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Radiation Effects on AlGaAs/GaAs Solar Cells Using 0.9 - 3.0 MeV Protons and 1.0 - 1.4 MeV Electrons

L.F. LOWE J.R. CAPPELLI L.W. JAMES R.L. MOON

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Unclassified SECURITY CLASSIFICATION OF THIS PAGE (When Date Entered) READ INSTRUCTIONS BEFORE COMPLETING FORM REPORT DOCUMENTATION PAGE REPORT NUMBER 2. GOVT ACCESSION RECIPIENT'S CATALOG NUMBER 9 Technical RADC-TR-77-332 rept TITLE (and Subtitle) RADIATION EFFECTS ON AlGaAs/GaAs SOLAR CELLS, USING Ø. 9 - 3.Ø MeV PROTONS AND 1.Ø - 1.4 MeV ELECTRONS 1 Jan, -31 Dec 2076 RIM THOR(+) 8. CONTRACT OR GRANT NUMBER(S) L. F. Lowe, J. R. Cappelli, L. W. James*, R. L. Moon T. PROJECT, TASK PERFORMING ORGANIZATION NAME AND ADDRESS 0. PROGRAM ELEMEN Deputy for Electronic Technology (RADC/ESR) Hanscom AFB 62702F Massachusetts 01731 46002001 1. CONTROLLING OFFICE NAME AND ADDRESS 2. REPORT DAT Deputy for Electronic Technology (RADC/ESR September 1977 Hanscom AFB 33 Massachusetts 01731 15. SECURITY CLASS. (of this report) MONITORING AGENCY NAME & ADDRESS(if different from Controlling Office) Unclassified 15. DECLASSIFICATION/DOWNGRADING 6. DISTRIBUTION STATEMENT (of this Report) Approved for public release; distribution unlimited. 17. DISTRIBUTION STATEMENT (of the abstract entered in Block 20, if different from Report) 18. SUPPLEMENTARY NOTES *Varian Associates, Palo Alto, California 19. KEY WORDS (Continue on reverse side if necessary and identify by block number) (1) Solar cells (2) AlGaAs **Radiation effects** (3) (4) Space environment 20. ABSTRACT (Continue on reverse side if necessary and identify by block number) Aluminum gallium arsenide so ar cells were irradiated with 1.0 and 1.4 MeV electrons, and with 0.9 and 3.0 MeV protons to determine radiation sensitivity. Electron fluences ranged from 1×10^{14} to 3×10^{10} electrons on ², and proton fluences from 5×10^{40} to 2.7×10^{12} cm⁻². A solar simulator and a tungsten lamp were used to evaluate changes in the current-voltage characteristic curves. In most cases, AlGaAs solar cells showed a greater resistance to radiation than silicon cells. pagen 10 to the 12th more DD 1 JAN 73 1473 EDITION OF I NOV 65 IS OBSOLETE Unclassified SECURITY CLASSIFICATION OF THIS PAGE (When Date Entered) 10 to the 14th power 10 to the 16th power 10 to the 10th power /sgcm 309050

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Preface

The authors thank Roger Little and John Menucci of Simulation Physics, Inc. for their assistance in the solar simulator and tungsten lamp measurements.

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Radiation Effects on A1GaAs/GaAs Solar Cells Using 0.9 – 3.0 MeV Protons and 1.0 – 1.4 MeV Electrons

1. INTRODUCTION

In order to evaluate the applicability of AlGaAs for use as a solar cell material, information is needed on its sensitivity to the radiation that they would be exposed to in space. To this end AlGaAs solar cells were exposed to 1.0 and 1.4 MeV electrons, and 0.9 and 3.0 MeV protons. The electron fluences ranged from 1×10^{14} to 3×10^{16} electrons cm⁻², and the proton fluences ranged from 5×10^{10} to 2.7×10^{12} protons cm⁻².

