The Theoretical and Numerical Determination
of the Radar Cross Section of a Finite Cone

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Purdue Research Foundation
Lafayette, Indiana

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THEORETICAL AND NUMERICAL DETERMINATION
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F. V. Schultz, G. M. Ruckgaber, S. Richter, and J. K. Schindler

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This investigation of the radar cross-section of a finite cone can be divided into three areas. First, the exact solution for the scattering of a plane electromagnetic wave by a finite cone is presented. Rigorous electromagnetic theory is used in the solution, and no approximations are made. Secondly, methods of obtaining numerical results for the radar cross-section from the analytic solution by using a digital computer are discussed. The third area is a presentation and discussion of the numerical results obtained.
LIST OF CONTRIBUTORS

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RELATED CONTRACTS AND PUBLICATIONS

The present contract is a continuation of Contract No. 19(604)4051. On this earlier contract the following publications were produced:


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THEORETICAL AND NUMERICAL DETERMINATION
OF THE RADAR CROSS-SECTION OF A FINITE CONE

1. **Statement of the Problem**

   The problem undertaken is the exact solution for the scattering of a plane electromagnetic wave by a finite perfectly conducting cone. We consider only "nose-on" incidence (Figure 1). In order that the entire surface of the cone can be expressed as a constant co-ordinate surface in spherical co-ordinates, the end cap of the cone is taken to be a segment of a spherical surface with center at the apex of the cone. Time variations are assumed to be given by $e^{i\omega t}$ and MKS units are used.

   A considerable amount of effort, both theoretical and experimental, has been devoted to the cone scattering problem by many workers. No attempt is made here to summarize the work, in view of the very excellent summary which appears in the report by Kleinman and Senior (1963). It should be noted that the present work is an extension of that done earlier by Rogers and Schultz (1960), and by Rogers, Schindler, and Schultz (1962).

   This scattering problem is treated herein as a boundary-value problem in electromagnetic theory and no physical approximations are made. The partial differential equation is, of course, the vector Helmholtz equation,

   $$ \nabla^2 \vec{c} + k^2 \vec{c} = 0 $$  \hspace{1cm} (1)

   where $k = 2\pi/\lambda$ and $\vec{c}$ may be either the electric field vector $\vec{E}$ or
Fig. 1. Cone Configuration
the magnetic field vector $\mathbf{H}$. Solutions of (1) are obtained in the form of infinite series containing unknown constants. To complete the solution of the problem, these constants are determined by satisfying the necessary boundary conditions for $\mathbf{E}$ and $\mathbf{H}$ on the surface of the perfectly conducting cone, the radiation condition at infinity, and the finite energy condition.

Numerical results have been obtained, and these are compared with experimental results obtained elsewhere, as well as with theoretical results obtained with the use of approximate methods.

2. Solution of Vector Helmholtz Equation

The procedure used here for obtaining the solutions of the vector Helmholtz equation is well known (Stratton, 1941).

Solutions of (1) are

$$I = \nabla \phi,$$
$$\mathbf{m} = \nabla \times (\phi \mathbf{r}),$$
$$\mathbf{n} = \frac{1}{k} \nabla \times \mathbf{m},$$

where $\mathbf{r}$ is the radial vector in spherical co-ordinates and $\phi$ is the solution of the scalar Helmholtz equation

$$\nabla^2 \phi + k^2 \phi = 0.$$  \hspace{1cm} (3)

In the region surrounding the cone, $\nabla \cdot \mathbf{E} = \nabla \cdot \mathbf{H} = 0$. Since $\nabla \cdot I \neq 0$, we use only the $\mathbf{m}$ and $\mathbf{n}$ solutions to represent $\mathbf{E}$ and $\mathbf{H}$. 

It is also well known that the solution of (3) is

\[ \phi_{n}(r, \theta, \phi) = z_{n}^{m}(kr) P_{v}^{m}(\cos \theta) \left[ \begin{array}{c} \cos m\phi \\ \sin m\phi \end{array} \right], \quad (4) \]

where \( n \) can have the values 1, 2, 3, or 4 to represent Bessel functions of the first kind \( (J_{n}(kr)) \), Bessel functions of the second kind \( (n_{n}(kr)) \), Hankel functions of the first kind \( (h^{1}_{n}(kr)) \), and Hankel functions of the second kind \( (h^{2}_{n}(kr)) \), respectively. \( P_{v}^{m}(\cos \theta) \) is an associated Legendre function of degree \( v \) and order \( m \), and we let \( e \) signify "even" and \( o \) signify "odd" for \( \cos m\phi \) and \( \sin m\phi \), respectively.

