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# TECHNICAL REPORT RE-CR-85-3

# EXACT DETECTION PROBABILITY AND FLUCTUATION LOSS FOR A PARTIALLY CORRELATED RAYLEIGH TARGET

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PARTIALLY CORRELATED RAYLEIGH TARGET

### IRVING KANTER

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tuation loss for a Gauss-Markov signal is determined as a function of number of pulses integrated, the correlation between pulses, and the specified detection and false alarm probabilities. This exact loss is compared to Barton's approximation.

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# PART ONE. EXACT DETECTION PROBABILITY

### I. INTRODUCTION

The determination of detection probability when the sum of N detected pulses of signal plus noise is compared to a threshold has been studied by Marcum, Swerling, Schwartz, and Vannicola, among others. Marcum<sup>1</sup> treated the non-fluctuating target; Swerling<sup>2</sup> treated two limiting fluctuating target situations (complete correlation and complete decorrelation) - he also presented a general method for treating arbitrary correlation<sup>3</sup>, but gave no results for particular correlation models; Schwartz<sup>4</sup> achieved an exact solution when there are only two pulses; Vannicola<sup>5</sup> has constructed a solution from Swerling's two extreme fluctuation models by considering M independent sets of N fully correlated pulses - his solution applies only to targets which are block correlated. This paper extends the previously referenced work and presents exact results for a Rayleigh target whose inphase (and quadrature) components have exponential correlation.

<sup>1,2,3,4,85</sup> See References at end of text.

### 11. ANALYSIS

buncte the inphase and quadrature components of signal and noise respectively by the Nx1 complex vectors A+iB and X+iY. We assume (i) A,B are i.i.d.; (ii) X,Y are i.i.d stationary Gaussian; (iii) the noise is additive and independent of the signal. Employing the Neyman-Pearson detection criterion, the probability of detection is given by

$$\mathbf{P}_{\mathrm{D}} \stackrel{\Delta}{=} \int_{\mathbf{V}_{\mathrm{T}}}^{\infty} \mathbf{P}_{\mathbf{v}}(\mathbf{v}) \, \mathrm{d}\mathbf{v} \tag{1}$$

where v is the integrator output normalized to the average per pulse noise power, i.e.,

$$\mathbf{v} \stackrel{\Delta}{=} \frac{|\mathbf{A} + \mathbf{x}|^2 + |\mathbf{B} + \mathbf{y}|^2}{26^2} \tag{2}$$

and  $V_{\rm T}$  is the normalized threshold determined by the specified probability of false alarm,

$$P_{FA} \stackrel{\Delta}{=} \int_{V_{T}}^{\infty} p_{v}(v ||A|^{2} + |B|^{2} = 0) dv$$
(3)

In order to determine  $P_D$  we require the probability density function (pdf) of v. We first write  $P_T(v)$  as an inverse Laplage transform

$$p_{v}(v) = \frac{1}{2\pi i} \int_{\gamma-i\infty}^{\gamma+i\infty} L_{v}(s) e^{SV} ds$$
(4)

Next recognizing

$$L_{v}(s) = \int_{0}^{\infty} p_{v}(v) e^{-sv} dv$$
 (5)

as the expectation of  $e^{-SV}$  we proceed to calculate this expectation over the domain of A,B,X,Y i.e.,

$$L_{v}(s) = \int \int \int_{-\infty}^{\infty} \int p(A, B, X, Y) e^{-s} \frac{|A+X|^{2} + |B+Y|^{2}}{2\beta^{2}} dA dB dX dY$$
(6)

In view of the three assumptions characterizing the signal and noise, the joint pdf of A,B,X,Y may be written as

$$p(A,B,X,Y) = p_{\bar{A}}(A)p_{\bar{A}}(B) \frac{1}{(2\pi\beta^2)} e^{-\frac{|x|^2 + |Y|^2}{2\beta^2}}$$
(7)

Thus (6) becomes

$$L_{v}(s) = \left[ \int_{-\infty}^{\infty} p_{A}(A) dA \int_{-\infty}^{\infty} \frac{a}{(2\pi\beta^{2})^{N/2}} dX \right]^{2}$$
(8)

or, completing the square in X,

$$L_{v}(s) = \left[\int_{-\infty}^{\infty} \frac{-\frac{s}{s+1} \frac{|A|^{2}}{2\beta^{2}}}{(s+1)^{N/2}} dA \int_{-\infty}^{\infty} \frac{-\frac{s+1}{2\beta^{2}} |x + \frac{s}{s+1} A|^{2}}{(2\pi\beta^{2}/(s+1))^{N/2}} dx\right]^{2}$$
(9)

Thus for any signal model we have the general formula

$$L_{v}(s) = \frac{1}{(s+1)^{N}} \left[ \int_{-\infty}^{\infty} -\frac{s}{s+1} \frac{|A|^{2}}{2\beta^{2}} dA \right]^{2}$$
(10)

