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EXTENDED STUDIES OF A QUADRILATERALIZED SPHERICAL CUBE EARTH DATA BASE

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ABSTRACT

This report presents the results of a study of mathematical formulae and algorithms necessary for the implementation of the Quadrilateralized Spherical Cube Earth Data Base preprocessor function. Topics covered include an improved formulation of the inverse equal-area transformation, f*; the development of a new, exact, equal-area transformation utilizing a polar coordinate system on the cube face; an improved computation of the coefficients of the equal-area transformation, f, and its inverse, f*; a rapid method of obtaining cube face coordinates from Keplerian elements and scanner angle; a new fast-filling algorithm; and the maintenance of multiple and variable resolution data bases.

This report addresses the properties and merits of the newly derived transformation and compares it with a previously derived form. Approximation methods for the equal-area transformation are presented. A revised, fast-filling algorithm is presented, which combines the new method of cube face coordinate computation with a recently suggested pixel sampling scheme. Applications of these methods and algorithms to an initial operational preprocessor and to refined future programs are discussed.

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SECTION 1 - INTRODUCTION

Previous studies (References 1 and 2) have indicated that a quadrilateralized spherical cube dissection of the Earth will provide a uniquely simple and useful set of coordinates in which to store remote sensor data on a global scale. This representation divides the globe into six square grids, allowing an extremely simple and elegant addressing scheme.

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A mapping function (Reference 1) has been developed which permits an equivalent (equal area) mapping of a sphere onto the faces of a cube. This mapping provides an ideal match to constant resolution sensors and allows the simplification of map factor calculations for meteorological applications. Figure 1-1 illustrates the cubic dissection of a sphere, and Figure 1-2 shows a map of the world as projected using the equivalent mapping function.

In studying the feasibility of implementing this quadrilateralized spherical cube Earth data base system, a number of areas were identified which were deemed worthy of further study. In the course of these studies, additional details and problems were brought to light. The initial objectives of this study were

- Improvement of the formulation of the data base mapping functions
- Study of methods of representing satellite motions in terms of the data base coordinate system
- Study of data base Input/Output optimization on the CYBER/175 computer
- Production of maps illustrating various geographic orientations of the data base cube



Figure 1-1. Cubic Dissection of a Sphere by an Inscribed Cube



Additional studies required were

- Study of algorithms required to implement a revised data base filling scheme
- Improvement in the accuracy of the data base mapping functions and the matching of direct and inverse functions

The general purpose in these studies has been twofold: (1) to optimize the existing data base scheme and (2) to assist in the timely development of a preliminary implementation by improving and clarifying previously developed formulae and algorithms. Most sections of this report deal with the maintenance of current data in the data base. The portion of the system which will fulfill these functions has been designated the preprocessor program.

This report presents several new and refined algorithms and mappings aimed at improving the speed and accuracy of an operational preprocessor program.

SECTION 2 - BACKGROUND

2.1 PRIOR RESULTS

A proposed set of algorithms for implementing the quadrilateralized spherical cube data base was presented in another report (Reference 1). That report addressed numerous problems: mapping points from the Earth's surface into the data base, determining serial addresses for data base locations, fast-filling and data retrieval, and determining the ground locations of scanner pixels. The results presented herein are derived from further study of these problems. For background information, Sections 2.2 and 2.3 present the previously derived transformation and method of fast-filling.

2.2 PREVIOUSLY DEVELOPED TRANSFORMATION

Section 2 and Appendix A of Reference 1 present the mathematical development of the mapping function, f, and its inverse, f^* . These functions relate the data base coordinates (x, y) to the rectangular coordinates (ξ, η) of the point on any cube face resulting from the central projection of a datum point on the sphere. For convenience, the mapping functions, f and f^* , are repeated here. The coordinates (ξ, η) are related to the data base coordinates (x, y) via

$$\xi = f(x, y)$$
(2-1)
$$\eta = f(y, x)$$

where the function f, called the direct function, is given by

$$f(\mathbf{x}, \mathbf{y}) = \gamma \mathbf{x} + \frac{(1 - \gamma)}{r_0^2} \mathbf{x}^3$$

+ $xy^2 \left(r_0^2 - x^2\right) \left[\delta + \left(r_0^2 - y^2\right) \sum_{\substack{i \ge 0 \\ j \ge 0}} c_{ij} \mathbf{x}^{2i} \mathbf{y}^{2j} \right]$ (2-2)
+ $x^3 \left(r_0^2 - x^2\right) \left[\omega + \left(r_0^2 - \mathbf{x}^2\right) \sum_{\substack{i \ge 0 \\ j \ge 0}} d_i \mathbf{x}^{2i} \right]$

where

$$\delta = \frac{1}{4r_0^4} \left[-(\mu + 2\gamma) + \sqrt{(\mu^2 - 4\mu\gamma + 4\gamma^2 + 16\sqrt{2\gamma^2})} \right]$$
(2-3a)

$$\omega = \frac{1}{2r_o^4} \left(3 - 2\gamma - \mu - 2r_o^4 \delta \right) = -1.2254 \ 41487 \ 984 \tag{2-3b}$$

$$\gamma = \sqrt{\pi/6} \tag{2-3c}$$

$$\mu = \sqrt{\frac{\sqrt{3}\pi}{2}} \tag{2-3d}$$

$$r_0 = 1/\sqrt{3}$$
 (2-3e)

and the coefficients $\begin{array}{c} c_{ij} \end{array}$ and $\begin{array}{c} d_i \end{array}$ are

$$c_{00} = -2.7217 \quad 05366 \quad 1814 \qquad d_{0} = 1.4833 \quad 12929 \quad 4187$$

$$c_{10} = -5.5842 \quad 16830 \quad 5430 \qquad d_{1} = 1.1199 \quad 72606 \quad 9742$$

$$c_{01} = 2.1711 \quad 17480 \quad 9423 \qquad d_{2} = 6.0515 \quad 38216 \quad 1464$$

$$c_{20} = -3.4578 \quad 62747 \quad 3390$$

$$c_{11} = -6.4160 \quad 15152 \quad 6783$$

$$c_{02} = 1.9736 \quad 26575 \quad 8872$$

The inverse function, f^* , relates (x, y) to (ξ, η) ; that is,

$$x = f^* (\xi, \eta)$$

 $y = f^* (\eta, \xi)$
(2-4)

and is given by

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$$f^{*}(\xi, \eta) = \gamma^{*}\xi + \frac{(1-\gamma^{*})}{r_{o}^{2}}\xi^{3} + \xi\eta^{2}\left(r_{o}^{2}-\xi^{2}\right)\left[\delta^{*}+\delta_{1}^{*}\left(r_{o}^{2}-\xi^{2}\right) + \left(r_{o}^{2}-\eta^{2}\right)\sum_{\substack{i\geq0\\j\geq0}}c_{ij}^{*}\xi^{2i}\eta^{2j}\right] + \xi^{3}\left(r_{o}^{2}-\xi^{2}\right)\left[\omega^{*}+\left(r_{o}^{2}-\xi^{2}\right)\sum_{\substack{i\geq0\\i\geq0}}d_{i}^{*}\xi^{2i}\right] + \xi^{3}\left(r_{o}^{2}-\xi^{2}\right)\left[\omega^{*}+\left(r_{o}^{2}-\xi^{2}\right)\sum_{\substack{i\geq0\\i\geq0}}d_{i}^{*}\xi^{2i}\right]$$

where

$$\mathbf{y}^* = \mathbf{1}/\mathbf{y} \tag{2-6a}$$

$$\mu^* = 1/\mu \tag{2-6b}$$

$$\gamma_1^* = 1/\left(\gamma + r_0^4 \delta\right) \tag{2-6c}$$

$$\mu_1^* = 1/(\mu + 2r_0^4 \delta)$$
 (2-6d)

$$\delta^* = (\mu_1^* - \mu^*)/2r_0^4$$
 (2-6e)

$$\omega^* = (1/2r_0^4) (3 - 2\gamma^* - \mu^* - 2r_0^4 \delta^*)$$
 (2-6f)

$$\delta_{1}^{*} = \frac{1}{r_{o}^{2}} \left[\frac{(\gamma_{1}^{*} - \gamma^{*})}{r_{o}^{4}} - \delta^{*} \right]$$
(2-6g)

and the coefficients c_{ij}^* and d_i^* are

$$c_{00}^{*} = 3.973 \quad 89249 \qquad d_{0}^{*} = 1.811 \quad 28250$$

$$c_{10}^{*} = 6.591 \quad 19476 \qquad d_{1}^{*} = 37.635 \quad 47857$$

$$c_{01}^{*} = -25.368 \quad 92536 \qquad d_{2}^{*} = 63.000 \quad 23655$$

$$c_{20}^{*} = -73.064 \quad 97000$$

$$c_{11}^{*} = 77.381 \quad 61133$$

$$c_{02}^{*} = 21.685 \quad 89623$$

The use of f^* in an algorithm for obtaining the data base coordinates (x, y) of any datum point on the Earth's surface is described in Section 3.3. The algorithm presented in Section 3.3 is the result of the present study to obtain a method of implementing f^* in a computationally efficient manner.

2.3 PREVIOUSLY DEVELOPED FAST-FILLING ALGORITHM

Sections 4 and 5 of Reference 1 present a "fast-filling" method for updating the data base. The method was developed to obtain, as directly as possible, the data base location corresponding to the location of a given scanner data element and to store the value associated with that element in the calculated data base location. Starting with a given satellite location and scanner angle, a series of five steps are necessary to calculate the corresponding data base location.

- 1. Determine the three-dimensional Cartesian coordinates of the point observed by the scanner.
- 2. Determine the face onto which the point is to be mapped
- 3. Project from the Cartesian (a, b, c) system into an intermediate system (ξ , η) as shown in Figure 2-1
- 4. Transform from (ξ, η) into two-dimensional Cartesian coordinates (x, y) on the face of the data base cube
- 5. Convert from face number and (x, y) to a serial location in the data base

This transformation need not be performed for every point, since both the projection onto the Earth's surface and the (ξ, η) to (x, y) transformation may be linearly approximated over a relatively long range. Thus, the complete computation is required only for widely spaced benchmark points, with intermediate points being determined by interpolation.

Input and output are handled, as described in Section 5 of Reference 1, by blocking both the input data and the data base into rectangular arrays. Methods for handling problems such as queuing data base I/O, splitting input blocks across face boundaries, and locating a vertex within an input block were also discussed in Reference 1.

2-5



Figure 2-1. Projection From the Cartesian (a, b, c) System onto the Intermediate (ξ, η) System

Production of the local

In Reference 3, this scheme was compared with an alternative scheme. The alternative scheme follows the above method through step 3; then, instead of storing from the input block directly into the data base, the following steps are performed:

- For each input data block, determine which data base blocks will be completely covered by the input data
- Determine, in the (ξ, η) system, a set of benchmark points corresponding to the corners of the data base blocks to be filled
- Use the above benchmarks to interpolate the (ξ, η) coordinates corresponding to each (x, y) data base location and store the closest point in the input data block into the corresponding serial location

This method offers the advantage of providing a completely filled data base block and eliminates data base read accesses during the update process. After a thorough examination of this method, CSC concluded that the advantages outweigh the slight additional computational cost incurred in finding the appropriate data base blocks. Consequently, some of the work of this report was oriented toward this approach.

SECTION 3 - NEW PROJECTION AND TRANSFORMATION METHODS

3.1 GROUND TRACE EQUATIONS IN KEPLERIAN ELEMENTS

This section presents general relations that yield the geocentric coordinates of the subsatellite point and of any point on the Earth's (spherical) surface lying on an arc scanned by an onboard satellite sensor. These geocentric coordinates are to be computed as an intermediate step in algorithms for computing data base coordinates (Sections 3.2 and 3.3).

The following subsections distinguish between two cases: (a) the computation of the coordinates of the subsatellite point, and (b) the computation of the coordinates of any point on the scan arc. Although the first case could be treated as a special case of the second, its development is separated in order to facilitate computer implementation. The following subsections summarize the results, which are developed in detail in Appendix A.

3.1.1 Subsatellite Point

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It is assumed that the satellite ephemeris is available in the form of Keplerian orbital elements. Expressions can be obtained in terms of these elements for the Earth-centered direction cosines (1, m, n) of the subsatellite point in the equatorial x, y, z coordinate system shown in Figure 3-1.

Since the subsatellite point lies on the geocentric position vector of the satellite (see Figure 3-1), 1, m, and n will be identical with the corresponding direction cosines for the satellite's orbital position. The desired expressions for 1, m, and n are developed in Appendix A to be

$$l = +\cos \Omega \cos (\omega + f) - \sin \Omega \sin (\omega + f) \cos i$$

$$m = +\sin \Omega \cos (\omega + f) + \cos \Omega \sin (\omega + f) \cos i$$
 (3-1)

$$n = +\sin (\omega + f) \sin i$$

3-1



Figure 3-1. Keplerian Orbital Elements and Geographic Coordinates

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The angular arguments are Keplerian elements and are referenced to the true equator and prime (Greenwich) meridian of date coordinate system. The direction cosines (l, m, n) in Equation (3-1) are obtained for the subsatellite point corresponding to any point in the orbit using the true anomaly, f, which reflects the motion of the satellite in its orbit, and the time-dependent node, Ω , which reflects the rotation of the Earth. The time dependence of the node is given by

$$\Omega = \Omega_0 - s \tag{3-2}$$

where Ω_0 is the initial (constant) value of Ω and s is the elapsed Greenwich sidereal time since the instant that $\Omega = \Omega_0$. For high precision, ω , ω , i, and s should be corrected for the Earth's nutation, precession, and polar motion.

