

AD-775 235

INVERSE SCATTERING

Norbert N. Bojarski

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Prepared for:

Naval Air Systems Command

February 1974

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AD-775 235

REPORT DOCUMENTATION PAGE		READ INSTRUCTIONS BEFORE COMPLETING FORM
1. REPORT NUMBER N00019-73-C-0312/F	2. GOVT ACCESSION NO.	3. RECIPIENT'S CATALOG NUMBER
4. TITLE (and Subtitle)  INVERSE SCATTERING	5. TYPE OF REPORT & PERIOD COVERED FINAL REPORT, 1973	
	6. PERFORMING ORG REPORT NUMBER	
7. AUTHOR(s)  Bojarski, N. N.	8. CONTRACT OR GRANT NUMBER(s)  N00019-73-C-0312	
9. PERFORMING ORGANIZATION NAME AND ADDRESS Norbert N. Bojarski 16 Circle Drive Moorestown, New Jersey 08057		10. PROGRAM ELEMENT, PROJECT, TASK AREA & WORK UNIT NUMBERS
11. CONTROLLING OFFICE NAME AND ADDRESS Naval Air Systems Command ATTN. AIR-310B Washington, D. C. 20361		12. REPORT DATE February 1974
14. MONITORING AGENCY NAME & ADDRESS (if different from Controlling Office)		13. NUMBER OF PAGES 30
		15. SECURITY CLASS. (of this report) Unclassified
16. DISTRIBUTION STATEMENT (of this Report)  APPROVED FOR PUBLIC RELEASE; DISTRIBUTION UNLIMITED		15a. DECLASSIFICATION/DOWNGRADING SCHEDULE
17. DISTRIBUTION STATEMENT (of the abstract entered in Block 20, if different from Report)		
18. SUPPLEMENTARY NOTES		
19. KEY WORDS (Continue on reverse side if necessary and identify by block number)  Inverse Scattering  Reproduced by NATIONAL TECHNICAL INFORMATION SERVICE U S Department of Commerce Springfield VA 22151		
20. ABSTRACT (Continue on reverse side if necessary and identify by block number) The Physical Optics and Exact Inverse Scattering Solutions of this author are summarized. For the Physical Optics Inverse Scattering Method, shown are computer-reconstructed images of a sphere and cylinder from computed synthetic scattering data, as well as a sphere from experimentally measured data. For the Exact Inverse Scattering Method, shown are computer-reconstructed source distributions (currents) of a half-wave dipole antenna, a point-source, and two point-sources separated by one-half-wavelength, from computed synthetic scattering data.		

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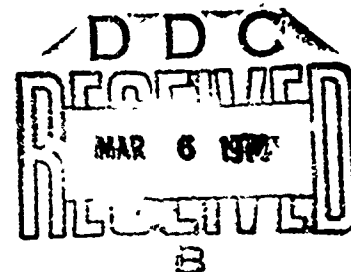
# INVERSE SCATTERING

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Final Report to  
Contract N00019-73-C-0312

February 1974



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WASHINGTON, D. C. 20360

## ABSTRACT

The Physical Optics and Exact Inverse Scattering Solutions of this author are summarized. For the Physical Optics Inverse Scattering Method, shown are computer-reconstructed images of a sphere and cylinder from computed synthetic scattering data, as well as a sphere from experimentally measured data. For the Exact Inverse Scattering Method, shown are computer-reconstructed source distributions (currents) of a half-wave dipole antenna, a point-source, and two point-sources separated by one-half-wavelength, from computed synthetic scattering data.

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## SECTION I

### INTRODUCTION

This report is a continuation of the final report [1] to the preceding contract, as well as to an earlier quarterly report [2] to this contract.

The theoretical results obtained to-date were reported on in detail in the above mentioned reports, and will thus be only briefly summarized in this report. The purpose of this report is to summarize the numerical-experimental results obtained to-date.

For the Physical Optics Inverse Scattering Method, shown are computer reconstructed images of a sphere and cylinder from computed synthetic scattering data, as well as a sphere from experimentally measured data.

For the Exact Inverse Scattering Method, shown are computer reconstructed source distributions (currents) of a half-wave dipole antenna, a point source, and two point sources separated by one-half wavelength, from computed synthetic scattering data.

## SECTION II.

### PHYSICAL OPTICS INVERSE SCATTERING

#### II.1. THEORETICAL RESULTS

The basic Physical Optics Inverse Scattering Identity [1], applicable to complete bandwidth and perspective (aspect angles) information, is

$$\gamma(x) = \text{Re} \int d^3\kappa e^{i\kappa \cdot x} \left[ \frac{\rho(\kappa)}{\kappa^2} \right] \quad (1)$$

where  $\kappa \equiv 2k$ , the range and phase normalized field-cross section  $\rho$  is related to the power cross section  $\sigma$  by  $\rho\rho^* \equiv \frac{\sigma}{4\pi^2}$ , and the characteristic function  $\gamma$  of the scatterer is

$$\gamma(x) \equiv \begin{cases} 1 & , \quad x \in V_S \\ 0 & , \quad x \in V_S^c \end{cases} \quad (2)$$

where  $V_S$  is the volume of the scatterer.

