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# FIREBALL PHENOMENOLOGY AND CODE DEVELOPMENT

Volume IV

SPUTTER Subroutines for Radiation Transport in Planes

General Atomic Division of General Dynamics Corporation

Special Nuclear Effects Laboratory

San Diego, California

Contract AF 29(601)-6492

TECHNICAL REPORT NO. AFWL-TR-65-143, Vol IV

July 1966

**AIR FORCE WEAPONS LABORATORY**

Research and Technology Division

Air Force Systems Command

Kirtland Air Force Base

New Mexico

Research and Technology Division  
AIR FORCE WEAPONS LABORATORY  
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FOREWORD

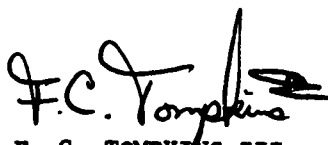
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This report is divided into six volumes as follows: Volume I, Summary and the Fireball Models; Volume II, Early Fireball Phenomena in the TIGHTROPE Event;\* Volume III, SPUTTER Subroutines for Radiation Transport in Spheres; Volume IV, SPUTTER Subroutines for Radiation Transport in Planes; Volume V, Material Properties; and Volume VI, Extensions of the Physics and Problem Areas.

The SPUTTER subroutines for radiation transport in planes described in Volume IV were developed by Dr. B. E. Freeman and Dr. C. G. Davis, Jr. The cooperation and contributions of Captains Milton Gillespie, William Whittaker, and George Spillman of the Air Force Weapons Laboratory are gratefully acknowledged.

This technical report has been reviewed and is approved.



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\* Volume II has been withdrawn, and will not be published.

ABSTRACT

The radiation-transport subroutines of the SPATTER code for plane slab geometry have been supplemented by alternative formulation based on integration along sampling ray paths through the slab. Angular integrations are performed by the Gaussian quadrature method which determines the ray angles. Options may be exercised to determine the number of angles and the nature of the radiation boundary condition at one boundary of the transport region. The characteristic ray code differs from the current integral method in performing problems having a large number of zones more rapidly and in having more general boundary conditions. For most applications a small number of angles give adequate accuracy. The numerical method used in the ray code is described. In addition, the organization of the code is discussed and subroutines are listed.

The SPUTTER code subroutines for radiation transport in planes described herein are as they existed on July 30, 1965. General Atomic has exercised due care in preparation, but does not warrant the merchantability, accuracy, and completeness of these subroutines or of their description contained herein. The complexity of this kind of program precludes any guarantee to that effect. Therefore, any user must make his own determination of the suitability of these subroutines for any specific use and of the validity of the information produced by their use.

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SECTION I  
INTRODUCTION

The new routines for radiation transport in planes closely parallel the spherical radiation transport subroutines (Volume III) in mathematical formulation and code organization. This parallelism is especially close between the SRADTN (for spheres) and PRADTN (for planes) subroutines in which calculations peripheral to the transport integration are performed. In fact, it is likely that these subroutines can be condensed to form a single subroutine for spheres and planes, although currently they are separate.

The routines reported here are to be considered as alternatives to the routines based on the integral formulation of the transport equation currently in use. By comparison, the current integral version is more accurate in the performance of angular integrations of the intensity, but for problems requiring a large number of zones it requires more computation time. Boundary conditions on the radiation intensity, however, are much more naturally incorporated into the new version.

Conditions suggesting a preference for using the new routines are:

- (1) a large number of zones and a desire to reduce calculation time and
- (2) the necessity of specifying a radiation intensity incident on the slab surface which has arbitrary angular and frequency dependence.

The numerical sequences used in solving the transport equation are discussed in Section II. A description of the diffusion approximation used in conjunction with the transport solution is given in Section III and a brief description of the methods of frequency integration is given in Section IV. Section V includes the actual code description in terms of code organization and economics. Section VI includes some initial studies on timing and accuracy in the angular integrations.

SECTION II  
NUMERICAL SOLUTION OF THE TRANSPORT EQUATION

The radiation routines described herein contain a formulation based on the numerical solution of the radiation transport equation along a selection of sampling rays through the slab. Relevant averages over the angular distribution are obtained by numerical quadrature, as described in Section 2.3, and the numerical solution of the transport equation along the photon ray is presented in Section 2.1. Criteria for selecting the sampling rays are discussed in Section 2.2. All of the derivations of this section apply to photons of a particular frequency; integration over frequency is discussed in Section IV.

The radiation transport equation in plane geometry that describes the changes in the specific intensity  $I_\nu$  of photons of frequency  $\nu$  resulting from pure absorption and emission according to the local thermodynamic equilibrium assumption is

$$\mu \frac{\partial I_\nu}{\partial x} = \sigma'_\nu (B_\nu - I_\nu) , \quad (2.1)$$

where

$$B_\nu = \frac{2h}{c^2} \frac{\nu^3}{e^{h\nu/\theta} - 1} ,$$

$$\sigma'_\nu = \sigma_\nu (1 - e^{-h\nu/\theta}) ,$$

and  $\sigma_\nu$  is the pure absorption coefficient. The scattering coefficient is assumed to be negligibly small compared to the absorption coefficient. Additionally, the retardation of the photons is neglected, as is valid when the radiation energy is small and temperatures change slowly. The resulting equation describes the quasi-steady intensity field resulting from the distribution of sources existing at a particular time.

Defining the monochromatic optical depth,  $\tau$ , as

$$\tau = \frac{1}{\mu} \int_0^x \sigma'_\nu dx , \quad (2.2)$$

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the values

$$B = a_- + b_- \tau, \quad \tau_{i-1} \leq \tau \leq \tau_{i-\frac{1}{2}},$$

where

$$a_- = \frac{B_{i-1} \tau_{i-\frac{1}{2}} - B_{i-\frac{1}{2}} \tau_{i-1}}{\tau_{i-\frac{1}{2}} - \tau_{i-1}}, \quad b_- = \frac{B_{i-\frac{1}{2}} - B_{i-1}}{\tau_{i-\frac{1}{2}} - \tau_{i-1}} \quad (2.4)$$

and

$$B = a_+ + b_+ \tau, \quad \tau_{i-\frac{1}{2}} \leq \tau \leq \tau_i,$$

where

$$a_+ = \frac{B_{i-\frac{1}{2}} \tau_i - B_i \tau_{i-\frac{1}{2}}}{\tau_i - \tau_{i-\frac{1}{2}}}, \quad b_+ = \frac{B_i - B_{i-\frac{1}{2}}}{\tau_i - \tau_{i-\frac{1}{2}}}.$$

For the case of a constant or step-function source, the source function  $B$  takes a value dependent on which interface of the zone is affected. If the left interface ( $\tau = \tau_{i-1}$ ) satisfies the criteria for a constant source,

$$B = B_{i-\frac{1}{2}}, \quad \tau_{i-1} \leq \tau \leq \tau_{i-\frac{1}{2}}.$$

If the right interface ( $\tau = \tau_i$ ) satisfies the criteria,

$$B = B_{i-\frac{1}{2}}, \quad \tau_{i-\frac{1}{2}} \leq \tau \leq \tau_i.$$

The integral of Eq. (2.3) can be evaluated with the interpolation function of Eq. (2.4) to give for the intensity

$$I_i = \alpha_{i-\frac{1}{2}} + [(I_{i-1} + \gamma_{i-\frac{1}{2}})e^{-\Delta/2} + \beta_{i-\frac{1}{2}}]e^{-\Delta/2}, \quad (2.5)$$

where

$$\begin{aligned} \alpha_{i-\frac{1}{2}} &= a_+ + b_+(\tau_i - 1), \\ \beta_{i-\frac{1}{2}} &= a_- - a_+ + (b_- - b_+)\left(\tau_i - \frac{\Delta}{2} - 1\right), \\ \gamma_{i-\frac{1}{2}} &= b_-(1 + \Delta - \tau_i) - a_-. \end{aligned}$$

In these expressions,  $\Delta = \tau_i - \tau_{i-1}$ . The coefficients of Eq. (2.5) can be re-expressed by using the definitions of Eq. (2.4):

$$\left. \begin{aligned}
 \alpha_{i-\frac{1}{2}} &= B_i - \frac{B_i - B_{i-\frac{1}{2}}}{\Delta/2}, \\
 \beta_{i-\frac{1}{2}} &= \frac{B_i - B_{i-\frac{1}{2}}}{\Delta/2} - \frac{B_{i-\frac{1}{2}} - B_{i-1}}{\Delta/2}, \\
 \gamma_{i-\frac{1}{2}} &= - \left( B_{i-1} - \frac{B_{i-\frac{1}{2}} - B_{i-1}}{\Delta/2} \right).
 \end{aligned} \right\} \quad (2.6)$$

The terms in Eq. (2.6) may be interpreted as containing combinations of numerical approximations to the values of the source function and the  $\tau$  derivative of the source function at the boundaries of the interval.

This form of the equation, in fact, can be obtained in another way starting from Eq. (2.3). Two successive integrations by parts transforms the expression for  $I_i$  into the following equivalent form:

$$I_i = \left( B - \frac{\partial B}{\partial \tau} \right)_i + \left[ I_{i-1} - \left( B - \frac{\partial B}{\partial \tau} \right)_{i-1} \right] e^{-\Delta} + \int_{\tau_{i-1}}^{\tau_i} \frac{\partial^2 B}{\partial \tau^2} e^{-(\tau_i - \tau)} d\tau, \quad (2.7)$$

in terms of values of the source function and the first two derivatives of the-source function with respect to  $\tau$ .

In an optically thin interval, the most important contribution arises from the terms  $I_{i-1}$  and  $B$ , which represent the transmitted intensity and the emission from the zone. The derivative terms cancel in this approximation; this is perhaps more directly indicated by Eq. (2.3). In the optically thick interval, which is the extreme opposite, only the first two terms evaluated at  $i$  are usually of significance. The terms from  $i-1$  are strongly attenuated and  $\partial^2 B / \partial \tau^2$  in the integral is usually small. In the limit, the diffusion approximation results from the term  $\partial B / \partial \tau)_i$ . Between limits, it is necessary to consider the integral term in Eq. (2.7).

If  $\Delta$  is not too large, a representative mean value of the exponential in the interval may be taken to give for the integral of Eq. (2.7)

$$\int_{\tau_{i-1}}^{\tau_i} \frac{\partial^2 B}{\partial \tau^2} e^{-(\tau_i - \tau)} d\tau \cong e^{-\Delta/2} \left[ \frac{\partial B}{\partial \tau} \right)_i - \frac{\partial B}{\partial \tau} \right)_{i-1} \right],$$

and thus the expression for intensity becomes

$$I_i = \left( B - \frac{\partial B}{\partial \tau} \right)_i + \left\{ \left[ I_{i-1} - \left( B - \frac{\partial B}{\partial \tau} \right)_{i-1} \right] e^{-\Delta/2} + \left[ \left( \frac{\partial B}{\partial \tau} \right)_i - \left( \frac{\partial B}{\partial \tau} \right)_{i-1} \right] \right\} e^{-\Delta/2} . \quad (2.8)$$

This expression has just the form of Eqs. (2.5) and (2.6) when the difference expressions are identified with the derivatives.

It is clear from the derivation of Eq. (2.5) that the resulting intensity is a positive quantity. With positive values for zone source functions, the linear interpolation expression assures that the integral contribution is always positive. Since the boundary intensity is always a positive quantity, the positivity of all intensities is assured.

In the diffusion approximation limit, only quantities at interface  $i$  will survive, and

$$I_i = \left( B_i - \frac{\partial B}{\partial \tau} \right)_i ,$$

which can be evaluated as

$$\frac{\partial B}{\partial \tau} = \mu \frac{\partial B}{\partial h} , \quad (2.9)$$

where

$$h = \int_0^x \sigma' dx .$$

The independent variable  $h$  depends only on  $x$ , so that angular integrations of  $I_i$  can be performed explicitly in the diffusion approximation, which takes account of the dependence on angle of Eq. (2.9). A difference approximation can also be based on this expression, assuming that  $B$  is linear in  $h$ , i. e.,

$$\left( \frac{\partial B}{\partial \tau} \right)_i \cong \frac{B_i - B_{i-\frac{1}{2}}}{h_i - h_{i-\frac{1}{2}}} \mu , \quad (2.10)$$

where  $\mu$  is the cosine of the angle which the ray makes with the slab normal. The corresponding equation for the intensity is Eq. (2.5), in which

$$\alpha_{i-\frac{1}{2}} = B_i - \nu \frac{B_i - B_{i-\frac{1}{2}}}{h_i - h_{i-\frac{1}{2}}},$$

$$\beta_{i-\frac{1}{2}} = \mu \frac{B_i - B_{i-\frac{1}{2}}}{h_i - h_{i-\frac{1}{2}}} - \mu \frac{B_{i-\frac{1}{2}} - B_{i-1}}{h_{i-\frac{1}{2}} - h_{i-1}}, \quad (2.11)$$

$$\gamma_{i-\frac{1}{2}} = - \left( B_{i-1} - \mu \frac{B_{i-\frac{1}{2}} - B_{i-1}}{h_{i-\frac{1}{2}} - h_{i-1}} \right)$$

### 2.1.2. Small-optical-depth Expansion

If the optical depth is very small, the intensity expression in Eq. (2.3) takes a much simpler form,

$$I_i = I_{i-1} + \left[ \frac{1}{4} B_i + \frac{1}{4} B_{i-1} + \frac{1}{2} B_{i-\frac{1}{2}} - I_{i-1} \right] \Delta. \quad (2.12)$$

Although this result is the limiting form of Eqs. (2.5) and (2.6), the terms must cancel through second order in an expansion in  $\Delta$  before the first surviving term, derived in part from the quadratic terms of the exponentials, is obtained. Consequently, for sufficiently small argument, the finite number of figures used in the exponential will render the result inaccurate. For the exponential from the IBM-7044 system, this restricts the argument to a number greater than  $\sim 2 \times 10^{-4}$ ; but with the lower-accuracy fast exponential (see Section V), the argument must be somewhat larger. Since the relative error approximately equals the argument of the exponential, the criterion for using Eq. (2.12) in the PTRANS subroutine is now set at  $\Delta \leq 2 \times 10^{-2}$ . With this value, the greatest relative error arising from the expansion and cancellation should be on the order of 1 percent.

### 2.1.3. Boundary Conditions

Integration of the transport equation to obtain intensities is performed through the thickness of a zone, called a "trans" region. At intersections of characteristic rays with the inner and outer surfaces of each layer it is necessary to supply the starting value of the intensity  $I_{i-1}$  required in Eq. (2.5). Three classes of boundary conditions occur:

1. The trans region outside boundary coincides with outside zones of the SPUTTER calculation and a prescribed function,  $I_0$ , is applied at the left boundary value:

$$I(X_{1A+1}, \mu) = I_0, \quad \mu \leq 0 \text{ or blackbody boundary condition.} \quad (2.13)$$



2. The right-hand boundary of the slab provides for reflective and transmittal boundary conditions as well as special routines to establish prescribed intensities for angles with  $\mu < 0$  at the boundary:

$$I(X_{IB}, -\mu) = I(X_{IB}, \mu), \quad I(B_{IB}, \mu) = 0,$$

or

$$I(B_{IB}, -\mu) = I_0(\mu, t), \quad (2.14)$$

where  $I_0$  is the prescribed negatively directed boundary intensity applied to the outer boundary as a function of angle, frequency, and time. Intensities at up to 50 frequencies and six angles can be accommodated in the table located in the array QINT1(N). The table entries are used as  $I_0$  and are formed in the subroutine QUE4 where they are stored in the QINT1 array. Since this subroutine is appropriate to the thermal interaction application, additional uses may require subroutines tailored to the specific application.