The desired solutions of the vector Helmholtz equation are then obtained from (2) and (4):

\[
\begin{align*}
\frac{m}{r} e_{m} & = \frac{m}{\sin \theta} z_{v}^{n}(kr) P_{v}^{m}(\cos \theta) \left[ \begin{array}{c} \sin m\phi \\ \cos m\phi \end{array} \right] \hat{a}_{\theta} \\
- z_{v}^{n}(kr) \frac{dP_{v}^{m}}{d\theta} \left[ \begin{array}{c} \cos m\phi \\ \sin m\phi \end{array} \right] \hat{a}_{\phi} \\
- \frac{m}{r} e_{m} & = \frac{v(v+1)}{kr} z_{v}^{n}(kr) P_{v}^{m}(\cos \theta) \left[ \begin{array}{c} \cos m\phi \\ \sin m\phi \end{array} \right] \hat{a}_{r} \\
+ z_{v}^{n}(kr) \frac{dP_{v}^{m}}{d\theta} \left[ \begin{array}{c} \cos m\phi \\ \sin m\phi \end{array} \right] \hat{a}_{\theta} \\
- \frac{m}{\sin \theta} z_{v}^{n}(kr) P_{v}^{m}(\cos \theta) \left[ \begin{array}{c} \sin m\phi \\ \cos m\phi \end{array} \right] \hat{a}_{\phi}
\end{align*}
\]

where \( z_{v}^{n}(kr) = \frac{1}{kr} \frac{d}{dr} \left[ r z_{v}^{n}(kr) \right] \), and \( \hat{a}_{r}, \hat{a}_{\theta}, \) and \( \hat{a}_{\phi} \) are the spherical unit vectors.
3. Space Sectionalization

One of the most important characteristics of the solution of this problem is that of dividing the space surrounding the cone into two regions to facilitate the field expansions and application of the boundary conditions. The $E$ and $H$ fields are then expanded in terms of the radial and spherical functions appropriate to each region.

Since the scattered fields must be spherically diverging waves for large values of the co-ordinate $r$, the use of Hankel functions is obvious since they possess the desired wave behavior as $r \to \infty$. In particular, since we assume a time variation of the form $e^{i\omega t}$, the use of $z_n^h(kr) = h_n^2(kr)$ functions is necessary to achieve an outward traveling wave. At the tip of the cone, however, the Hankel functions possess a singularity the order of which is too large to satisfy the finite energy condition. This characteristic of the radial functions suggests a division of the two regions at a finite value of $r$.

The behavior of the associated Legendre functions indicates a division of the two regions at $r = b$. This is then the surface that we use to separate regions I and II (Figure 2). In region II, the fields exist and are bounded everywhere in the complete $\theta$ domain of $0 \leq \theta \leq \pi$, requiring the use of only associated Legendre functions of integral degree. In region I, however, $\theta = \pi$ is not in the domain of interest, allowing the use of associated Legendre functions of non-integral degree. It will be seen that the boundary conditions will determine the non-integral degree of each associated Legendre function to be used in region I.
Fig. 2. Space Sectionalization
The reader may wish to refer to Rogers and Schultz (1960) for a more complete discussion of the selection of the various space divisions possible.

4. Field Expansions

In region I the total fields are designated by \( E_i^t \) and \( H_i^t \). In region II we wish to keep the incident and scattered fields separate, and designate the incident fields by \( E_{II}^i \) and \( H_{II}^i \) and the scattered fields by \( E_{II}^s \) and \( H_{II}^s \).

In region II the incident electric field may be expressed (Stratton, 1941) as

\[
E_{II}^i = e^{i k z} \hat{a}_x = e^{i k r \cos \theta} \hat{a}_x = \sum_{n=1}^{\infty} \left( H_n \frac{m}{\sin \theta} + \Gamma_n \frac{n}{\sin \theta} \right),
\]

where

\[
\gamma_n = i^{n} \frac{2n+1}{n(n+1)} , \quad \Gamma_n = -i^{n+1} \frac{2n+1}{n(n+1)},
\]

and \( \hat{a}_x \) is a unit vector in the x direction. The \( \phi \) variation in the incident field requires that \( m=1 \) and forces us to use odd \( m \) functions and even \( n \) functions in all expansions of the electric field.