3.1.2 Nonzero Scan Angle

It is assumed that the satellite collects data through an onboard sensor which scans the surface of the Earth through an arc of fixed size in such a way that its sweep carries it through the subsatellite point in a direction at right angles to the direction of the satellite's orbital motion. The trace of the scan arc on the surface of the (spherical) Earth is a segment of a great circle subtending a fixed central angle, ψ_0 , as illustrated in Figure 3-2. The introduction of the central angle, ψ_0 , though generally applicable in defining the ground location of a scan element (datum point), is particularly suitable when scan elements are spaced equally along the great circle arc.

Expressions are presented below for obtaining the geocentric longitude, θ' , and colatitude, ϕ' , of any pair of points (datum points) on the scan arc which are located symmetrically with respect to the subsatellite point. Details of their development are given in Appendix A. These expressions are then

3-3



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Figure 3-2. The Scan Arc as a Central Angle

employed in Sections 3.2 and 3.3 in describing the two alternative algorithms for obtaining the data base coordinates (x, y) of the datum points.

First, the sines and cosines of the geocentric longitude, θ , and colatitude, ϕ , of the subsatellite point are required in terms of the direction cosines available from Equation (3-1), i.e.,

$$\sin\phi\cos\theta=1$$
 (3-3a)

 $\sin\phi\sin\theta = m \tag{3-3b}$

$$\cos \phi = n \tag{3-3c}$$

Actually, Equation (3-3c) yields both $\cos \phi$ and $\sin \phi$, since $0 \le \phi \le 180$ degrees. Consequently, $\sin \theta$ and $\cos \theta$ can be determined from Equations (3-3a) and (3-3b).

Returning to the scan arc, it is convenient to view $\psi/2$ as a general variable describing the positions of any pair of points along the arc. The following expressions yield two unique values for both θ' and ϕ' , depending on the algebraic sign of $\psi/2$. Thus,

$$\cos \phi' = \cos \phi \cos \psi/2 - \cos i \sin \psi/2 \qquad (3-4)$$

where $\cos \phi$ and $\cos i$ are already available. Equation (3-4) gives ϕ' uniquely, since $0 \le \phi' \le 180$ degrees, from which $\sin \phi'$ is then computed and used in

$$\sin (\theta' - \theta) = \frac{\cos (\omega + f) \sin i \sin \psi/2}{\sin \phi \sin \phi'}$$
(3-5)

$$\cos (\theta' - \theta) = \frac{\cos \psi/2 - \cos \phi \cos \phi'}{\sin \phi \sin \phi'}$$
(3-6)

to obtain θ' , where all of the terms to the right of the equalities have previously been determined.

Caution is necessary when applying Equations (3-5) and (3-6) near the poles, since these relations are singular for subsatellite points or datum points near the poles. This is expected, since the geocentric longitudes θ and θ' are undefined at the poles. No problem need arise in computations, however, since the satellites are not expected to have exactly 90-degree inclinations and since, for the case of an equatorial orientation of the data base cube, it is a simple matter to regard the cube face coordinates of such datum points as zero within an arbitrary, but nonzero, distance from the pole.

It is emphasized that Equations (3-4), (3-5), and (3-6) yield two unique values for both θ' and ϕ' , depending on the choice of $\pm \psi/2$; therefore, both halves of the scan arc are treated efficiently using the same auxiliary quantities (sin i, sin ϕ , etc.).

The expressions just presented will be used in Section 3.2 in an algorithm employing a newly developed transformation to obtain data base coordinates of datum points. These results are also used in Section 3.3 in an algorithm designed to employ the older transformation of Section 2.2 to compute the data base coordinates of datum points in a computationally efficient manner.

3.2 POLAR FORMULATION OF THE AREA-PRESERVING TRANSFORMATION

Section 2 of Reference 1 presents an area-preserving transformation from the spherical square to the inscribed plane square. That transformation utilizes a central projection onto the plane square and a transformation from the projected

3-6

coordinates (ξ, η) to the data base coordinates (x, y). This transformation, designated by f*, is given in Equation (2-4) as

x = f*(
$$\xi$$
, η)
y = f*(η , ξ)

and is based on the Jacobian determinant. It was noted in Reference 1 that the solutions obtained to differential equations based on this determinant are not fully constrained by requiring the interior and edges of the spherical square to map onto the plane square. Additionally, it was found that, despite adding further constraints, no analytic solution could be found for f*. An approximate solution was arrived at using a numerical steepest descent method. Under the current contract, an alternative solution was studied; the results are presented below.

Two sets of polar coordinate systems are considered (Figures 3-3 and 3-4). Figure 3-3 shows a spherical square of radius R, produced by the central projection of an inscribed square upon the sphere. The polar coordinates (θ, ϕ) are defined by an axis of the sphere passing through the center of the spherical square and by a great circle arc passing through the axis and bisecting one edge of the square. The azimuthal angle, θ , is measured from this "meridian" and the central angle, ϕ , is measured from the axis. For the case where the coordinate axis and meridian coincide with those of the sphere, θ equals the longitude and ϕ is the colatitude.

Figure 3-4 shows a plane square of side $2r_{o}$. The axis in this system is normal to the plane, passes through the center of the plane, and has its origin at a distance r_{o} below the plane. Coordinates (μ, ν) are defined in the same sense as (θ, ϕ) .

3-7



Figure 3-3. Polar Coordinate System on Spherical Square



Figure 3-4. Polar Coordinate System on an Inscribed Plane

If the axes of the two systems are aligned and the plane square is inscribed within the spherical square, it can be shown that

$$\mathbf{r}_{0} = \mathbf{R}/\sqrt{3} \tag{3-7}$$

The area of the sphere is

$$A_s = 4\pi R^2$$
 (3-8)

and the surface area of the inscribed cube is

$$A_c = 24r_o^2$$
(3-9)

The ratio of the areas of the spherical and plane squares is given by

$$\lambda = \frac{A_{s}}{A_{p}} = \frac{4\pi R^{2}}{24r_{o}^{2}}$$
(3-10)

Substituting Equation (3-7) into Equation (3-10) yields

$$\lambda = \pi/2 \tag{3-11}$$

To further constrain the mapping of the spherical square upon the plane, consider the condition $\mu = f(\theta)$, where μ is independent of ϕ . This condition implies the useful property that any great circle on the sphere passing through the center of the spherical square will map as a straight line passing through the center of the plane square. This function must be area preserving in the sense that sectors of equal area on the sphere must map into equivalent areas on the plane. If the function f exists, then it is possible to relate the areas of differential elements $d\theta$ and $d\mu$ of the spherical and plane squares, as shown below.

To obtain the function $\mu = f(\theta)$, consider the right spherical triangle ABC in Figure 3-5, and let one side, B, be the meridian and A, the azimuth angle, θ . For a unit sphere, R = 1 and b = $\pi/4$.

From the general relation for the area of a spherical triangle

$$S = R(A + B + C - \pi)$$
 (3-12)

and the area of the right spherical triangle is

$$S_{s} = \theta + \beta - \pi/2 \qquad (c \equiv \pi/2) \qquad (3-13)$$

By Napiers' identity, $B = \cos^{-1}(\sin A \cos b)$ and

$$S_{s} = \theta + \cos^{-1}(\sin \theta \cos \pi/4) - \pi/2 \qquad (3-14)$$

This triangle may represent any segment of the spherical square in the range $-\pi/4 \le \theta \le \pi/4$.

The area of the corresponding right triangle on the plane (Figure 3-6) is

$$S_{p} = \frac{ar_{o}}{2}$$
(3-15)



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Figure 3-5. Right Spherical Triangle ABC



Figure 3-6. Plane Right Triangle Corresponding to ABC

or, since $a = r_0 \tan \mu$,

$$S_{p} = \frac{r_{o}^{2}}{2} \tan \mu \qquad (3-16)$$

For $\mu = f(\theta)$ to preserve area, the condition must hold that, for any θ ,

$$S_{g}/S_{p} = \lambda$$
(3-17)

Thus,

 $S_{s} = \pi/2 S_{p} = \frac{r_{o}^{2}}{4} \tan \mu$ (3-18)

and

$$\mu = \tan^{-1} \left(\frac{4}{\pi r_o^2} S_s \right) \tag{3-19}$$

In the case of a unit sphere, $r_0 = 1/\sqrt{3}$, and

$$\mu = \tan^{-1} \left(\frac{12}{\pi} S_{\mathbf{s}} \right) \tag{3-20}$$

$$\mu = \tan^{-1} \left\{ \frac{12}{\pi} \left[\theta - \cos^1 \left(\sin \theta \cos \frac{\pi}{4} \right) - \frac{\pi}{2} \right] \right\}$$
(3-21)

It can be shown that $\mu = 0$ at $\theta = 0$ and $\mu = \pm \pi/4$ at $\theta = \pm \pi/4$. The function f is defined over the quadrant $-\pi/4 \leq (\theta, \mu) \leq \pi/4$ and, by reflection, over the entire spherical and plane squares.

To relate ν and ϕ , expressions for the differential area elements, $\phi d\theta$ and $\nu d\mu$, must be obtained. Figure 3-7 illustrates a spherical zone of one base, having area

$$S_s = 2\pi Rh \tag{3-22}$$

Substituting $h = R - R \cos \phi$ and taking $R \equiv 1$ yields

$$S_s = 2\pi (1 - \cos \phi)$$
 (3-23)

Since

$$S_{s} = \int_{0}^{2\pi} 1 - \cos \phi \, d\theta \qquad (3-24)$$
$$dS_{s} = (1 - \cos \phi) \, d\theta$$

The area of the corresponding sector on the plane is given by

$$ds_{p} = r^{2} d\mu \qquad (3-25)$$

In order to apply the area-preserving condition, Equation (3-17) is differentiated to yield

$$dS_{s} = \pi/2 \ dS_{p} \tag{3-26}$$

Substituting Equations (3-24) and (3-25) into Equation (3-26) gives

$$r^{2}du = 2/\pi (1 - \cos \phi) d\theta$$
 (3-27)

3-13



Figure 3-7. Spherical Zone of One Base

Since $r = r_0 \tan \nu$, Equation (3-27) can be rewritten as

$$\tan^2 \nu = \frac{2}{\pi r_0^2} (1 - \cos \phi) \frac{d\theta}{du}$$
(3-28)

For fixed values of θ and u, Equation (3-28) becomes

$$\tan^2 \nu = C(1 - \cos \phi)$$
 (3-29)

where

$$C = \frac{2}{\pi r_{o}^{2}} \frac{d\theta}{du} |_{\theta, u}$$
(3-30)

Clearly, C can be evaluated from boundary conditions:

1.	When	$\phi=0, \nu=0$	points at the center map onto themselves
2.	When	$ \phi = \phi_{\max} $ $ v = v_{\max} $	points on the boundary of the spherical square must map onto points on the boundary of the plane square

To develop these boundary conditions, it is necessary to obtain ϕ_{\max} and μ_{\max} as functions of θ and μ , respectively. Figures 3-8 and 3-9 illustrate these angles. It should be noted that to obtain the boundary of the spherical square, a central projection to the inscribed plane is employed. This projection preserves both θ and ϕ and reduces the problem to the case of the plane square.

The loci of points describing these edges can be expressed as

$$\mathbf{r}_{\mathbf{s}} = \mathbf{r}_{\mathbf{o}} \sec \theta \tag{3-31}$$

and

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U

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I

$$\mathbf{r}_{\mathbf{p}} = \mathbf{r}_{\mathbf{o}} \sec \mathbf{u} \tag{3-32}$$

and the central angles can be expressed as

$$\phi_{\max} = \tan^{-1} \sec \theta \qquad (3-33)$$

and

$$\nu_{\max} = \tan^{-1} \sec u \qquad (3-34)$$

Substituting ϕ_{\max} and ν_{\max} in Equation (3-27)

$$\tan^2 \nu_{\max} = C(1 - \cos \phi_{\max}) \tag{3-35}$$



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Figure 3-8. Relationship of ϕ_{\max} to θ



Figure 3-9. Relationship of ν_{max} to μ

Therefore

$$C = \frac{\tan^2 \nu_{\max}}{1 - \cos \phi_{\max}} = \frac{\tan^2 (\tan^{-1} \sec u)}{1 - \cos (\tan^{-1} \sec \theta)} = \frac{\sec^2 u}{1 - \cos (\tan^{-1} \sec \theta)}$$
(3-36)

Substituting C into Equation (3-35) yields

$$\tan^2 \nu = \frac{\sec^2 u}{1 - \cos \tan^{-1} \sec \theta} (1 - \cos \phi) \tag{3-37}$$

and

$$\nu = \tan^{-1} \sqrt{\frac{\sec^2 u(1 - \cos \phi)}{1 - \cos \tan^{-1} \sec \theta}}$$
(3-38)

This satisfies the boundary condition $\phi = 0$, $\nu = 0$, and additionally has the property that $\mu = 0$ at $\phi = 0$ and $\phi_{\max} = \nu_{\max}$, where $\theta = \pm \pi/4$, $u = \pm \pi/4$. This relation holds from $\phi = 0$ to $\phi = \pi/2$, and because ϕ_{\max} is never greater than $\tan^{-1}\sqrt{2} = 0.95$, it is correct for the entire spherical square.