3. All other trans boundaries are bounded by regions in which the diffusion approximation is valid (see Section III). Consequently, the boundary surface intensities on contiguous trans regions inside or outside of a diffusion region are given by the diffusion approximation intensity derived in Section III:

$$I_{i-1} = B_{i-1} - \mu \left. \frac{\partial B}{\partial h} \right|_{i-1} \quad (2.15)$$

## 2.2. ANGULAR INTEGRATION

Integrals over the polar angle of the intensity are required, as described in Section 2.2.1, to carry the calculation forward in time and to provide edits of informative derived quantities. These are formed by numerical quadrature using the intensities evaluated at a series of discrete values of polar angle by the integrations described in Section 2.2. Since in the plane calculation the value of the polar angle remains fixed along a characteristic ray and enters only parametrically in the equations, it is possible to exercise a choice of polar-angle values in order to optimize the accuracy of the resulting integrals.

The numerical quadrature method used for the PTRANS subroutine is the so-called double Gaussian.<sup>(1)</sup> In this method the integrals of the radiation quantities (flux, energy, pressure, etc.) are approximated by

$$\int_0^1 f(\mu) d\mu = \sum_0^n A_m (If)_m,$$

where  $(I_f)_m$  is the known value of the integrand at a chosen value  $\mu_m$  of the cosine of the polar angle,  $\mu$ . In the method of Gaussian, not only are the coefficients  $A_m$  determined but the values of  $\mu_m$  are fixed to minimize the difference between the integral and the approximation. The result of this minimization is to relate the  $\mu_m$  to the zeros of the Legendre polynomial of order  $n + 1$ .

For those integrals having the range  $-1 \leq \mu \leq 1$  it is frequently advantageous to treat the forward and backward hemispheres separately to allow for the possibility of a discontinuity in  $I$  at  $\mu = 0$ . Such discontinuities or very abrupt changes in the values of the intensity between forward and backward directions may occur in systems which are transparent enough that strong source regions are accessible. In these cases, a better fit to the integrand is obtained by the two approximating functions which permit a discontinuity at  $\mu = 0$  than by a single approximating function which imposes a smooth behavior near  $\mu = 0$ . The method used in PTRANS, based on separate integration regions for  $-1 \leq \mu \leq 0$  and  $0 \leq \mu \leq 1$ , is called the double Gaussian quadrature method. Values of  $A_m$  and  $\mu_m$  are derived by a simple transformation from those for the single integration region. Since the angles for a single integration region are arranged symmetrically about the interval midpoint, for double region integration it is possible to identify pairs of angles  $\pm\mu_m$  having the same weight  $A_m$ . In Table 2.1, the values<sup>(2)</sup> for the  $0 \leq \mu \leq 1$  interval are recorded for values of  $n = 1, 2, 3, 4, 5$ . The total number of forward and backward angles,  $2n + 2$ , for each  $n$  (also equal to the total number of entries in the table of  $\mu_m$  and  $A_m$  for each  $n$ ) is also listed in the table.

The backward and the corresponding forward ray integrations in the PTRANS subroutine are performed sequentially. Since the same absolute values of  $\mu_m$  are required for these two calculations, many of the quantities formed in the backward integration pass can be used for the forward pass as well, and hence these quantities are saved to increase calculation efficiency. Contributions of the pair of forward and backward intensities to the weighted sums corresponding to the angular integrals are tallied at the same time that the forward integration pass is being calculated.

Table 2.1

GAUSSIAN WEIGHTS

$n = \text{LMDA}(37)-1$	Total No. of Angles, $2n+2$	$\mu_m = \text{RR}(\text{NMMU})$	$A_m$	$\mu_m A_m = \text{RR}(\text{NCS})$
1	4	0.2113248	0.5	0.1056624
		0.7886752	0.5	0.3943376
2	6	0.1127017	0.2777778	0.03130603
		0.5	0.4444444	0.2222222
3	8	0.8872983	0.2777778	0.2464718
		0.0694318	0.1739274	0.0120761
4	10	0.3300095	0.3260726	0.1076071
		0.6699905	0.3260726	0.2184655
5	12	0.9305682	0.1739274	0.1618513
		0.0469101	0.1184634	0.00555713
6	14	0.2307653	0.2393143	0.055225436
		0.5	0.2844444	0.1422222
7	16	0.7692347	0.2393143	0.1840889
		0.9530899	0.1184634	0.1129063
8	18	0.0337652	0.0856622	0.00289240
		0.1693953	0.1803808	0.03055566
9	20	0.3806904	0.2339570	0.0890652
		0.6193096	0.2339570	0.1448918
10	22	0.8306047	0.1803808	0.1498251
		0.9662348	0.0856622	0.0827698

SECTION III  
THE DIFFUSION APPROXIMATION

The radiation transport equation in the limiting case of an optically thick medium admits of the diffusion approximation in which the expression for the radiation intensity is greatly simplified; only the local properties affect the radiation intensity at the point in question. An expansion of the radiation source function  $B_\nu$  about the point  $r$  permits the intensity  $I_\nu(\mu)$  of the radiation field in the direction making an angle, whose cosine is  $\mu$ , with the linear direction to be formed.

3.1. DIFFERENTIAL FORM OF THE DIFFUSION FLUX

The general solution of the transport equation forms the starting point of the derivation. The integral expression for the intensity applicable to all geometries is

$$I(\tau) = \int_{-\infty}^{\tau} B(\tau') e^{-(\tau-\tau')} d\tau' ,$$

where  $\tau = \int_0^x \kappa_\nu \rho ds$ , in which  $\kappa_\nu$  is the monochromatic absorption coefficient (in  $\text{cm}^2/\text{g}$ ) at frequency  $\nu$ . By expanding  $B(\tau')$  in series about the point  $\tau$ , i. e.,

$$B(\tau') = B(\tau) + \frac{\partial B}{\partial \tau} (\tau' - \tau) + \frac{1}{2} \frac{\partial^2 B}{\partial \tau^2} (\tau' - \tau)^2 + \dots ,$$

the intensity becomes

$$I = B - \frac{\partial B}{\partial \tau} + \frac{1}{2} \frac{\partial^2 B}{\partial \tau^2} - \dots$$

or

$$I = B - \frac{\mu}{\kappa \rho} \frac{\partial B}{\partial x} + \frac{\mu^2}{\kappa \rho} \frac{\partial}{\partial x} \left( \frac{1}{\kappa \rho} \frac{\partial B}{\partial x} \right) - \dots$$

for plane slab geometry.

The diffusion approximation results from retention of only the first two terms, so that the diffusion intensity is

$$I = B - \frac{\mu}{\kappa\rho} \frac{\partial B}{\partial x} \quad (3.1)$$

and the monochromatic diffusion flux  $\phi_r$  and radiation energy  $E_R$  are

$$\phi_r = 2\pi c \int_{-1}^1 I \mu \, d\mu = -\frac{4\pi c}{3} \frac{1}{\kappa\rho} \frac{\partial B}{\partial x}, \quad (3.2)$$

$$E_R = 2\pi \int_{-1}^1 I \, d\mu = 4\pi B \dots$$

### 3.2. CRITERIA FOR THE SELECTION OF DIFFUSION REGIONS

The criteria for the validity of the diffusion approximation can be obtained by examination of the above derivation--namely, that the expansion of the source function be justified and that the expansion converge rapidly so that the neglect of all but the leading terms is valid. If the source function is linear in  $\tau'$  at the point in question and is also linear for a distance of the order of one mean free path on either side of the point, the criteria are satisfied. These criteria are difficult to quantify since they refer to a finite region containing the point in question. If all of the terms (or a large number of them) were checked for rapid convergence, this would imply (making a smoothness assumption) that the diffusion criterion is met. It is not possible with finite differences, however, to form the higher-order local derivatives approximations.

In the SPUTTER subroutine PRADTN, criteria designed to give an indication of both the local and nonlocal behavior have been employed. First, at the zone interface at which the intensity and flux are to be evaluated, the inequality

$$\left| \frac{\partial B}{\partial h} \right| \ll B \quad (3.3)$$

is required. In this expression  $h = \int \kappa\rho \, dx$  is the optical depth normal to the slab; the derivative is approximated by the centered first difference of  $B$  between adjacent zones. The resulting expression, of course, contains some nonlocal aspects resulting from the finite difference approximation, which ensures that when neighboring zones are optically thick, no nonlocal source perturbation is close enough to invalidate the diffusion approximation. However, to provide for the cases when a source perturbation is located a fraction of an optical depth from an interface meeting the condition of Eq. (3.3), the diffusion region is constricted. Starting from the closest

interfaces outside the diffusion region (where Eq. (3.3) is not satisfied), all of those interfaces lying within a prescribed number of mean free paths are removed from the diffusion region.

The criteria used in SPUTTER are controlled by input numbers. The criterion of Eq. (3.3) uses the input number HCB:

$$|TG| < HCB \times Y2, \quad (3.4)$$

where TG is the difference approximation to the gradient and Y2 is the source function evaluated at the interface by interpolation. The second criterion uses the input number HVB (in mean free paths). If

$$|Q3(I) - Q3(J)| > HVB, \quad (3.5)$$

then the interface with index J which satisfies Eq. (3.4) is removed from the diffusion region. In Eq. (3.5); Q3 is the normal optical depth and I is the index of the nondiffusion interface adjoining the diffusion region.

Although the diffusion calculation is considerably faster than the transport, the establishment of two transport regions separated by the single zone requires still more calculation to set up characteristic rays and perform bookkeeping operations. To avoid the duplicate setup calculations required for an additional transport region, a test is made to eliminate a diffusion region consisting of a single zone.

### 3.3. DIFFERENCE FORM OF THE DIFFUSION FLUX

The diffusion intensity derived above is

$$I = B - \frac{\mu}{\kappa\rho} \frac{\partial B}{\partial x}$$

In the group frequency approximation of SPUTTER, the intensity integrated over a frequency interval  $(\nu_j, \nu_{j+1})$  is required:

$$\int_{\nu_j}^{\nu_{j+1}} I \, d\nu = \int_{\nu_j}^{\nu_{j+1}} B \, d\nu - \frac{\mu}{\rho} \frac{\partial \theta^4}{\partial x} \int_{\nu_j}^{\nu_{j+1}} \frac{\partial B}{\partial \theta^4} \frac{d\nu}{\kappa_\nu}$$

In terms of the partial Rosseland mean absorption coefficient

$$\kappa_j = \frac{\int_{\nu_j}^{\nu_{j+1}} \frac{\partial B}{\partial \theta^4} d\nu}{\int_{\nu_j}^{\nu_{j+1}} \frac{\partial B}{\partial \theta^4} \kappa_\nu d\nu}$$

the frequency group intensity becomes

$$I_j = \int_{\nu_j}^{\nu_{j+1}} I d\nu = \int_{\nu_j}^{\nu_{j+1}} B d\nu - \frac{\mu}{\rho} \frac{\partial \theta^4}{\partial x} \frac{\int_{\nu_j}^{\nu_{j+1}} \frac{\partial B}{\partial \theta^4} d\nu}{\kappa_j} \quad (3.6)$$

It is desired to evaluate this quantity at each zonal interface in the mesh. Since the known quantities are the zone temperatures and densities, the absorption coefficients  $\kappa_j$  and the integrated source functions  $X_6 = \int B d\nu$  are first evaluated, not at the interfaces but at positions representative of each zone.

The question remains as to how best to approximate the derivatives and interpolate for the coefficients in Eq. (3.6) at the interfaces from the quantities available at zone positions. The answer depends on the temperature and density profile across the interface from which these terms could be calculated directly. Since the profile is not known, we must select a reasonable approximation which will permit the calculation to be carried out. In fact, the appropriate profile depends on the events which have taken place in the calculation and on the energy transport mechanisms of greatest importance in it. As extreme examples, a problem dominated by hydrodynamics might have quantities determined by passage of a strong shock and subsequent linearization in mass coordinates of the pressure behind the shock, whereas a radiation-dominated diffusion problem is characterized by linearity of the radiation potential, which, in turn, depends on the Rosseland opacity. Of course, such detailed information about the progress of a problem is generally unavailable, so, at best, an approximation based on over-all accuracy is needed.

Since the terms under consideration are the radiation diffusion equations, the interpolation is performed in a way to give greatest accuracy when the diffusion terms are most important--namely, when the profile is being determined entirely by radiation diffusion. It is also desirable to reduce the number of coefficients requiring interpolation. This can be done by noting the identity

$$\frac{\partial}{\partial r} \int_{\nu_j}^{\nu_{j+1}} B \, d\nu = \frac{\partial \theta^4}{\partial x} \int_{\nu_j}^{\nu_{j+1}} \frac{\partial B}{\partial \theta^4} \, d\nu$$

and by forming the variable  $\tau = \int \rho \kappa_j \, dx$ . In terms of these quantities, the intensity can be written as

$$I_j = \int_{\nu_j}^{\nu_{j+1}} B \, d\nu - \mu \frac{\partial \int_{\nu_j}^{\nu_{j+1}} B \, d\nu}{\partial \tau} .$$



SECTION IV  
FREQUENCY INTEGRATION

Equations derived in Sections II and III which are applicable to a particular frequency of the radiation field are of limited usefulness in the SPUTTER calculations. Although in principle a calculation at a particular frequency might be valuable for comparison with high-resolution spectroscopy, in practice no such data have been available. Of much more use are intensities averaged over a wide frequency band. These quantities can be compared with data from wide-band measurements and, most important of all, can be summed for use in the energy integration in the SPUTTER code. The quantities to be summed are the frequency-integrated radial flux component, the radiation energy density, and the radiation pressure. For performing interaction calculations, it is also valuable to form other components of the radiation flux.

