range of integration and may be factored out. Then s/R = sin 0 and we have

-4-

$$B = \frac{v^2}{cc_0} \frac{\sin \theta}{R} \int \rho_0^{-1}(u) dz^{-1}.$$
 (5)

The main difference between the calculations below and those done previously<sup>5</sup> is that the range of  $z^2$  is finite, from 0 to Z, to represent a finite length (Z) of radiation source. The results of the integration depend on the relation of various parameters.

#### A. Outside the Cerenkov Cone, Z Long

Fig. 2 represents the situation involved in evaluating Eq. (5). The u(z') curve moves down in time (Eq. 3). For early times, the integrand is always zero. Later the u curve moves down and intersects the dotted rectangle representing the region of u and z' where the integrand is constant. In the situation shown, the integral builds up to a peak value in a time interval a/v, and the integral saturates at the value  $\rho_0\Delta z = \rho_0 a/slope$ of u-curve. From Eq. (3), the slope is

$$\frac{\partial u}{\partial z} = 1 - \frac{v}{c} \frac{z - z'}{\sqrt{s^2 + (z - z')^2}}$$
$$= 1 - \frac{v}{c} \cos \theta \quad . \tag{6}$$

Furthermore  $\rho_0'a$  is the peak value of  $\rho_0$  which is  $I_0c_0/v_{\star}$  where  $I_0$  is the peak current.

Thus the integral saturates at

$$B_{\max} = \frac{v}{c} \frac{\sin \theta}{R} I_0 \frac{1}{1 - \frac{v}{c} \cos \theta}$$
(7)

where R is measured from the start of the source, z' = 0, to the observer, and it makes an angle 0 to the z-axis.

The rise time was a/v, a similar fall time occurs, and the duration of the pulse is given by Z(slope of u curve)/v or  $Z(1-\frac{v}{c} \cos \theta)/v$ . Because this combination appears often, we define the effective length

$$Z_{\theta} = Z(1 - \frac{v}{2} \cos \theta)$$

(8)

The significance of  $Z_e$  is that it yields the time difference for two signals emitted by a given charge at two points separated by a distance Z in the lab. Note that  $Z_e = 0$  at the Cerenkov angle, as expected, because signals emitted at  $\theta_c$ from all parts of the path reach a distant observer at the same time.

Thus we have, for the leading ramp of the current pulse, a field at the observer of value given by Eq. (7) with lengths shown in Fig. 3a. Here the lengths are times multiplied by v. Also shown is the negative field pulse caused by the back ramp of the current pulse. The two pulses combine to give the symmetric pulse of Fig. 3b, with a separation that is the larger of  $Z_e$  or b, and a duration that is the smaller of  $Z_e$  or b.

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# B. On the Cerenkov Cone

If the observer is in the Cerenkov region (region B in Fig. 1), the rectangular region of integration in Fig. 1 is centered about the minimum in the  $u(z^{-})$  curve. If Z is large, the integral for the field is the same as in our earlier paper. The pulse starts when the  $u(z^{-})$  curve, as it advances down in time, becomes tangent to the rectangular integration region. The integral increases as  $t^{1/2}$  until the u curve is tangent to the lower boundary, and then decreases, again proportional to  $t^{1/2}$ . The result is

$$B_{\max} = \frac{v^2}{c^2} \frac{1}{R_m} \frac{\sin \theta_c}{\tan^{3/2} \theta_c} I_0 \frac{2\sqrt{2}}{\sqrt{3/a}}$$
(9)

The positive and negative pulses have rise and fall times of a/v and are separated by b/v. The field falls as  $1/R_m$  as one proceeds outward at a fixed angle with s =  $R_m \sin \Theta_c$ .

If the path is short, the B field pulse, as a function of time, increases as  $t^{1/2}$  as noted above, but the maximum value of the integral for Eq. (5) occurs when the  $u(z^{*})$  curve first intersects the vertical limits (0 and Z) of integration of the rectangular region in the u-z<sup>\*</sup> plane. Then the maximum magnetic field becomes

$$B_{\text{max}} = \frac{v^2}{cc_0} \frac{\sin \theta_c}{R_m} \rho_0^2 (u)Z$$
$$= \frac{v^2}{cc_0} \frac{\sin \theta_c}{R_m} \rho_0^2 = \frac{Z}{cc_0}$$

(10)

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Eliminating  $\rho_0$  in terms of  $I_0$  yields:

$$B_{\max} = \frac{v}{c} \frac{\sin \theta_c}{R_m} I_o \frac{Z}{a} .$$
 (11)

In this case, positive and negative pulses of width a have a separation b, which is the same as for the long radiator, but the pulses have flat tops.

If the observer moves out along the Cerenkov cone, the long path case changes into the short path case. In the former, the source is long and the observer sees a field associated with cylindrical symmetry. But as the observer moves out, the radiator of length Z appears to be short, leading to the  $R^{-1}$ field of Eq. (10) instead of the  $R^{-1}$  (cylindrical) field of Eq. (9). The transition from the cylindrical wave to the spherical wave occurs when  $2\Delta z$  (defined in Eq. (18) of Ref. 5) is equal to Z. This yields for the value of s, denoted by s<sub>s</sub>

$$a s_s \frac{v^2}{c^2} \frac{2}{\tan^3 \theta_c} = z^2$$
 (12)

That is, for  $s > s_s$  [where  $s_s$  satisfied Eq. (12)] the wave becomes spherical and decreases as  $R^{-1}$  or  $s^{-1}$ .