If the information aperture is incomplete (i.e., only finite bandwidth and/or incomplete perspective data are available), then (1) yields the following Fredholm Convolution Integral Equation of the First Kind [5]

$$\gamma(x) * \text{Re} a(x) = \text{Re} \int_A d^3\kappa e^{i\kappa \cdot x} \left[ \frac{\rho(\kappa)}{\kappa^2} \right] \quad (3)$$

where  $a(x) \leftrightarrow A(\kappa)$ , and where  $A(\kappa)$  is the characteristic information function, i.e.,

$$A(\kappa) \equiv \begin{cases} 1 & , \quad \rho(\kappa) \text{ known} \\ 0 & , \quad \rho(\kappa) \text{ unknown} \end{cases} \quad (4)$$

This integral equation (3) can be solved exactly numerically by a variety of existing techniques such as the matrix methods of Ritz-Galerkin [6], the associated Least Square Best Estimate method [7], and the associated moments method of Harrington [8], the Eigen-function expansion method of Toraldo Di Francia [9], leading to so-called super-resolution, and the k-space method of this author [10], which also leads to super-resolution (This solution will be summarized in Sect. III.1. of this report, since similar Fredholm convolution integral equations of the first kind arise in the exact inverse problem). Several closed-form solutions of (3) for apertures of specific geometry have been obtained by Lewis [11].

Since the unknown function  $\gamma(x)$  in the integral equation (3) is not a completely arbitrary function, but a characteristic function restricted to the form (2), the following closed-form approximate solutions [12] to (3) are obtained

$$\chi(x) \cong \text{Im} \int_A d^3\kappa e^{i\kappa \cdot x} \left[ \frac{\kappa_3 \rho(\kappa)}{\kappa^2} \right] \quad (5)$$

$$z(x_1, x_2) \approx \int dx_3 x_3 \chi(x) \quad (6)$$

where  $\chi(x)$  is a three-dimensional resolution density function which is a measure of the location of the surface of the scatterer, and  $z(x_1, x_2)$  is the geometrical function of this surface.



## II.2. NUMERICO-EXPERIMENTAL RESULTS

The solution to the integral equation (3) proposed by Lewis [13] was successfully numerically tested for a sphere by this author in 1969. This test consisted of a computer implementation of a special case version of this solution, applicable only to scatterers about which only a priori knowledge of cylindrical symmetry exists. This test essentially confirms the correctness of the basic inverse scattering identity (1) and the finite aperture integral equation (3). This solution was, however, not pursued further because of its inherent practical limitations. These limitations are the lack of generality of the required k-space aperture (i.e., the required aperture is impractical for physically realizable radar systems; which is not the case with this author's solution 5 and 6), the error enhancement introduced by the process of numerical differentiation of noisy data (vis-a-vis the error reduction resulting from the process of integration of such data in solutions 5 and 6), and the unapplicability of the Fast Fourier Transform (FFT) to this solution (which is essential if large amounts of data are to be processed in reasonable time by existing computers, yielding three-dimensional high-resolution descriptions of arbitrarily shaped scatterers about which no a priori knowledge of special geometry exists).

Solutions (5) and (6) were computer implemented with the aid of the FFT for arbitrarily shaped apertures, realizable with existing radar systems. This computer program was tested with the exact solution of Mie for scattering by a sphere, with a variety of band limited aspect angles and fractional frequency bandwidths, with the results shown in fig. 1 and 3. The ripples in fig. 1 are due to the very small three-dimensional raster of  $16^3$  data points for  $z(x_1, x_2)$ ; vis-a-vis the much larger two-dimensional raster of  $128^2$  data points for  $\chi(x_1, x_2, 0)$  in fig. 3, for which these ripples disappear. A full bandwidth three-dimensional display of a reconstructed cylinder is shown in fig. 2.

Solution (5) was also tested against experimentally measured (in anechoic chamber by The General Dynamics Corp., Fort Worth, Texas) scattering data from a test sphere; the results are shown in Fig. 4 and 5; in the latter figure the

the correct size of the sphere was added (it is this author's opinion that the small faint circular images are due to interference between the test sphere and its supporting strut).