Basically, the quantity which is required for each of the above applications is the frequency-group intensity  $I_{ij}$ ,

$$I_{ij} = \int_{\nu_j}^{\nu_{j+1}} I_i \, d\nu \quad (4.1)$$

Then, for example, this quantity can be integrated over angles to form  $\phi_{ij}$ , the contribution to the flux at position  $i$  of frequency group  $j$ :

$$\phi_{ij} = \int_{-1}^1 (I_{ij}^+ - I_{ij}^-) \mu \, d\mu \quad ,$$

and thus the total radiant flux at position  $i$  is

$$\phi_i = \sum_j \phi_{ij} \quad .$$

Equation (2.8) gives the expression for the frequency-dependent intensity to be used in Eq. (4.1). The frequency integration of Eq. (2.8) has been reported recently,<sup>(4)</sup> but the current SPUTTER code does not include the transmission functions. The first two terms of Eq. (2.8) which form the diffusion limit can be integrated, as in Section 3.3, to give

$$I_{ij} = B_{ij} - \frac{\mu}{\sigma_{Rj}} \frac{\partial B_{ij}}{\partial x} \quad (\text{diffusion limit}) , \quad (4.2)$$

in which the first term

$$B_{ij} = \int_{\nu_j}^{\nu_{j+1}} B_i(\nu) d\nu$$

is the frequency-group Planck function and the second term contains the frequency-group Rosseland mean absorption coefficient  $\sigma_{Rj} = \rho \kappa_j$ . In this form, Eq. (4.2) correctly gives the frequency-group intensity for the optically thick limiting case. The remaining  $B_i$  and  $\partial B / \partial \tau_i$  terms of Eq. (2.8) are formed in the same way. Thus,

$$I_{ij} = B_{ij} - \left( \frac{\mu}{\sigma_R} \frac{\partial B}{\partial x} \right)_{ij} + \left[ \left( \frac{\mu}{\sigma_R} \frac{\partial B}{\partial x} \right)_{ij} - \left( \frac{\mu}{\sigma_R} \frac{\partial B}{\partial x} \right)_{i-1,j} \right] e^{-\Delta/2} + \left[ I_{i-1,j} - B_{i-1,j} + \left( \frac{\mu}{\sigma_R} \frac{\partial B}{\partial x} \right)_{i-1,j} \right] e^{-\Delta} . \quad (4.3)$$

In Eq. (4.3), mean values of the exponentials have been extracted from the frequency integrals and the outstanding problem is to specify their values. Two options are available; they differ in the absorption coefficient used to calculate the optical depth. The first is

$$\overline{e^{-\Delta}} = e^{-\sigma_R \delta} \quad (4.4)$$

and the second is

$$\overline{e^{-\Delta}} = e^{-\sigma_P \delta} ,$$

where

$$\sigma_P = \frac{\int_{\nu_j}^{\nu_{j+1}} \sigma_\nu B_\nu d\nu}{B_{ij}}$$

and

$$\delta = x_i - x_{i-1} .$$

For small optical depth, the correct result makes use of the Planck mean absorption coefficient. From Eq. (2.12) the frequency integration then gives

$$I_{ij} = I_{i-1,j} + \left[ \frac{1}{4} B_{ij} + \frac{1}{4} B_{i-1,j} + \frac{1}{2} B_{i-\frac{1}{2},j} - I_{i-1,j} \right] \sigma_P \delta . \quad (4.5)$$

The above prescriptions for frequency-group means are far from satisfying and call for further work. Considerable economies can be made through reductions in the number of frequency groups if a more accurate means of averaging within groups can be found. Presently used choices of frequency groups appear to give a reasonably accurate result, however, as indicated by comparisons between calculations with the nominal number of frequency groups and calculations with a very large number of frequency groups. (It is expected that a unique correct result will be obtained as the number of frequency groups is increased, irrespective of the choice of the weighting function in the frequency-group-average absorption coefficient.) Consequently, a very few frequency groups should be adequate if a suitable averaging procedure were developed.

Even with a crude averaging scheme, considerable improvement in accuracy results from choice of frequency-group boundaries so as to reduce the variation of the absorption coefficient within the group.

Work on the absorption coefficient for air indicates that approximately 20 groups, carefully selected as to their locations, afford quite adequate resolution. Enough information is known about air to make this selection appear quite reasonable. Air absorption coefficient tapes (DIANE)\* have been prepared for 18, 20, and 90 groups. The 90-group tape is used to check on the frequency integrations at selected times. The proper averages to use are difficult to decide on at this time. There are provisions for reading into storage from the DIANE tapes both the Rosseland and Planck averages, which are used at present in the thick or thin limits, respectively.

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\* See Section VI of Volume V.

SECTION V  
SUBROUTINE ORGANIZATION AND ECONOMICS

The present plane transport subroutines were written with the idea of removing unnecessary calculations from inside the frequency loop and characteristic ray integrations. These improvements required an increase in storage for the subroutines to attain a decrease in calculational time. The reorganized subroutines will be discussed in two sections, corresponding to the two major subroutines: (1) the radiation subroutine (PRADTN) in which most of the preliminary setup and the diffusion calculation is completed and (2) the transport subroutine (PTRANS) in which the intensity calculation and angular integrations are performed. The subroutines which execute the opacity interpolations (KAPPA), Planck function (PLNKUT), and fast exponential (FREXP) will be discussed in Section 5.4. The input numbers and the output edits will be presented in Sections 5.5 and 5.6.

5.1. THE PRADTN SUBROUTINE

In PRADTN, the high-frequency groups are merged, a source region is established, boundary sources and derivatives are calculated, regions for transport and diffusion are formed, diffusion fluxes are calculated, frequency integration is performed, and the radiation time-step control is evaluated. Each of these activities in PRADTN will be discussed in subsequent paragraphs.

5.1.1. Merge Frequency Groups

Frequency groups that are too far out on the Planck tail for a "maximum" temperature in the mesh are merged. The criterion used is as follows: If the lower frequency boundary  $h\nu_1$  of the group in question ( $h\nu_1, h\nu_2$ ) is greater than ten times the maximum temperature (THMAX) in the mesh, this group will be merged with the next lower group. Merging will continue until over half the groups have been merged; at this point, either the calculation is terminated or a second DIANE tape is called. On merging, Rosseland and Planck averages are formed by using the following equation for  $dB/d\theta^4$  and the appropriate sums:

$$\frac{dB_\nu}{d\theta^4} \cong \frac{0.0384974}{\theta^4} \left[ \left( \frac{h\nu_2^4}{1 - e^{-h\nu_2/\theta}} \right) e^{-h\nu_2/\theta} - \left( \frac{h\nu_1^4}{1 - e^{-h\nu_1/\theta}} \right) e^{-h\nu_1/\theta} \right],$$

$$\sum b_j \theta^4, \sum \frac{dB_\nu}{d\theta^4}, \sum b_j \theta^4 \kappa_P, \text{ and } \sum \frac{dB_\nu}{d\theta^4 \kappa_R}. \quad (5.1)$$

The Planck weighting functions ( $b_j$ ) are obtained from PLNKUT, as described later. On completing the merging, the merged opacities are formed:

$$\overline{\kappa_R} = \sum dB_\nu / d\theta^4 / \sum dB_\nu / (d\theta^4 \times \kappa_R) \quad (\text{CAPAR}),$$

$$\overline{\kappa_P} = \sum b_j \theta^4 \kappa_P / \sum b_j \theta^4 \quad (\text{CAPAC}). \quad (5.2)$$

### 5.1.2. Set Up Sources and Derivatives

The frequency-dependent sources must be established at the interfaces from the zonal quantities  $b_j \theta_{i+\frac{1}{2}}^4$  (X6(i)) and  $\tau_{i+\frac{1}{2}}$  (H3(i)). The difference equations used were given in Section 2.1. Before the calculation of the Planck function ( $b_j$ ) is made, i. e., before calling PLNKUT, a test is made to see if  $u_1$  (i. e., the reduced frequency  $h\nu_1/\theta$ )  $\geq 19$ ; if so,  $b_j = 0$  (i. e., the source X6(i) = 0.0). If  $u_1 < 19.0$  and  $u_2 \leq 0.01$ , then  $b_j = 0$  also, assuming that for  $\theta^4 < 10^5$ , the small  $b_j$  ( $b_j \sim 10^{-5}$ ) will produce a negligibly small source contribution. An index (ICX) is set equal to the last zone that contains a source. This source index is used to limit the transport calculation to the region containing sources. While setting up the sources and derivatives, tests are made on their discontinuous nature to use either a linear or constant form in the intensity integrations. The initial check is on the minimum optical depth of adjacent zones to ensure that both are transparent (less than 0.3). If this condition holds and if both the sources and optical depths are changing rapidly in  $x$  (change greater than a factor of two), the derivative at that interface (TC(i)) is set equal to zero. The zero source derivative is used in PTRANS, as a test, to set up the constant source terms. For the intensity integration, special boundary sources and derivatives are also established at the edge of the source region ( $I = \text{ICX}$ ) and at the outside of the mesh ( $I = \text{IM}$ ) (see Section 2.1.3).

### 5. 1. 3. Determine Diffusion Region

The principal criterion for defining a diffusion region is that the first derivative of the source function (TG) be small compared to the source (Y2) (see Section 3. 2). When the zone is found to be diffusion, the boundary is tagged by setting (X3 = -1. ). Before incorporating this interface into a diffusion region, the possible influence from sources on either side is considered and a further test is made. From the last diffusion boundary, a test is made for an optical depth in succeeding zones to the left. If more than HVB optical depths appear in the next zone, then this zone is calculated by transport and removed from the diffusion region (set X4(i) = -1. 0). HVB is an input number, which is usually around 5. When x = 0 is reached after testing each zone, zones out to the right of the present transport region are tested in the same manner. The above test buffers the transport region with an (HVB) mean-free-path-thick diffusion boundary. If the zone boundary stays diffusion, i. e. , X3(i) = -1. 0 and X4(i) = 0. 0, a diffusion flux is calculated from the source gradients, as described in Section 3. 1. The regions where X3(i) = 0. or X3(i) = -1. and X4(i) = -1. have been established as transport regions because they did not meet the diffusion criteria or they reverted to transport regions by the optical-depth test described above. This transport region is then identified by setting the left boundary to IAX and the right boundary to IBX. More than one trans region may be set up in PRADTN, and if so, a PTRANS calculation will be made for each region. No one-zone diffusion region is allowed and the region outside the sources (I > ICX) is always considered a transport region.

### 5. 1. 4. Time-step Control and Monofrequency Calculation

These two aspects of the new code are related since the "grey" absorption coefficients from the DIANE tape are used to estimate a radiation time step as well as to form the monofrequency time-dependent calculation. In the multifrequency calculation, after all groups have been processed, an additional call for KAPPA is made to read in the grey absorption coefficients. These averages were obtained by integrating the frequency-dependent absorption coefficients for both Planck ( $\kappa_P$ ) and Rosseland ( $\kappa_R$ ) in the DIANE code. The actual time step for radiation transfer is then obtained from the formula

$$\Delta t_R = (0.5 + 1.5 H3(i)^2) / (ac \kappa_R \theta^3) \times CV(i) , \quad (5.3)$$

where CV(i) is the specific heat and  $ac = 4.12 \times 10^{12}$ . The mass point in question is also checked to ensure that it will not gain or lose more than half its original energy:

$$\Delta t_R = 0.5 \times CV(i) \times \theta(i) \times G(i) / |ER(i)| , \quad (5.4)$$

where  $\bar{E}R(i)$  is the divergence of the flux and  $G(i)$  is the mass in the zone. The minimum of these values is compared to the hydro time step (Courant) and if smaller,

$$NRAD = \text{FIX}(DTH2/DTRMIN) \quad \text{and} \quad DTR = DTH2/NRAD \quad (5.5)$$

is set to cycle NRAD times through the radiation routine.

The monofrequency calculation also uses the grey absorption coefficients from the DIANE tape. If  $KMAX = 0.0$  and  $S15 = 1.0$ , the frequency-averaged opacities are bypassed on the tape and only the grey absorption coefficients are read into storage. For succeeding cycles,  $S15$  is set equal to zero and the interpolations for  $\kappa_R$  and  $\kappa_P$  are performed in KAPPA using the stored opacities originally read into KAPPA's common storage. When the problem is restarted it is therefore necessary to reload  $S15$  equal to one. If the DIANE tape is not designated (the tape unit assigned must be stored in  $AMASNO(J+17)$ , where  $J$  is the material number), then the KAP routine is called (KAP8 for air) and used for the monofrequency calculation.

## 5.2. THE PTRANS SUBROUTINE

The subroutine PTRANS is called by PRADTN to carry out the intensity integration between IAX and IBX, saving various quantities on the inward pass that will be used on the outward pass as well as the angular integration of the flux between rays ( $\int_{-1}^1 I_\mu d\mu$ ). After the intensity transport along a typical ray in the outward direction ( $iA \rightarrow iB$ ) is done, the flux is calculated while the inward pass of the intensity calculation is being completed. The angular integration is based on a linear interpolation of the intensities between rays. The logic in PTRANS is described in detail in the following sections.

### 5.2.1. Selection of Angles

At present, only five sets of Gaussian angles and weights are stored in the subroutine. These can be selected by setting an input number (LMDA(37), the number of angles with  $\mu > 0$ ) to the desired  $n+1$ . The selection from storage is made from the following indices

$$NY = \text{LMDA}(37) - 1$$

$$NMU = (NY-1) \times (NY+2) + 1$$

$$NGS = NMU + NY + 1$$

$NMU$  selects the cosine of the angle ( $\mu_m$ );  $NGS$  selects the relation ( $\mu_m A_m$ ), the cosine of the angle times the Gaussian weights for the flux formulation.

### 5.2.2. Intensity Integration along Characteristic Rays

The integration using Eq. (2.11) starts at the left boundary with the appropriate boundary condition and proceeds outward, storing the exponentials  $e^{-\Delta\tau}$  in (H4(i)), the derivatives  $\mu \partial B/\partial h)_{i+\frac{1}{2}}$  in X8(i), and the calculated intensities in sum X3(i). The more general boundary conditions are established (see Section 2.1) and the stored quantities are now used except for the change of sign of  $\mu \partial B/\partial h)_{i+\frac{1}{2}}$  in Eq. (2.1) to calculate the intensities I(F2), on the outward pass.

The regions where constant sources, and therefore zero boundary derivatives, should be used in the intensity integrations were established in PRADTN by setting TG(i) equal to zero. In the integration along a particular ray, a test is made on TG(i) at each interface; if zero, the source terms Y2(i - 1) and Y2(i) are set equal to X6(i -  $\frac{1}{2}$ ) respectively (see Fig. 2.1).

As discussed in Section 2.1.1, the accuracy of the exponential term and the effect of truncating errors mean that the general formula will not reduce in the limit of small optical depths to the transparent case. To correct this situation, a test is made on  $\tau_i$  (the half optical depth  $\tau$  is stored in H2(i)), and if  $\tau < 10^{-2}$  a switch is made to the limiting form of the transport equation (Eq. (2.12) developed in Section II).