A. Marcum Model

The non-fluctuating target may be accommodated by choosing

$$p_{A}(a_{1},\ldots,a_{N}) = \prod_{1}^{N} \int_{0}^{N} \int_{0}^{1} \int$$

where  $\delta$  (+) is the Dirac delta function. Then (10) yields Marcum's result

$$L_{v}(s) = \frac{1}{(s+1)^{N}} e^{-\frac{s}{s+1}N\chi}$$
(12)

where  $\boldsymbol{\chi}$  is the per pulse ratio of signal power to average noise power, i.e.,

$$\chi \stackrel{\Delta}{=} \frac{\alpha^2}{\beta^2} \tag{13}$$

B. Swerling Models

Typical radar targets are composed of a large number of scatterers whose distances relative to the radar change with body vibration and varying target aspect. The echoes from the individual scatters contribute amplitude and phase terms which combine at the radar frequency to produce a fluctuating signal. Swerling has bounded the effects of target correlation by considering two limiting situations: The slow fluctuation model (unity amplitude correlation on a scan, scan-to-scan independence) and the fast fluctuation model (zero amplitude correlation on a scan, pulse-topulse independence). These may be accommodated by choosing respectively

$$p_{A} = (a_{1}, \dots, a_{N}) = p_{a} (a_{1}) = \prod_{n=1}^{N} (a_{n} - a_{1})$$
 (14)

and

$$p_{a} (a_{1}, \dots, a_{N}) = \prod_{n} p_{a} (a_{n})$$
 (15)

Measurements<sup>6</sup> of echoes from many types of radar targets confirm that the best characterization of the first order amplitude statistics is given by the Rayleigh distribution. (The best characterization of the autocorrelation function of the radar cross section is given by an exponential function.) Thus we assume

$$p_{\sqrt{a^2+b^2}}(r) = \frac{r}{\alpha^2} e^{-\frac{r^2}{2\alpha^2}} \qquad (0 \le r < \infty)$$
(16)

where  $2a^2$  is the average target cross section. In the remainder of this paper we confine our attention to this "Rayleigh target".

The inphase (or quadrature) components have first order statistics given by

$$p_{a}(a) = \frac{1}{\sqrt{2\pi} \alpha} e^{-\frac{a^{2}}{2\alpha^{2}}}$$
 (17)

 $^{6}$  See References at end of text.

Thus the general formula (10) specializes to Swerling's slow fluctuation result (case 1)

$$L_{v}(s) = \frac{1}{(s+1)^{N-1} [(1+N\chi)s+1]}$$
(18)  
p=1

when (14) is employed as the correlation model and to Swerling's fast fluctuation result (case 2)

$$L_{v}(s) = \frac{1}{[(1+\chi)s+1]^{N}}$$
(19)

when (15) is employed as the correlation model.

# C. Partially Correlated Model

In order to accommodate Rayleigh targets which give rise to partially correlated signals  $(0 \le p \le 1)$  we introduce the general correlation model

$$P_{A}(a_{1},...,a_{N}) = \frac{e^{-\frac{1}{2\alpha^{2}}A^{T}C^{-1}A}}{(2\pi)^{N/2}|\alpha^{2}C|^{\frac{1}{2}}}$$
(20)

where "T" denotes transpose and C is the covariance matrix of  $\epsilon$ For this correlated Rayleigh model, (10) yields

$$L_{v}(s) = \frac{1}{(s+1)^{N}} \left[ \int_{-\infty}^{\infty} \frac{-\frac{1}{2\alpha^{2}} A^{T} (C^{-1} + \frac{s}{s+1} \chi I) A}{(2\pi)^{N/2} |\alpha^{2}C|^{\frac{1}{2}}} dA \right]^{2}$$
(21)

In (21) I is the identity matrix and  $\chi$  [c.f.(13)] is the average signal-tonoise ratio. In order to evaluate the integral we assume that C is positive definite; then so also is C<sup>-1</sup>; hence when  $s \ge 0$ , C<sup>-1</sup> +  $\frac{s}{s+1}$  I is nonsingular and (21) may be written as

$$L_{v}(s) = \frac{1}{(s+1)^{N}} \frac{1}{|c| |c^{-1} + \frac{s}{s+1} \chi I|} \left[ \int_{-\infty}^{\infty} \frac{-\frac{1}{2\alpha^{2}} A^{T} \left(c^{-1} + \frac{s}{s+1} \chi I\right)^{A}}{(2\pi)^{N/2} \left| \alpha^{2} \left(c^{-1} + \frac{s}{s+1} \chi I\right)^{-1} \right|^{\frac{1}{2}}} dA \right]^{2}$$
(22)

Since the integral equals unity we have

$$L_{v}(s) = \frac{1}{|(s+1)I + s\chi C|}$$
 (23)

Expressing the determinant in terms of the (positive) eigenvalues  $\lambda_1, \ldots, \lambda_N$  of C the Laplace transform of  $p_v(v)$  is given by

$$L_{v}(s) = \frac{1}{N}$$
(24)  
$$\prod_{i} [(1+\chi\lambda n)s+1]$$

Swerling's results may be obtained from this formula by choosing  $\lambda_1 = N$ ,  $\lambda_n = 0$ (n≠1) for the slow fluctuation case [c.f.(18)], and  $\lambda_n \equiv 1$  for the fast fluctuation case [c.f.(19)]. These correspond respectively to the singular covariance matrix

$$c\Big|_{\rho=1} = \begin{bmatrix} N_0 \\ O \\ O \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \end{bmatrix}$$

in which the first pulse contains all the average signal power of the pulse train and the remaining pulses contain no signal, and to the identity matrix

$$C |_{p=0} = \begin{bmatrix} 1 \\ 0 \\ 1 \end{bmatrix}$$
(26)

in which the total average signal power is uniformly distributed among the N pulses.