Thus, Equations (3-38) and

$$u = \tan^{-1}\left\{\frac{12}{a}\left[\theta + \cos^{-1}\left(\sin \theta \cos\frac{\pi}{4}\right) - \frac{\pi}{2}\right]\right\}$$
(3-39)

represent an area-preserving transformation from the spherical square to the inscribed plane square for the quadrant.

To treat the remaining quadrants, each is successively rotated into the quadrant of definition by subtracting multiples of $\pi/2$, as shown in Table 3-1.
Tab e 3-1. Rotations of θ and u' for a Given ${\theta'}^a$

Range	θ	u'
$-\frac{\pi}{4} \leq \theta' \leq \frac{\pi}{4}$	θ'	u
$\frac{\pi}{4} < \theta' \leq \frac{3\pi}{4}$	$\theta' - \frac{\pi}{2}$	$u + \frac{\pi}{2}$
$\frac{3\pi}{4} < \theta' < \frac{5\pi}{4}$	θ'-π	u + 1
$\frac{5\pi}{4} < \theta' < \frac{7\pi}{4}$	$\theta' - \frac{3\pi}{2}$	u + 31

^aThe quantities θ' and u' are input and output, respectively; and θ and u are as used in the computation of Equations (3-38) and (3-39).

3.3 REFINEMENT OF THE CARTESIAN TRANSFORMATION

Section 4.2 of Reference 1 outlines a method for obtaining the data base coordinates (x, y) of any datum point from Keplerian orbital elements, time, a scan angle, and parameters describing the attitude of the satellite using the inverse transformation, $f^*(\xi, \eta)$. This method has been modified in order to improve computational efficiency. The results are presented below in two parts: the first part treates the computation of the coordinates (x, y) of the subsatellite point (zero scan angle); the second part develops the necessary added formalism for datum points lying along a scan arc, as defined in Section 3.1.2. Details and derivations are included in Appendix C.

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3.3.1 Data Base Coordinates of the Subsatellite Point

Computations for obtaining the data base coordinates (x, y) of the subsatellite point are performed as follows:

- 1. Evaluate Equation (3-1) for 1, m, and n, and note the appropriate cube face by selecting the largest of 1, m, and n and observing its sign. Table 3-2 establishes the necessary conventions. The general variables ξ' , η' , and ρ are the geocentric rectangular coordinates of the subsatellite point deduced from 1, m, and n by entering Table 3-2, where ρ is the largest in absolute magnitude of 1, m, and n.
- 2. Compute the variables α and β defined by:

$$\alpha = \xi' / \rho$$

$$\beta = \eta' / \rho$$
(3-40)

Face		J	
Number	٤'	η'	<u>ρ</u>
1	m	n	1
2	-m	n	-1
3	-1	n	m
4	1	n	-m
5	-1	-m	n
6	-1	m	-n

Table 3-2. Cube Face Identification

3. Compute x and y from

$$x = f^*(\alpha, \beta)$$

$$y = f^*(\beta, \alpha)$$
(3-41)

where $f^*(\alpha, \beta)$ is developed in Appendix C. For convenience, the functional relationship is repeated below.

$$f^{*}(\alpha, \beta) = r_{0} \begin{cases} \alpha \gamma^{*} + \alpha^{3} (1 - \gamma^{*}) + \alpha \beta^{2} (1 - \alpha^{2}) \\ \times \left[\Gamma + (\Gamma - M) (1 - \alpha^{2}) + (1 - \beta^{2}) \sum_{\substack{i \ge 0 \\ j \ge 0}} C_{ij} \alpha^{2i} \beta^{2j} \right] \quad (3-42) \\ + \alpha^{3} (1 - \alpha^{2}) \left[\Omega_{1} + (1 - \alpha^{2}) \sum_{\substack{i \ge 0}} D_{i} \alpha^{2i} \right] \end{cases}$$

where α and β are defined by Equations (3-40), $r_0 = 1\sqrt{3}$, and the remaining constants are defined in Appendix C (Equations (C-10) and (C-11)).

3.3.2 Data Base Coordinates for Nonzero Scan Angle

The data base coordinates (x, y) for any pair of datum points lying on a scan arc at an angular separation of $\pm \psi/2$ from the subsatellite point are obtained by performing the following computational steps:

1. Obtain the geocentric longitude, θ , and colatitude, ϕ , of the subsatellite point using Equations (3-1) and (3-3). (If the ground trace has already been determined, its direction cosines will already have been computed from Equations (3-1).)

- Compute the sines and cosines of the geocentric longitudes, θ', and colatitudes, φ', of the datum points, using Equations (3-4), (3-5), and (3-6), and the results of step 1.
- 3. Using the sines and cosines of θ' and ϕ' in Equations (3-3), obtain the two sets of direction cosines (1', m', and n') of the datum points.
- 4. Identify the appropriate cube face(s) and variables ξ' , η' , and ρ using Table 3-2, and compute values for α and β using Equations (3-40) for each point.
- 5. Calculate the (x, y) coordinates using Equations (3-41).

3.4 APPROXIMATE FORMS

The transformations outlined in the previous subsections share the common disadvantage of being complex and difficult to evaluate. While studying the transformations f and f* for the previous report, it was determined that a linear interpolation scheme was sufficient to find points between known benchmark grid points (Reference 1, Section 4.3). It is therefore evident that the transformation equation may be approximated by low-order polynomials. This subsection presents a method for accurately approximating the transformations over the entire cube face using polynomials of varying degree with locally evaluated and tabulated coefficients.

Because of symmetry properties of the transformations, it is necessary to evaluate the coefficients of the approximation polynomial over only one octant of one face. In addition, since

$$x = f (\xi, \eta); \ y = f (\eta, \xi)$$

$$\xi = f^*(x, y); \ \eta = f^*(y, x)$$
(3-43)

only one set of coefficients is necessary to determine both x and y, and a second set to determine ξ and η .

From Table 4-2 of Reference 1, it can be seen that a double linear interpolation scheme will allow computation of intermediate points to an accuracy of 0.25 pixel over a range of at least 16 pixels (and up to 1024 pixels in the central region) on a 4096 × 4096 square. Consequently, it is reasonable to assume that a higher order polynomial would allow one-pixel accuracy over considerably greater distances. As a first approximation, it is assumed that determination and tabulation of coefficients on a grid with 64-pixel intervals will suffice. This interval is convenient, since it corresponds to the proposed standard data base record size.

To obtain suitable interpolating polynomials, the standard bivariate interpolation formulae were examined for three, four, and six points (Reference 4, Sections 25.2.65 to 25.2.67). From Figure 3-10, the formulae are

• Three Points (Linear)

$$f_a (x_0 + ph, y_0 + qk) = (1 - p - q) f_{0,0} + pf_{1,0} + qf_{0,1} + O(h^2)$$
 (3-44)

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• Four Points

$$f_{a} (x_{0} + ph, y_{0} + qk) = (1 - p) (1 - q) f_{0,0} + p(1 - q) f_{1,0} + q(1 - p) f_{0,1} + pqf_{1,1} + O(h^{2})$$
(3-45)







Figure 3-10. Grid Points Used for Interpolation

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Six Points

$$f_{a} (x_{0} + ph, y_{0} + qk) = \frac{q(q - 1)}{2} f_{0, -1} + \frac{p(p - 1)}{2} f_{-1, 0}$$

$$+ (1 + pq - p^{2} - q^{2}) f_{0, 0}$$

$$+ \frac{p(p - 2q + 1)}{2} f_{1, 0}$$

$$+ q(q - 2p + 1) f_{0, 1} + pqf_{1, 1} + O(h^{3})$$
(3-46)

Equations (3-44) and (3-45) can be written as follows:

$$f_{a}(x,y) = a_{10} x + a_{01} y + a_{00}$$
(3-47)

$$f_{a}(x, y) = a_{11} xy + a_{10} x + a_{01} y + a_{00}$$
(3-48)

The coefficients a_{00}^{0} , a_{10}^{0} , ..., are evaluated by setting up a set of linear equations corresponding to the interpolation benchmark points and solving them simultaneously.

The six-point formula is less convenient to apply systematically, since it requires knowledge of points outside its range for solution. A compromise form

$$f_{a}(x,y) = a_{20} x^{2} + a_{02} y^{2} + a_{10} x + a_{01} y + a_{00}$$
(3-49)

can be evaluated on five points, as shown in Figure 3.11. This form eliminates the xy cross-product term and approximates the functional form with a paraboloid of revolution whose axis is perpendicular to the xy plane.



Figure 3-11. Five-Point Solution

A computer program was written to solve the system of simultaneous equations arising from each of the above approximate forms on 64×64 pixel intervals on one octant of the cube face (Figure 3-12). A standard IBM Scientific Subroutine Package (Reference 5) routine, SIMQ¹, was used for solution of the equations. To test the accuracy of each approximation, a set of evaluation points was chosen (Figure 3-13). The approximations were used successively to determine the coordinates of each evaluation point and a residual was computed for each point by solving the full transformation. Any 64×64 region in which the largest residual exceeded one pixel was flagged and the next higher degree approximation was tried. Table 3-3 shows the number of coefficients necessary for one pixel accuracy as a function of location.

The table size required is not excessive because, as stated above, only one octant (plus diagonal elements) must be stored. For a 64×64 grid, this would be

$$M \times (31^2 + 16) = M \times 977$$

¹SIMQ employs the Gaussian elimination method using largest pivotal divisors.





Table 3-3.Number of Coefficients Necessary for One-PixelAccuracy as a Function of Position



where M is the largest number of coefficients required. For the case where M = 5, this requires 4885 words.

The program which evaluates the coefficients for the approximate direct transformation f_a is shown in Appendix E. Appendix D tabulates the coefficients derived and indicates the necessary number of coefficients for one-pixel accuracy. The use of these approximations will be discussed in Section 4.1.



Figure 3-13. Arrangement of Evaluation Points

3.5 COMPARISON OF TRANSFORMATIONS

In the previous subsections, four different formulations were developed for the f* function (inverse transformation):

- 1. The original Cartesian form (Equation (2-5))
- 2. The revised Cartesian form (Equation (3-42))
- 3. The polar form (Section 3.2)
- 4. The approximate form (Section 3.4)

In choosing between forms 1 and 2, which produce the same mapping, form 2 is clearly arithmetically simpler and has proved to produce a more accurate determination of its coefficients.

Form 4 may be used to approximate the Cartesian or the polar forms and is the most efficient form computationally. It does, however, require core storage for the coefficient tables.

Figures 3-14 and 3-15 illustrate the differences in the nature of the mappings produced by the Cartesian and polar forms. The salient difference is that the polar form exhibits nondifferentiability at crossings of the diagonals of the face. This directional discontinuity arises from the condition that all great circles pass through the center of a face map as straight lines. This problem, in conjunction with the somewhat slower computation time shown in Table 3-4, favors the Cartesian form for most applications. The polar form retains the advantage of exactness in that it produces equal area to any desired accuracy



Figure 3-14. Distortions Arising From the Cartesian Form of the Transformation



Figure 3-15. Distortions Arising From the Polar Form of the Transformation

Table 3-4.	Execution Times for Transformation	
	Forms on the IBM S/360-91 (One	
Ordered Pair)		

Transformation	Execution Time (microseconds)
Cartesian	300
Polar	1000
Approximate	40

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without recomputing the coefficients. It is difficult to compute coefficients that produce Cartesian transformations which will match retransformed points below the 2-nautical-mile level.

Both the Cartesian and polar formulations have their own uses: the Cartesian system at moderate resolutions and the polar system at the highest resolutions over small areas not including a diagonal. For benchmark transformations, the speed of either transformation is adequate, given a reasonably great interpolation distance. The high computational efficiency of the approximate forms tends to favor this approach for direct computation of data base coordinates (or for computation of corresponding points in the scan line buffer). This technique could also be particularly useful in the corner regions where interpolation is not accurate over useful distances.

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SECTION 4 - I/O SYSTEM DESIGN CONSIDERATIONS

4.1 ALTERNATIVE METHOD OF FAST-FILLING

Section 2.3 briefly outlines a proposed alternative fast-filling scheme for the preprocessor program. This method, described in Reference 3, partially solves the problem of edge and vertex crossings and eliminates the necessity for explicit data spreading. By combining this approach with the approximation techniques descussed in this report, the overall efficiency of the preprocessor function can be increased still further.

Sections 3.1 and 3.3 present a method for obtaining data base coordinates (x, y) directly from Keplerian orbital elements and scanner angle. By omitting the last step of the process outlined in Section 3.3, positions in the (α, β) system can be obtained, which is simply a rescaled (ξ, η) system $(\alpha = \xi/r_0, \beta = \eta/r_0)$. Using this algorithm, a benchmark grid in (α, β) can be obtained and a procedure can then be followed which consists of a modification to that described in Reference 3. As outlined in Section 2.3 of Reference 3, filling is performed only for complete data base records (thus avoiding the necessity to read and update partially covered records). The procedure employs a new rotating Earth-centered coordinate system (x^R, y^R, z^R) in which both scanner and data base benchmarks are defined. CSC proposes modifying this procedure as follows:

- 1. Instead of using the x^R , y^R , z^R system, both scanner and benchmark grids in a coplanar system should be defined, projecting the scan line grid into the inscribed cube face and transforming the data base coordinates to the same (α, β) system.
- 2. The data base benchmarks should be defined in terms of the projected scanner benchmarks (i.e., the (E, L) coordinates of the data base benchmarks should be defined as shown in Figure 4-1).