### 5.2.3. Angular Integration

The only integral over angle formed in PTRANS, at present, is the flux; ( $\int I \mu d\mu$ ) the formula for energy ( $\int I d\mu$ ) is included for possible use later. These integrals are formed on the outward pass from the intensity (sum X3(i)) stored on the inward pass and the intensity being calculated (F2). The difference forms of the equations are

$$X2(i) = \sum (F2 - \text{sum } X3(i)) \times \mu_m A_m ,$$

$$ER(i) = \sum (F2 + \text{sum } X3(i)) \times A_m .$$

### 5.3. DIFFERENCES WITH INTEGRAL FORMULATION

The principal difference in the subroutines is in replacing the integration of angle done explicitly in the integral formulation by a sampling scheme of a double Gaussian nature. It is expected that accuracy can be achieved with a minimum number of rays (presumably less than  $n = 6$ , see Section 2.2). This result is in logical agreement with the use of the S4 approximation in the neutron-transport work. The advantage, therefore, will appear in problems with many zones, since the integral method will



increase as the number of zones squared whereas the present method will only increase linearly with zones. Furthermore, the present method makes it possible to have special boundary conditions depending on angle (see Section 2. 1. 3).

#### 5. 4. AUXILIARY SUBROUTINES

In addition to the two new basic subroutines PRADTN and PTRANS, some changes have been made in the auxiliary subroutines EXP, PLNKUT, and KAPPA. These changes include (1) a fast exponential (FREXP), (2) a two-argument Planck function, and (3) the use of the average opacities from KAPPA ( $\theta$  and  $\rho$  interpolation) for the monofrequency calculation as well as for the Planck opacities.

The new fast exponential routine FREXP uses table lookup and interpolation rather than the normal expansion methods. The routine is written in machine language but uses the library routine EXP(X) for positive X or  $X > -10$ . An over-all gain in speed of a few percent was achieved in one comparison SPUTTER calculation.

The PLNKUT routine, with its associated tables PLNKTT, has been corrected and made more efficient by using a two-argument call which now calculates from either the analytic form or from the tables the difference in

$$b(u_1, u_2) = \frac{1}{B} \int_{\nu_1}^{\nu_2} \frac{2}{c^2} \frac{h\nu^3}{e^{-h\nu/kT} - 1} d\nu = b_j \quad (5.6)$$

The accuracy is improved since now not only differences of nearly equal numbers are subtracted.

The subroutine KAPPA, which calls in the group-averaged absorption coefficients from the DIANE tape and performs a bilinear log interpolation in temperature and density, has been modified to obtain the grey absorption coefficients as well as the Planck averages. At present, the format of the DIANE (absorption coefficient) tape includes a BCD record for tape identification, the Rosseland and Planck averages for a selected set of temperatures and densities from 0.25 eV to 50 eV and from 10 normal to  $10^{-6}$  normal, and the actual integration,  $\int \kappa_\nu d\nu$ , for the grey case. The grey or frequency-integrated averages are also used for an estimate of the time steps in PRADTN. KAPPA reads in first the tape name, the number of frequency groups, and the size of the records. If the sentinel for multifrequency is set to KMAX = 1, then the first frequency group,  $h\nu_1$ , and its absorption coefficients are read into storage. The interpolations in  $\log \theta_j$  and  $\log \rho_j$  are performed and a return to PRADTN is

made. If KMAX = 0, then KAPPA skips over the frequency-dependent absorption coefficients and reads into storage the grey averages. A signal, S15 = 0, is subsequently set, and for further cycles the interpolations are made on the stored quantities; the tape is not called again.

#### 5. 5. INPUT NUMBERS

The input quantities used in the radiation-transport subroutines and their functions are listed in Table 5. 1. The entries in it are as follows: column 1 is the storage location number used for entering the quantity into storage with the CARDS subroutine, column 2 lists the FORTRAN name of the stored quantity, column 3 gives the range of admissible values of the input quantity, column 4 describes its function and identifies special values it may assume, and column 6 records a set of values of the quantities which might be typical of those for a normal problem. Included is a set of values for the input quantities selected for solving typical problems.

#### 5. 6. EDITS

The editing of such frequency-dependent quantities as H3, the optical depth (Rosseland), X6, the source ( $b_j \theta^4$ ), X2, the flux (in  $\text{ergs}/4/3 \pi \text{ sec}$ ) X2/DHNU, the flux divided by the frequency group, THETA, the temperature (eV), and EI, the energy (ERGS/G) versus radius is accomplished by setting S12 to the desired number of cycles between prints. These multi-frequency edits have been used to evaluate the criteria for the subroutines as well as for diagnostics during the calculations.

A list of sample editing for a particular frequency group is given on page 27. The HNU is in electron volts. The quantities found useful to display for each frequency group and for a characteristic ray are listed on page 28. The format statements, in the listings appended, have been revised for the debug print from those used on page 28.

Table 5.1  
PLANE RADIATION INPUT QUANTITIES

Card No.	Quantity	Range of Values	Description	Typical Value
37	LMDA(37)	2, 3, 4, 5, 6	Number of angles with $\mu > 0$ .	2
44	KMAX	----	$\neq 0$ , performs multigroup frequency approximation; $= 0$ , performs single-group frequency approximation.	0
81	HVB	$\geq 0$	Number of optical mean free paths by which a transport region is extended at the expense of each adjacent diffusion region. (See Section 3.2.)	5
83	HCB	$\geq 0$	Criterion to define a diffusion region in terms of relative gradient of the source function. Diffusion regions are eliminated if 0. (See Section 3.2.)	0.1
87	CB	$\geq 0$	Criterion to combine frequency groups. If the lower frequency of the group is more than CB times the temperature of the hottest zone, that group is combined with the adjacent group of lower frequency. A half-integer value presents termination of the problem when half or more of the groups have been combined. (See Section 5.1.1.)	10.5
88	GA	$\geq 0$	One of two criteria for choice of linear or stepwise constant source within a zone. (See Section 5.1.2.)	0.333
90	GL	Neg., 0, $\frac{1}{2}$ , pos. integer	Indicator for radiation boundary condition at IB. GL = negative, total reflection; GL = 0, intensity for $\mu < 0$ is zero; GL = $\frac{1}{2}$ , blackbody intensity based on temperature located in THETA(1B) for $\mu < 0$ ; GL = positive integer, intensity for $\mu < 0$ obtained from source routine. GL must equal number of frequency groups. (See Section 2.1.3.)	0
121	AC	$\geq 0$	One of two criteria for choice of linear or stepwise constant source. Minimum value for using a linear source. (See Section 5.1.2.)	0.3
127	ACO3T4	----	Transport debug edit criterion. Edit occurs if $\neq 0$ and $<$ cycle number.	9
147	S12	----	Number of cycles between multifrequency edits.	10
150	S15	0, 1	Trigger controlling call of DIANE tape. Must have value $\neq 0$ on starts or restarts.	1
8466	TELM(25)	$\geq 0$	Constant multiplying the radiation time step. Can be used to modify the stability criterion.	1
8478	TELM(37)	$\geq 0$	Maximum permissible fractional energy in any zone due to radiation. Time step may be reduced to meet this requirement.	0.05
8858	SOLID(10)	----	Thick-thin criterion. If 0, Planck mean is used to form HZ; otherwise, Rosseland mean. (See Section 4.)	1

Table 5.1  
 SAMPLE EDIT FOR A CHARACTERISTIC DIRECTION FOR PTRANS

FREQUENCY BAND FROM 17.5000 TO 20.1000		IBR = 31		ICX = 31		ICV = 23	
JJ = 0	IAX = 1	X6	(Z)	H3	SUMX3	SUMX4	X2
1	-7.5994869-08	0.	2.4339197+02	2.4339197+02	0.	0.	-0.
2	9.9999999+00	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
3	1.0010000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
4	1.0020000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
5	1.0030000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
6	1.0040000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
7	1.0050000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
8	1.0060000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
9	1.0070000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
10	1.0080000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
11	1.0090000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
12	1.0100000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
13	1.0110000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
14	1.0120000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
15	1.0130000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
16	1.0140000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
17	1.0150000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
18	1.0160000+01	0.	2.4339192-01	2.4339192-01	0.	0.	-0.
19	1.0170000+01	0.	2.4339191-01	2.4339191-01	8.2503304-24	0.	-4.5847912-26
20	1.0180000+01	0.	2.4339191-01	2.4339191-01	8.5070054-15	0.	-1.4722185-21
21	1.0190000+01	0.	2.4339191-01	2.4339191-01	8.7716644-06	0.	-4.7274279-17
22	1.0199993+01	0.	2.4314869-01	2.4314869-01	0.	0.	-1.5180203-12
23	1.0209993+01	0.	1.1762804-01	1.1762804-01	2.7876050-01	0.	-4.8745017-08
24	1.0216361+01	8.7425114+00	2.7366146-03	2.7366146-03	1.2072228+01	5.4142551+00	-3.6999020-02
25	1.0219025+01	3.0955258+01	1.7085025-03	1.7085025-03	1.2082294+01	6.1653042+00	-3.2881306-02
26	1.0219959+01	3.2643747+01	1.6747771-03	1.6747771-03	1.0801617+01	7.7731507+00	-1.6829438-02
27	1.0220899+01	3.2132795+01	1.6698545-03	1.6698545-03	9.2048018+00	9.4688165+00	1.4671562-03
28	1.0221687+01	3.1355411+01	1.6399733-03	1.6399733-03	7.5149186+00	1.1024201+01	1.9501436-02
29	1.0222843+01	3.3702363+01	1.4826168-03	1.4826168-03	5.7588749+00	1.2424946+01	3.7044026-02
30	1.0223844+01	4.5108789+01	1.0629491-03	1.0629491-03	3.8479576+00	1.3815454+01	5.5390377-02
31	1.0224930+01	6.0067694+01	7.6399395-04	7.6399395-04	1.8744850+00	1.5234913+01	7.4245233-02
32	1.0225205+01	0.	0.	0.	0.	1.6622066+01	9.2370485-02

Table 5.2

SAMPLE MULTIFREQUENCY EDIT FOR PTRANS

CYCLE = 63. TIME = 0.305089-07 MWU FROM 17.8000 TO 20.1000

I	R	M3	X6	X2	X2/DMMU	RNU	IMETA	E1
1	7.81073952-08	2.28637886-02	-0.	-0.	0.	2.499999813-02	2.20139597+09	
2	9.9999998+00	2.28637835-01	0.	-1.08559455+07	-4.71997637+06	0.	2.49999925-02	2.20139750+09
3	1.00000000+01	2.28637835-01	0.	-1.81788317+07	-7.90384000+04	0.	2.49999918-02	2.20139741+09
4	1.00199999+01	2.28637831-01	0.	-3.04437489+07	-1.32364126+07	0.	2.49999922-02	2.20139747+09
5	1.00299999+01	2.28637831-01	0.	-5.09894950+07	-2.21433660+07	0.	2.49999922-02	2.20139747+09
6	1.00399998+01	2.28637833-01	0.	-8.54158400+07	-3.71373225+07	0.	2.49999927-02	2.20139754+09
7	1.00499998+01	2.28637829-01	0.	-1.43122512+08	-6.22271800+07	0.	2.49999934-02	2.20139743+09
8	1.00599998+01	2.28637831-01	0.	-2.39907566+08	-1.04307639+08	0.	2.49999929-02	2.20139757+09
9	1.00699998+01	2.28637833-01	0.	-4.02371692+08	-1.74946216+08	0.	2.49999927-02	2.20139754+09
10	1.00799998+01	2.28637833-01	0.	-6.75427098+08	-2.93643956+08	0.	2.49999936-02	2.20139766+09
11	1.00899998+01	2.28637833-01	0.	-1.15205550+09	-4.93567620+08	0.	2.49999939-02	2.20139770+09
12	1.00999998+01	2.28637833-01	0.	-1.91150912+09	-8.31090936+08	0.	2.49999939-02	2.20139770+09
13	1.01099998+01	2.28637831-01	0.	-3.22768322+09	-1.40325794+09	0.	2.49999963-02	2.20139776+09
14	1.01199998+01	2.28637831-01	0.	-5.47133190+09	-2.37884000+09	0.	2.49999948-02	2.20139782+09
15	1.01299998+01	2.28637831-01	0.	-9.32924768+09	-4.03618462+09	0.	2.49999966-02	2.20139779+09
16	1.01399997+01	2.28637831-01	0.	-1.60407853+10	-6.97425959+09	0.	2.49999962-02	2.20139786+09
17	1.01499997+01	2.28637829-01	0.	-2.79074358+10	-1.21336680+10	0.	2.49999953-02	2.20139789+09
18	1.01599997+01	2.28637829-01	0.	-4.93469158+10	-2.14551811+10	0.	2.49999962-02	2.20139802+09
19	1.01699997+01	2.28637829-01	0.	-8.91642235+10	-3.87670543+10	0.	2.49999964-02	2.20139808+09
20	1.01799997+01	2.28637841-01	0.	-1.65617304+11	-7.20075244+10	0.	2.49999994-02	2.20139837+09
21	1.01899997+01	2.28602695-01	0.	-3.18095679+11	-1.38302672+11	0.	2.50133509-02	2.20266403+09
22	1.01999997+01	2.2925232-02	4.91292107-04	-6.36769354+11	-2.76856246+11	0.	9.64961499-02	9.86790720+09
23	1.02099997+01	2.33086449-03	8.66041410+00	-1.42616981+12	-4.26073837+11	0.	1.22273718+00	4.73666470+11
24	1.02157248+01	1.76119019-03	3.10579107+01	-6.76098961+11	-3.60912599+11	0.	3.36150387+00	2.19259465+12
25	1.02179935+01	1.65541358-03	3.27618031+01	-6.73073798+11	-2.92640784+11	0.	4.36233952+00	2.96451945+12
26	1.02189915+01	1.63309816-03	3.22924628+01	1.52772108+11	4.66226570+10	0.	4.39482421+00	2.99672643+12
27	1.02198487+01	1.51176248-03	3.15173852+01	5.69157353+11	2.47459723+11	0.	4.37915815+00	3.00061431+12
28	1.02208086+01	1.51176248-03	3.15173852+01	5.69157353+11	2.47459723+11	0.	4.35496644+00	2.99155453+12
29	1.02218024+01	1.51176248-03	3.15173852+01	5.69157353+11	2.47459723+11	0.	4.42670870+00	3.11306738+12
30	1.02228336+01	1.05740126-03	4.51018038+01	1.40871439+12	6.12484730+11	0.	4.74113518+00	3.45431513+12
31	1.02239041+01	7.48561593-04	6.00003300+01	1.80893424+12	7.86483161+11	0.	5.10125035+00	4.14872179+12
32	1.02251422+01	6.19309597-01	0.	2.16810211+12	9.42653104+11	0.	0.	0.

SECTION VI  
TIMING STUDIES AND ACCURACY IN  
ANGULAR INTEGRATION

6.1. TIMING CALLS

The comparison in calculation time was obtained by using timing calls at selected locations in the logic of the code. To use the timing calls, it was necessary to establish a fiducial time from the system clock and then print the location of the time call, the time, and the difference in time between calls for each call. The subroutine that carries out these steps is CLOCK.

In the calculations described above, the subroutine CLOCK was called at the following locations in PRADTN and PTRANS:

PRADTN

- 13.105 - Before frequency loop
- 13.140 - After call KAPPA on merge
- 13.701 - After call KAPPA on main frequency loop
- 13.151 - After calculating general sources
- 13.180 - Before calling PTRANS
- 13.292 - After EDIT (normal) end of frequency loop
- 13.286 - After last frequency start time step
- 13.239 - End of cycle (return to main program)

PTRANS

- 14.708 - Before debug print

The following calculation was timed in units of 1/60 sec for the above breakdown in computing time.