The inverse Laplace transform of (24) now yields the pdf of the integrator output. When  $\chi=0$  we obtain

$$p_v(v|\chi=0) = e^{-v} \frac{v^{N-1}}{(N-1)!}$$
 (27)

and when  $\chi \neq 0$ , (assuming distinct eigenvalues) we obtain

(25)

$$p_{v}(v) = \sum_{\substack{n=1 \ k \neq 1 \\ k \neq n}}^{N} \left[ \left( 1 - \frac{1 + \chi \lambda k}{1 + \chi \lambda n} \right) \right]^{-1} \frac{e}{1 + \chi \lambda_{n}}$$
(28)

Thus [c.f.(3)] the threshold,  $V_{T'}$  is given in terms of the specified P FA by

$$P_{FA} = e^{-V_{T}} \sum_{n=0}^{N-1} \frac{V_{T}^{n}}{n!}$$
(29)

and [c.f.(1)] the probability of detection by

$$P_{D} = \sum_{n=1}^{N} \prod_{\substack{k \neq n \\ k \neq n}}^{N} \left[ \left( 1 - \frac{1 + \chi \lambda_{k}}{1 + \chi \lambda_{n}} \right) \right]^{-1} e^{-\frac{V_{T}}{1 + \chi \lambda_{n}}}$$
(30)

In order to calcula e  $P_D$  for any particular correlation model we must provide a description of the correlation matrix. If the signal arises from a stationary process, then C is a symmetric Toeplitz matrix with N distinct elements. We should like to associate the eigenvalues with a single correlation parameter, p, in terms of which  $P_D$  may be conveniently characterized. For this purpose we assume that the signal is described by a first order Markov process. This is consistent with Edrington's measurements which show that the radar cross section is exponentially correlated.

Consider a signal which consists of a train of N pulses with uniform spacing T. Let the k,n element of C be taken as

$$C_{kn} = e^{-|k-n|T_{v}} \Delta_{p}|k-n| \qquad (0 \le p \le 1) \qquad (31)$$

 $\mathbf{C} = \begin{bmatrix} \mathbf{1} & \rho & \cdots & \rho^{N-1} \\ \rho & \ddots & \ddots & \rho \\ \vdots & \ddots & \ddots & \rho \\ \rho & \cdots & \rho & 1 \end{bmatrix}$ 

The eigenvalues of C provide a non-trivial solution to the matrix equation

$$[C - \lambda I]U = 0 \tag{33}$$

Since the sum of the eigenvalues equals the trace of C we have the relation

$$\sum_{i=1}^{N} \lambda_{n} = N$$
(34)

which we make use of later.

Consider first the special case of a low prf (pulse repetition frequency) waveform for which the interpulse spacing, T, is so large that the correlation of noncontiguous pulses may be neglected. Thus C becomes tri-diagonal and (33) is equivalent to a homogeneous by (boundary value problem). This consists of a set of homogeneous second order difference equations

$$\rho u_{n-1} + (1-\lambda)u_n + \rho u_{n+1} = 0 \quad (n=1,...,N)$$
(35)

together with the homogeneous boundary conditions

$$u_0 = u_{N+1} = 0$$
 (36)

Since the equation is linear and has constant coefficients, there are

then

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(32)

two solutions in the form

$$u_n = \gamma^n$$
 (37)

where

$$\gamma = \frac{\lambda - 1}{2\rho} \pm \sqrt{\left(\frac{\lambda - 1}{2\rho}\right)^2 - 1}$$
(38)

The condition  $\left|\frac{\lambda-1}{2\rho}\right| \ge 1$  implies either  $\lambda \ge 1+2\rho$  or  $\lambda \le 1-2\rho$ , each of which leads to a contradiction of (34). Hence we must have  $\left|\frac{\lambda-1}{2\rho}\right| \le 1$ , i.e.,  $\gamma = e^{\pm i\theta}$  where

$$\cos \theta \stackrel{\Delta}{=} \frac{\lambda - 1}{2\rho} \tag{39}$$

We write the general solution to (35) in terms of  $e^{\pm i n \theta}$  as

$$u_n = K_1 \cos n\theta + K_2 \sin n\theta$$
(40)

Since  $u_0 = 0$ , we have

$$u_n = K_2 \sin n\theta \tag{41}$$

and since  $u_{N+1} = 0$ ,  $\theta$  satisfies the transcendental equation

$$\sin (N+1)\theta = 0 \tag{42}$$

Further since (41) is a non-trivial solution,  $\theta$  cannot equal 0 or  $\pi$ . Thus the roots of (42) are

$$\theta_n = \frac{n}{N+1} \pi \quad (n=1,\ldots,N) \tag{43}$$

Note that they are equally spaced in the open interval  $(0,\pi)$ . Equation (39) then yields the distinct eigenvalues

$$1-2\rho < \lambda_n = 1 + 2\rho \cos\theta_n < 1+2\rho$$
 (n=1,...,N) (44)

so that  $P_D$  is given by (30).