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Figure 4-1. Coplanar Representation of Scanner and Data Base Grids

3. The (E, L) coordinates of each data base location should be interpolated and the nearest pixel stored in that location. (This may be done by offsetting the (E, L) system 1/2 pixel in both coordinates and obtaining E and L indices of the desired pixel by truncation.)

The objective of this modification is to avoid the necessity of searching for the nearest pixel in a three-dimensional space.

Neither this method nor the procedure in Reference 3 completely eliminates the edge and vertex crossing problems. It is necessary to predetermine, from the benchmarks of the scanner array, which data base records will be completely covered by the scanner data currently in the input buffer. This process must check for edge and vertex crossings.

Figure 4-2 illustrates the relationship between the scanner data input buffer and the data base records. The scan buffer must have sufficient length (in the along-track direction) to cover the longest diagonal of a data base record. For the 64 x 64 pixel case, this will be approximately 128 pixels in the corners of

the face. The scan line buffer must be organized as a creeping buffer, and no more lines may be added in the forward direction (and deleted in the backward direction) than the number of lines by which the buffer exceeds the minimum needed to cover the longest diagonal. A typical scanner input buffer would have a width (along the scan direction) of approximately 1600 nautical miles (about 1344 pixels), with a minimum length (in the ground track direction) of 128 pixels. This will require sufficient buffer storage for 172,032 pixels. Two methods emerge for handling this large array. The first is to pack more than one pixel per 60-bit word. This would allow the entire buffer to be maintained in central memory on the CYBER/175 but would introduce a large packing and unpacking cost. A second method would be to maintain the buffer in ECS, segmented so that it could be brought into core in blocks as required. The scheduling of these blocks, along with the ordering of output records, could be determined from





the (x, y) locations of the projected and transformed scanner input benchmark grid, ordering the filling of data base records by their corresponding along-scan locations.

Because the objective of this new method is to reduce redundant stores, it could also be useful to consider windowing down the along-scan length of the input buffer as a function of latitude to reduce scan path overlap at high latitudes. Similarly, at low latitudes, it may be necessary to subdivide data base records into two or more longitude bands and partially update the records rather than fully rewriting them.

4.2 MAINTENANCE OF MULTIPLE- AND VARIABLE-RESOLUTION DATA BASES

Considerable savings in computation and I/O time and in data base storage can be realized by reducing the resolution of the data base in areas of lesser interest (such as, perhaps, over land and in the southern hemisphere). It may also be desirable to be able to maintain higher resolution for other selected areas. One method for providing such variable resolution coverage consists of maintaining a single, fully global data base at the lowest desired resolution, plus separate higher resolution partial data bases. A second method consists of maintaining a single data base of variable resolution.

In either case, it will be necessary to provide a system for indicating the desired resolution as a function of global position. This can be provided for by establishing a data block properties table (for example at the 128×128 or 256×256 element level) which would contain the desired resolution described as a multiple of the finest usable resolution or by the number of bits required for the serial address. This would allow the preprocessor program to determine the desired resolution at any point.

For the multiple data base scheme, the preprocessor program would first determine which elements of the resolution table are covered by the data currently in the scan line buffer and then access the table to determine what actions must

be taken. This table would consist of entries for each of the desired resolution levels indicating

- The desired resolution
- A base address for the block of records corresponding to the table element
- The size of the data base records (in elements)

The resolution and record size for all records in a block corresponding to one table element can both be expressed as a number of bits to be added to the corresponding binary division levels of the basic global data base, i.e., if the desired resolution is twice the linear resolution of the basic data base, two more address bits must be added and, if the records are to have the same number of elements, the record/element division of the serial string must take place two bits lower.

The serial string for each resolution level would be kept separate, and the use of a base address for each record block would obviate any need for unused filler records in the data base. For maximum flexibility, each resolution level could be maintained as a separate file.

With multiple data bases, the size of a data base record would be adjustable to allow an optimum match of record size to scanner input buffer size and data item size (where one data item might be all of the simultaneously available, colocated scanner readings). To maintain this match, a separate pass through the data will be required for each resolution level.

In the variable resolution approach, only one pass through the data is required and all data records cover the same area. This means that the ratio of the scanner buffer size to the data record area remains fixed, as do the total number of I/O operations. The total number of data records remains constant; thus, a consistent record numbering and ordering can be maintained, avoiding

the necessity of a table lookup for a block address, as in the multiple data base scheme. In this case, the data block lookup table need contain only the resolution (which will also imply the record length).

Both of the above schemes have advantages and disadvantages:

- The multiple data base scheme requires redundant stores and computation
- The variable resolution scheme requires stringing together unlike data base records, thus requiring a variable record length capability
- The multiple data base scheme allows easy addition or deletion of nonstandard resolution areas for coverage of new areas of interest
- The variable resolution scheme allows more systematic control over the data base storage allocations

A possible resolution of these conflicts is to use a hybrid system in which the variable resolution format would be used for resolutions coarser than a set standard resolution level, and the multiple data base approach for higher resolution levels (and for other subdata bases required for special purposes or limited durations).

A unified block properties table would be maintained which would allow display of the current status and availability of a data record in the system. This table could be expanded to include such information as the last update time, satellite ID, dayside/nightside information, etc. The advantages of such a scheme would be that multiple passes through the data would be restricted to selected regions of highest resolution, special changes in data base content could be accomplished without restructuring the main data base as a whole, systematic control over the entire data base could be maintained, and the serial string addressing scheme would be maintained for the standard resolution data.

SECTION 5 - GEOGRAPHIC ORIENTATION

5.1 INTRODUCTION

In order to maximize the usefulness and efficiency of the spherical cube data base, it may be necessary to orient the inscribed cube in some arbitrary manner with respect to the Earth. The simplest orientation of the cube aligns one Cartesian axis of the cube coordinate system with the Earth's axis of rotation and points another cube axis toward the prime meridian. A simple variation is to have one cube axis aligned with the Earth's axis of rotation and another pointing to a specified longitude. This section will discuss these orientations and will present an algorithm for producing arbitrary cube rotations.

5.2 SIMPLE ORIENTATIONS

In Section 4.2 and Appendix E.6 of Reference 1, algorithms are presented for the transformation of Earth-centered Cartesian coordinates (a, b, c) to data base coordinates (x, y). For the case where the Earth-centered (a, b, c)system is aligned as shown in Figure 5-1, the Cartesian coordinates can be represented as:

$$a = \cos (\text{latitude}) \cos (\text{longitude})$$
 (5-1)

$$b = \cos (latitude) \sin (longitude)$$
 (5-2)

$$c = sin (latitude)$$
 (5-3)

To accomplish a simple rotation about the Earth's polar axis, the following algorithms can be used:

$$a = \cos (\text{latitude}) \cos (\text{longitude} + \delta)$$
 (5-4)

b = cos (latitude) sin (longitude +
$$\delta$$
) (5-5)

$$c = sin (latitude)$$
 (5-6)

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where δ represents the angle of rotation.



Figure 5-1. Orientation of the Axes of the Cube (Unrotated)

5.3 THE GENERAL ROTATION CASE

In order to conveniently generate arbitrary rotations of the data base cube with respect to the sphere, a method was adopted for making successive rotations of the sphere about each of the axes of the cube. By choosing the order in a suitable manner, this yields a conceptually simple way of arriving at a given orientation. Taking a general set of orthogonal axes, G_1 , G_2 , G_3 , as shown in Figure 5-2, the rotation about axis G_1 can be expressed as follows:

Let α be the desired angle of rotation about G_1 . For any arbitrary point, g_1 , g_2 , g_3 , the position may be expressed in terms of the following three variables:

- g_1 , the G_1 coordinate
- β , the angular position of the projection of g_1 , g_2 , g_3 on the G_2G_3 plane measure from G_3 in the clockwise sense

• ρ , the perpendicular distance from g_1 , g_2 , g_3 to the G_1 axis For a pure rotation about G_1 (Figure 5-3), g_1 and ρ are constant and the final orientation, g'_1 , g'_2 , g'_3 , is determined by the new orientation angle, γ , where

$$\boldsymbol{\gamma} = \boldsymbol{\alpha} + \boldsymbol{\beta} \tag{5-7}$$

and since

$$\boldsymbol{\beta} = \tan^{-1} \left(g_3^{\prime} / g_2^{\prime} \right) \tag{5-8}$$

then

$$\gamma = \alpha + \tan^{-1} \left(g_3^{/} g_2^{} \right) \tag{5-9}$$

where

$$\rho = g_2^2 + g_3^2 \tag{5-10}$$



Figure 5-2. Point g_1 , g_2 , g_3 in the General Reference Frame, G_D , G_2 , G_3



Figure 5-3. Rotation of the Point (g_1, g_2, g_3) About the G_1 Axis

yielding

$$g'_1 = g_1$$
 (5-11)

$$\mathbf{g}_{\mathbf{p}}' = \boldsymbol{\rho} \cos \boldsymbol{\gamma} \tag{5-12}$$

$$\mathbf{g}_{\mathbf{p}}' = \boldsymbol{\rho} \, \sin \, \boldsymbol{\gamma} \tag{5-13}$$

Using this general approach, a rotation through an angle α_1 is first made about the axis coincident with the Earth's axis of rotation. This is functionally equivalent to the rotation described in Section 5.1 and results in a rotation of the α_1 longitude line into the center of face number 1.

The second rotation is made about the axis normal to the pole and emerging through a face normal to face number 1 (axis b in Figure 5-1). If this rotation is through an angle α_2 , then the point with longitude α_1 and latitude α_2 will be rotated into the center of face 1.

A third rotation about the axis emerging from the center of face 1 will produce a "position angle" change on face 1 but will not translate the point at the center of the face.

Using this method, any point whose latitude and longitude are known may be translated to the center of face 1 and the rotational aspect of that face may be altered without changing the latitude and longitude of the center. Referring to Figure 5-1, the order of these rotations is about axes c, b, and a, respectively.

Appendix E contains an algorithm for the application of the general transformation to successive axes through the interchange of components.

Figures 5-4 and 5-5 illustrate the continental outlines after the indicated rotations and are typical of maps generated by this process.



Figure 5-4. Data Base Map of Earth With Cube Oriented 90 Degrees West of Greenwich



Figure 5-5. Data Base Map of Earth With Cube Oriented 90 Degrees West of Greenwich and 30 Degrees South of the Equator

SECTION 6 - SUMMARY AND CONCLUSIONS

6.1 NEW RESULTS

The purpose of the research reported herein is to improve the procedures and algorithms required for the efficient maintenance of data in a quadrilateralized spherical cube Earth data base. The new results obtained relate to the fastfilling of the data base. A technique was developed for rapidly finding the ground trace and scanner-sensed locations from Keplerian elements and scanner angle. The previously derived form of the inverse transformation, f*, was improved by a change of variables. A new inverse transformation was derived using polar coordinates, and a technique was developed for approximating the transformation functions using low-order polynomials.

All of these results will contribute to improving the accuracy and speed of the preprocessor function. The general functioning of the preprocessor input/output algorithms was re-examined in light of an alternative scheme (Reference 3). This scheme and the previously developed fast-filling method (Reference 1) were compared, and a modification of the scheme described in Reference 3 is recommended.

In addition, improved world maps in data base coordinates were produced for use in determining an optimum orientation of the data base cube. Additional efforts were devoted to improving the accuracy of the transformations described in Reference 1. The revised coefficients for these transformations are included in Appendix B of this report.

6.2 RECOMMENDATIONS FOR FURTHER STUDY

A number of topics considered during this study are worthy of additional attention. The polar formulations of the transformation functions should be further developed using other constraints to maximize conformality of the mapping.

The role of the MAP III processor in implementing the preprocessor function was excluded from this study because documentation is not available. The role which MAP III and the high-speed communications system's processor could assume in evaluating the transformation equations, particularly the approximate transformation, should be explored further.

The various CYBER I/O access methods should be studied further, particularly to ascertain whether or not they may cause fragmentation of the data base.

Consideration should be given to producing display output in data base coordinates with a geographic grid overlay, thereby saving a major computation cost in retransforming the data.

APPENDIX A - GROUND TRACE EQUATIONS

This appendix presents the derivation of the expressions summarized in Section 3.1 and referred to elsewhere in this document. Two problems are addressed: (1) the computation of the geocentric direction cosines (l, m, n) of the subsatellite point in terms of the Keplerian orbital elements of the satellite and (2) the computation of the geocentric longitudes and colatitudes of any pair of datum points on a scan arc in terms of previously computed quantities (e.g., l, m, n). A spherical Earth is assumed throughout.