The scattered field in region II is written as

\[
E_{II}^s = \sum_{n=1}^{\infty} \left( c_n \frac{m}{\sin \theta} + d_n \frac{n}{\sin \theta} \right),
\]

where \( c_n \) and \( d_n \) are expansion coefficients to be determined from the boundary conditions. Here we have selected \( E_n^i(kr) = h_n^e(kr) \) and the \( m \) and \( n \) functions as previously discussed.
in region I the total electric field is expressed as

\[ \mathbf{E}_I^t = \sum_{\nu} a_{\nu} \mathbf{n}^{\mathbf{ol}}_{\nu} + \sum_{\mu} b_{\mu} \mathbf{n}^{\mathbf{el}}_{\mu}. \]  

(10)

Here \( a_{\nu} \) and \( b_{\mu} \) are expansion coefficients to be determined, and \( \mu \) and \( \nu \) are the non-integral degrees of the associated Legendre functions, which are also yet to be determined.

The analogous representations for the magnetic field are obtained from Maxwell's equations,

\[ \nabla \times \mathbf{E} = -i \omega \mu \mathbf{H}, \quad \nabla \times \mathbf{H} = i \omega \varepsilon \mathbf{E}, \]  

(11)

and the relations,

\[ \nabla \times \mathbf{m} = \mathbf{k} \mathbf{n}, \quad \nabla \times \mathbf{n} = \mathbf{k} \mathbf{m}. \]  

(12)

By using (7) through (10), in addition to (11) and (12), and noting that \( k = \omega \mu \varepsilon \), one obtains the expressions for the magnetic fields:

\[ \mathbf{H}_I^t = \frac{1}{\eta} \left[ \sum_{n=1}^{\infty} \left( c_n \mathbf{n}^{\mathbf{ol}}_{\mathbf{n}} + d_n \mathbf{n}^{\mathbf{el}}_{\mathbf{n}} \right) \right], \]  

(13)

\[ \mathbf{H}_I^s = \frac{1}{\eta} \left[ \sum_{n=1}^{\infty} \left( c_n \mathbf{n}^{\mathbf{ol}}_{\mathbf{n}} + d_n \mathbf{n}^{\mathbf{el}}_{\mathbf{n}} \right) \right], \]  

(14)

\[ \mathbf{H}_I^r = \frac{1}{\eta} \left[ \sum_{\nu} a_{\nu} \mathbf{n}^{\mathbf{ol}}_{\nu} + \sum_{\mu} b_{\mu} \mathbf{n}^{\mathbf{el}}_{\mu} \right], \]  

(15)

where \( \eta \) is the intrinsic impedance of free space, \( \sqrt{\mu_0/\varepsilon_0} \).

For future reference, the field quantities are now expanded in their entirety:
\[ \mathbf{E}_t^i = \left[ \sum u \left( \frac{u(u+1)}{kr} \right) J_u(kr) P_u^1(\cos \theta) \right] \cos \phi \hat{a}_r \\
+ \left[ \sum a J_u(kr) \frac{P_u^1(\cos \theta)}{\sin \theta} + \sum b J_u(kr) \frac{dP_u^1}{d\theta} \right] \cos \phi \hat{a}_\theta \\
- \left[ \sum a J_u(kr) \frac{dP_u^1}{d\theta} + \sum b J_u(kr) \frac{P_u^1(\cos \theta)}{\sin \theta} \right] \sin \phi \hat{a}_\phi . \]  

(16)

\[ \mathbf{E}_t^{II} = \sum_{n=1}^{\infty} \left\{ \left[ \Gamma_n \frac{n(n+1)}{kr} J_n(kr) P_n^1(\cos \theta) \right] \cos \phi \hat{a}_r \\
+ \left[ \gamma_n J_n(kr) \frac{P_n^1(\cos \theta)}{\sin \theta} + \Gamma_n J_n(kr) \frac{dP_n^1}{d\theta} \right] \cos \phi \hat{a}_\theta \\
- \left[ \gamma_n J_n(kr) \frac{dP_n^1}{d\theta} + \Gamma_n J_n(kr) \frac{P_n^1(\cos \theta)}{\sin \theta} \right] \sin \phi \hat{a}_\phi \right\} . \]