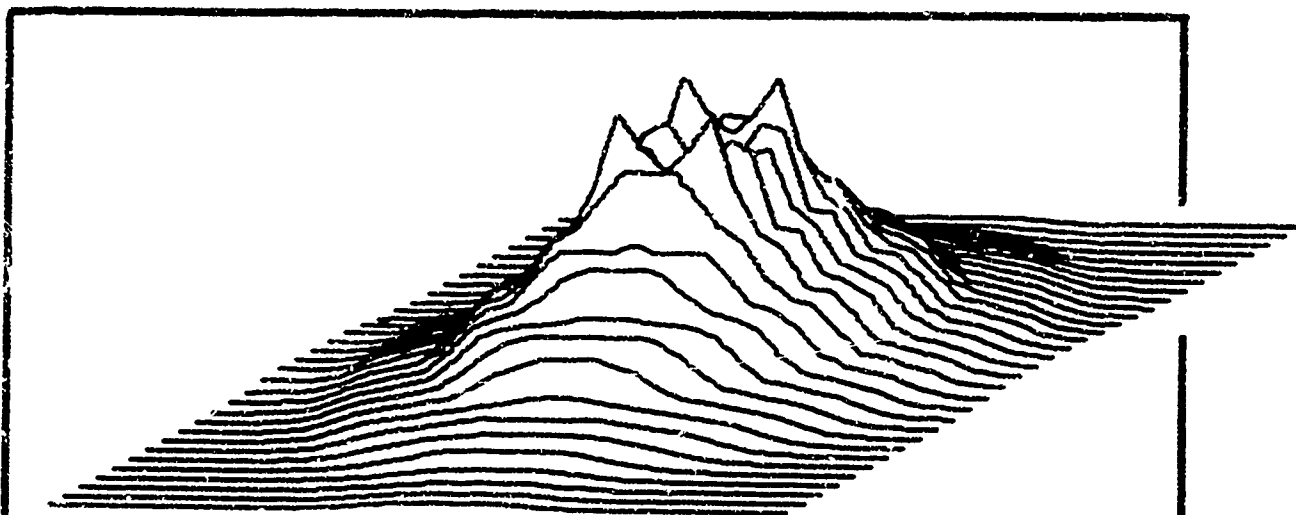


FIGURE 1  
THREE-DIMENSIONAL DISPLAY OF RECONSTRUCTED SPHERE