A.1 THE SUBSATELLITE POINT

Figure A-1 shows the geometric relationship of the satellite, the satellite's orbit, and the Earth's equator to the (rotating) rectangular coordinate system (x, y, z). The coordinate system is fixed in the Earth and, therefore, rotates at the diurnal rate. The Keplerian orbital elements whose definitions are given by the figure are the inclination, i; the argument of perigee, ω , shown as part of the angle $(\omega + f)$, where f is the satellite's true anomaly; and the node, Ω , which is not the usual fixed quantity, but varies with the *i* idereal time (diurnal rate), s, as follows:

$$\Omega = \Omega_{0} - s \tag{A-1}$$

where Ω_{Ω} is the value of Ω at an arbitrary initial time.

Figure A-1 shows the subsatellite point having geocentric coordinates longitude, θ , and colatitude, ϕ , and lying, by definition, on the line joining the satellite and the Earth's center. Therefore, the direction cosines (l, m, and n) describing the direction to the satellite in the x, y, z system are identical with those for the subsatellite point. The straightforward relationship between the direction cosines and θ and ϕ will be exploited in Section A.2. First, the

A-1



Figure A-1. Satellite Orbit Orientation and Geographic Coordinates

expressions for l, m, and n in terms of Ω , ω , i, and f are derived below.

Consider the geocentric position vector, \vec{r} , of the satellite in its orbit. Express \vec{r} in the two-dimensional rectangular coordinate system (\hat{p}, \hat{q}) in the orbital plane, where the unit vector \hat{p} is in the direction of the node and the unit vector \hat{q} is in the direction 90 degrees in advance of \hat{p} (see Figure A-1)

$$\vec{\mathbf{r}} = \mathbf{r} \cos (\omega + \mathbf{f}) \, \hat{\mathbf{p}} + \mathbf{r} \sin (\omega + \mathbf{f}) \, \hat{\mathbf{q}}$$

where r is the magnitude of \vec{r} . The coordinates (r_x, r_y, r_z) of the satellite in the geocentric equatorial rectangular coordinate system are, then,

$$\mathbf{r}_{\mathbf{x}} = \mathbf{\vec{r}} \cdot \hat{\mathbf{x}} = \mathbf{r} \cos (\omega + \mathbf{f}) \, \hat{\mathbf{p}} \cdot \hat{\mathbf{x}} + \mathbf{r} \sin (\omega + \mathbf{f}) \, \hat{\mathbf{q}} \cdot \hat{\mathbf{x}}$$
$$\mathbf{r}_{\mathbf{y}} = \mathbf{\vec{r}} \cdot \hat{\mathbf{y}} = \mathbf{r} \cos (\omega + \mathbf{f}) \, \hat{\mathbf{p}} \cdot \hat{\mathbf{y}} + \mathbf{r} \sin (\omega + \mathbf{f}) \, \hat{\mathbf{q}} \cdot \hat{\mathbf{y}} \qquad (A-2)$$
$$\mathbf{r}_{\mathbf{z}} = \mathbf{\vec{r}} \cdot \hat{\mathbf{z}} = + \mathbf{r} \sin (\omega + \mathbf{f}) \, \hat{\mathbf{q}} \cdot \hat{\mathbf{z}}$$

where the term in $\hat{p} \cdot \hat{z}$ for r_z is omitted, since $\hat{p} \cdot \hat{z} = 0$. Now,

$$\hat{\mathbf{p}} \cdot \hat{\mathbf{x}} = \cos \Omega$$

$$\hat{\mathbf{q}} \cdot \hat{\mathbf{x}} = \cos (90^{\circ} + \Omega) \cos i = -\sin \Omega \cos i$$

$$\hat{\mathbf{p}} \cdot \hat{\mathbf{y}} = \sin \Omega \qquad (A-3)$$

$$\hat{\mathbf{q}} \cdot \hat{\mathbf{y}} = \sin (90^{\circ} + \Omega) \cos i = \cos \Omega \cos i$$

$$\hat{\mathbf{q}} \cdot \hat{\mathbf{z}} = \cos (90^{\circ} - i) = \sin i$$

Substituting Equations (A-3) into Equations (A-2), r_x , r_y , and r_z become

$$r_{x} = r \cos \Omega \cos (\omega + f) - r \sin \Omega \sin (\omega + f) \cos i$$

$$r_{y} = r \sin \Omega \cos (\omega + f) + r \cos \Omega \sin (\omega + f) \cos i$$
 (A-4)

$$r_{z} = r \sin (\omega + f) \sin i$$

Since the satellite position vector, \vec{r} , in the $(\hat{x}, \hat{y}, \hat{z})$ system is

$$\vec{\mathbf{r}} = \mathbf{r}_{\mathbf{x}} \hat{\mathbf{x}} + \mathbf{r}_{\mathbf{y}} \hat{\mathbf{y}} + \mathbf{r}_{\mathbf{z}} \hat{\mathbf{z}}$$
(A-5)

the direction cosines (l, m, n) of its position are defined by

$$\vec{\mathbf{r}} \cdot \hat{\mathbf{x}} = |\vec{\mathbf{r}}| |\hat{\mathbf{x}}| \mathbf{1}$$

$$\vec{\mathbf{r}} \cdot \hat{\mathbf{y}} = |\vec{\mathbf{r}}| |\hat{\mathbf{y}}| \mathbf{m} \qquad (A-6)$$

$$\vec{\mathbf{r}} \cdot \hat{\mathbf{z}} = |\vec{\mathbf{r}}| |\hat{\mathbf{z}}| \mathbf{n}$$

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Now, from Equation (A-5),

$$\vec{r} \cdot \hat{x} = r_x \hat{x} \cdot \hat{x} = r_x$$
$$\vec{r} \cdot \hat{y} = r_y \hat{y} \cdot \hat{y} = r_y$$
$$\vec{r} \cdot \hat{z} = r_y \hat{z} \cdot \hat{z} = r_y$$

A-4

which, when substituted into Equations (A-6), yields

$$1 = r_{x} / |\vec{r}| |\hat{x}| = r_{x} / r$$

$$m = r_{y} / |\vec{r}| |\hat{y}| = r_{y} / r$$

$$n = r_{z} / |\vec{r}| |\hat{z}| = r_{z} / r$$
(A-7)

since $|\hat{\mathbf{x}}| = |\hat{\mathbf{y}}| = |\hat{\mathbf{z}}| = 1$, and $|\hat{\mathbf{r}}| = \mathbf{r}$.

Equations (A-4) and (A-7) then yield

$$l = \cos \Omega \cos (\omega + f) - \sin \Omega \sin (\omega + f) \cos i$$

$$m = \sin \Omega \cos (\omega + f) + \cos \Omega \sin (\omega + f) \cos i$$
 (A-8)

$$n = \sin (\omega + f) \sin i$$

Equations (A-8) provide one way of defining the "ground trace" of the satellite, since 1, m, and n may be regarded as the "unit" geocentric coordinates of the subsatellite point. That is, assuming a spherical Earth, the actual coordinates of the subsatellite point are obtained by multiplying 1, m, and n by the Earth's radius in whatever units are required (e.g., kilometers). Equation (A-8) is given as Equation (3-1) in the text.

A.2 THE LOCATION OF DATUM POINTS ON A SCAN ARC

The satellite's onboard sensors are assumed to sweep out a scan arc having ground trace geometric properties as defined in Section 3.1.2 and as illustrated in Figure A-2. It is convenient to regard the ground trace of the scan arc as being composed of pairs of datum points symmetrically placed on either side of the subsatellite point, and to regard $\psi/2$ as a general variable locating the points by assigning alternate algebraic signs to $\psi/2$ (e.g., the ends of the

A-5



Figure A-2. The Scan Arc as a Central Angle ψ_0

scan arc are located at $\pm \psi_0/2$ from the subsatellite point). It is desired to obtain pairs of values for the geocentric longitude, θ' , and colatitude, ϕ' , of the datum points in terms of known quantities in a computationally efficient manner.

The sines and cosines of the geocentric longitude, θ , and colatitude, ϕ , of the subsatellite point are required as intermediate quantities in terms of 1, m, and n from Equations (A-8). By inspection of Figure A-1,

$$\cos \theta \sin \phi = 1$$

 $\sin \theta \sin \phi = m$ (A-9)
 $\cos \phi = n$

Care must be taken in using Equations (A-9) to obtain $\sin \phi$, $\cos \phi$, $\sin \theta$, and $\cos \theta$, since the algebraic signs of these quantities must be preserved for later use. This is readily achieved by noting that $\cos \phi$ yields ϕ uniquely, since $0 \le \phi \le 180$ degrees. Straightforward manipulation yields the remaining quantities. Equation (A-9) is given in the text as Equation (3-3).

To obtain the values of ϕ' , the necessary relations may be derived by considering the two spherical triangles PSQ and PQR shown in Figure A-2 and having the common side PQ, which is the colatitude, ϕ , of the subsatellite point. These triangles are shown in detail in Figure A-3, where it is noted that the arc PS is 90 degrees, that the scan arc intersects the ground trace at right angles, and that point R' may be regarded as any point on the scan arc at a distance of $\psi/2$ from the subsatellite point, Q. (Figure A-2 shows only the end point of the scan arc, R, at a distance $\psi_0/2$ from the subsatellite point.) Also shown in Figure A-3 are two auxiliary angles, A and β , such that A + β + 90 degrees = 360 degrees.



Figure A-3. Two Adjoining Spherical Triangles and the Subsatellite Point Q
The cosine law applied to triangle PQR' yields

$$\cos \phi' = \cos \phi \cos \psi/2 + \sin \phi \sin \psi/2 \cos \beta \qquad (A-10)$$

Since $\psi/2$ is a quantity given a priori, all of the quantities to the right of the equality in Equation (A-10) are known except $\cos \beta$, which is obtained as follows.

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$$\beta = 270^{\circ} - A$$

Therefore

$$\cos \beta = \cos (270^{\circ} - A) = \cos (270^{\circ}) \cos A + \sin (270^{\circ}) \sin A$$

= -sin A (A-11)

Applying the sine law to triangle PSQ,

$$\frac{\sin A}{\sin 90^{\circ}} = \frac{\sin (90^{\circ} - i)}{\sin \phi}$$
(A-12)

or

 $\sin A = \cos i / \sin \phi$

Employing Equations (A-11) and (A-12) in Equation (A-10) yields

$$\cos \phi' = \cos \phi \cos \psi/2 - \cos i \sin \psi/2 \qquad (A-13)$$

which is the desired relation for $\cos \phi'$. Equation (A-13) yields ϕ' uniquely, since $0 \le \phi' \le 180$ degrees, so that $\sin \phi'$ follows immediately. By appropriately redrawing Figure A-3 and repeating the above analysis, it can be readily seen that ϕ' for the datum point corresponding to $-\psi/2$ on the other half of the scan arc is obtained from Equation (A-13) simply by replacing $\psi/2$ with $-\psi/2$. Next, to obtain the values of θ' , Figure A-3 is again used to note that the cosine law applied in triangle PQR' yields

$$\cos \psi/2 = \cos \phi \cos \phi' + \sin \phi \sin \phi' \cos (\theta' - \theta)$$

(A-14)

(A-15)

 \mathbf{or}

$$\cos (\theta' - \theta) = (\cos \psi/2 - \cos \phi \cos \phi') / \sin \phi \sin \phi'$$

Since θ and all other quantities in Equation (A-14) are previously known through Equations (A-9), it remains only to obtain an independent relation for sin ($\theta' - \theta$) to uniquely specify the quadrant of θ' . The sine law applied to triangle PQR' yields

$$\frac{\sin\beta}{\sin\phi'} = \frac{\sin(\theta'-\theta)}{\sin\psi/2}$$

or

$$\sin (\theta' - \theta) = \sin \beta \sin \psi/2/\sin \phi'$$

To obtain $\sin \beta$, it should be noted that, by definition,

$$\boldsymbol{\beta} = 270^{\rm O} - \mathrm{A}$$

Therefore,

$$\sin \beta = \sin (270^{\circ} - A) = \sin (270^{\circ}) \cos A - \cos (270^{\circ}) \sin A$$

= -\cos A (A-16)

Applying the cosine law for angles to triangle PQS,

$$\cos A = -\cos (90^{\circ} - i) \cos (\theta - \Omega) + \sin (90^{\circ} - i) \sin (\theta - \Omega) \cos (90^{\circ})$$

$$= -\sin i \cos (\theta - \Omega)$$
(A-17)

However, the cosine law for sides in triangle PQS yields

$$\cos (\omega + f) = \cos (90^{\circ}) \cos \phi + \sin (90^{\circ}) \sin \phi \cos (\theta - \Omega)$$

or

$$\cos (\theta - \Omega) = \cos (\omega + f) / \sin \phi$$

Therefore, Equation (A-17) becomes

$$\cos A = -\sin i \cos (\omega + f) / \sin \phi \qquad (A-18)$$

Substituting Equation (A-18) into Equation (A-16), and then substituting the result into Equation (A-15) yields

$$\sin (\theta' - \theta) = \cos (\omega + f) \sin i \sin \psi/2/\sin \phi \sin \phi' \qquad (A-19)$$

Equations (A-14) and (A-19) are the desired expressions that uniquely determine θ' from previously known quantities. Alternative forms for Equation (A-19) were found, but the one given here offers the simplest functions of the quantities, all of which are previously known and is the most rapid to compute, since no additional evaluations of trigonometric functions are required. Again, by appropriately redrawing Figure A-3 and repeating the above analysis, it can be readily seen that Equations (A-14) and (A-19) uniquely yield θ' for

the datum point corresponding to $-\psi/2$ on the other half of the scan arc. This is accomplished by entering Equations (A-19) with $-\psi/2$, and entering both Equations (A-14) and (A-19) with the appropriate ϕ' from Equation (A-13).