The calculation described here did not use the TG criteria (see Section 5.1.2) nor the special boundary conditions (see Section 5.2.4). Three ray passes were completed for each frequency group for a total of six angles, forward and back, and with 32 active zones. The total time for 21 frequency groups was ~14.9 sec. The breakdown in time for a

single frequency group (in units of 1/60 sec) are the following:

Call KAPPA for absorption coefficient	14.
After sources	1.
Characteristic ray passes (3)	3.
After call PTRANS with EDIT	23.
Average time required	~41/hv
The start and merge of KAPPA is	~32
Total time with EDIT	~869 (1/60 sec)
Total time without EDIT	~515 (1/60 sec)

## 6.2. ACCURACY IN ANGULAR INTEGRATIONS

Comparisons have been made between calculations for two sets of angles for a radiation shock problem. The problem consists of a hot (5 ev) shock moving into cold low-density air. The effect on the flux versus linear zones for a set of four and eight angles is given in Fig. 6.1. This result indicates that as few as six and probably even four angles would be sufficient for reasonable accuracy.

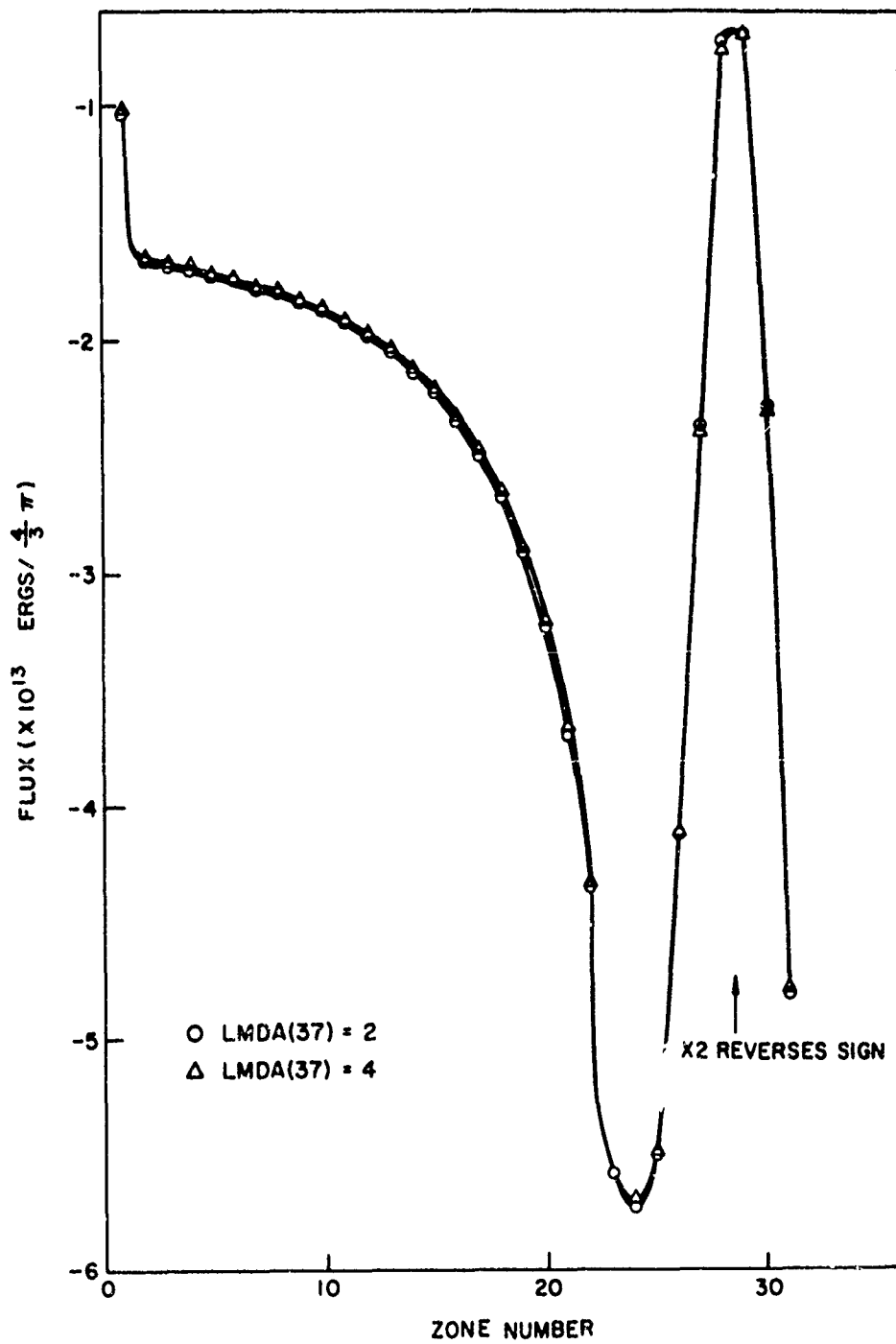


Fig. 6.1--Flux vs linear zones for six and ten angles



## REFERENCE

1. Margenall, H., and G. Murphy, The Mathematics of Physics and Chemistry, D. Van Nostrand Co. Inc., New York, 1943.

Appendix A

PRADTN

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$IBF: C PRADTN FULIST,DECK,REF PRAD0000
SUBROUTINE RADTN PRAD0010
CCMPILD OCTOBER 7, 1965 WBL
C PLANE CHARACTERISTIC TRACE WITH DOUBLE GAUSSIAN INTEGRATION PRAD0030
C*****PRAD0040
C* S P U T T E R C O M M O N **PRAD0050
C *PRAD0060
COMMON LMDA(37), NR , NSMLR , IA , IB , ICA , ICB , PRAD0070
1 KMAX , BLANK1, BLANK2, BLANK3, IAP1 , IBP1 , ICAP1 , ICBP1 , PRAD0080
2 II , IG , NRAD , BLANK4, IAM1 , IBM1 , ICAM1 , ICBM1 , PRAD0090
3 IIP1 , IGM1 , IALPHA, BLANK5, TH , TMAX , BLANK6, DELPRT, PRAD0100
4 FREQ , CNTMAX, AR , ASMLR , PUSHA , PUSHB , BOILA , BOILB , PRAD0110
5 CVA , CVB , SLUG , ALPHA , HVA , HVB , HCA , HCB , PRAD0120
6 EMINA , EMINB , CA , CB , GA , GB , GL , GR , PRAD0130
7 RHOL , RHOR , EPIO , EPSI , RIA , RIB , RDIA , RDIB , PRAD0140
8 RPIA , RPIB , RPDIA , RPDIB , TPRINT, TA , TB , TC , PRAD0150
COMMON TD , TE , DTH2 , DTH1 , DTRMIN, DTMAX , PRAD0160
1 DTMAX1, DTMAX2, DTMAX3, DTR , SWITCH, CO , CMIN , DELTA , PRAD0170
2 GAMA , WCRIT , SIGMAQ, AC , ACO3T4, CNVRT , SUMRA , SUMRB , PRAD0180
3 ROIA , ROIAM1, ROIB , ROIBP1, GMS , S1 , S2 , S3 , PRAD0190
4 S4 , S5 , S6 , S7 , S8 , S9 , S10 , S11 , PRAD0200
5 S12 , S13 , S14 , S15 , S16 , S17 , S18 , S19 , PRAD0210
6 S20 , EO , FO , TAU , ZERO , R (152), DELTAR(152), PRAD0220
7 ASQ (152), RD (152), VD (152), RDD (152), SMLR (152), PRAD0230
8 DELR ( 37), P (152), P1 (152), PB (152), PB1 (152), PRAD0240
COMMON P2 (152), SV (152), RHO (152), THETA (152), PRAD0250
1 W (152), E (152), EI (152), EK (152), A (152), PRAD0260
2 V (152), G (152), D (152), C (152), X2 (152), PRAD0270
3 X3 (152), X4 (152), X5 (152), X6 (152), X7 (152), PRAD0280
4 SMLA (152), SMLB (152), SMLC (152), SMLD (152), SMLE (152), PRAD0290
5 EC (152), ER (152), SMLQ (152), SMLH (152), BIGA (152), PRAD0300
6 BIGB (152), CV (152), BC (152), BR (152), CHIC (152), PRAD0310
7 CHIR (152), CAPAC (152), CAPAR (152), CRTC (152), CRTR (152), PRAD0320
8 CRTPC (152), GOFR (152), FEW (152), CAR (152), OKLM ( 37), PRAD0330
COMMON TELM ( 37), EKLM ( 37), ELM ( 37), FCLM ( 37), PRAD0340
1 FRLM ( 37), WLM ( 37), QLM ( 37), AMASNO( 37), CHRNO ( 37), PRAD0350
2 ZP1 ( 37), ZP2 ( 37), SOLID ( 37), ECHCK ( 37), RK (104), PRAD0360
3 RL ( 37), RHOK (104), ROK (104), THETA(104), TEMP ( 16), PRAD0370
4 HEAD ( 12), MAXL , MAXLM PRAD0380
C* **PRAD0390
C*****PRAD0400
DIMENSION Q3(1),TG(1),H2(1),Q1(1),X8(1),SUMX3(1),SUMX4(1) PRAD0410
DIMENSION H4(1),Y2(1),H(1),SUMX2(1),Q2(1) PRAD0420
DIMENSION Q37(1), Q38(1), H3(1) PRAD0430
COMMON /LINDLY/ HNU,SGNL,IHNU,NHNU,HNUP,NT,IM,IN,DHNU,THICK,NY PRAD0440
COMMON /CNTRL/ SCYCLE, JMUL PRAD0450
COMMON /DAVIS/ ICX, ICY PRAD0460
COMMON /TQ/ QINT1(300), QINT2(300), TITLE(12) PRAD0470
C PRAD0480
EQUIVALENCE(SMLA,H4),(SMLD,Y2) PRAD0490
EQUIVALENCE (BC,TG),(BIGB,H),(CRTR,SUMX2),(CHIC,SUMX3) PRAD0500
EQUIVALENCE (SMLH,X8),(CAR,Q37),(CHIR,Q38),(SMLC,H3) PRAD0510
EQUIVALENCE (ACO3T4,TROBG),(S12,EDITMF) PRAD0520
EQUIVALENCE (EC,Q1),(X7,H2),(BIGA,SUMX4),(GOFR,Q3) PRAD0530
C PRAD0540

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C*****PRADO550
C                                     *PRADO560
C           EDITMF SAME AS   S12      *PRADO570
C           H       SAME AS   BIGB     *PRADO580
C           H2 SAME AS X7              PRADO590
C           H3       SAME AS   SMLC    *PRADO600
C           H4       SAME AS   SMLA    *PRADO610
C           Q1       SAME AS   EC      *PRADO620
C           Q3       SAME AS   GOFR    *PRADO630
C           Q37      SAME AS   CAR     *PRADO640
C           Q38      SAME AS   CHIR    *PRADO650
C           SUMX2    SAME AS   CRTR    *PRADO660
C           SUMX3    SAME AS   CHIC    *PRADO670
C           SUMX4    SAME AS   BIGA    *PRADO680
C           TG       SAME AS   BC      *PRADO690
C           TRDBG    SAME AS   ACO3T4  *PRADO700
C           X8       SAME AS   SMLH    *PRADO710
C           Y2       SAME AS   SMLD    *PRADO720
C*****PRADO730
C                                     PRADO740
C                                     PRADO750
C                                     PRADO760
C           FEX, FM, SUMRHO, CSQD, XSQD, Y, YSQD, Q2, NOT USED PRADO770
C                                     PRADO780
C*****PRADO790
C                                     *PRADO800
C           P L A N E S   O N L Y      *PRADO810
C                                     *PRADO820
C*****PRADO830
C                                     *PRADO840
C           B O U N D A R Y   C O N D I T I O N S   *PRADO850
C                                     *PRADO860
C*****PRADO870
C                                     PRADO880
C                                     PRADO890
C           L.H. BLACKBODY IF SOLID ON LEFT PRADO900
C                                     PRADO910
C           NO VAPOR ZONE HAS FLUX OUT FROM SOLID PRADO920
C                                     PRADO930
C           NTIMES=80ILB PRADO940
C           IM=IBM1 PRADO950
C           IN=IA PRADO960
C           IF(ZP1(26).EQ.0.) GO TO 15
C           SAVE STUFF FROM EIONX FOR NONEQ AND RESET IN OR IM
C           IF (PUSHA .LT. 0.0) GO TO 100
C           IM = NR - 1
C           WSZ2=BC(IM+1)
C           WSZ3=BR(IM+1)
C           WSZ4 = CRTC(IM+1)
C           WSZ5=RHO(IM+1)
C           GO TO 15
100 IN = NR
C           WSZ2 = BC(IN-1)
C           WSZ3 = BR(IN-1)
C           WSZ4 = CRTC(IN-1)
C           WSZ5 = RHO(IN-1)

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15	CONTINUE	PRAD0990
	IMP1=IN+1	PRAD1000
	INM1=IN-1	PRAD1010
	CALL CVCHK (K000FX)	PRAD1020
	IF (IMP1-IN) 190,190,125	PRAD1030
C		PRAD1040
C	NO VAPOR ZONES	PRAD1050
C		PRAD1060
190	X2(IMP1) = 1.0283E12 * A(IM) * (THETA(IM)**4 - THETA(IMP1)**4)	PRAD1070
	ER(IM)=-X2(IMP1)	PRAD1080
	GO TO 1300	PRAD1090
125	IR=IN	PRAD1100
	THTAMX=.025	PRAD1110
	IF (IALPHA-1) 130,140,130	PRAD1120
130	S1 = 13.0130	PRAD1130
	CALL UNCLE	PRAD1140
140	DO 180 I=IN,IM	PRAD1150
	X3(I)=0.	PRAD1160
	X4(I)=0.	PRAD1170
	X5(I)=0.	PRAD1180
	X6(I)=0.	PRAD1190
	CRTRI(I)=0.	PRAD1200
C		PRAD1210
C	SET UP FOR KAPPA INTERPOLATION	PRAD1220
C		PRAD1230
	Q1(I)=THETA(I)**4	PRAD1240
	Q37(I)=ALOG(THETA(I))	PRAD1250
	Q38(I)=ALOG(SV(I))	PRAD1260
C		PRAD1270
C	FIND IR, RIGHTMOST ZONE WITH THETA GREATER THAN 0.05 EV	PRAD1280
C		PRAD1290
	IF (THETA(I)-THTAMX) 160,160,150	PRAD1300
150	THTAMX=THETA(I)	PRAD1310
160	IF (THETA(I)-0.05) 180,170	PRAD1320
170	IR=I	PRAD1330
180	CONTINUE	PRAD1340
	IF (THTAMX .LT. THETA(IR)) THTAMX = THETA(IR)	PRAD1350
C		PRAD1360
C		PRAD1370
C	*****	PRAD1380
C		*PRAD1390
C	BEGIN FREQUENCY LOOP	*PRAD1400
C		*PRAD1410
C	*****	PRAD1420
200	HNUP = 3.E3	PRAD1430
C		PRAD1440
C	SET UP MAX FREQ BOUNDARY	PRAD1450
C		PRAD1460
	HNUP4 = 8.1E13	PRAD1470
	IHNU=1	PRAD1480
	DO 210 I=IN,IMP1	PRAD1490
210	SUMX2(I)=0.0	PRAD1500
	IF (KMAX.EQ.0) GO TO 280	PRAD1510
C		PRAD1520
C	THIS CODING WONT WORK IF HNU NOT EVALUATED	PRAD1530