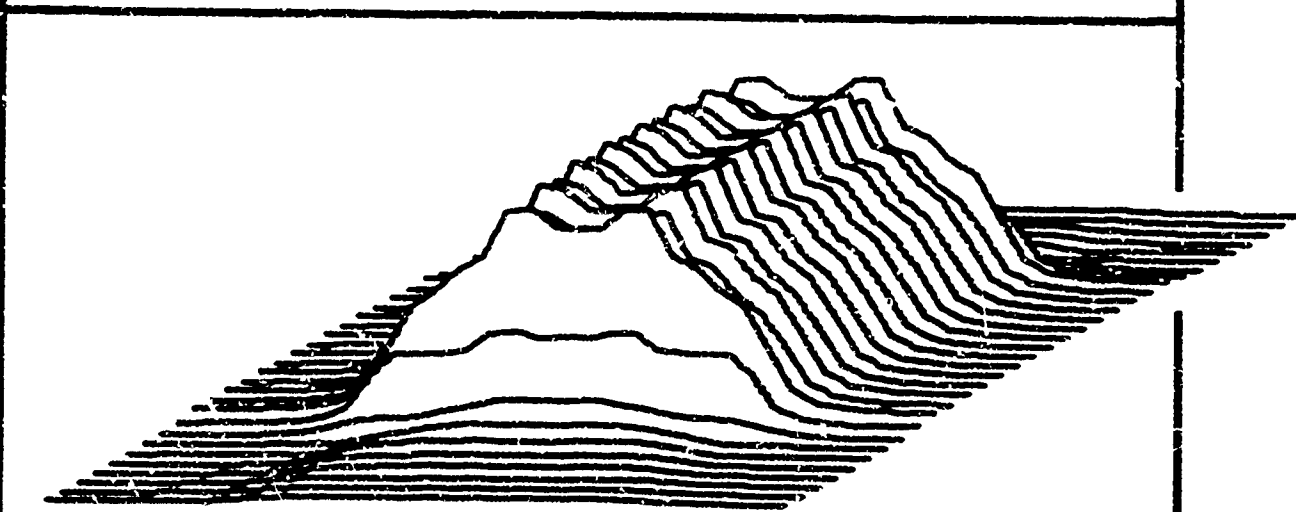
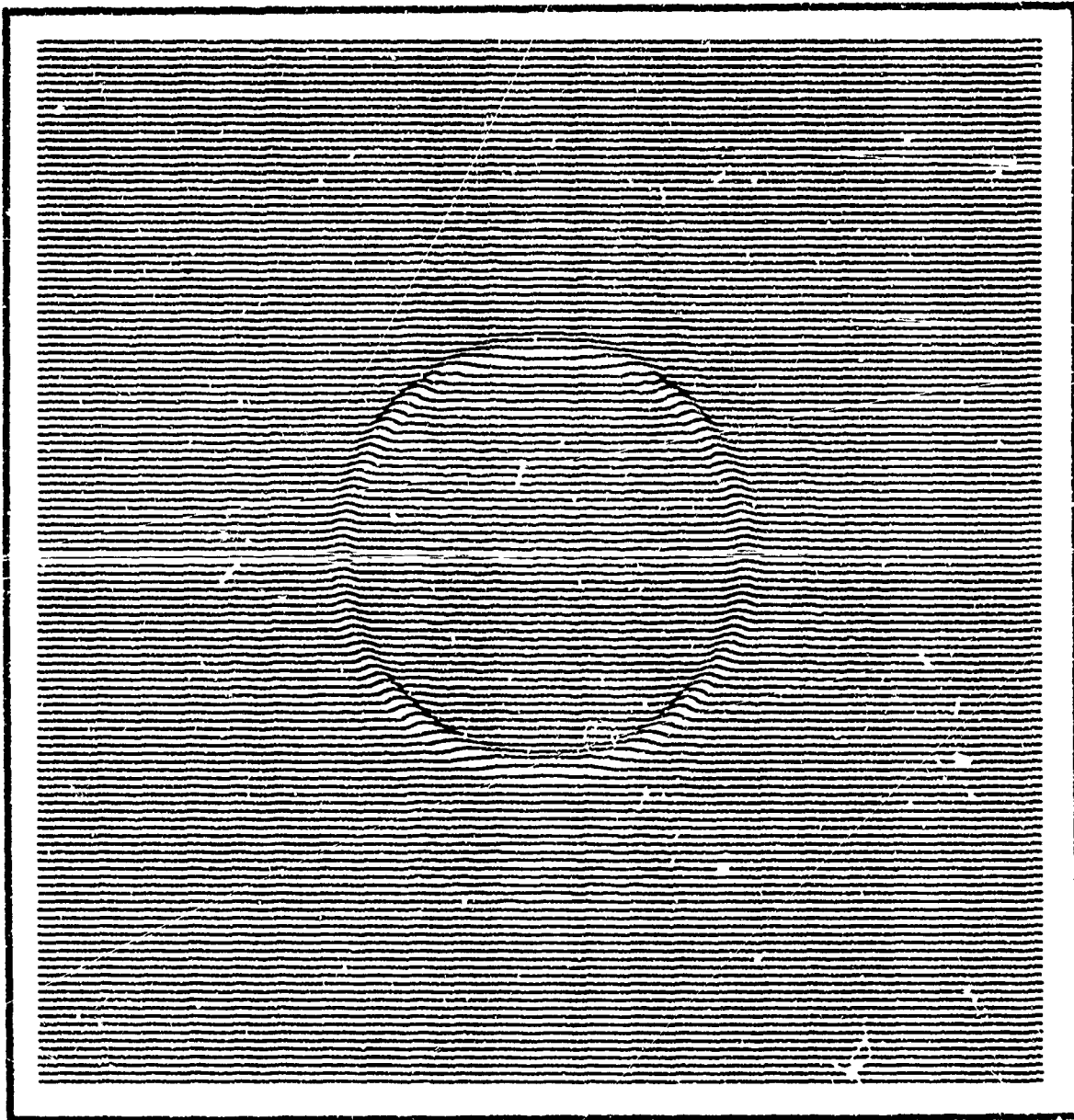