The important results just derived are Equations (A-13), (A-14) and (A-19), which yield ϕ' and θ' for pairs of datum points on any scan arc, and are cited as Equations (3-4), (3-5), (3-6), respectively, in the text. Caution is necessary when applying Equations (A-14) and (A-19) near the poles, since these relations are singular for subsatellite points or datum points at the poles. This is expected, since the geocentric longitudes θ and θ' are undefined at the poles. No problem need arise in computations, however, since the satellites are not expected to have exactly 90-degree inclinations and since, for the case of an equatorial orientation of the data base cube, it is a simple matter to regard the cube face coordinates of such datum points as zero within an arbitrary, but nonzero, distance from the poles.

APPENDIX B - BEST VALUES FOR THE COEFFICIENTS OF f*

For the form of the inverse mapping function, f^* , given in Equation (3-42) the best derived coefficients are

$\gamma * = 1.3682$	6932	2 362
$\Gamma = -0.1268$	8142	8338
M = -0.0024	3108	1471
$\Omega_1 = -0.1583$	3565	5509
$C_{00} = 0.1014$	5485	8689
$C_{10} = 0.0160$	1398	776 3
$C_{01} = -0.0904$	3332	1481
$C_{11} = 0.0574$	8768	517 3
$C_{20} = -0.0272$	7606	9952
$C_{02} = -0.0342$	1969	6290
$D_0 = 0.0292$	5271	6867
$D_1 = 0.0059$	2537	5028

APPENDIX C - REFINEMENT OF THE CARTESIAN TRANSFORMATION

Section 3.3 of the text presented a numerically efficient and compact algorithm for employing the Cartesian transformation, f^* , defined in Section 2.2 to obtain the data base coordinates (x, y) of any datum point or pair of datum points on the (spherical) Earth's surface. This Appendix discusses the rationale for the new algorithms and provides a derivation of Equation (3-42).

C.1 OVERVIEW

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The algorithm presented in Reference 1 (Section 4, Appendix B) already provides the essential information needed to implement f^* , but this study has found that the algorithm may be made more efficient computationally and better suited to computer programming by the following modifications:

- Take advantage of the nondimensional analysis presented in Appendix A that removes the need to compute the three-dimensional spatial coordinates of datum points by relying on simple direction cosines (1, m, n) of the subsatellite point and on the longious and colatitudes of the datum points.
- Allow for the Earth's rotation by using a varying node (see Appendix A) in computing the direction cosines instead of performing a separate transformation.
- Reduce by about half the computations required in cases of nonzero scan angles by viewing the scan arc as a central angle ψ_0 (see Appendix A) and computing the data base coordinates for points in symmetrically placed pairs instead of treating each datum point independently.
- Remove a step from the algorithm, saving computations as well as yielding desirable numerical properties (see Section 3.4), by

C-1

introducing a change of variable that leads to a new formulation of f^* . The results of this step are summarized in Equations (3-40), (3-41), and (3-42) of Section 3.3.1. The mathematical details are presented in Section C.2 below.

C.2 REFORMULATION OF f*

Table 3-2 (Section 3.3.1) establishes the conventions that define variables ξ' , η' , and ρ . These variables also appear in Table 4-1 of Reference 1, differing only in that 1, m, and n are dimensionless (see Appendix A), whereas a, b, and c (of Reference 1) are spatial coordinates. Mathematically, ξ' , η' , and ρ behave identically here and in Reference 1.

In Reference 1, once ξ' , η' , ρ are found in Table 4-1, they are transformed to (ξ, η) by

where $P = r_0 / \rho$ and $r_0 = 1/\sqrt{3}$. The new variables ξ and η are then used in f* to obtain the data base coordinates (x, y) by

$$x = f^* (\xi, \eta)$$

 $y = f^* (\eta, \xi)$
(C-2)

The transformation given by Equations (C-1) can be eliminated by defining new variables α and β by

$$\alpha = \xi' / \rho$$

$$\beta = \eta' / \rho$$
(C-3)

Comparison of Equations (C-1) and (C-3) shows that the new variables α and β eliminate the computation of the intermediate quantity P and reduce the transformation to simple ratios of previously known quantities. The introduction of α and β will now be completed by deriving $f^*(\alpha, \beta)$, the transformed equivalent of $f^*(\xi, \eta)$, which was given in Section 2.2 and Reference 1 as

$$f^{*} (\xi, \eta) = \gamma^{*} \xi + \frac{1 - \gamma^{*}}{r_{o}^{2}} \xi^{3} + \xi \eta^{2} (r_{o}^{2} - \xi^{2}) \left[\delta^{*} + \delta_{1}^{*} (r_{o}^{2} - \xi^{2}) + (r_{o}^{2} - \eta^{2}) \sum_{\substack{i \ge 0 \\ j \ge 0}} c_{ij}^{*} \xi^{2i} \eta^{2j} \right] + \xi^{3} (r_{o}^{2} - \xi^{2}) \left[\omega^{*} + (r_{o}^{2} - \xi^{2}) \sum_{\substack{i \ge 0 \\ i \ge 0}} d_{i}^{*} \xi^{2i} \right]$$

where $r_0 = 1/\sqrt{3}$ and

$$\gamma^{*} = \frac{1}{\gamma} \qquad \delta^{*} = \frac{(\mu_{1}^{*} - \mu^{*})}{2r_{0}^{4}}$$

$$\mu^{*} = \frac{1}{\mu} \qquad \omega^{*} = \frac{1}{2r_{0}^{4}} \qquad (3 - 2\gamma^{*} - \mu^{*} - 2r_{0}^{4} \delta^{*})$$

$$\gamma_{1}^{*} = \frac{1}{\gamma + r_{0}^{4} \delta} \qquad \delta_{1}^{*} = \frac{1}{2r_{0}^{2}} \qquad (C-5)$$

$$\mu_{1}^{*} = \frac{1}{\mu + 2r_{0}^{4} \delta} \qquad \delta_{1}^{*} = \frac{1}{r_{0}^{2}} \left[\frac{(\gamma_{1}^{*} - \gamma^{*})}{r_{0}^{4}} - \delta^{*} \right]$$

and γ , μ , δ , ω , and the c_{ij}^* and d_i^* have the numerical values given in Section 2.2.

First, Equation (C-4) is expressed in terms of ξ' , η' , and ρ by using Equations (C-1)

$$\begin{aligned} f^{*}(\xi', \eta', \rho) &= \gamma^{*} r_{o} \xi' / \rho + \frac{1 - \gamma^{*}}{r_{o}^{2}} r_{o}^{3} (\xi' / \rho)^{3} \\ &+ r_{o}^{3} (\xi' \eta'^{2} / \rho^{3}) \left[r_{o}^{2} - r_{o}^{2} (\xi' / \rho)^{2} \right] \left\{ \delta^{*} + \delta_{1}^{*} \left[r_{o}^{2} - r_{o}^{2} (\xi' / \rho)^{2} \right] \\ &+ \left[r_{o}^{2} - r_{o}^{2} (\eta' / \rho)^{2} \right] \sum_{i \ge 0} c_{ij}^{*} r_{o}^{2i+2j} (\xi' / \rho)^{2i} (\eta' / \rho)^{2j} \right\} \end{aligned}$$
(C-6)
$$&+ r_{o}^{3} (\xi' / \rho)^{3} \left[r_{o}^{2} - r_{o}^{2} (\xi' / \rho)^{2} \right] \\ &\times \left\{ \omega^{*} + \left[r_{o}^{2} - r_{o}^{2} (\xi' / \rho)^{2} \right] \sum_{i \ge 0} d_{i}^{*} r_{o}^{2i} (\xi' / \rho)^{2i} \right\} \end{aligned}$$

Next, α and β are introduced as defined in Equations (C-3), and powers of r are canceled in Equation (C-6), where appropriate, to obtain

$$\begin{aligned} \mathbf{f}^{*}(\alpha, \ \beta) &= \mathbf{r}_{o} \ \gamma^{*} \alpha + \mathbf{r}_{o} \ (1 - \gamma^{*}) \ \alpha^{3} \\ &+ \alpha \beta^{2} \ (1 - \alpha^{2}) \ \mathbf{r}_{o}^{5} \left[\delta^{*} + \delta_{1}^{*} \ \mathbf{r}_{o}^{2} \ (1 - \alpha^{2}) \\ &+ \mathbf{r}_{o}^{2} \ (1 - \beta^{2}) \ \sum_{\substack{i \ge 0 \\ j \ge 0}} \ \mathbf{c}_{ij}^{*} \ \mathbf{r}_{o}^{2i+2j} \ \alpha^{2i} \ \beta^{2j} \right] \\ &+ \alpha^{3} \ (1 - \alpha^{2}) \ \mathbf{r}_{o}^{5} \left[\omega^{*} + \mathbf{r}_{o}^{2} \ (1 - \alpha^{2}) \sum_{\substack{i \ge 0 \\ i \ge 0}} \ \mathbf{d}_{i}^{*} \ \mathbf{r}_{o}^{2i} \ \alpha^{2i} \right] \end{aligned}$$
(C-7)

Equation (C-7) may be simplified ultimately by introducing the definitions (Equations (C-5)) of δ^* , ω^* , and δ_1^* to obtain:

$$\begin{aligned} \mathbf{f}^{*}(\alpha, \ \beta) &= \mathbf{r}_{0} \gamma^{*} \alpha + \mathbf{r}_{0} (1 - \gamma^{*}) \ \alpha^{3} + \alpha \beta^{2} (1 - \alpha^{2}) \ \mathbf{r}_{0}^{5} \\ &\times \left[\frac{\mu_{1}^{*} - \mu^{*}}{2\mathbf{r}_{0}^{4}} + \frac{\mathbf{r}_{0}^{2}}{\mathbf{r}_{0}^{2}} \left(\frac{\gamma_{1}^{*} - \gamma^{*}}{\mathbf{r}_{0}^{4}} - \frac{\mu_{1}^{*} - \mu^{*}}{2\mathbf{r}_{0}^{4}} \right) (1 - \alpha^{2}) \\ &+ \mathbf{r}_{0}^{2} (1 - \beta^{2}) \sum_{\substack{i \ge 0 \\ j \ge 0}} \mathbf{c}_{ij}^{*} \ \mathbf{r}_{0}^{2i+2j} \ \alpha^{2i} \ \beta^{2j} \right] \\ &+ \alpha^{3} (1 - \alpha^{2}) \ \mathbf{r}_{0}^{5} \left\{ \frac{3 - 2\gamma^{*} - \mu^{*} - 2\mathbf{r}_{0}^{4} \left[(\mu_{1}^{*} - \mu^{*})/2\mathbf{r}_{0}^{4} \right]}{2\mathbf{r}_{0}^{4}} \right. \end{aligned}$$
(C-8)
$$&+ \mathbf{r}_{0}^{2} (1 - \alpha^{2}) \ \mathbf{r}_{0}^{5} \left\{ \frac{3 - 2\gamma^{*} - \mu^{*} - 2\mathbf{r}_{0}^{4} \left[(\mu_{1}^{*} - \mu^{*})/2\mathbf{r}_{0}^{4} \right]}{2\mathbf{r}_{0}^{4}} \right\} \end{aligned}$$

Equation (C-8) is simplified by factoring the expression by r_0 and carrying r_0^4 through the terms in box brackets to obtain

$$\begin{aligned} \mathbf{f}^{*}(\alpha, \ \beta) &= \mathbf{r}_{0} \begin{cases} \gamma^{*} \alpha + (1 - \gamma^{*}) \ \alpha^{3} + \alpha \beta^{2} \ (1 - \alpha^{2}) \left[\frac{1}{2} (\mu_{1}^{*} - \mu^{*}) \right. \\ &+ \left(\gamma_{1}^{*} - \gamma^{*} - \frac{1}{2} (\mu_{1}^{*} - \mu^{*}) \right) (1 - \alpha^{2}) + (1 - \beta^{2}) \sum_{\substack{i \ge 0 \\ j \ge 0}} \mathbf{c}_{ij}^{*} \ \mathbf{r}_{0}^{2(i+j+3)} \ \alpha^{2i} \ \beta^{2j} \end{bmatrix} \end{aligned} (C-9) \\ &+ \alpha^{3} \ (1 - \alpha^{2}) \left[\frac{1}{2} (3 - 2\gamma^{*} - \mu^{*} + \mu^{*} - \mu_{1}^{*}) + (1 - \alpha^{2}) \right. \\ &\times \sum_{i \ge 0} \ \mathbf{d}_{i}^{*} \ \mathbf{r}_{0}^{2(i+3)} \ \alpha^{2i} \end{bmatrix} \Biggl\}$$

Three new constants, M, Γ , and Ω_1 , are introduced to replace δ^* , ω^* , and δ_1^* through the following definitions

$$M = \frac{1}{2} (\mu_1^* - \mu^*)$$

$$\Gamma = \gamma_1^* - \gamma^* \qquad (C-10)$$

$$\Omega_1 = \frac{1}{2} (3 - 2\gamma^* - \mu_1^*)$$

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By comparing Equation (C-10) with Equation (C-5), it is seen that these new constants are considerably simpler in form.