2. EXPERIMENTAL

Cells were fabricated by conventional LPE processes, where the p-type AlGaAs (AlAs - 0.80) was grown on an n-type GaAs(111)B substrate. During growth, the p-dopant diffused into the n-layer to form a p-n junction approximately 0.3 to 0.8 micron deep in the GaAs. Doping densities of the p-AlGaAs + n-GaAs substrate were 3×10^{18} cm⁻³ and 8×10^{17} cm⁻³, respectively. A p-type GaAs layer was grown on top of the AlGaAs in order to facilitate contacting, and later removed by a selective etch using the contact grid as a mask. Cell area was 1.613 cm². These cells were not optimized for air mass zero (AMO) performance,

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since the AlGaAs were generally too thick, reducing the blue response. SiO₂ deposited to a thickness of 750 Å was used as an antireflection coating. AMO efficiencies were 10-12 percent. Proton irradiation experiments were conducted at the Van de Graaff facility at RADC/ET, Massachusetts. After each irradiation to a given fluence, the I-V curve was measured using an AMO solar simulator and a tungsten lamp as light source. Proton energies of 0.9 and 3 MeV were used, up to maximum fluences of $5 \times 10^{11} \text{ p/cm}^2$ and $3 \times 10^{12} \text{ p/cm}^2$, respectively.

Electron bombardment was carried out at the Dynamitron accelerator at RADC/ET, Massachusetts. Electron energies of 1 and 1.4 MeV were used up to maximum fluence of 3×10^{16} e/cm² and 3×10^{15} e/cm², respectively. AMO data were taken using Spectrolab solar simulators at these locations. All measurements were taken at 25° C.

After initial characterization, all samples were evaluated for radiation induced changes in their current-voltage (i-v) characteristic curves using both a solar simulator at AMO and also a tungsten lamp. After each tungsten run, the tungsten intensity was adjusted so as to bring the short circuit current back to what it was before any irradiation (I_{sco}). In addition after the final irradiation, the spectral response was determined for each cell. The individual irradiation procedures were as follows.

3. ELECTRONS

The electron irradiations were made using a 1.5 MeV Dynamitron electron accelerator. Two samples each were mounted on either side of the aperture of a Faraday cup so that all fluences were measured directly. The cells were open circuited during the irradiation and the exposures were carried out at room temperature in air. The electron flux was approximately 10^{12} electrons cm⁻² s⁻¹. Two energies were used, 1.0 and 1.4 MeV.

4. PROTONS

A 3.0 MeV Van de Graaff accelerator was used for the proton exposures. Because of the short range of protons these experiments were carried out in vacuum, one cell at a time. Each cell was mounted in the center of a Faraday cup which was part of the vacuum system. A collimator, electrically isolated from the Faraday cup, defined the exposure area. The Van de Graaff is equipped with a beam sweeping device which leads to an exposure uniformity of better than 10 percent. Proton flux levels were 10^{10} protons cm⁻² s⁻¹, and runs were made open circuited at 0.9 and 3.0 MeV.

5. RESULTS

Figures 1-8 give the individual I-V curves as a function of fluence using the solar simulator at AMO. Figures 9-16 show tungsten lamp data for the same exposures. Figures 17-21 illustrate the effect of increasing the tungsten lamp intensity so as to bring the cell output current back to its preirradiation value (the I_{sco} case). Contact problems were encountered with 3 cells. Pertinent data from Figures 1-16 are given in Tables 1-8, including the incomplete data for the three cells with contact problems. The output power data is shown in Figures 22-26. After two weeks, the cells were checked for annealing. None was observed. Spectral response data are shown in Figures 27-30.