FIGURE 2  
THREE-DIMENSIONAL DISPLAY OF RECONSTRUCTED CYLINDER



**FIGURE 3**

**TWO-DIMENSIONAL DISPLAY OF RECONSTRUCTED SPHERE FROM SYNTHETIC  
SCATTERING DATA**

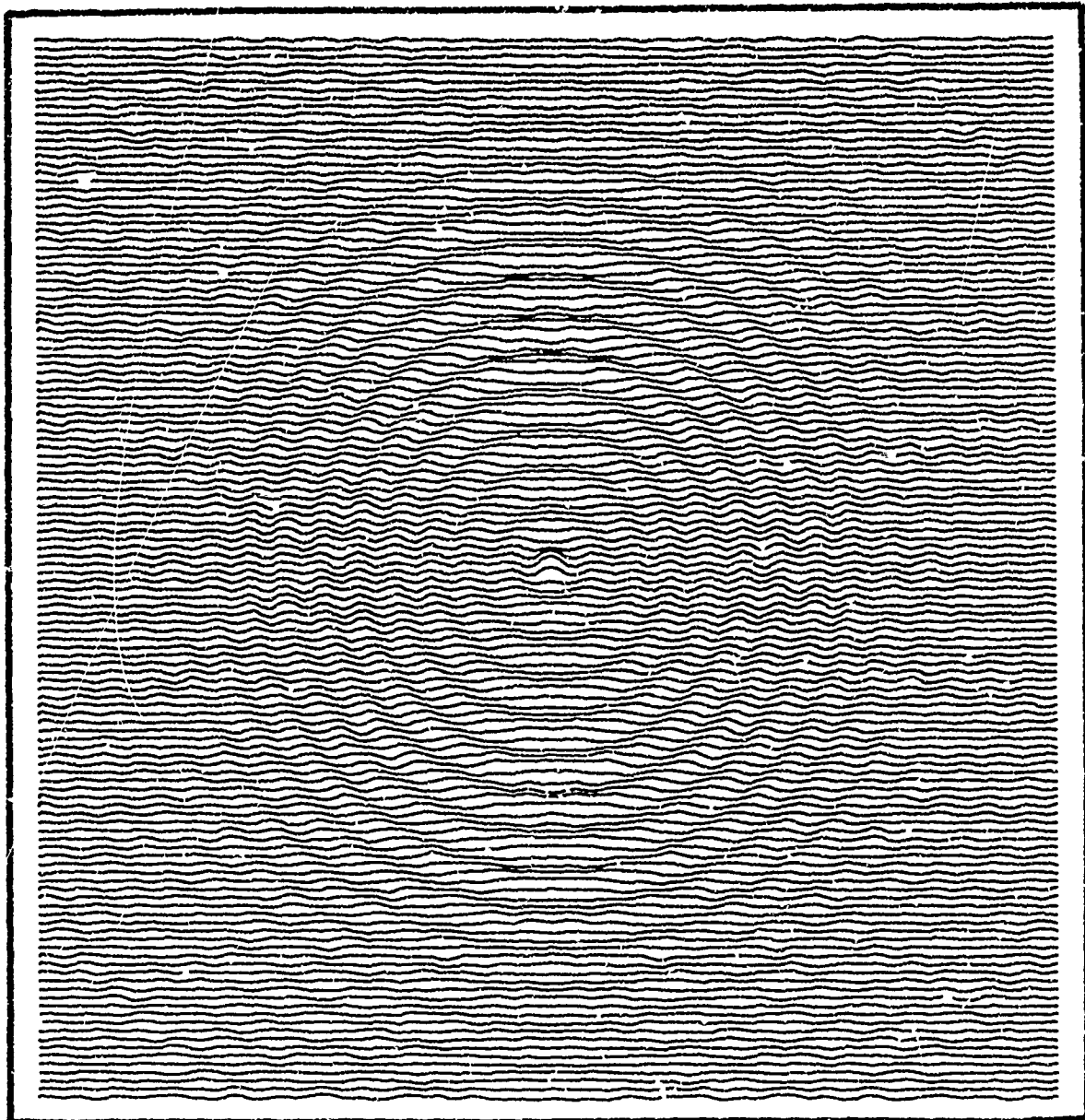


FIGURE 4

TWO-DIMENSIONAL DISPLAY OF RECONSTRUCTED SPHERE FROM MEASURED  
SCATTERING DATA

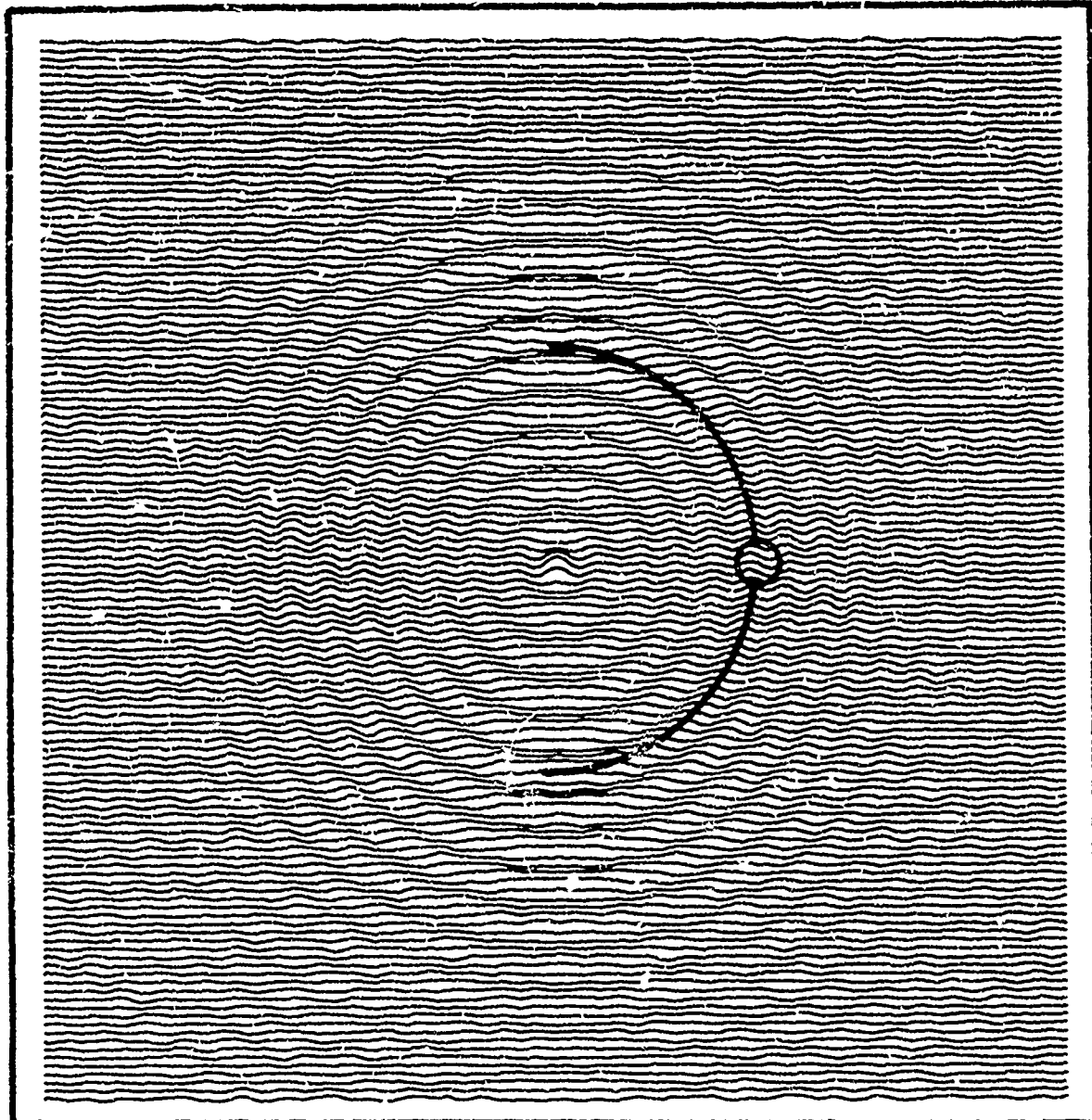


FIGURE 5

TWO-DIMENSIONAL DISPLAY OF RECONSTRUCTED SPHERE FROM MEASURED  
SCATTERING DATA

## SECTION III

### EXACT INVERSE SCATTERING

#### III.1. THEORETICAL RESULTS

The basic Exact Monochromatic Scalar Field Inverse Scattering Integral Equation [14], applicable to complete perspective information, is

$$\int_V dv \operatorname{Im} G \rho = \frac{1}{2i} \oint_S ds \cdot (G^* \nabla \phi - \phi \nabla G^*) \quad (7)$$

which, when  $\phi$  is measured in the far-field, reduces to

$$\int_V dv \operatorname{Im} G \rho = \int_{4\pi} d\Omega e^{-ik_s \cdot x} k_s \psi(k_s) \quad (8)$$

where  $\psi$  is the range and phase normalized scattered far-field.

For the incomplete perspective information aperture, (8) reduces to the following Double Fredholm Convolution Integral Equation of the First Kind