C	220 CALL KAPPA(IN,IM)	PRAD1540
	HNNU4=HNNU**4	PRAD1550
	DHNUP = DHNU	PRAD1560
	DHNU = HNUP - HNNU	9/29/65
C		PRAD1570
C	MERGE GROUPS WITH HNNU MORE THAN CB TIMES LARGEST THETA	PRAD1580
C		PRAD1590
C	IF (THTAMX- HNNU/CB) 240,300,300	PRAD1600
C		PRAD1610
C	REJECT TAPE IF MORE THAN HALF OF GROUPS MERGE	PRAD1620
C		PRAD1630
C	240 IF (IHNU+IHNU-NHNU) 260,250,250	PRAD1640
	250 IF (AMOD(CB,1.) .EQ. 0.5) GO TO 260	PRAD1650
	S1=13.0250	PRAD1660
	CALL UNCLE	PRAD1670
	260 DO 270 I=IN,IM	PRAD1680
	BETA=HNNU/THETA(I)	PRAD1690
	BETAP=HNUP/THETA(I)	PRAD1700
	DFB=PLNKUT(BETA,BETAP)	PRAD1710
	IF (DFB.EQ.0.) GO TO 270	PRAD1720
	TEMP(1)=DFB*Q1(I)	PRAD1730
	EMB1=EXP(-BETA)	PRAD1740
	EMB2=EXP(-BETAP)	PRAD1750
	TEMP(2)=DFB+0.0384974/Q1(I)*(HNNU4/(1.0-EMB1)	PRAD1760
	1*EMB1-HNUP4/(1.0-EMB2)*EMB2)	PRAD1770
C		PRAD1780
C	FORM NUMERATORS AND DENOMINATORS OF MERGED KAPPAS	PRAD1790
C		PRAD1800
C	X6(I)=X6(I)+TEMP(1)	PRAD1810
	X4(I)=X4(I)+TEMP(2)	PRAD1820
	X5(I)=X5(I)+CAPAC(I)*TEMP(1)	PRAD1830
	X3(I)=X3(I)+TEMP(2)/CAPAR(I)	PRAD1840
	270 CONTINUE	PRAD1850
	IF (GL .LT. 1. .OR. IHNU .EQ. 1) GO TO 275	PRAD1860
C	MERGE FREQUENCY-DEPENDENT EXTERNAL INPUT INTENSITIES	9/29/65
	NMU = LMDA(37)	PRAD1880
	IQNT = NMU * (IHNU - 2)	PRAD1890
	DO 272 I = 1, NMU	9/29/65
	IQNT1 = IQNT + I	PRAD1910
	IQNT2 = IQNT1 + NMU	PRAD1920
	IF (IHNU.GT.2) DHNUP=1.	PRAD1930
	272 QINT1(IQNT2) = QINT1(IQNT2)*DHNUP + QINT1(IQNT1)*DHNUP	9/29/65
	275 HNUP=HNNU	9/29/65
	IHNU=IHNU+1	PRAD1950
	HNUP4=HNNU4	PRAD1960
	IF (THTAMX- HNNU/CB) 220,310,310	PRAD1970
C		PRAD1980
C	MONOFREQUENCY CALCULATION	PRAD1990
C		PRAD2000
C	280 NHNU=1	PRAD2010
	CALL KAPPA (IN,IM)	PRAD2020
	DO 290 I=IN,IM	PRAD2030
	X5(I)=1.	PRAD2040
	290 X6(I)=Q1(I)	PRAD2050
		PRAD2060

DFB=1.0	PRAD2070
HNH = .001	PRAD2080
ICX=IR	PRAD2090
IF (GL .GT. 0.0) ICX = IM	
ICY=IN	PRAD2100
GO TO 480	PRAD2110
300 IF (IHNU-1) 550,370,260	PRAD2120
C	PRAD2130
C	PRAD2140
C	PRAD2150
310 DC 350 I=IN, IM	PRAD2160
IF (X6(I)) 320,350,330	PRAD2170
320 S1=13.0320	PRAD2180
CALL UNCLE	PRAD2190
330 CAPAR(I)=X4(I)/X3(I)	PRAD2200
340 CAPAC(I)=X5(I)/X6(I)	PRAD2210
350 CONTINUE	PRAD2220
HNUP = 3.E3	PRAD2230
HNUP4 = 8.1E13	PRAD2240
DHNU = HNUP - HNU	PRAD2250
IHNU=IHNU-1	PRAD2260
IF(GL.LT.1.) GO TO 370	9/29/65
DO 355 I=1,NMU	9/29/65
IQNT2 = IQNT + I + NMU	9/29/65
355 QINT1(IQNT2) = QINT1(IQNT2)/DHNU	9/29/65
GO TO 370	PRAD2270
C	PRAD2280
C	PRAD2290
C	PRAD2300
360 CALL KAPPA(IN,IM)	PRAD2310
DHNU=HNUP-HNU	PRAD2320
HNU4=HNU**4	PRAD2330
370 IF (GL-1.) 390,380,380	PRAD2340
380 IF (HNU.NE.RDK(IHNU+52)) GO TO 490	PRAD2350
IF (GL.NE.FLOAT(NHNU)) GO TO 490	PRAD2360
C A L C U L A T E I C X , I C Y	PRAD2370
390 ICX=IN	PRAD2380
ICY = IN	PRAD2390
IF (GL .LE. 0.) GO TO 395	PRAD2400
ICX = IM	PRAD2410
DO 392 I = IN, IM	PRAD2420
DFB = PLNKUT(HNU / THETA(I), HNUP / THETA(I))	PRAD2430
392 X6(I)=DFB*Q1(I)	PRAD2440
GO TO 480	PRAD2450
395 DO 470 I=IN,IR	PRAD2460
BETA=HNU/THETA(I)	PRAD2470
C	PRAD2480
C	PRAD2490
C	PRAD2500
IF (BETA-19.0) 400,410,410	PRAD2510
400 BETAP=HNUP/THETA(I)	PRAD2520
C	PRAD2530
EMB2=EXP(-BETAP)	PRAD2540
IF (BETAP-0.01) 410,410,460	PRAD2550
410 IF (ICX-IR) 430,420,420	PRAD2560
420 ICX=I-'	

430	IF (1-ICY) 440,440,450	PRAD2570
440	ICY=ICY+1	PRAD2580
C		PRAD2590
C	ICX IS INDEX OF LAST ZONE WITH SIGNIFICANT SOURCE	PRAD2600
C		PRAD2610
C	ICY IS INDEX OF FIRST ZONE WITH SIGNIFICANT SOURCE	PRAD2620
C		PRAD2630
450	X6(I)=0.0	PRAD2640
	X5(I)=0.0	PRAD2650
	GO TO 470	PRAD2660
C		PRAD2670
C	FORM SOURCES X6 AND X5	PRAD2680
C		PRAD2690
460	DFB=PLNKUT(BETA,BETAP)	PRAD2700
	X6(I)=DFB*Q1(I)	PRAD2710
C	TEMP(2)=0.0384974/Q1(I)*(HNU4/(EXP(BETA)-1.0)	PRAD2720
C	1 -HNU4/(1.0-EMB2)*EMB2)	PRAD2730
C	X5(I)=DFB+TEMP(2)	PRAD2740
	ICX=IR	PRAD2750
470	CONTINUE	PRAD2760
480	IF (INM1) 490,520,500	PRAD2770
C		PRAD2780
C	SET BLACKBODY CONDITION FOR IA GREATER THAN 1	PRAD2790
C		PRAD2800
490	S1=13.0490	PRAD2810
	CALL UNCLE	PRAD2820
500	DFB = PLNKUT (HNU/THETA(INM1),HNU/THETA(INM1))	PRAD2830
	X6(INM1) = DFB * THETA(INM1)**4	PRAD2840
C	SET BLACKBODY CONDITION IF DESIRED FOR IMP1	PRAD2850
520	IF (GL.NE.0.5) GO TO 530	PRAD2860
	DFB = PLNKUT(HNU / THETA(IMP1), HNU / THETA(IMP1))	PRAD2870
	X6(IMP1) = DFB * THETA(IMP1)**4	PRAD2880
530	Q31=0.0	PRAD2890
C		PRAD2900
C	FORM ROSSELAND AND PLANCK OPTICAL DEPTHS	PRAD2910
C		PRAD2920
	DO 590 I=IN,IM	PRAD2930
	IF (CAPAR(I)) 550,550,540	PRAD2940
540	IF (CAPAC(I)) 550,550,560	PRAD2950
550	S1=13.0550	PRAD2960
	CALL UNCLE	PRAD2970
C		PRAD2980
C	CHOOSE ALL ROSSELAND IF SOLID 10 IS POSITIVE	PRAD2990
C		PRAD3000
560	IF (SOLID(10).EQ.0.) GO TO 570	PRAD3010
	H(I)=CAPAR(I)/SV(I)	PRAD3020
	GO TO 580	PRAD3030
570	H(I)=CAPAC(I)/SV(I)	PRAD3040
580	H2(I)=H(I)*DELTAR(I)	PRAD3050
	IF (ALPHA .GT. 1.) GO TO 586	PRAD3060
	H3(I) = CAPAR(I) * G(I)	PRAD3070
	GO TO 588	PRAD3080
CAVEAT.	ASYNCHRONISMS IN SV AND DELTAR LEAD TO ERRONEOUS FLUCTUATIONS	PRAD3090
C	IN H3. THIS CAN BE FIXED BY SUBSTITUTING G IN PLANES, BUT SPHERES	PRAD3100
C	WILL STILL HAVE THIS TROUBLE.	PRAD3110



586	H3(I)=CAPAR(I)/SV(I)*DELTAR(I)	PRAD3120
588	Q31=Q31+H3(I)	PRAD3130
	Q3(I+1)=Q31	PRAD3140
	H(I)=0.5*H(I)	PRAD3150
	H2(I)=0.5*H2(I)	PRAD3160
	H3(I)=0.5*H3(I)	PRAD3170
C		PRAD3180
C	ZERO DIFFUSION INDICATORS AND X2	PRAD3190
C		PRAD3200
	X2(I)=0.0	PRAD3210
	X3(I)=0.0	PRAD3220
	X4(I)=0.0	PRAD3230
590	RHC(I)=0.0	PRAD3240
	X2(IMP1)=0.0	PRAD3250
	X3(IMP1)=0.0	PRAD3260
	X4(IMP1)=0.0	PRAD3270
	IF (ICY .GT. ICX) GO TO 990	PRAD3280
C		PRAD3290
C	STEP-LINEAR CRITERION AT ICY	PRAD3300
C	UNCONDITIONAL STEP AS BOUNDARY CONDITION IF ICY = IN	PRAD3310
C		PRAD3320
	IF (ICY-IN) 675,600,610	PRAD3330
600	Y2(IN)=X6(IN)	PRAD3340
	TG(IN)=0.0	PRAD3350
	GO TO 620	PRAD3360
610	TEMP(1)=H3(ICY-1)+H3(ICY)	PRAD3370
	TG(ICY)=X6(ICY)/TEMP(1)	PRAD3380
	Y2(ICY)=TG(ICY)*H3(ICY-1)	PRAD3390
C		PRAD3400
C	FORM Y2 AND TG SET X3=-1 IF A DIFFUSION CRITERION MET USING HCB	PRAD3410
C		PRAD3420
620	ICXM1=ICX-1	PRAD3430
	IF (ICY .GT. ICXM1) GO TO 672	PRAD3440
	DO 670 I=ICY,ICXM1	PRAD3450
	TEMP(1)=H3(I+1)+H3(I)	PRAD3460
	IF (AMAX1(X6(I), X6(I+1)) .LE. 0.) GO TO 640	PRAD3490
	IF (AMIN1(H3(I), H3(I+1)) .GT. AC) GO TO 650	PRAD3500
	IF (ABS((H3(I)-H3(I+1))/TEMP(1)) .GT. GA) GO TO 640	PRAD3510
	IF (ABS((X6(I)-X6(I+1))/(X6(I)*X6(I+1))) .GT. GA) 650,650,640	PRAD3520
640	TG(I+1)=0.0	PRAD3530
	GO TO 670	PRAD3540
650	TG(I+1)=(X6(I+1)-X6(I))/TEMP(1)	PRAD3550
C		PRAD3560
C	TG(I+1)=(Q1(I+1)-Q1(I))/TEMP(1)*(X5(I+1)+X5(I))/2.0	PRAD3570
C		PRAD3580
	Y2(I+1)=(X6(I+1)*H3(I)+X6(I)*H3(I+1))/TEMP(1)	PRAD3590
	IF (ABS(TG(I+1))-HCB*Y2(I+1)) 660,670,670	PRAD3600
660	X3(I+1)=-1.0	PRAD3610
670	CONTINUE	PRAD3620
C		PRAD3630
C	RADIATION BOUNDARY CONDITION AT ICX	PRAD3640
C	(VACUUM IF ICX = IM AND GL NOT 1/2)	PRAD3650
C		PRAD3660
672	IF (ICX-IM) 680,690,675	
675	S1 = 13.0675	