Noting that r_0 is strictly a constant, $r_0 = 1/\sqrt{3}$, new coefficients C_{ij} and D_i are defined as follows

$$C_{ij} = c_{ij}^{*} r_{o}^{2(i+j+3)}$$

$$D_{i} = d_{i}^{*} r_{o}^{2(i+3)}$$
(C-11)

Using Equations (C-10) and (C-11) in (C-9), the following is obtained

$$f^{*}(\alpha, \beta) = r_{0} \begin{cases} \gamma^{*} \alpha + (1 - \gamma^{*}) \alpha^{3} + \alpha \beta^{2} (1 - \alpha^{2}) \\ \times \left[M + (\Gamma - M) (1 - \alpha^{2}) + (1 - \beta^{2}) \sum_{i \ge 0} C_{ij} \alpha^{2i} \beta^{2j} \right] \\ + \alpha^{3} (1 - \alpha^{2}) \left[\Omega_{1} + (1 - \alpha^{2}) \sum_{i \ge 0} D_{i} \alpha^{2i} \right] \end{cases}$$
(C-12)

This completes the derivation of the transformed expression for f^* . Equation (C-12) is given as Equation (3-42) in Section 3.3.1 and its numerical advantages are discussed in Section 3.5.

APPENDIX D - APPROXIMATION OF THE DIRECT AND INVERSE TRANSFORMATIONS

This appendix presents tables of coefficients for the approximate forms of the transformations f and f* as presented in Section 3.4. These tables are presented for one octant of one face. Each table contains 977 sets of coefficients (ordered by column), determined by using intersection points of a 32×32 triangular grid. Each set of coefficients represents the approximate form of the transformation within one element of the 32×32 grid. Each of these elements corresponds to one 64×64 element data face record at the 1.2 nautical mile scale.

Figure 3-13 illustrates the table octant (octant 1) and the seven other octants. Table D-1 is used to transform from (x, y) to (ξ, η) and Table D-2 is used to transform from (ξ, η) to (x, y). The steps outlined below are followed to transform from (x, y) to (ξ, η) :

- 1. Add 2048.0 to x if x is less than 2048.0
- 2. Add 2048.0 to y if y is less than 2048.0
- 3. If y is greater than x, interchange x and y values
- 4. Truncate the x and y values obtained, divide each by 32, and add one, use these integers and IX and IY indices and look up the coefficients and number of terms in the table
- 5. Evaluate $f_{\alpha}(x, y)$ to obtain ξ
- 6. Reverse steps 3, 2, and 1 to place ξ in its proper octant

7. Interchange x and y and repeat steps 2 through 7 to obtain η Alternatively, steps 1, 2, 3, and 7 may be avoided by constructing a full 64×64 table containing redundant data. Table D-1 (1 of 7)

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00 -4.7
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0-= 00
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366P3C*3-00 11
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11810 30
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00
-3+8+08-0-00 -1.1511
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Table D-1 (2 of 7)

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	2024.00	CONSTRACTOR	20.00	00-0002	00.0145	5488.00	2688.00	5438.30	2752.00	00-25/2	00.244	00.25.12	2752.00	2752.00	2416.00	2916-00	2010-011-2	2010-0112		SK16.00	00-9192	2410-00	2440.09	2 160.00	00-0442	2460-00	00.00th	00-0052	200000000	C-0422 00-4442	00.444	00***62	2444.00	00.4445	00.44455	2000-000	3008-00	00-000	3008.00	00-9000	3008-00
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Table D-1 (3 of 7)

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Table D-1 (5 of 7)

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COPY AVAILABLE TO DDC DOES NOT Permit fully legible production Table D-1 (6 of 7)

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COPY AVAILABLE TO EDG DOES NOT PERMIT FULLY LEGIBLE PRODUCTION

Table D-1 (7 of 7)

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COPY AVAILABLE TO DDC DOES NOT PERMIT FULLY LEGIBLE PRODUCTION

Table D-2 (1 of 7)

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COPY AVAILABLE TO DDC DOES NOT PERMIT FULLY LEGIBLE PRODUCTION Table D-2 (2 of 7)

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COPY AVAILABLE TO DDC DCES NOT PERMIT FULLY LEGIBLE PRODUCTION

Table D-2 (3 of 7)

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COPY AVAILABLE TO DDC DOES NOT PERMIT FULLY LEGIBLE PRODUCTION Table D-2 (4 of 7)

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0.712605	0.726776	341111-0	0.74331E	0.763526	10.110.00	30508.0	0.430715	344436.0	0.82094E	0.82347E	361228-0	0.629015	305766.0	0.94 370E	0.34994E	3061 A8-0	3-8585AE		3.97 303E	162 16 .0	193466	0- 73406E		3.211296	0.735455	0.942995	351656-0	0.964105	0.9730*E	0-9751AF	0. 37733E	556696-0	0-10456	0-10431E	0-10392E		0.104156	0.104776	0-105655	311001-0	0-107405	0-107676	0-107465	0-107026
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COPY WALLY LEGIBLE PRODUCTION

Table D-2 (5 of 7)

COPY AVAILABLE TO BDC BUES NOT PERMIT FULLY LEGIBLE PRODUCTION

Table D-2 (6 of 7)

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00 - 222	776.00	225-00	776.00	176.00	775-00	775.30	776-30	776-00	770.00	00-04	776.00	773.00	00-922	776.00	00.04	775.00	775.00	340.00	60°646		340.00	00-041	00.000	140-00	00-0-6	00-046	140.00	00.010			840-00	340.00	00.004	00-045	00.000	0000	904.00	004-00	00. 100	00.004	904.00	904.00	00-100	00-100	00- 106	00-+00		00-00
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COPY AVAILABLE TO DDC DOES NOT Permit fully legible production

Table D-2 (7 of 7)

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A 304E-02 0	20-39198				C 10-39003	07215-02 0	4344E-02 0	C 20-34664	0 20-35996		20-35-11	13635-02 0	20-12550	19256-02 0	5 144E-02 0	20-1650		0 20-35-00 4	0 20-39620	20-34146	20-11640	0020E-02	20435-02		0 20-31.449	20-27194	20-1104 S	0 - 0 - 19 19 19 19 19 19 19 19 19 19 19 19 19	20-3055S	0 00-31900	0 10-10001-0 00		75445-33 0	27855-03 0	CO-12016	01105-02 0	0 20-31810	C-19653	0	53256-97 0	0 60-35+99	1 10-341Ct	79356-03 0		12 11 10 10 10 10 10 10 10 10 10 10 10 10	77.345-93 9	0 - 0 - 0 - 0 - 0	04105-01 0	79206-03 0		5407E-03 0		2424E-02 0		143	L G B X		KEPT	
0-00	B-0-00				A.D 00	00 0.1	00 0.1	1.0 00	1.0 00		N-0 00	2.2	1.0 00		100 00			0-00	1.0-00	1-0-00		-C-00	S-2-00	-0-0C	OC	6*0-0C		0-00	4-0-00	0000			9.0 00	2.0 00	6.00	00 00	1.0 00			40 2.5			2-0-00	0-00		-0-00				0-00	6-0-00				0000				1
0.42340	E 12.3.0	0.02200	1001	A SCOT TO	WACKE -0	0. 392.01	0. 391 665	9	r I		04	n 1	n	n		nr		3-3-5-5-6	0.582545	S. 1585 *0		ALL 495 -0	0-4954.70	THE ADD TO THE	32+141-0	15-005-0		100000	0.61 34-0	0.417426	0.576.57		0. 7755 35	0.1.2.	224.44	0.572406	0-571455		0.55.651	0. 36730	167.47.00	10.000.*0	0.566836	0.1702-0	0.556286	0-56307	0.57369	0-321-010	200 - 10 - 0	0.575.00	0.57558	0.575975							
-	•••				•	•	•			1.		•		0					*				•		•	+ -	6r		•	1	• •	1		-	•	1	-	, ,	•	•	• •	е. Т		9		1	r. 41	+	13 F	•••	P.	P) P			CNC2 - 0	.1=4			1-0.
3643-20	00-211		100.00		112.03	2175.00	224 3. 30	2334.00	1363.70			2307.30	0	00"11.02				CC-HOCT	CC. CTOE		00-0022	11.1.13			1341.00	CC*6.446		100.0.00	1504.70			04.94.42	274 0.30	210.00	1364.10		236 3.03	CC-8261		20. 7.00	00*0662	60-560E	3672.00	25	22	4	3456.10	2	3648.70		00-04EC	3534-30	00-2004	- 1		INDS= M2DRWI	024 450	11.10 1015 1015 1015 1015 1015 1015 1015	
3204-00				00-1001	CO.ESET	1068.00	1108.00	1458.00	00.595		00-19-1	00-E3+1	CO	CO. 544	00-101			1455.00	1967.00	CC . 295.		0095	3.53.00		0-1-0-1			126.4.0	0.8601			CO.CLC.	C0.280.	A032.23		00-7664	0.52.04	• 232 • 0		+032.00	.032.00	00-200	-032-00			9		3	20	4032.00	9				50	VOL SER N	3	SVS1 FORT	
-			0				•		e 1			-		185 195	1	Ľ	1		-	* :		Ľ		1		::		2	2	i	- *	1	•	*	**	1			- 14		-		1												•		_	-	_
2			-P		F		2	-	-		-						-		E			-	=	1				1	5	1			P.						-, nu - (9)	1	n,		1				- 1997 11 (Pr)		3 (1) h Mh		0				LEF2A	16F285	16.53	LEF2L	LET S

COPY AVAILABLE TO DDC DOES NOT PERMIT FULLY LEGIBLE PRODUCTION

In Tables D-1 and D-2, IX and IY (IP, IE) are the x and y (ξ, η) indices of the table, x_0 and y_0 (ξ_0, η_0) are the coordinates of the origin of the area of definition for the coefficient set, NTERMS is the number of terms to be evaluated, $a_0 \dots a_5$ are the coefficients, and RESID is the largest residual found in the evaluation of the coefficients (in elements) as illustrated in Figure 3-12.

APPENDIX E - LISTINGS OF TEST AND EVALUATION PROGRAMS

E.1 SUMMARY

This appendix contains listings of FORTRAN programs used for testing and evaluating the transformations in this report and plotter programs used for generating the world maps.

Listing	Figure
Program used for evaluating the coefficients of $f^*(\xi,\eta)$	E-1
Program used for verifying the f* transformations	E-2
Program used for verifying the polar transformation	E-3
Program used for evaluating coefficients of approximation	
polynomials	E-4
Program used for generating world maps	E-5
Alternative subroutines for E-5 using the polar transformation	E-6

THIS PROGRAM EVALUATES THE COFFETCIENTS OF THE INVERSE TRANSFORMATION F STAR BY FORCING FSTAR TO MAP POINTS TRANSFORMED BY THE DIRECT TRANSFORMATION BACK ONTO THEMSFLVFS PROGRAMMER - E. NICHAEL O'NEILL COMPUTER SCIENCES CORPORATION 8728 COLESVILLE RO. SILVER SPRING, MD.

 Implifing the implified interview in the implified interview in the implified interview in the implified interview in the implified interview int MFMATEA MATEA MEI MEA CALL STEPIT(FUNK) CALL DISP(N) STOP END

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Figure E-1 (1 of 3)

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SUBBRILTINE FILME (A-M, N-Z)

(MMRM/RILAN/ DIFF.VAR(2C)

(MMRM/RILAN/ DIFF.VAR(2C)

(MMRM/RILAN/ DIFF.VAR(2C)

(MMRM/RILAN/ DIFF.VAR(2C)

(MMRM/RILAN/ DIFF.VAR(2C)

(MMRM/RILAN/ DIFF.

(MMR/RILAN/ DIFF.

(MMRM/RILAN/ DIFF.

(MMRM/RILAN/ DIFF.

(MMR/RILAN/ DIFF.