$$a(x) * \operatorname{Im} G(x) * \rho(x) = \int_A d\Omega e^{-ik_s \cdot x} k_s \psi(k_s) \quad (9)$$

The time-domain integral equation associated with (7) is

$$\theta(\mathbf{x}, t) = \frac{1}{2i} \int_V d^3x' \frac{\rho(\mathbf{x}, t-r/c)}{4\pi r} - \frac{1}{2i} \int_V d^3x' \frac{\rho(\mathbf{x}, t+r/c)}{4\pi r} \quad (10)$$

The basic equations (7), (8), and (9) are Fredholm Convolution Integral Equations of the First Kind, similar to (3), and can be solved by the methods mentioned in the paragraph subsequent to (3). This author's solution [15] to this integral equation is

$$\rho_{n+1}(\mathbf{x}) = \rho_n(\mathbf{x}) - \frac{1}{\lambda(\mathbf{x})} \int_{V_0} d^3x' \operatorname{Im} G(\mathbf{x}|\mathbf{x}') \rho_n(\mathbf{x}') + \frac{\theta(\mathbf{x})}{\lambda(\mathbf{x})} \quad (11.1)$$

$$\theta \equiv \frac{1}{2i} \oint_S ds \cdot (G^* \nabla \phi - \phi \nabla G^*) \quad (11.2)$$

Preliminary investigations of the time-domain integral equation (10) have been initiated by Prof. N. Bleistein [16].