When the correlation of non-contiguous pulses may not be neglected we require the eigenvalues of the full C matrix. As is easily verified, the inverse matrix,  $C^{-1}$ , is the tri-diagonal matrix



Thus the eigenvalue problem to be solved is

$$(1-\rho^2) \left( c^{-1} - \frac{1}{\lambda} I \right) V = 0$$
(46)



In the special case N=2, this becomes

$$\begin{bmatrix} 1 - \frac{1-\rho^2}{\lambda} & -\rho \\ -\rho & 1 - \frac{1-\rho^2}{\lambda} \end{bmatrix} \begin{bmatrix} v_1 \\ v_2 \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \end{bmatrix}$$
(48)

which leads immediately to the distinct eigenvalues

$$\lambda = 1 \pm \rho \tag{49}$$

Equation (30) then yields  $P_D$  which may be written in the form

$$P_{\rm D} = e^{-\frac{(1+\chi)^{\rm v} {\rm T}}{(1+\chi)^{\rm 2} - (\rho\chi)^{\rm 2}}} \left[ \frac{1+\chi}{\rho\chi} \sinh \frac{\rho\chi {\rm v}_{\rm T}}{(1+\chi)^{\rm 2} - (\rho\chi)^{\rm 2}} + \cosh \frac{\rho\chi {\rm v}_{\rm T}}{(1+\chi)^{\rm 2} - (\rho\chi)^{\rm 2}} \right]$$
(50)

which is the same result as that of Schwartz.

To solve (46) when N>2, we employ the technique developed for the low prf waveform; i.e., we again formulate and solve an equivalenc homogeneous byp. From (47) we have the equation

$$-\rho v_{n-1} + \left(1 + \rho^2 - \frac{1 - \rho^2}{\lambda}\right) v_n - \rho v_{n+1} = 0 \qquad (n = 1, \dots, N)$$
(51)

and the boundary conditions

 $\mathbf{v}_{0} - \rho \mathbf{v}_{1} = 0 \tag{52}$ 

$$\mathbf{v}_{\mathbf{N}+1} - \rho \mathbf{v}_{\mathbf{N}} = 0 \tag{53}$$

Again there is a solution of the form (37), provided that

$$\gamma = \left[1 + \rho^{2} - \frac{1 - \rho^{2}}{\lambda} \pm \sqrt{\left(1 + \rho^{2} - \frac{1 - \rho^{2}}{\lambda}\right)^{2} - 4\rho^{2}}\right] / (2\rho)$$
(54)

Since 
$$\left| \left( 1 + \rho^2 - \frac{1 - \rho^2}{\lambda} \right) / (2\rho) \right| \ge 1$$
 implies either  $\lambda \ge \frac{1 + \rho}{1 - \rho}$  or  $\lambda \le \frac{1 - \rho}{1 + \rho}$ ,

equation (34) cannot be satisfied. Thus we introduce the real angle  $\theta$  by means of

$$\cos \theta \stackrel{\Delta}{=} \frac{1 + \rho^2 - \frac{1 - \rho^2}{\lambda}}{2\rho}$$
(55)

and write the solution in the form (40). Application of the boundary conditions then yields the pair of equations

$$[1-\rho \cos\theta]K_{1} - [\rho \sin\theta]K_{2} = 0$$
(56)

$$[\cos (N+1)\theta - \rho \cos N\theta]K_1 + [\sin (N+1)\theta - \rho \sin N\theta]K_2 = 0$$
(57)

whose determinant must vanish. Thus  $\theta$  obeys the transcendental equation [c.f.(42)],

$$\sin(N+1)\theta - 2\rho \sin N\theta + \rho^2 \sin(N-1)\theta = 0$$
(58)

Since the values  $\theta=0$ ,  $\pi$  do not permit a non-trivial solution to the bvp, the roots of (58) again lie in the open interval  $(0,\pi)$ . Since it has not been possible to solve (58) analytically, the following remarks allow a numerical solution to be easily achieved:

To show that there are exactly N roots between 0 and  $\pi$  we first write (58) as

$$[(1+\rho^{2})\cos\theta - 2\rho] \sin N\theta + [(1-\rho^{2})\sin \theta] \cos N\theta = 0$$
 (59)

then introduce the function  $\varphi\left( \boldsymbol{\vartheta}\right)$  by means of

$$\sin \phi(\theta) \stackrel{\Delta}{=} \frac{(1-\rho^2) \sin \theta}{1+\rho^2 - 2\rho \cos \theta}$$
(60)

$$\cos \phi(\theta) \stackrel{\Delta}{=} \frac{(1+\rho^2)\cos\theta - 2\rho}{1 + \rho^2 - 2\rho \cos\theta}$$
(61)

so that (59) becomes

sin 
$$[N\theta + \phi(\theta)] = 0$$
 (62)  
[N.B. as  $\rho + 0$ ,  $\phi(\theta) + \theta$  and (62) + (42)]

Since

$$\frac{d\phi}{d\theta} = \frac{1-\rho^2}{1+\rho^2 - 2\rho\cos\theta} > 0$$
 (63)

we have

$$\phi(\pi) = \phi(0) + \int_{0}^{\pi} \frac{1 - \rho^{2}}{1 + \rho^{2} - 2\rho \cos\theta} d\theta = \pi$$
(64)

Since (62) represents a modulated sinusoid whose total phase increases monotonically from 0 at  $\theta=0$  to  $(N+1)\pi$  at  $\theta=\pi$ , it has exactly N distinct zero crossings in the open interval  $(0,\pi)$ .

Further, since 
$$\frac{d^2\phi}{d\theta^2} < 0$$
 in  $(0,\pi)$  and  $\frac{d\phi}{d\theta} = 1$  at  $\cos\theta = \rho$ , the function

sin  $[N\theta + \phi(\theta)]$  oscillates more rapidly than does  $\sin(N+1)\theta$  in the domain  $0 < \theta < \cos^{-1}\rho$  and more slowly (but still more rapidly than does sin  $N\theta$ ) in the domain  $\cos^{-1}\rho < \theta < \pi$ . Using these observations concerning the spacing of the roots, it is an easy matter to accurately locate the roots by means of a Newton-Raphson method.