(MMR/
```

Figure E-1 (2 of 3)

THIS PROGRAM VERIFIES THE CORRECTNESS OF THE INVERSE TRANSFORMATION BY COMPARING POSITIONS ON A TRANSFORMED AN PETRANSFORMED GRID PROGRAMMER - F. MICHAEL O'NEILL COMPUTER SCIENCES CORPORATION 8728 COLESVILLE RD. SILVER SPRING, MD. 1MPL1CIT RFAL#R (A-H,O-Z) COMMON RSPH,RO,RO2 RSPH=1,000 PO=PSPH/DSORT(3,000) PO2#RO#RO OMAYE-1,000 YEO,000 ISKIP=16 SKIP=16 ISK ID=14 CY ID=15K ID NN 5N IY=1,2049,ISK IP Y=0,000 YY=0,047,2048,000 P=FNMAD[XY,90 F=FNMAD[XY,90 F=FNMAD[YY,XX] Xn=[FMAD2[F,D]#2048,000/R0] Yn=fmAD2[F,D]#2048,000/R0] Yn=FMAD2[F,D]#2048,000/R0] Va=(FMAP2(F;P)+2048,000/280) NY=X-X8 NY=X-X8 N=NSNRT(NX+#2+0Y##2) [F(N=1T.0MAX) GO TO 50 WPTTF(4,10) X;Y;P,F;X8+Y8+0X+0Y;0 NPTTF(4,10) X;Y;P,F;X8+Y8+0X+0Y;0 WPTTF(1,15) X;Y,P,F) WPTTF(1,15) X;Y,P,0 15 FOPMAT(1X,3F15,8) DMAYD 15 ENDMAT() NMAY=0 60 Y=Y+SV IP 100 X=X+SK IP STOP

 Atom of the second A2=A+A A3=A+A7 A4=A3+A7 A4=A3+A7 R2=R+R P4=R2+R7 MA2#1,000-A2 FMAP2#04(A+GAMMAS+A3+0MGMS+A*R2*0MA2*(CGAM+A2*(EM-CGAM)+().00-R2) *(C00+C10+A2+C01#R2+C))*A2*R2+C20*A4+C02*R4)) +A3*0MA2*(C0MG1-0MA2*(00+01*A2)) #FT10RM A4=A50+4A50 R50=B#R R4=R50+R50 FDMAP=CAMA#A+DMCAM/R50#A3+ A3#RMA#{DMFGA+RMA#{D0+D}#A50+D2#A4}}+ A3#RMA#{DFLTA+{R50+R5}} A#R50+RMA#{DFLTA+{R50-R50}} RETURN RETURN END

Figure E-2

E-5

COCOCCC THIS PROGRAM VERIFIES THE POLAR FORM OF THE TRANSFORMATION PROGRAMMER - F. MICHAFL DIMETLL A728 COLESVILLE RD. SILVER SPRING, MD. COMPUTER SCIENCES CORPORATION IMPLICIT REAL #R (A-H. N-7) NIMENCIAN ((11,11), (11,1), (07/11,1) ₽, SPPOJ(A, R, C, D, E)=A+(B+R+C)/((C+C+D+D+E+E')++).5001 ٢, 00=),000/05087(3,000) Y0=-00 Y0=y0 Y1M= 00/5,000 Y1M=710 FY=0.10-6 DY=0Y e. Y=Y0 DO 10 ,=1+11 wDITE(A,1001) 1,1,4,5P1(1,1),5P2(1,3),5S n01 EnpMAT(2]5,4022,14) \$ v=v+vIN 10 x=v+xIM cTnD EAN FINCTION AREA(P1,P2,P3) TWOIT(TT REAL #8 [A-H,0-7) DIMENCION P1(2), P2(2), P3(2) ACOMP(1)=P3(1)=P2(2), P3(2) ACOMP(1)=P3(1)=P2(2), P3(2)=P3(2)]=P2 CCOMP(1)=P3(1)=P2(2)=P3(2)=P3(2)]=P2 CCOMP(1)=P3(1)=P2(2)=P3(2)=P3(2)]=P2 CCOMP(1)=P3(1)=P2(2)=P3(2)=P3(2)]=P2 CCOMP(1)=P3(1)=P2(2)=P3(2)=P3(2)]=P2 CCOMP(1)=P3(1)=P3(2)=P3(2)]=P3(2) ſ.

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Figure E-3 (1 of 2)



Figure E-3 (2 of 2)

```
THIS PROGRAM USES THE IRM SSP ROUTINE SIMO TO SOLVE A SYSTEM
OF SIMUTAMEDUS FOUNTIONS TO OBTAIN COEFFICIENTS FOR THE
APPPOXIMATE TRANSFORMATION FA
                              FUNCTION THAP HAY BE HODIFIED TO CALL ANY DESIRED TRANSFORMATION 5. E STAP. OR THE POLAR TRANSFORMATION SUBROUTINE MAP
                              PROGRAMMER - F. HICHAFL O'MEILL
COMPUTER SCIENCES CORPORATION
R728 COLESVILLE RD. SILVER SPRING, MD.
                    R72R CDLESVILLE RD. SILVER SPRING,

DIMENSION RVECT(5)

DATA NUTTA/, MPIC/512/

DEFN(5,10) NOUT.NPIC.MAXD

GROWAT(215)

DATA TOLEQ/1.0/

INTED=2048/MPIC

DV=MV

V=2048.0

DD 100 IX=1.NITER

V=2048.0

DD 100 IX=1.NITER

V=2048.0

DD 100 IX=1.NITER

V=2048.0

DD 100 IX=1.NITER

V=2048.0

DT 100 IX=1.0

   10
     40
100
                       10
                              END

SUBDOUTINE VER3(X,Y,NX,NY,RVECT,TOLER,RESID,KV)

DIMENSION XP(4),YP(4),RVECI(5)

DIMENSION XP(4),YP(4),RVECI(5)

DATA XP/0.25,0.75,0.75,0.25/, YP/0.25,0.25,0.75,0.75/
      NUEA

KVEO

RCTOEON

DO TO TEL.NV

VYEY+OYAXP(I)

VYEY+OYAXP(I)

XAEYY*RVFCT(I)+YY*RVECT(2)+RVFCT(3)

XTETHAPIXY.YY)

DEGITEARG(XA-XT)

TE(DEGIT.GT.PESID) RESID=RESIT

TE(DEGIT.GT.TOLER) KV=1

DETIDY
                               NIV=4
                                END
```

Figure E-4 (1 of 3)

E-8

```
PETIDA
                            ENO
                          SINPROLITINE VER4(X,Y,NX,NY,RVECT,TOLER,RESID,KV)
DIMENSION XP(4),YP(4),RVECT(5)
DATA XP/0,25,0,75,0,75,0,25/, YP/0,25,0,25,0,75,0,75/
  NV=4

KV=7

RECTD#D.D

DD 10 T=1.NV

Y = Y + DY #XD { 1

Y = Y + DY #XD { 1

Y = Y + DY #YV F(T(1)+XX #RVFCT(2)+YY#RVECT(3)+BVFCT(4)

YT=TWAP(XX.YY)

PEC(1T=ARC(YA-XY)

TE(DFCTT_GT_RECTO) RESID=RESIT

10 (FWT1NUE

16(DFCTT_GT_TO(FR) KV=1)
                           ALV= 4
                           TEIRESIN.GT.TOLER) KV=1
[[[]RESID.GT.ITLER; KV=;
Cr/IDN
FHD
SINDONTIFIE DEGSINC.AVECT.X.Y.DX.DY.TDLER.RESID;
DIMENCID4 AVECT[5]. AVECT[25]
DIMENCID4 AVECT[5]. AVECT[25]
DIMENCID4 AVECT[5].YP[5]
DIMENCID4 AVECT[5].YP[5]
DIMENCID4 AVECT[15].YP[5]
AVECT[15].YP[5].YV
AVECT[15].YP[5].YV
AVECT[15].YP[5].YV
AVECT[15].YP[5].YV
AVECT[15].YP[5].YV
AVECT[15].YP[5].YV
AVECT[15].YP[5].YV
AVECT[15].YV
AVECT[15].
                         ETTERN

ETTERN

TITENSTON XP (4), YP (4), AVECT.TOLFR.RESID.KV)

DATA XP/0, 25.0.75.0.75.0.25/0.25.0.25.0.75.0.75/

NY=4
 NV=4

VV=0

RFC[N=n_0

PO 10 T=1,NV

VX=Y+NY*Y*P[]

VA=Y+NY*Y*P[]

VA=Y+NY*Y*P[]

YA=YY*YX*KVE[T(1]+YY*YY*RVE[T(2]+XX*RVE[T(3]+YY*RVF[T(4]+RVF[T(5]

XT=TMAP[XX,YV]

PFC[T=ARC[V4=XT]

IF(DFC[T_GT_RESID] RESID=RFS[T

IF(DFC[T_GT_RESID] RESID=RFS[T

IF(DFC[T_GT_RESID] RESID=RFS[T

RFT(RM)

FAD
                         END
```

Figure E-4 (2 of 3)

E-9

all the second of Dear stines to

Contraction and the second states

r. PETIIDN END

all a

Figure E-4 (3 of 3)

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          THIS PROGRAM PRODUCES WORLD MAPS IN DATA BASE COORDINATES
         PROGRAMMER - F. MICHAEL O'NFILL
COMPLITER SCIENCES CORPORATION
4728 COLESVILLE RD. SILVER SPRING, MD.
 5 PD(.1)=DP(.1)
10 FF(.1)=FF(.)
                                $XR(IEASE)
        ALL PLOT (PP(I), EF(I), NN, * *)
CALL PLOT (PP(I), EF(I), NN, * *)
CONT TAUE
CONT TAUE
BEWIND 1
RETURN
  101
        FND
```

Figure E-5 (1 of 2)

SUR ROUTINE ROT (COORD. THETA) THIS SURROUTINE ROTATESONE DATUM POINT ABOUT THREE AXES DATA UAX/3,1.2,2.1.3.1.2.3/ DO TAX (S=1,3) Y=COORD(UAX(2,1AXIS)) TE(Y.EO.0.0) X=1.0E-6 Y=COORD(UAX(3,1AXIS)) IE(Y.EO.0.0) Y=1.0E-6 PUI=ATAN2(Y.X) PUI=PHI+THETA(1AXIS) P=SOPT(X+X+Y+Y) COOPD(UAX(3,1AXIS))=R*COS(PHI) COOPD(UAX(3,1AXIS))=R*SIN(PHI) POINTINUE PETIRN END SUBPOUTINE OTBAN(COORD.X.Y.IFACE) r, FURDOUTINE OTRANSCOORD, X, Y, JEACE) THIS SUBROUTINE TAKES AS INPUT THE CARTESIAN COORDINATE ARRAY (COORD) AND RETURNS DATA BASE COORDINATES (X AND Y) AND THE FACE NUMBER (JEACE) OF THE X, Y SYSTEM 00000000 PROGRAMMER - E. MICHAEL DINEILL COMPUTER SCIENCES CORPORATION 8728 COLESVILLE RD. SILVER SPRING, MD. C. C. C С, FNO END FUNCTION FWAP (AA, BB) 'S IS FUNCTION FMAP2 PROGRAMMER - E. MICHAEL O'NEILL COMPUTER SCIENCES CORPORATION A728 COLESVILE RD. SILVER SPRING, NO. THT 00000 1M0L1C1T REAL #A (A-H, N-7) NATA RONT3/0.1732050R0D1/ RO/0.57735026900/ NATA GAMMAS(N.1374R484773201/ NATA GAMMAS(N.1374R484773201/ NATA GAMMAS(N.1374R484773201/ NATA GAM/0.1374R484773201/ NATA GON/0.1411896311520/. CJ0/0.8097012865250-1/ NATA CON/0.1411896311520/. CJ0/0.8097012865250-1/ NATA CON/0.1411896311520/. CJ0/0.8097012865250-1/ NATA CON/0.178519673555700/. CJ0/0.16959546931400/ NATA CO/.-0.7591962004670-1/. D1/-0.217762490699D-1/ NATA DN/0.759196200467D-1/. D1/-0.217762490699D-1/ A=AA#PONT3 R=RA#RONT3 A=SA#A A7=A#A A3=A#A7 A4=A7#A7 R7=R#A A2=R⊅A A∠=P3#R2 Π¹¹A2=1,∩DD=A2 F¹¹AP2=R0#[A=GAMMAS+A3*ΠMGMS+A#R2#ΠMA2#[CGAM+A2#[FM=CGAM]+[],DD=R2] F¹¹AP2=R0#[A=GAMMAS+A3*ΠMGMS+A#R2#GOAA2#[CGAM+A2#[FM=CGAM]+[],DD=R2] F¹¹A2=1,A0 + A3*NMA2#[CDMG1=DMA2#[OD+D1#A2]]) FFTHRM FMD *

Figure E-5 (2 of 2)

E-12

```
SUBROUTINE OTRAN (COORD +X+Y+IFACE)
 THIS SUBROUTINE TAKES AS INPUT THE CARTESIAN COORDINATE APRAY
(COORD) AND RETURNS DATA BASE COORDINATES (X AND Y) AND THE
FACE NUMBER (IFACE) OF THE X+Y SYSTEM
                                                            MODIFIED TO USE POLAR FORM OF TRANSFORMATION
                                                           PROGRAMMER - E. MICHAEL DINEILL
COMPUTER SCIENCES CORPORATION
8728 COLESVILLE RD. SILVER SPRING, MD.
                                                    PFAL#A P.F.H.V

DIMENSION COORD(3), C(3), RUT(3,6), IROT(3,6)

DATA POT/1.0.1.0., -1.0.1.0.-1.0., -3.0.1.0.1.0.,

1.0.1.0.1.0., 1.0.-1.0.1.0., 1.0.1.0.1.0.

DATA ROT /2.3.1.2.3.1.1.3.2.1.3.2.7.3.2.1.3.2.

DATA ROTO /2.3.1.2.3.1.1.3.2.1.3.2.7.3.1.3.2.1.3.2.

DATA ROTO /2.3.1.2.3.2.1.3.2.7.3.1.3.2.1.3.2.1.3.2.

DATA ROTO /2.3.1.2.3.2.1.3.2.1.3.2.7.3.3.2.1.3.2.

DATA ROTO /2.3.1.2.3.2.1.3.2.7.3.1.3.2.1.3.2.1.3.2.1.3.2.

DATA ROTO /2.3.1.2.3.2.1.3.2.1.3.2.7.3.1.3.2.1.3.2.1.3.2.

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Figure E-6

E-13

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