6. SUMMARY AND DISCUSSION

Solar cells of p-AlGaAs/n-GaAs were subjected to 0.9 and 3 MeV proton radiation, and 1 and 1.4 MeV electron radiation. Values of critical fluence \emptyset_c at which the output power was reduced by 25 percent were determined. As expected protons with energies of 0.9 MeV caused the most severe degradation, and a $\emptyset_c = 4 \times 10^{10} \text{ p/cm}^2$ was determined. Fortunately because of the short range of low energy protons, this type of radiation can be protected against by cover slips. For 3 MeV protons, $\emptyset_c = 5 \times 10^{11} \text{ p/cm}^2$.

Critical fluences for 1.4 MéV electrons were $7-9 \times 10^{14} \text{ e/cm}^2$. A total of five cells were irradiated at electron energies of 1 MeV. \emptyset_{C} varied between 7×10^{14} and $7 \times 10^{15} \text{ e/cm}^2$. At fluences $< 10^{15} \text{ e/cm}^2$, the relative cell parameters followed the characteristic radiation equation. At fluences greater than this the relative parameters varied linearly with $\ln \emptyset$, as did the relative diffusion length.

These results show that the radiation resistance of AlGaAs/GaAs solar cells is equal to or better than that observed in conventional and violet Si cells for all the radiation fluxes investigated here, except 0.9 MeV protons. In the case of 1 MeV electrons, the value of $\emptyset_c = 7 \times 10^{15} \text{ e/cm}^2$ is believed to be the highest observed for conventional cells. The spread in \emptyset_c we observed is thought to be caused by variations in junction depth, which can be optimized. So even for 1 MeV electrons, cells of AlGaAs/GaAs are superior to those of Si.

Computer calculations show that the variation in I_{sc} as a function of irradiation level is sensitive to junction depth. Thus a shallow junction solar cell should be more radiation-resistant than the experimental data shown here, since I_{sc} does not change as rapidly with decreasing diffusion length as in a deep junction device.

Calculations also demonstrated that the AMO efficiency increases with AlAs concentration in the AlGaAs for a given thickness. This will allow operation in

concentrated sunlight in the space environment because the sheet resistance of the AlGaAs can be reduced without substantially degrading performance through optical absorption of high energy protons in the contact layer.

Combining the radiation resistance characteristics with the ability to obtain AMO efficiencies of 16 percent, and potentially 18 percent, makes the AlGaAs/GaAs solar cells attractive for space applications.





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actor	г	0, 698	0.711	0.709	0.723	0.736	0.711	0. 705
Fill F	s s	0.708	0.729	0.743	0.741	0.752	0.737	0.707
(mm)	гч	8.1	8.1	7.8	7.3	6.4	5.4	4.3
Pmax	s s	22.8	22.6	22.0	20.3	17.9	15.7	11.6
(mm)	г	11.6	11.4	11.0	10.1	8.7	7.6	6.1
P _{ep} (S S	32.2	31.0	29.6	27.4	23.8	21.3	16.4
volts)	гч	0, 89	0.88	0.87	0.85	0.81	0.78	0.74
V _{oc} (s s	0.96	0.94	0.91	0.88	0.85	0.82	0.77
nA)	Tungsten Lamp	13.0	13.0	12.6	11.9	10.7	9.7	8.2
I _{sc} (r	Solar Simulator	33.5	33.0	32.5	31.1	28.0	26.0	21.3
	e ⁻ cm ⁻²	0	1.0×10^{14}	4.0×10^{14}	1.4×10^{15}	5.0×10^{15}	1.0×10^{16}	3.0×10^{16}

Table 2. Cell No. 9-23-75B

	Isc	(mA)	V _{oc} (volts)	Pep	(mm)	Pmax	(mm)	Fill Fa	ictor
e ⁻ cm ⁻²	Solar Simulator	Tungsten Lamp	s s	гч	s s	нJ	w w	нJ	s s	нJ
0	41.5	16.5	0.99	0.92	41.1	15.2	30.0	10.7	0.730	0.704
1.0×10^{14}	40.5	16.3	0.96	06.0	38.9	14.7	28.9	10.7	0.743	0.728
4.0×10^{14}	39.5	15.6	0.91	0.88	35.9	13.7	24.1	10.0	0.671	0.730
1.4×10^{15}	35.5	14.5	:	0.82	1	11.9	1	;	1	1
2.5×10^{16}	19.5	1	0.78	:	15.4	;	11.3	:	0.734	:

13

Table 1. Cell No. 10-7-75B

Table 3. Cell No. 1-26-76B

	I _{sc} (m	(Ar	Voc	(volts)	Pep ((mm)	Pmax	(mm)	Fill F	actor
e cm -2	Solar Simulator	Tungsten Lamp	s s	нJ	s so	타니	w w	нч	s s	нJ
0	32.6	15.9	1.0	0.97	32.6	15.4	25.8	12.1	0.791	0.786
5.0×10^{14}	28.5	1	0.90	1	25.7	1	20.5	1	0.798	1
8.0×10^{14}	26.6	12.7	0.89	0.86	23.7	10.9	19.0	8.7	0.802	0. 798
1.6×10^{15}	22.6	;	0.85	1	19.2	;	15.1	;	0.786	1
3.0×10^{15}	18.4	9.2	0.83	0.81	15.3	7.5	11.9	5.7	0.778	0.760

Table 4. Cell No. 1-22-76B

	I _{sc} (m	(Ar	V _{oc} (volts)	Pep	(mm)	Pmax	(mm)	Fill F	actor
e ⁻ cm ⁻²	Solar Simulator	Tungsten Lamp	w w	нч	w w	FЛ	s s	нл	s s	г Г
0	26.5	13.8	0.99	0.96	26.2	13.2	19.8	9.8	0.756	0.742
5.0×10^{14}	23.0	1	0.89	1	20.5	1	16.1	1	0.785	1
8.0×10^{14}	22.0	11.2	0.89	0.86	19.6	9.6	15.4	7.5	0.786	0.781
1.6×10^{15}	19.5	1	0.86	1	16.8	1	13.0	1	0.774	1
3.0×10^{15}	16.7	8.8	0.83	0.81	13.9	7.1	10.7	5.4	0.770	0.761

Table 5. Cell No. 9-15-75C

T				
actor	нј	0.644	1	0.633
Fill F	s s	0.534	0.505	1
(mm)	г	8.5	1	6.9
Pmax	νv	18.9	16.5	;
mw)	нл	13.2	1	10.9
volts) Pep (s s	35.4	32.7	1
	гч	0.91	!	0.81
V _{oc} (s s	0.95	0.89	1
(A)	Tungsten Lamp	14.5	1	13.5
I _{sc} (m	Solar Simulator	37.3	36.7	33.5
	e ⁻ cm ⁻²	0	5.0×10^{10}	9.6 \times 10 ¹⁰

Table 6. Cell No. 2-4-76B

	I _{sc} (m	(Ar	V _{oc} (volts)	P _{ep} ((mm)	Pmax	(mm)	Fill F	actor
.m-2	Solar Simulator	Tungsten Lamp	s s	гı	s s	нл	s s	нJ	w w	гл
0	29.7	15.3	1.0	0.98	29.7	15.0	23.5	11.8	0.791	0.787
× 10 ¹⁰	25.5	1	0.87	:	22.2	1	17.2	1	0.775	;
× 10 ¹⁰	23.7	12.0	0.86	0.81	20.4	9.7	15.5	7.4	0.760	0.763
× 10 ¹¹ -	21.9	:	0.82	1	18.0	1	13.4	1	0.744	
× 10 ¹¹	20.1	10.6	0.80	0.77	16.1	8.2	12.1	6.0	0.752	0.732
× 10 ¹¹	17.0	1	0.75	1	12.8	1	9.3	!	0.727	:

1 61 1

Table 7. Cell No. 1-30-76D

ictor	T L	0.808	1	0.761
Fill Fa	νv	0.811	0.778	;
(mm)	гJ	10.5	1	8.9
Pmax	SS	26.6	23.8	1
(mu	ГЧ	13.0	1	11.7
P _{ep} (r	S S	32.8	30.6	;
volts)	гл	0.93	1	0.89
V _{oc} (s s	0.97	0.94	:
(An	Tungsten Lamp	14.0	1	13.1
I _{sc} (1	Solar Simulator	33.8	32.5	1
	e ⁻ cm ⁻²	0	1.3×10^{11}	2.6×10^{11}

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	I _{sc} (n	(Ar	V _{oc} (volts)	P _{ep} ((mm)	P _{max} ((mm)	Fill Fa	ctor
e ⁻ cm ⁻²	Solar Simulator	Tungsten Lamp	ω w	нIJ	s s	нл	s s	нл	s s	нIJ
0	27.5	13.2	1.0	0.96	27.5	12.7	21.0	9.6	0.764	0.756
1.3×10^{11}	26.4	1	0.92	1	24.3	1	18.4	1	0.757	!
2.5 \times 10 ¹¹	25.5	11.8	0.91	0.86	23.2	10.1	17.5	7.7	0.754	0.762
4.9×10^{11}	23.0	ł	0.83	ļ	19.1	1	14.4	1	0.754	1
9.7×10^{11}	22.1	10.1	0.82	0.78	18.1	7.9	13.6	5.8	0.752	0.734
2.7×10^{12}	20.5	1	0.77	1	15.8	1	11.7	1	0.741	1





Figure 23. Power Curves of Cell No. 1-26-76B



Figure 24. Power Curves of Cell No. 1-22-76B



Figure 25. Power Curves of Cell No. 2-4-76B







METRIC SYSTEM

BASE UNITS:			
Quantity	Unit	SI Symbol	Formule
length	metre	m	
mass	kilogram	kg	***
time	second	5	
electric current	ampere	٨	
thermodynamic temperature	kelvin	ĸ	
amount of substance	mole	mol	***
luminous intensity	candela	cd	
SUPPLEMENTARY UNITS:			
nlane angle	radian	rad	
solid angle	steradian	sr	
DERIVED UNITS:			
Acceleration	metre per second squared		m/s
activity (of a radioactive source)	disintegration per second		(disintegration)/s
angular acceleration	radian per second squared		rad/s
angular velocity	radian per second		rad/s
area	square metre		m
density	kilogram per cubic metre		kg/m
electric canacitance	farad	F	A-s/V
electrical conductance	siemens	S	AN
electric field strength	volt per metre	***	V/m
electric inductance	henry	н	V-s/A
electric notential difference	volt	v	W/A
electric resistance	ohm		VA
electromotive force	volt	v	WA
energy	ioule	1	N-m
entrony	joule ber kelvin		J/K
force	newton	N	kg·m/s
frequency	hertz	Hz	(cycle)/s
illuminance	lux	lx	lm/m
luminance	candela per square metre		cd/m
luminous flux	lumen	lm	cd-sr
magnetic field strength	ampere per metre		A/m
magnetic flux	weber	Wb	V-s
magnetic flux density	tesla	т	Wb/m
magnetomotive force	ampere	٨	
magnetomotive force	watt	w]/s
power	nascal	Pa	N/m
quantity of electricity	coulomb	С	A-s
quantity of heat	ioule	1	N-m
radiant intensity	watt per steradian		W/sr
enocific heat	ioule per kilogram-kelvin		J/kg-K
specific field	nascal	Pa	N/m
thermal conductivity	walt per metre-kelvin		W/m·K
merman conductivity	metre per second		m/s
velocity	nascal-second		Pa-s
viscosity, uynamic	square metre per second		m/s
viscosity, kinematic	volt	v	W/A
voltage	cubic metre		m
volume	cubic metre		(wave)/m
work	joule	Ï	N·m
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