Denoting the roots by  $\theta_1, \ldots, \theta_N$  the eigenvalues [c.f.(55)] are given by

$$\frac{1-\rho}{1+\rho} < \lambda_n = \frac{1-\rho^2}{1+\rho^2 - 2\rho\cos\theta_n} < \frac{1+\rho}{1-\rho} \quad (n=1,\ldots,N)$$
(65)

and the detection probability by (30).

Note that expansion of (65) into a power series in  $\rho$  yields a first order approximation which is identical to (44) - a useful check.

**III.** RESULTS

Figures 1-7 present detection probability versus per pulse average signal-to-noise ratio in dB for false alarm probability of  $10^{-6}$  and N=2,4, 6,10,15,20,30 pulses respectively.

The dashed curve is the Marcum result which was obtained by employing (11) in (10), performing the Laplace inversion and integrating the resulting pdf over the threshold; this leads to

$$P_{D} = P_{FA} + e^{-(V_{T} + N\chi)} \sum_{n=1}^{\infty} \frac{(N\chi)^{n}}{n!} \sum_{k=N}^{N-1+n} \frac{V_{T}^{k}}{k!}$$
(66)

The closest dark curve is the Swerling case 2 result (p=0) which was obtained by replacing V<sub>T</sub> by V<sub>T</sub>/(1+ $\chi$ ) in (29).

The next (lighter) curves are the results for  $\rho=0.40$ , 0.60, 0.80, 0.90, 0.95, 0.99 respectively.

The last (dark) curve is the Swerling case 1 result ( $\rho=1$ ) which was obtained by performing the inverse Laplace transform of (18) and integrating the resulting pdf over the threshold; this leads to

$$P_{D} = P_{FA} + \frac{1}{1+N\chi} e^{-V_{T}} \sum_{n=1}^{\infty} \left(\frac{N\chi}{1+N\chi}\right)^{n} \sum_{k=N}^{N-1+n} \frac{V_{T}}{k!}$$
(67)

For N=20, the numerical instability exhibited at small values of  $\chi$  when  $\rho = 0.99$  is caused by the small eigenvalues [c.f.(65)] which cluster together; i.e., by indexing the eigenvalues according to size, it may be seen that the individual terms in (30) alternate in sign and increase in magnitude as the difference  $\chi\lambda_n - \chi\lambda_k$  becomes smaller. This instability becomes more pronounced and occurs at even smaller values of  $\rho$  as N increases (it occurs at  $\rho = 0.95$  when N = 30). In order to present a clean figure when N = 30, the curve for  $\rho = 0.99$  has been deleted from figure 7.

The figures show that as  $\rho$  increases from zero to unity, more per pulse average signal-to-noise is required to achieve the same  $P_D$ . This increase, however, is smaller than one would intuitively expect; for example, increasing  $\rho$  from zero to one-half requires an additional increase of less than a dB in signal-to-noise at any  $0.50 \leq P_D \leq 0.99$  and N > 1 while, for example at  $P_D = 0.95$ , increasing  $\rho$  from zero to one requires an increase of 5.2 dB for N = 2 or 10.6 dB for N = 30.



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Figure 1.



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Figure 2.

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Figure 3.







Figure 5.



Figure 6.



Figure 7.

PART TWO. FLUCTUATION LOSS

# 1. INTRODUCTION

The exact fluctuation loss for a Gauss-Markov signal is determined as a function of the number of integrated pulses, the correlation between contiguous pulses and the specified detection and false alarm probabilities. This exact loss is compared to Barton's<sup>7</sup> approximation which calculates fluctuation loss by assuming a Swerling case 2 target and a reduced number of "independent" pulses. Two expressions for the number ( $N_e$  or  $N_I$ ) of equivalent independent pulses are derived. The domain of validity of the approximations is established.

<sup>7</sup>See references at end of text.

### 11. BACKGROUND

Figures 1-7 of part 1 show that at fixed  $P_{FA}$  and N a greater per pulse average signal-to-noise ratio (detectability factor) is required to achieve a specified  $P_D$  ( $\geq$  0.50) when the target fluctuates than when it does not. This required increase in detectability factor is called the fluctuation loss. For fixed  $P_D$  and  $P_{FA}$  it depends on the number of detected pulses and their correlation and will be denoted by  $L_{r}(N, p)$ .

Table I presents the detectability factors  $D_0(N)$  for a non-fluctuating target,  $D_2(N)$  for a Swerling case 2 (p=0) target and  $D_1(N)$  for a Swerling case 1 (p=1) target at various  $P_D$ 's of interest; also shown is the normalized threshold voltage required to achieve the common  $P_{FA}$  of 10<sup>-6</sup>. This table has been calculated from eqs. (66), (29) with  $V_T$  replaced by  $V_T/(1+\chi)$ ; and (67) of part J.

For the  $P_n$ 's of interest the fact that for N>1,

 $L_{f}(N,1) \stackrel{\Delta}{=} D_{1}(N) - D_{0}(N) > D_{2}(N) - D_{0}(N) \stackrel{\Delta}{=} L_{f}(N,0)$ 

is well known to radar systems engineers from the previously published work of Marcum and Swerling.