	CALL UNCLE		PRAD3670
680	TEMP(1)=H3(ICX+1)+H3(ICX)		PRAD3680
	TG(ICX+1)=-X6(ICX)/TEMP(1)		PRAD3690
	Y2(ICX+1)=-TG(ICX+1)*H3(ICX+1)		PRAD3700
	GO TO 700		PRAD3710
690	IF (GL .EQ. 0.5) GO TO 700		PRAD3720
	Y2(IMP1)=X6(ICX)		PRAD3730
	TG(IMP1)=0.0		PRAD3740
C			PRAD3750
C	EXTEND TRANSPORT REGION BOUNDARIES, IF NEEDED, TO PROVIDE HVB MEAN		PRAD3760
C	FREE PATHS		PRAD3770
C			PRAD3780
700	I=IN+1		PRAD3790
710	IF (X3(I)) 720,740,730		PRAD3800
720	I=I+1		PRAD3810
	IF (I-ICX-1) 710,740,820		PRAD3820
730	S1=13.0730		PRAD3830
	CALL UNCLE		PRAD3840
740	J=I-1		PRAD3850
750	IF (Q3(I)-Q3(J)-HVB) 760,760,770		PRAD3860
760	X4(J)=-1.0		PRAD3870
	J=J-1		PRAD3880
	IF (J-IN) 770,750,750		PRAD3890
770	I=I+1		PRAD3900
	IF (I-ICX-1) 780,780,820		PRAD3910
780	IF (X3(I)) 790,770,730		PRAD3920
790	J=I		PRAD3930
800	IF (Q3(J)-Q3(I-1)-HVB) 810,810,720		PRAD3940
810	X4(J)=-1.0		PRAD3950
	J=J+1		PRAD3960
	IF (J-ICX-1) 800,720,720		PRAD3970
820	I=IN+1		PRAD3980
C			PRAD3990
C	TEST TO FORM TRANSPORT REGIONS		
C			PRAD4010
830	IAX=IN		PRAD4020
840	IF (X3(I)) 850,860,730		PRAD4030
850	IF (X4(I)) 860,870,730		PRAD4040
C			PRAD4050
C	REMOVE ONE ZONE DIFFUSION REGION		
C			PRAD4070
860	I=I+1		PRAD4080
	IF (I-ICX-1) 840,950,950		PRAD4090
870	I=I+1		PRAD4100
	IF (I-ICX-1) 890,950,950		PRAD4110
880	IF (X3(I)) 890,840,730		PRAD4120
890	IF (X4(I)) 840,900,730		PRAD4130
900	I=I-3		PRAD4140
	GO TO 960		PRAD4150
910	IF (X3(I)) 920,940,730		PRAD4160
920	IF (X4(I)) 940,930,730		PRAD4170
C	FORM X2 FOR DIFFUSION ZONES IN ORDER		
930	X2(I) = -1.37E12 * TG(I)		PRAD4190
	I=I+1		PRAD4200
	IF (I-ICX-1) 910,980,980		PRAD4210

C		PRAD4220
C	DO TRANSPORT TO IM IN REGION OF NO SOURCE	
C		PRAD4240
	940 IAX=I	PRAD4250
	GO TO 860	PRAD4260
	950 IBX=IM	PRAD4270
	960 CALL TRANS(IAX,IBX)	PRAD4280
	IF (IBX-IM) 970,990,990	PRAD4290
	970 I=IBX+2	PRAD4300
	GO TO 930	PRAD4310
	980 IF (I .GT. IM) GO TO 990	
	IAX=I	
	GO TO 950	
C		PRAD4330
C	OPTIONAL EDIT OF X2 ETC.	PRAD4340
C		PRAD4350
	990 IF (EDITMF) 1020,1020,1000	PRAD4360
	1000 IARG1=SOLID(18)+0.001	PRAD4370
	IARG2=EDITMF+0.001	PRAD4380
	IF (MOD(IARG1,IARG2)) 1020,1010,1020	PRAD4390
	1010 WRITE (3) HNU, IN, IM, IMP1, SOLID(18), TH, DHNU	PRAD4400
C	IN SPHERICAL VERSION, REDIT IS GIVEN RHO.	PRAD4410
C	REPLACED HERE BY CAPAR.	PRAD4420
	WRITE (3) (C(I), I=IN,IMP1), (H3(I), I=IN,IM), (X6(I), I=IN,IMP1),	PRAD4430
	1(X2(I),I=IN,IMP1),	PRAD4440
	2 (CAPAR(I), I=IN,IMP1), (THETA(I), I=IN,IMP1), (EI(I), I=IN,IM)	PRAD4450
	XX=-2.0	PRAD4460
	WRITE (3) XX,XX,XX,XX,XX,XX,XX	PRAD4470
	BACKSPACE 3	PRAD4480
	JMULT=1	PRAD4490
	1020 DO 1030 I=IN,IMP1	PRAD4500
	SUMX2(I)=SUMX2(I)+X2(I)	PRAD4510
	1030 CCNTINUE	PRAD4520
C		PRAD4530
C	ADVANCE FREQ, STORE EMERGENT FLUX, TEST FOR COMPLETION OF GROUPS	PRAD4540
C		PRAD4550
	HNU4=HNU	PRAD4560
	HNU4=HNU4	PRAD4570
	IHNU = IHNU + 1	PRAD4580
	IF (IHNU-NHNU) 1040,1040,1060	PRAD4590
	1040 CALL DVCHK (K000FX)	PRAD4600
	GO TO (1050,360), K000FX	PRAD4610
		PRAD4620
C	*****	PRAD4630
C		*PRAD4640
C	END FREQUENCY LOOP	*PRAD4650
C		*PRAD4660
C	*****	PRAD4670
	1050 S1 = 13.1050	PRAD4680
	CALL UNCLE	PRAD4690
	1060 SUMX2(INM1) = 0.0	PRAD4700
	DO 1070 I=INM1,IMP1	PRAD4710
	X2(I) = SUMX2(I)	PRAD4720
	1070 ER(I) = SUMX2(I) - SUMX2(I+1)	PRAD4730
C		PRAD4740
C	FORM MAGNOFREQUENCY QUANTITIES AND FIND MIN TIME STEP	PRAD4750

C	WSB = 0.0	PRAD4760
	DO 1075 I = 1, MAXLM	PRAD4770
1075	WSB = WSB + ELM(I)	PRAD4780
	DTR1=1.E10	PRAD4790
	DTR2=1.E10	PRAD4800
	IF (KMAX.EQ.0) GO TO 1080	PRAD4810
	CALL KAPPA(IN,IM)	PRAD4820
1080	DO 1230 I=IN,IM	PRAD4830
C		PRAD4840
C	IF ROSS IS ZERO EXIT	PRAD4850
C		PRAD4860
	IF (CAPAR(I)) 1090,1090,1100	PRAD4870
1090	S1=13.1090	PRAD4880
	CALL UNCLE	PRAD4890
1100	TEMP(3)=CAPAR(I)	PRAD4900
	IF (SCLID(10)) 1110,1120,1110	PRAD4910
1110	TEMP(1)=CAPAR(I)	PRAD4920
	GO TO 1130	PRAD4930
1120	TEMP(1)=SQRT(CAPAR(I)*CAPAC(I))	PRAD4940
1130	IF (TEMP(1)) 1090,1090,1140	PRAD4950
1140	H(I)=0.5*TEMP(1)/SV(I)	PRAD4960
	H3(I) = H(I) * DELTAR(I)	PRAD4970
	IF (.001-THETA(I)) 1160,1230,1230	PRAD4980
1160	IF (H3(I).GT.0.1) GO TO 1170	PRAD4990
	IF (ER(I).EQ.0.) GO TO 1170	PRAD5000
	WSBB = E(I) * G(I)	PRAD5010
	IF (TELM(37) .EQ. 0.0) GO TO 1170	PRAD5020
	IF (WSBB - TELM(37) * WSB) 1170, 1165, 1165	PRAD5030
1165	TEMP(2)=.5*CV(I)*THETA(I)*G(I)/ABS(ER(I))	PRAD5040
	GO TO 1180	PRAD5050
1170	TEMP(2)=(.5+1.5*H3(I)**2)*CV(I)/(4.1132E12*TEMP(3)*THETA(I)**3)	PRAD5060
	TEMP(2)=TEMP(2)*TELM(25)	PRAD5070
C	*****	PRAD5080
C		PRAD5090
C	FIND MINIMUM TIME STEP	*PRAD5100
C		*PRAD5110
C		*PRAD5120
C	*****	PRAD5130
1180	IF (TEMP(2)) 1230,1230,1190	PRAD5140
1190	CONTINUE	PRAD5150
	IF (TEMP(2)-DTR1) 1200,1210,1210	PRAD5160
1200	DTR2=DTR1	PRAD5170
	IMN2=IMN1	PRAD5180
	DTR1=TEMP(2)	PRAD5190
	IMN1=I	PRAD5200
	GO TO 1230	PRAD5210
1210	IF (TEMP(2)-DTR2) 1220,1230,1230	PRAD5220
1220	DTR2=TEMP(2)	PRAD5230
	IMN2=I	PRAD5240
1230	CONTINUE	PRAD5250
	DTRMIN=DTR1	PRAD5260
	EO=IMN1	PRAD5270
C		PRAD5280
C	PRINT MINIMUM TIME STEPS BETWEEN EDITS	PRAD5290
C		PRAD5300

IF (DTR1-TELM(26)) 1240,1250,1250	PRAD5310
1240 TELM(26)=DTR1	PRAD5320
TELM(27)=IMN1	PRAD5330
TELM(28)=DTR2	PRAD5340
TELM(29)=IMN2	PRAD5350
TELM(30)=SOLID(18)+1.0	PRAD5360
1250 CONTINUE	PRAD5370
C	PRAD5380
C DETERMINE IF RADIATION OR HYDRO WILL SUBCYCLE	PRAD5390
C	PRAD5400
IF (DTRMIN-DTR) 1280,1300,1260	PRAD5410
1260 BLANK3=TH+AMIN1(DTRMIN,GR*DTH2)	PRAD5420
IF (S17) 1300,1270,1300	PRAD5430
1270 S9 = 1.0	PRAD5440
GO TO 1300	PRAD5450
C*****	PRAD5460
C	*PRAD5470
C R E D U C E T I M E S T E P	*PRAD5480
C	*PRAD5490
C*****	PRAD5500
1280 NRAD=ZP1(18)/DTRMIN+1.0	PRAD5510
DTR=ZP1(18)/FLOAT(NRAD)	PRAD5520
IF (NRAD-NTIMES) 1300,1300,1290	PRAD5530
1290 S1=13.1290	PRAD5540
CALL UNCLE	PRAD5550
C ZERO OUT STRAY QUANTITIES FROM PREVIOUS CYCLES	PRAD5560
1300 DG 1310 I = IMP1, IG	PRAD5570
CAPAR(I) = 0.	PRAD5580
CAPAC(I) = 0.	PRAD5590
X2(I+1) = 0.	PRAD5600
X3(I+1) = 0.	PRAD5610
X4(I+1) = 0.	PRAD5620
SUMX2(I+1) = 0.	PRAD5630
SUMX3(I+1) = 0.	PRAD5640
SUMX4(I+1) = 0.	PRAD5650
1310 ER(I+1) = 0.	PRAD5660
IF (ZP1(26) .EQ. 0.0) GO TO 1400	
C RESTORE GOODIES FOR NONEQ	
. (PUSHA .LT. 0.0) GO TO 1350	
BC (IMP1)=WSZ2	
BR (IMP1)=WSZ3	
CPTR (IMP1) = WSZ4	
RHO (IMP1)=WSZ5	
GO TO 1400	
1320 BC (INM1) = WSZ2	
BR (INM1) = WSZ3	
CRTR (INM1) = WSZ4	
RHC (INM1) = WSZ5	
1400 RETURN	
END	PRAD5680

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Appendix B

PTRANS

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$IBFTC PTRANS FULIST,DECK,REF                                PTRAO000
SUBROUTINE TRANS(N,M)                                        PTRAO010
C   CCMPILED JULY 1, 1965   WBL                               PTRAO020
C   PLANE CHARACTERISTIC TRACE WITH DOUBLE GAUSSIAN INTEGRATION PTRAO030
C   DECK DBLGAU REQUIRED FOR INTEGRATION COEFFS.              PTRAO040
C*****PTRAO050
C*   S P U T T E R   C O M M O N                               **PTRAO060
C   **PTRAO070
COMMON LMDA(37), NR , NSMLR , IA , IB , ICA , ICB , PTRAO080
1 KMAX , BLANK1, BLANK2, BLANK3, IAP1 , IBP1 , ICAP1 , ICBP1 , PTRAO090
2 II , IG , NRAD , BLANK4, IAM1 , IBM1 , ICAM1 , ICBM1 , PTRAO100
3 IIP1 , IGM1 , IALPHA, BLANK5, TH , TMAX , BLANK6, DELPRT , PTRAO110
4 FREQ , CNTMAX, AR , ASMLR , PUSHA , PUSHB , BOILA , BOILB , PTRAO120
5 CVA , CVB , SLUG , ALPHA , HVA , HVB , HCA , HCB , PTRAO130
6 EMINA , EMINB , CA , CB , GA , GB , GL , GR , PTRAO140
7 RHCL , RHOR , EPIO , EPSI , RIA , RIB , RDIA , RDIB , PTRAO150
8 RPIA , RPIB , RPDIA , RPDIB , TPRINT , TA , TB , TC , PTRAO160
COMMON TD , TE , DTH2 , DTH2P , DTH1 , DTRMIN , DTMAX , PTRAO170
1 DTMAX1, DTMAX2, DTMAX3, DTR , SWITCH, CO , CMIN , DELTA , PTRAO180
2 GAMA , WCRIT , SIGMAQ, AC , ACO3T4, CNVRT , SUMRA , SUMRB , PTRAO190
3 ROIA , ROIAM1, ROIB , ROIBP1, GMS , S1 , S2 , S3 , PTRAO200
4 S4 , S5 , S6 , S7 , S8 , S9 , S10 , S11 , PTRAO210
5 S12 , S13 , S14 , S15 , S16 , S17 , S18 , S19 , PTRAO220
6 S20 , EO , FO , TAU , ZERO , R (152), DELTAR(152), PTRAO230
7 ASQ (152), RD (152), VD (152), RDD (152), SMLR (152), PTRAO240
8 DELR ( 37), P (152), P1 (152), PB (152), P81 (152) PTRAO250
COMMON P2 (152), SV (152), RHO (152), THETA (152), PTRAO260
1 W (152), E (152), EI (152), EK (152), A (152), PTRAO270
2 V (152), G (152), D (152), C (152), X2 (152), PTRAO280
3 X3 (152), X4 (152), X5 (152), X6 (152), X7 (152), PTRAO290
4 SMLA (152), SMLB (152), SMLC (152), SMLD (152), SMLE (152), PTRAO300
5 EC (152), ER (152), SMLQ (152), SMLH (152), BIGA (152), PTRAO310
6 BIGB (152), CV (152), BC (152), BR (152), CHIC (152), PTRAO320
7 CHIR (152), CAPAC (152), CAPAR (152), CRTC (152), CRTR (152), PTRAO330
8 CRTPC (152), GOFR (152), FEW (152), CAR (152), OKLM ( 37) PTRAO340
COMMON TELM ( 37), EKLH ( 37), ELM ( 37), FCLM ( 37), PTRAO 350
1 FRLM ( 37), WLM ( 37), QLM ( 37), AMASNO( 37), CHRNO ( 37), PTRAO360
2 ZP1 ( 37), ZP2 ( 37), SOLID ( 37), ECHCK ( 37), RK (104), PTRAO370
3 RL ( 37), RHOK (104), RDK (104), THETAK(104), TEMP ( 16), PTRAO380
4 HEAD ( 12), MAXL , MAXLM , PTRAO390
C*   **PTRAO400
C*****PTRAO410
DIMENSION Q3(1),TG(1),H2(1),Q1(1),X8(1),SUMX3(1),SUMX4(1) PTRAO420
DIMENSION H4(1),Y2(1),H(1),SUMX2(1),Q2(1) PTRAO430
DIMENSION Q37(1), Q38(1), H3(1) PTRAO440
COMMON /LINDLY/ HNU,SGNL, IHNU, NHNU, HNUP, NT, IM, IN, DHNU, THICK, NY PTRAO450
COMMON /CNTRL/ SCYCLE, JMULT PTRAO460
COMMON /DAVIS/ ICX, ICY PTRAO470
COMMON /TQ/ QINT1(300), QINT2(300), TITLE(12) PTRAO480
EQUIVALENCE (SMLA,H4), (SMLD,Y2) PTRAO490
EQUIVALENCE (BC,TG), (BIGB,H), (CRTR,SUMX2), (CHIC,SUMX3) PTRAO500
EQUIVALENCE (SMLH,X8), (CAR,Q37), (CHIR,Q38), (SMLC,H3) PTRAO510
EQUIVALENCE (ACO3T4,TRDBG), (S12,EDITMF) PTRAG520
EQUIVALENCE (EC,Q1), (X7 ,H2), (BIGA,SUMX4), (GOFR, Q3) PTRAO530
C*****PTRAO540

```