#### C. Outside the Cerenkov Cone, Z Short

In the previous sections, cases were considered in which  $Z_e > b > a$ , but other situations are possible. Because a is the rise

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and fall length of the current pulse, and b is the width measured between the half-maximum points, we must have  $b \ge a$ . We could then have  $b > Z_e > a$  and  $b > a > Z_e$  cases. For the latter the field is again obtained by evaluating Eq. (5). The horizontal range of integration is short so that the range of the variable  $z^{-1}$ is always Z, if we are off the Cerenkov cone. The result is 
$$B = \frac{v^2}{cc_0} \frac{\sin \theta}{R} c_0^2 .$$
 (13)

Noting that  $\frac{v}{c_0} \rho_0 a = I_0$ , the peak current in the pulse, we find

$$B = \frac{v}{c} \frac{\sin \theta}{R} I_0 \frac{Z}{a}$$
 (14)

The field pulse at the observer will have a rise and fall time  $Z_e/v$ , the pulse length is a/v, and it will be followed by a similar negative pulse at a time b/v later.

#### III. SEMI-INFINITE PATH

Let the beam emerge from the accelerator window at z = 0and traverse a path, which is idealized to be infinitely long. Starting from the point where the beam emerges, define a cone with apex angle  $0_{\rm C}$  relative to the beam. If the observer is anywhere inside the cone, the minimum of the  $u(z^2)$  curve will

-9-

intersect the  $\rho_0(u)$  pulses, and the radiation field will be the same as found in Ref. 5 and described in Sec. II. B. The radiation fields will have a positive and negative pulse separated by a distance b. This region is dominated by Cerenkov radiation.

Now let the observer be outside the Cerenkov cone. The situation is somewhat like that described in Sec. II., except that the horizontal range of integration is infinite in the positive direction. The rising part, or held. of the current pulse leads to fields which are essentially constant, whereas the following tail gives an opposite field delayed by a time b/v. We thus have a field pulse in the region outside the Cerenkov cone of length b but only of one sign. This is EMP, and in the above model, both EMP and Cerenkov radiation exist. The field lines are shown qualitatively in Fig. 4.

## IV. JUSTIFICATION OF MODEL

The model using a finite path is supposed to represent radiation from a bunch of electrons, emerging from an accelerator system at z' = 0, and stopped by some means at z' = Z. No specific account has been made for these boundaries; radiation by return currents has been neglected. Consider the following model: the charge that suddenly appears at z' = 0 in all the cases is furnished by a source electron pulse of lower velocity  $v_s$  which moves in the region z' < 0 and meets the previously specified pulse at z' = 0. To conserve charge at all times at z' = 0,  $b_s = b \frac{v_s}{v}$ ,  $a_s = a \frac{v_s}{v}$  and  $I_{OS} = I_O$ , where  $I_{OS}$ ,  $a_s$ , and  $b_s$ 

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are the current, rise, and width parameters for the slow pulse approaching the boundary from negative  $z^2$ .

The radiation pulse from the slow incoming electrons is given by Eq. (7) with the appropriate lower velocity  $v_s$ substituted. By inserting a very low value of  $v_s$  the field becomes small, and because  $b_s$  becomes small, the time in which the field pulse occurs becomes short, so that the radiated energy becomes very small. Thus we conclude that inclusion of source and return currents, required to conserve charge, contribute little to radiation fields calculated, and may be neglected.

#### V. DISCUSSION

Cerenkov radiation has been described above and in Ref. 5 in terms of time dependence of the radiation fields caused by finite charge distributions such as are realized by bunches emitted by an accelerator. In the usual Fourier-expansion formalism, either a finite size of the charge or dispersion in the medium limits the radiated power at the high-frequency end of the spectrum, and a finite length of path in the medium produces diffraction of the angle of the emitted radiation about the Cerenkov angle.

The present time-dependent field formulation reveals the following properties of Cerenkov radiation: A. The radiation is associated with dI/dt at the leading and trailing parts of the pulse. B. The Cerenkov radiation (for an infinite path) consists of positive and negative pulses separated by a distance, which is the pulse length. C. For a semi-infinite path, Cerenkov radiation appears within a cone of angle  $\Theta_c$ , whose apex

-11-

is at the start of the path. An EMP pulse appears outside that cone. D. For a finite path of length Z, both Cerenkov and EMP appear, the latter dominating as Z becomes smaller. For Z short and  $\beta$  small, the sin  $0/(1 - \beta \cos \theta)$  dependence of the EMP pulse becomes sin  $\theta$ ; thus the fields at low  $\beta$  becomes essentially a single pulse of dipole radiation.

Finally it should be noted that Cerenkov radiation is not a different radiation to be added on to other forms of radiation when v > c (medium) but should appear naturally in a correct calculation of the radiation. If v < c, radiation also occurs but without the characteristic shockwave-like character of Cerenkov radiation.

#### ACKNOWLEDGMENT

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Fig. 1. For a finite path length Z, an observer in regions A or C finds EMP radiation, shown in Fig. 4 and developed in Secs. II.A and II.C. In region B, a Cerenkov field pulse described in Sec. II.B occurs.

Fig. 2. The variable u plotted as a function of z'. As time increases, the curves are displaced downward. The derivative of the current pulse is shown. When the u(z') first intersects the  $p_0'$  pulse, the integrand of Eq. (4) becomes nonzero, and the Cerenkov pulse starts (Sec. II).

Fig. 3. The field pulse outside the Cerenkov cone for finite path length. Curve (a) shows the positive and negative pulses separately, whereas (b) shows the composite field. The separation S is the larger of  $Z_e$  or b, whereas the duration D is the smaller (Sec. II.A and II.C.)

Fig. 4. A qualitative representation of field lines for an electron bunch traversing a semi-infinite path. As the position of the observe changes, the pulse form changes as shown in the inset. Cerenkov radiation occurs at the lower right; EMP pulses occur to the upper left.

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