Whenever a surveillance radar encounters the situation  $\rho \approx 1$ , the target may lie in a deep fade for all N pulses received on a given scan, thus causing a detection to be missed. In order to avoid this situation some radar systems decorrelate the N signal returns by transmitting pulse-to-pulse frequency diverse waveforms. This certainly decreases the fluctuation loss from  $L_f(N,p)$  to  $L_f(N,0)$ (a considerable reduction at high P<sub>D</sub> and N) but involves increased system complexity and cost.

Fluctuation loss depends not only on  $P_{FA}$  and  $P_D$  but also on the fluctuation model and the correlation model. For  $P_{FA} = 10^{-6}$ , a Gaussian fluctuation model (Rayleigh envelope), and a first order Markov correlation model the analysis of part 1 yields the appropriate detectability factors for each N and  $\rho$ . By subtracting the detectability factor  $D_O(N)$ , we obtain the fluctuation losses,  $L_f(N, \gamma)$ , of table II.

It is observed that the fluctuation loss increases monotonically with  $P_D$ and, except for the singular case p = 1, decreases monotonically with N. Of greater significance however is the nature of the monotonic increase in fluctuation loss with p and in particular its rate of increase as p approaches 1.

			_				
ſ	N	P_=0.50	P0.70	P0.90	P_=0.95	P_=0.99	V <sub>T</sub> (dB)
		D <sub>0</sub> (N)					
	1	11.24	12.09	13.18	13.66	14.50	11.40
	2	8.80	9.61	10.65	11.11	11.91	12.22
	4	6.49	7.26	8.25	8.68	9.44	13.29
	6	5.21	5.94	6.90	7.32	8.06	14.05
	10	3.65	4.35	5.27	5.67	6.38	15.15
	15	<b>~ 4</b> 7	3.14	4.03	4.41	5.10	16.13
	20	1.66	2.31	3.17	3.55	4.21	16.89
ļ	30	0.54	1.18	2.00	2.37	3.01	18.03
		D <sub>2</sub> (N)					
1	1	12.77	15.77	21.14	24.29	31.37	
i	2	9.52	11.53	14.83	16.62	20.47	
:	4	6.83	8.28	10.51	11.65	13.97	
· !	6	5 4	6.65	8.49	9.41	11.22	
i	10	3.7	4.80	6.29	7.02	8.40	
	15	2.5	3.46	4.75	5.36	6.52	
	20	1.71	2.55	3.73	4.29	5.32	
	30	0.58	1.34	2.39	2.88	3.78	
		D <sub>1</sub> (N)					
	1	12.77	15.77	21.14	24.29	31.37	
	2	10.33	13.32	18.69	21.84	28.91	
	4	8.03	11.00	16.36	19.51	26.60	ļ
	6	6.74	9.71	15.07	18.21	25.30	
	10	5.19	8.15	13,50	16.64	23.73	
	15	4.00	6.96	12.31	15.45	22.54	
	20	3.19	6.14	11.49	14.63	21.72	
	30	2.08	5.03	10.37	131	20.57	

TABLE I

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DETECTABLILITY FACTORS AT  $P_{FA} = 10^{-6}$ 

PD	N	ρ <b>=</b> 0.00	ρ <b>=0.4</b> 0	ρ <del>=</del> 0.60	ρ <b>=</b> 0.80	p <b>=0.90</b>	£=0.95	p <b>=0.99</b>	ρ=1.0
	2	0.72	0.85	1.04	1.28	1.41	1.48	1.51	1.53
	6	0.21	0.31	0.47	0.81	1.12	1.32	1.47	1.53
0.50	10	0.12	0.19	0.29	0.58	0.92	1.19	1.43	1.54
	15	0.07	0.11	0.19	0.42	0.73	1.04	1.39	1.53
	20	0.05	0.08	0.14	0.32	0.60	0.93	1.34	1.53
	30	0.04	0.06	0.10	0.22	0.44	0.65	1.26	1.54
		1							
	2	1.92	2.13	2.44	2.95	3.31	3.51	3.64	3.71
	4	1.02	1.24	1.58	2.26	2.86	3.22	3.57	3.74
0.70	0		0.90	1.21	1.86	2.54	3.06	3.53	3.//
0.70	10		0.59	0.83	1.40	2.08	2.72	3.44	3.00
	20	0.32	0.42	0.01	1.00	1./1	2.35	2.35	3.02 7 93
	30	0.16	0.22	0.33	0.68	1.12	1 74	3.04	3,85
			0.11	0.33					
	2	4.18	4.49	4.93	5.79	6.59	7.21	7.88	8.04
	4	2.26	2.59	3.13	4.23	5.36	6.36	7.73	8.11
	6	1.59	1.90	2.39	.3.48	4.66	5.79	7.65	8.17
0.90	10	1.02	1.27	1.68	2.65	3.78	5.00	7.44	8.23
	15	0.72	0.91	1.25	2.11	3.15	4.33	-7.24	8.28
	20	0.56	0.72	1.01	1.73	2.72	3.87	6.98	8.32
	30	0.39	0.52	0.73	1.32	2.17	3.23	6.58	8.37
						l			
ł	2	5.51	5.84	6.33	7.33	8.32	9.18	10.52	10.73
	4	2.97	3.36	3.96	5.24	6.60	7.86	10.32	10.83
0.05	10	2.09	2.44	3.03	4.28	2.68	7.08	10.22	10.89
0.95	15	0.05	1 10	1 60	2 59	3.84	5 26	9.35	11 04
	120	0.52	0.94	1 29	2.50	3.34	4.69	9.32	11 08
	30	0.51	0.67	0.94	1.65	2.66	3.92	8.79	11.14
	+		1	1	1	1		1	
	2	8.56	8.91	9.47	10.61	11.83	13.04	16.70	17.00
1	4	4.53	4.98	5.70	7.22	8.88	10.56	16.38	17.16
	6	3.16	3.60	4.32	5.84	7.56	9.34	16.22	17.24
0.99	10	2.02	2.40	3.02	4.44	6.13	7.93	15.77	17.35
1	15	1.42	1.74	2.27	3.52	5.11	6.87	15.34	17.44
}	20	, 1.11	1.38	1.86	3.02	4.46	6.16	14.80	17.51
· ·	30	0.77	0.99	1.38	2.31	3.61	5.10	13.94	17.56
1	<u>ا</u>		1	L		L	!	<u> </u>	<u> </u>