```

DIMENSION RR(40)
DATA RR/2.113248E-01,7.886752E-01,1.056624E-01,3.943376E-01,
1 1.127017E-01,5.000000E-01,8.872983E-01,3.130600E-02,
2 2.222222E-01,2.464718E-01,6.943180E-02,3.300095E-01,
3 6.699905E-01,9.305682E-01,1.207610E-02,1.076071E-01,
4 2.184655E-01,1.618513E-01,4.691010E-02,2.307653E-01,
5 5.000000E-01,7.692347E-01,9.530899E-01,5.557100E-03,
6 5.522540E-02,1.422222E-01,1.840889E-01,1.129063E-01,
7 3.376520E-02,1.693953E-01,3.806903E-01,6.193096E-01,
8 8.306047E-01,9.662348E-01,2.892400E-03,3.055570E-02,
9 8.906520E-02,1.448918E-01,1.498251E-01,8.276980E-02/

C
C EDITMF SAME AS S12 *PTRAO660
C H SAME AS BIGB *PTRAO670
C H2 SAME AS X7 *PTRAO680
C H3 SAME AS SMLC *PTRAO690
C H4 SAME AS SMLA *PTRAO700
C Q1 SAME AS EC *PTRAO710
C Q3 SAME AS GOFR *PTRAO720
C Q37 SAME AS CAR *PTRAO730
C Q38 SAME AS CHIR *PTRAO740
C SUMX2 SAME AS CRTR *PTRAO750
C SUMX3 SAME AS CHIC *PTRAO760
C SUMX4 SAME AS BIGA *PTRAO770
C TG SAME AS BC *PTRAO780
C TROBG SAME AS AC03T4 *PTRAO790
C X8 SAME AS SMLH *PTRAO800
C Y2 SAME AS SMLD *PTRAO810
C *****PTRAO820
C PTRAO830
C PTRAO840
C PTRAO850
C PTRAO860
C FEX, FM, SUMRHO, CSQD, XSQD, Y, YSQD, Q2, NOT USED *PTRAO870
C PTRAO880
C PTRAO890
C *****PTRAO900
C *PTRAO910
C PLANES ONLY *PTRAO920
C *PTRAO930
C *****PTRAO940
C IAX=N *PTRAO950
C IBX=M *PTRAO960
C IN=IA *PTRAO970
C INM1=IN-1 *PTRAO980
C IMP1 = IM + 1 *PTRAO990
C CALL DVCHK(K000FX) *PTRA1000
C GO TO (100,110), K000FX *PTRA1010
100 S1=14.0100 *PTRA1020
CALL UNCLE *PTRA1030
110 IBXP1=IBX+1 *PTRA1040
IALPHA=ALPHA *PTRA1050
C *PTRA1060
C ERROR IF NOT PLANE *PTRA1070
C *PTRA1080
C GO TO (130,120,120), IALPHA *PTRA1090

```

120	S1=14.0120	PTRA1100
	CALL UNCLE	PTRA1110
130	NY = LMDA(37) - 1	PTRA1120
	NMU = (NY - 1) * (NY + 2) + 1	PTRA1130
	NGS = NMU + NY + 1	PTRA1140
	JJ = 0	PTRA1150
C		PTRA1160
C	DO POSITIVE ANGLES FIRST	PTRA1170
C		PTRA1180
140	I=IAX	PTRA1190
	F2=0.0	PTRA1200
C		PTRA1210
C	IF IAX=IN TRANSFER TO 150 TO SET SPECIAL BOUNDARY CONDITIONS	PTRA1220
C		PTRA1230
	IF (IAX-IN) 360,150,180	PTRA1240
C		PTRA1250
C	CALCULATE BOUNDARY SOURCE INTENSITY	PTRA1260
C		PTRA1270
150	IF (INM1) 160,310,170	PTRA1280
C		PTRA1290
C	SET BLACKBODY CONDITION FOR PUSHER	PTRA1300
C		PTRA1310
160	S1=14.0160	PTRA1320
	CALL UNCLE	PTRA1330
170	F2=X6(INM1)	PTRA1340
	GO TO 310	PTRA1350
C		PTRA1360
C	DIFFUSION BOUNDARY CONDITION AT IAX	PTRA1370
C		PTRA1380
180	IF (TG(I) .EQ. 0.) Y2(I) = X6(I-1)	PTRA1390
	TEMP(1)=H2(I-1)/RR(NMU)	PTRA1400
	X8(I)=TG(I)*RR(NMU)	PTRA1410
	IF (TEMP(1)-1.E-2) 190,190,200	PTRA1420
190	F2X = Y2(I-1) - TG(I-1) * RR(NMU)	PTRA1430
	F2 = ((Y2(I) + Y2(I-1)) * 0.5 + X6(I-1) - F2X-F2X) * TEMP(1) + F2X	PTRA1440
	GO TO 250	PTRA1450
200	H4(I-1)=FREXP(-TEMP(1))	PTRA1460
	F2=Y2(I)-X8(I)+(X3(I)-RR(NMU)*TG(I-1))*H4(I-1)	PTRA1470
	GO TO 250	PTRA1480
210	IF (TG(I-1) .EQ. 0.) Y2(I-1) = X6(I-1)	PTRA1490
	X8(I-1)=TG(I-1)*RR(NMU)	PTRA1500
C		PTRA1510
C	REGULAR INTEGRATION STEP FOR F2, POSITIVE MU	PTRA1520
C		PTRA1530
220	IF (TG(I) .EQ. 0.) Y2(I) = X6(I-1)	PTRA1540
	X8(I)=TG(I)*RR(NMU)	PTRA1550
	TEMP(1)=H2(I-1)/RR(NMU)	PTRA1560
	IF (TEMP(1)-1.E-2) 230,230,240	PTRA1570
230	F2 = ((Y2(I) + Y2(I-1)) * 0.5 + X6(I-1) - F2 - F2) * TEMP(1) + F2	PTRA1580
	GO TO 250	PTRA1590
240	H4(I-1)=FREXP(-TEMP(1))	PTRA1600
	F2=Y2(I)-X8(I)+((F2-Y2(I-1)+X8(I-1))*H4(I-1)	PTRA1610
	1+X8(I)-X8(I-1))*H4(I-1)	PTRA1620
250	IF (F2.LT.0.0) GO TO 280	PTRA1630
260	SUNX3(I)=F2	PTRA1640

	IF (TG(I) .EQ. 0.) Y2(I) = X6(I)	PTRA1650
	I=I+1	PTRA1660
	IF (I-IBXP1) 270,270,320	PTRA1670
270	IF (I-ICX-1) 220,220,290	PTRA1680
C		PTRA1690
C	NEGATIVE F2 ERROR	PTRA1700
C		PTRA1710
280	S1=14.0280	PTRA1720
	CALL UNCLE	PTRA1730
C		PTRA1740
C	NO SOURCE IN ZONE GREATER THAN ICX	PTRA1750
C		PTRA1760
290	IF (F2.EQ.0.0) GO TO 260	PTRA1770
	TEMP(1)=H2(I-1)/RR(NMU)	PTRA1780
	H4(I-1)=FREXP(-TEMP(1)-TEMP(1))	PTRA1790
	F2=F2+H4(I-1)	PTRA1800
	GO TO 260	PTRA1810
300	IF (F2.EQ.0.0) GO TO 310	PTRA1820
	TEMP(1)=H2(I-1)/RR(NMU)	PTRA1830
	H4(I-1)=FREXP(-TEMP(1)-TEMP(1))	PTRA1840
	F2=F2+H4(I-1)	PTRA1850
310	SUMX3(I) = F2	PTRA1860
	I=I+1	PTRA1870
	IF (I-ICY) 300,300,210	PTRA1880
C		PTRA1890
C	DO NEGATIVE ANGLES SECOND	PTRA1900
C		PTRA1910
320	I=IBXP1	PTRA1920
	IF (IBX-IM) 370,330,360	PTRA1930
330	IF (GL) 480,520,340	PTRA1940
C	GL = 1/2 MEANS BLACKBODY CONDITION SET AT IMP1	PTRA1950
C	GL = POSITIVE INTEGER MEANS INTENSITIES FROM QINT1 TABLE AT IMP1	PTRA1960
C	GL = 0 MEANS VACUUM AT IMP1	PTRA1970
C	GL NEGATIVE MEANS REFLECTIVE CONDITION AT IMP1	PTRA1980
340	IF (GL.NE.0.5) GO TO 350	PTRA1990
	F2 = X6(IMP1)	PTRA2000
	GO TO 480	PTRA2010
350	IQNT = JJ + 1 + (NY + 1) * (IHNU - 1)	PTRA2020
	F2 = QINT1(IQNT) / 68.5 * DHNU	PTRA2030
	GC TO 480	PTRA2040
C		PTRA2050
C	ERROR IF INDEX EXCEEDS NORMAL RANGE	PTRA2060
C		PTRA2070
360	S1=14.0360	PTRA2080
	CALL UNCLE	PTRA2090
C		PTRA2100
C	DIFFUSION BOUNDARY CONDITION AT IBXP1	PTRA2110
C		PTRA2120
370	IF (TG(I) .EQ. 0.) Y2(I) = X6(I)	PTRA2130
	TEMP(1)=H2(I)/RR(NMU)	PTRA2140
	IF (TEMP(1)-1.E-2) 380,380,390	PTRA2150
380	F2X = Y2(I+1) + TG(I+1) * RR(NMU)	PTRA2160
	F2 = ((Y2(I) + Y2(I+1)) * 0.5 + X6(I) - F2X - F2X) * TEMP(1) + F2X	PTRA2170
	GO TO 430	PTRA2180
390	H4(I)=FREXP(-TEMP(1))	PTRA2190

	F2=Y2(I)+X8(I)+(RR(NMU)*TG(I+1)-X8(I))*H4(I)	PTRA2200
	GO TO 430	PTRA2210
399	IF (TG(I+1) .EQ. 0.) Y2(I+1) = X6(I)	PTRA2220
C		PTRA2230
C	REGULAR INTEGRATION STEP FOR F2, NEGATIVE MU	PTRA2240
C		PTRA2250
400	IF (TG(I) .EQ. 0.) Y2(I) = X6(I)	PTRA2250
	TEMP(1)=H2(I)/RR(NMU)	PTRA2270
	IF (TEMP(1)-1.E-2) 410,410,420	PTRA2280
410	F2 = ((Y2(I) + Y2(I+1)) * 0.5 + X6(I) - F2 - F2) * TEMP(1) + F2	PTRA2290
	GO TO 430	PTRA2300
420	F2=Y2(I)+X8(I)+((F2-Y2(I+1)-X8(I+1))*H4(I)+X8(I+1)-X8(I))*H4(I)	PTRA2310
430	IF (F2.LT.0.) GO TO 460	PTRA2320
440	SUMX4(I)=F2	PTRA2330
C		PTRA2340
C	FORM CONTRIBUTION TO X2	PTRA2350
C		PTRA2360
	X2(I)=X2(I)-(F2-SUMX3(I))*RR(NGS)	PTRA2370
C		PTRA2380
C	RHO(I)=RHO(I)+(F2+SUMX3(I))*RR(NGS)/RR(NMU)	PTRA2390
C		PTRA2400
	IF (TG(I) .EQ. 0.) Y2(I) = X6(I-1)	PTRA2410
	I=I-1	PTRA2420
	IF (I-1AX) 530,450,450	PTRA2430
450	IF (I-1CY) 500,400,400	PTRA2440
460	S1=14.0460	PTRA2450
	CALL UNCLE	PTRA2460
C		PTRA2470
C	NO SOURCE IN ZONE LESS THAN ICY	PTRA2480
C		PTRA2490
470	IF (F2.EQ.0.0) GO TO 480	PTRA2500
	TEMP(1)=H2(I)/RR(NMU)	PTRA2510
	H4(I)=FREXP(-TEMP(1)-TEMP(1))	PTRA2520
	F2=F2*H4(I)	PTRA2530
480	SUMX4(I)=F2	PTRA2540
490	X2(I)=X2(I)-(F2-SUMX3(I))*RR(NGS)	PTRA2550
	I=I-1	PTRA2560
	IF (I-1-ICX) 399,470,470	PTRA2570
C		PTRA2580
C	NO SOURCE IN ZONE LESS THAN ICY	PTRA2590
C		PTRA2600
500	IF (F2.EQ.0.0) GO TO 510	PTRA2610
	TEMP(1)=H2(I)/RR(NMU)	PTRA2620
	H4(I)=FREXP(-TEMP(1)-TEMP(1))	PTRA2630
	F2=F2*H4(I)	PTRA2640
510	SUMX4(I)=F2	PTRA2650
	X2(I)=X2(I)-(F2-SUMX3(I))*RR(NGS)	PTRA2660
	I=I-1	PTRA2670
	IF (I-1AX) 530,500,500	PTRA2680
520	F2=0.0	PTRA2690
	GO TO 480	PTRA2700
530	CONTINUE	PTRA2710
	IF (TRDBG .EQ. 0.0 .OR. TRDBG .GE. SOLID(18)) GO TO 539	PTRA2720
C	STORE INTENSITIES FOR DEBUG PRINT	PTRA2730
	XX = -0.5	PTRA2740

```
JJJ = JJ + 1
WRITE (3) XX, IAX, JJJ, IBXP1, SOLID(18), TH, RR(NMU)
WRITE (3) (SUMX3(I), SUMX4(I), I = IAX, IBXP1)
XX = -2.
WRITE (3) (XX, I = i, 7)
BACKSPACE 3
539 DHNU=HNUP-HNU
540 JJ = JJ + 1
    NMU = NMU + 1
    NGS = NGS + 1
    IF (JJ-NY) 140,140,550
550 DO 560 I=IAX,IBXP1
560 X2(I) = X2(I)* 2.052E12
RETURN
END
```

```
PTRA2750
PTRA2760
PTRA2770
PTRA2780
PTRA2790
PTRA2800
PTRA2810
PTRA2820
PTRA2830
PTRA2840
PTRA2850
PTRA2860
PTRA2870
PTRA2880
PTRA2890
```

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