For the electromagnetic vector fields, equations (7) through (11) have the following respective analogue

$$\int_V d^3x' \operatorname{Im} \nabla G(\mathbf{x}|\mathbf{x}') \times \mathbf{J}(\mathbf{x}') = \theta(\mathbf{x}) \quad (12)$$

$$\theta \equiv \frac{1}{2i} \oint_S ds [ \nabla G^* \times (\mathbf{n} \times \mathbf{H}) - \nabla G^* (\mathbf{n} \cdot \mathbf{H}) + i\omega \epsilon G^* (\mathbf{n} \times \mathbf{E}) ] \quad (13)$$

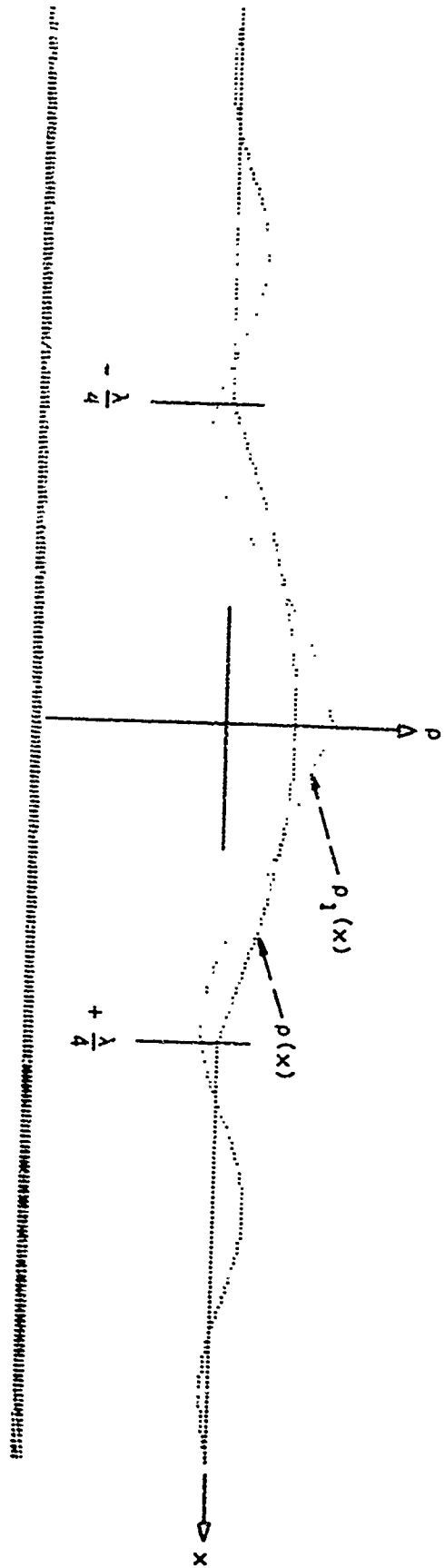
### III.2. NUMERICO-EXPERIMENTAL RESULTS

Figure 6 shows the first iteration synthetic computer reconstruction (computer and programmer provided by Dr. G. Tricoles, General Dynamics Corp., San Diego, California) of the source distribution (current) in a half-wave dipole antenna, as per (11). Since the errors are reduced by a factor of two per iteration, an order of ten iterations should suffice for most practical problems. It should be noted, however, that the first iteration already yields reconstruction of the source distribution beyond the Classical Rayleigh Diffraction resolution limit.

Figure 7 and 8 show similar results for the 8th, 16th, and 32nd iterations for a point source, and two point sources one-half wavelength apart, respectively.



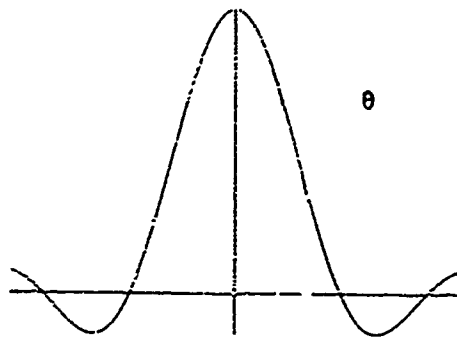
FIGURE 6  
RECONSTRUCTED HALF-WAVE  
DIPOLE ANTENNA



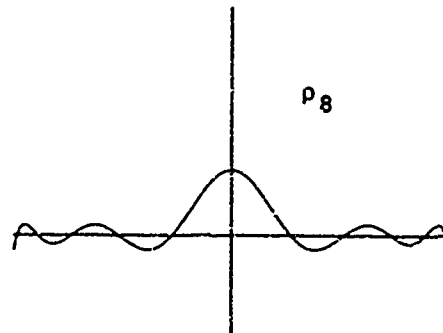
$\rho(x)$  True Source Distribution (Half-wave dipole antenna current).

$\rho_1(x)$  First iteration Reconstructed (Beyond classical Rayleigh diffraction resolution limit) source distribution.