		TABLE II		
$L_{z}(N,\rho)$	-	FLUCTUFTION	LOSS	(dB)

For example, suppose all we know concerning the target returns is that they are not independent; if we make the usual assumption that they are fully correlated (i.e.,  $\rho$ =1) and integrate 30 pulses to achieve required P<sub>D</sub>'s of say 0.90, 0.95, 0.99 respectively, then a system which employs pulse-to-pulse frequency divorsity could recover the fluctuation losses  $[L_f(30,1) - L_f(30,0)]$  of respectively, 8.0, 10.6, and 17.8 dB. If however  $\rho$  is not in fact equal to 1.0 but is instead 0.9 then the recoverable fluctuation losses are only 1.8, 2.2, 2.8 dB respectively (at  $\rho$ =0.8 the recoverable losses are even smaller, viz., 0.9, 1.1 and 1.5 dB). Thus the trade-off of system complexity versus recoverable fluctuation loss on signal correlation - a dependence established in part 1 of this paper.

In summary, the usually employed assumption,  $\rho=1$ , should not be facilely invoked simply because the signal returns are known to lack total independence. Doing so results in an extremely optimistic estimate of the bonefits which can be achieved by employing frequency diversity.

### III. ANALYSIS

e next consider two approximations which greatly simplify the determination of fluctuation loss for any correlation model. Earton<sup>7</sup> introduces the concept of the number of equivalent "independent" pulses; this is defined by

$$N_{e} \stackrel{\Delta}{=} Min. \left[ N, 1 + \frac{t_{o}}{t_{c}} \right]$$
(67)

In (67)  $t_0$  is the observation time and  $t_c$  is the correlation time (which Barton defines for any correlation model as the reciprocal of the effective noise bandwidth of the two sided fluctuation spectrum). The definition, (67) is assumed to be valid independent of  $P_D$  or  $P_{FA}$ . Barton's two conjectures state;

$$L_{f}(N,\rho) \approx L_{f}(N_{e},0)$$
 (FIRST CONJECTURE) (68)

$$L_{f}(N_{e}, 0) \approx \frac{1}{N_{e}} L_{f}(1, p)$$
 (SECOND CONJECTURE) (69)

The validity of these conjectures will now be tested against the exact results available from our analysis of the first order Markov correlation model. We begin by calculating Barton's  $N_{a}$ .

In (67) we interpret t as the correlation time of the process at the detec-

See references at end of text.

tor output. Thus letting x and y denote i.i.d. stationary Gaussians which represent the inphase and quadrature components of signal at the detector input, the detector output becomes

$$z(c) \stackrel{\Delta}{=} x^{2}(t) + y^{2}(t)$$
 (70)

Its average (denoted by an overbar) is

$$\overline{z}(t) = \overline{x^2}(t) + \overline{y^2}(t) = 2R_x(0)$$
 (71)

where [c.f.(31)]

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$$R_{x}(\tau) = R_{x}(0) e^{-\nu |\tau|}$$
 (72)

is the autocorrelation of the inphase (or quadrature) Markov process.

The covariance function at the detector output is

$$R_{z-z}(\tau) = [z(t)-\overline{z}(t)][z(t+\tau)-\overline{z}(t+\tau)] = 2[R_{x}^{2}(\tau)-R_{x}^{2}(0)]$$
(73)

which, since x(t) is a Gaussian process, yields

$$R_{z-\overline{z}}(\tau) = 4R_{x}^{2}(\tau)$$
 (74)

The "two sided fluctuation spectrum" is interpreted as the Fourier transform of the covariance function, i.e.,

$$S_{x}(f) = \int_{-\infty}^{\infty} 4R_{x}^{2}(0)e^{-2\nu|\tau| - i2\pi f\tau} d\tau = 4R_{x}^{2}(0)\frac{4\nu}{4\nu^{2} + (2\pi f)^{2}}$$
(75)

This has an effective noise bandwidth,  $\mathbf{f}_{_{\mathrm{B}}},$  given by

$$f_{B} \stackrel{\Delta}{=} \frac{\int_{-\infty}^{\infty} S_{x}(f) df}{S_{x}(0)} = \frac{R_{z-\overline{z}}(C)}{S_{x}(0)} = v$$
(76)

Since the observation time is (N-1)T and [c.f.(31)]  $\rho = e^{-\nabla T}$  the correlation time t<sub>c</sub>, is  $T/\ln \frac{1}{\rho}$ . Thus

$$N_{e}(N,\rho) = Min. [N, 1+(N-1)ln \frac{1}{\rho}]$$
(77)

A more common definition of the number of independent integrated pulses is the increase in the ratio of squared mean to variance when N pulses are summed, i.e., define

$$v_{N} \stackrel{\Delta}{=} \sum_{1}^{N} (x_{n}^{2} + y_{n}^{2})$$
 (78)

then

$$\mathbf{N}_{\mathbf{I}} \stackrel{\Delta}{=} \frac{\overline{\mathbf{v}}_{\mathbf{N}}^{2} / \sigma^{2} \mathbf{v}_{\mathbf{N}}}{\overline{\mathbf{v}}_{\mathbf{I}}^{2} / \sigma^{2} \mathbf{v}_{\mathbf{I}}}$$
(79)

It is easily shown that

$$N_{I} = N^{2} \frac{\sigma^{2} V_{I}}{\sigma^{2} V_{N}} = \frac{N^{2}}{N + 2 \sum_{m=1}^{N-1} (N-m) R_{X}^{2}(m) / R_{X}^{2}(0)}$$
(80)

Since for the first order Markov model

$$R_{x}(m) = R_{x}(0) \rho^{|m|}$$
 (81)

eq. (80) yields

$$N_{I}(N,\rho) = \frac{N}{1 + \frac{2\rho^{2}}{1 - \rho^{2}} \left(1 - \frac{1}{N} - \frac{1 - \rho^{2N}}{1 - \rho^{2}}\right)}$$
(82)

Note that  $N_e \ge N_I$  except at N = 1 or  $\rho = 1$ , or  $\rho = 0$  where equality obtains. Table III compares  $N_e$  with  $N_T$  for the N and  $\rho$  of table II.

Figures 8-12 are plots of the exact fluctuation losses of table II versus  $N_e$ ,

the number of equivalent independent pulses as defined by Barton. The solid curves represent  $L_f(N_e, 0)$ , i.e., the left side of (69). Aside from a negligible positive curvature which is most pronounced at low  $N_e$  and high  $F_D$ , these do (as Barton conjectures) exhibit a linear decrease with  $N_e$ . The dashed curves represent the right side of (69), i.e.,  $\frac{1}{N_e}L_f(1,\rho)$ . We see that the solid and

dashed curves are within 1/2 dB of each other; further we note that

$$L_{f}(N_{e}, C) < \frac{1}{N_{e}} L_{f}(1, 0)$$
 (P<sub>D</sub> = 0.50) (83)

$$L_{f}(N_{e},0) > \frac{1}{N_{e}} L_{f}(1,p)$$
 (P<sub>D</sub> = 0.70, 0.90, 0.95, 0.99) (84)

In Figures 10-12 the upper solid curve is the continuation of  $L_f(N_e, 0)$  into the region 10  $\leq N_e \leq$  30 (divide the ordinate by 10 and multiply the abscissa by 10). The dashed curve,  $\frac{1}{N_e} L_f(1, \rho)$ , is its own continuation.

The symbols  $\cdot$ , x,  $\bigcirc$ ,  $\triangle$ ,  $\Box$ ,  $\nabla$ , identify values of  $L_f(N,\rho)$  for  $\rho = 0.40$ , 0.60, 0.80, 0.90, 0.95, 0.99 respectively.

We see that for  $P_D = 0.50$  and 0.70, Barton's first conjecture underestimates the exact fluctuation loss by a fraction of a dB. For  $P_D = 0.90$ , 0.95, 0.99 the first conjecture overestimates the exact fluctuation loss at small  $N_e$  (by more than a dB at  $P_D = 0.95$  and by several dB at  $P_D = 0.99$ ) and underestimates the exact fluctuation loss by less than a dB at large  $N_e$ . Since the fluctuation loss monotonically decreases with  $N_e$ , Barton's first conjecture gives a result which is conservative when the losses are large, and optimistic when the losses are small.

If one were to use  $N_I$  instead of  $N_e$  to represent the number of independent pulses, the symbols representing  $L_f(N,C)$  would shift to the left so that the approximations would become even more conservative at  $P_D = 0.90$ , 0.95, 0.99 when the losses are large. Thus Barton's  $N_e$  is the preferred approximation for  $P_D \ge 0.90$  (large losses), while the approximation using  $N_I$  is to be preferred for  $P_D \le 0.70$  (small losses).

٥	N=	2	N=	4	N-	6	N=	10	N=	•15	N=	20	N=	•30
	N.	NI	N	NI	N	NI	N	NI	N	NI	N	NI	N.	NI
0.40	1.92	1.72	3.75	3.16	5.58	4.60	9.25	7.49	13.83	11.11	18.41	14.72	27.57	21.96
0.60	1.51	1.47	2.53	2.36	3.55	3.27	5,60	5.13	8.15	7.47	10.71	9.82	15.81	14.52
0.80	1.22	1.22	1.67	1.60	2.12	1.99	3.01	2.79	4.12	3.85	5.24	4.92	7.47	7.10
0.90	1.10	1.10	1.32	1.28	1.53	1.44	1.95	1.79	2.48	2.25	3.00	2.73	4.06	3.73
0.95	1.05	1.05	1.15	1.13	1.26	1.21	1.46	1.36	1.72	1.57	1.97	1.78	2.49	2.23
0.99	1.01	1.01	1.03	1.03	1.05	1.04	1.09	1.07	1.14	1.10	1.19	1.14	1.29	1.21

EQUIVALENT NUMBER OF INDEPENDENT PULSES

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TABLE III

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Figure 8.



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Figure 9.





Figure 10.



NUMBER OF INDEPENDENT PULSES

Figure 11.



NUMBER OF INDEPENDENT PULSES

Figure 12.

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