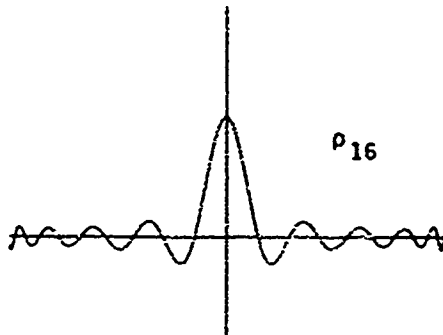
$$\rho(x) = \begin{cases} \cos(kx_1) \delta(x_2) \delta(x_3) & , \quad x_1 \in (-\frac{\lambda}{4}, +\frac{\lambda}{4}) \\ 0 & , \quad x_1 \notin (-\frac{\lambda}{4}, +\frac{\lambda}{4}) \end{cases}$$



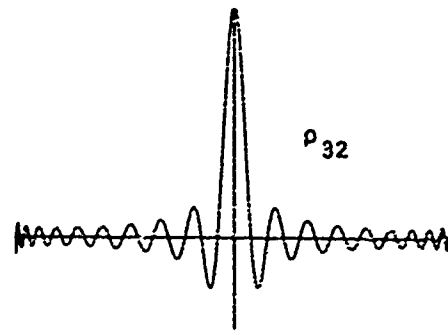
$\theta$



$\rho_8$



$\rho_{16}$



$\rho_{32}$

FIGURE 7

RECONSTRUCTED POINT SOURCE

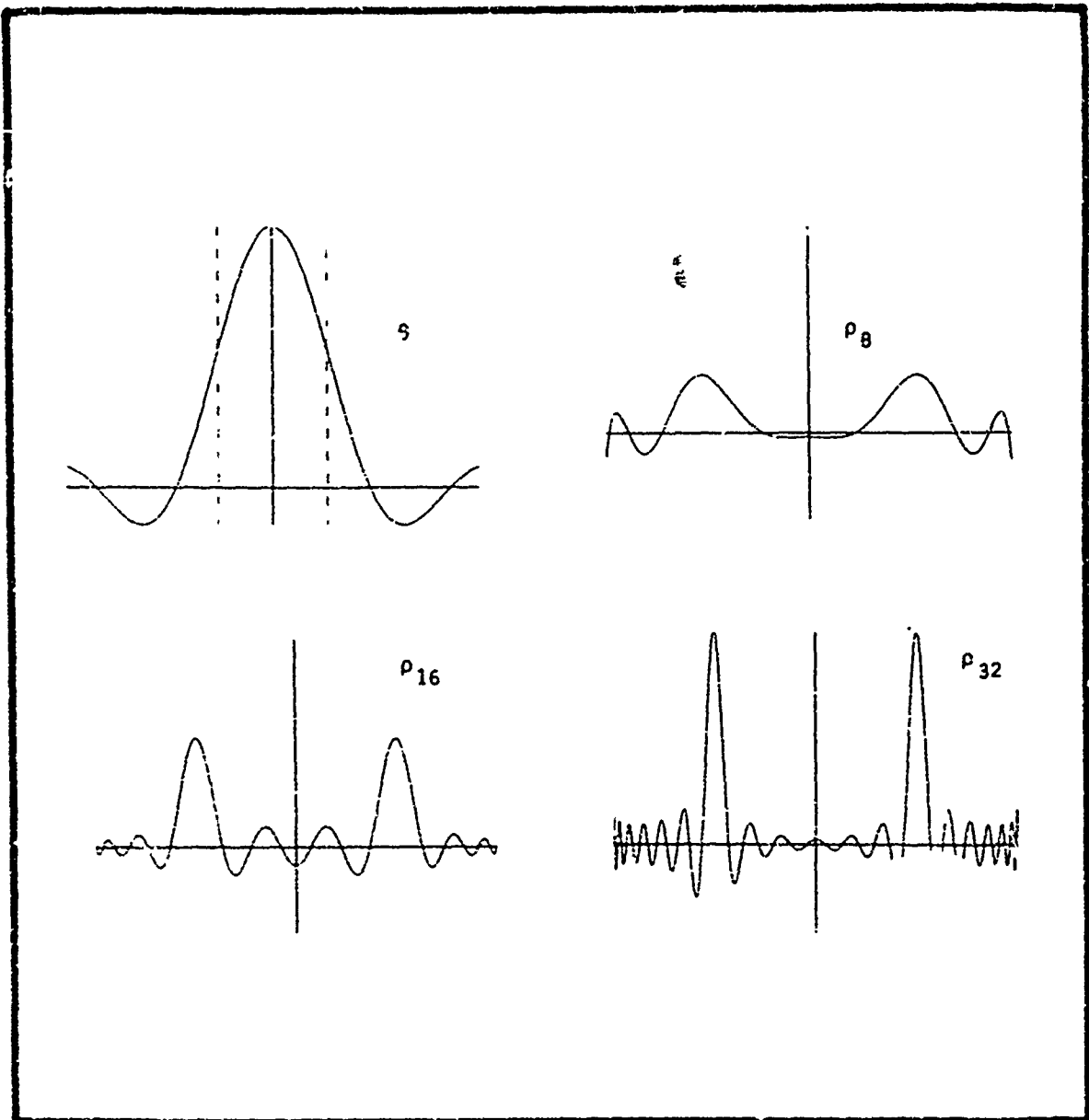


FIGURE 8

RECONSTRUCTED POINT-SOURCE-PAIR

## SECTION IV

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