(17)

\[ \mathbf{E}_t^{II} = \sum_{n=1}^{\infty} \left\{ \left[ \alpha_n \frac{n(n+1)}{kr} h_n(kr) P_n^1(\cos \theta) \right] \cos \phi \hat{a}_r \\
+ \left[ \alpha_n h_n(kr) \frac{P_n^1(\cos \theta)}{\sin \theta} + d_n h_n(kr) \frac{dP_n^1}{d\theta} \right] \cos \phi \hat{a}_\theta \\
- \left[ \alpha_n h_n(kr) \frac{dP_n^1}{d\theta} + d_n h_n(kr) \frac{P_n^1(\cos \theta)}{\sin \theta} \right] \sin \phi \hat{a}_\phi \right\} . \]

(18)

\[ \mathbf{H}_t^{I} = \frac{i}{\eta} \left\{ \left[ \sum a J_0(kr) P_0^1(\cos \theta) \right] \sin \phi \hat{a}_r \\
+ \left[ \sum a J_0(kr) \frac{dP_0^1}{d\theta} - \sum b J_0(kr) \frac{P_0^1(\cos \theta)}{\sin \theta} \right] \sin \phi \hat{a}_\theta \\
+ \left[ \sum a J_0(kr) \frac{P_0^1(\cos \theta)}{\sin \theta} - \sum b J_0(kr) \frac{dP_0^1}{d\theta} \right] \cos \phi \hat{a}_\phi \right\} . \]  

(19)
\[ h_{II}^1 = \frac{1}{\pi} \sum_{n=1}^{\infty} \left\{ \left[ \gamma_n \frac{n(n+1)}{kr} J_n(kr) P_n^1(\cos \theta) \right] \sin \theta \hat{a}_r + \left[ \gamma_n J_n'(kr) \frac{dP_n^1}{d\theta} - \Gamma_n J_n(kr) \frac{P_n^1(\cos \theta)}{\sin \theta} \right] \sin \theta \hat{a}_\theta \right\}. \]

\[ h_{IIIV}^1 = \frac{1}{\pi} \sum_{n=1}^{\infty} \left\{ \left[ c_n \frac{n(n+1)}{kr} h_n(kr) P_n^1(\cos \theta) \right] \sin \theta \hat{a}_r + \left[ c_n h_n'(kr) \frac{dP_n^1}{d\theta} - d_n h_n(kr) \frac{P_n^1(\cos \theta)}{\sin \theta} \right] \sin \theta \hat{a}_\theta \right\}. \] (20)

Equations (16) through (21) contain six sets of unknown constants, \( u, v, a, b, c, \) and \( d \). These are to be determined by satisfying the boundary conditions.

5. **Boundary Conditions**

We have already satisfied the finite energy condition at the tip of the cone and the radiation condition at infinity by the proper choice of radial functions in each region. The following boundary conditions remain to be satisfied:

(a) \[ \left[ \bar{E}^t_{II} \right]_{r, \theta} = 0 \quad \text{for} \quad \theta = \theta_c, \quad r \leq b \] (22a)

(b) \[ \left[ \bar{E}^t_{II} + \bar{E}^s_{II} \right]_{\theta, \phi} = \left\{ \begin{array}{ll} \left[ \bar{E}^t_{II} \right]_{\theta, \phi} & \text{for} \quad 0 \leq \theta < \theta_c \quad \text{for} \quad r = b \\ 0 & \text{for} \quad \theta_c \leq \theta \leq \pi \end{array} \right. \] (22b)
(c) \[ \left[ \bar{H}_{IJ}^r + \bar{H}_{II}^r \right]_{\theta,\phi} = \left[ \bar{H}_{I}^r \right]_{\theta,\phi} \quad \text{for } r = b, \quad 0 \leq \theta < \theta_0 \] \quad (22c)

(d) The finite energy condition at the edge of the cone
\( (r = b, \theta = \theta_0), \) \quad (22d)
where b is the radius of the spherical cap and \( \theta_0 \) is half of the exterior apex angle.

6. The Solution

To satisfy boundary condition (22a) we first equate the \( r \)-component of \( \bar{E}^t_I \) to zero at \( \theta = \theta_0 \),
\[ \sum_b \mu(\mu + 1) \frac{dJ_\mu(kr)}{kr} P_\mu^l(\cos \theta_0) \cos \phi = 0 , \] \quad (23)
and thus set
\[ P_\mu^l(\cos \theta_0) = 0 . \] \quad (24)

This equation determines the values of \( \mu \). Equating the \( \phi \)-component of \( \bar{E}^t_I \) to zero at \( \theta = \theta_0 \) gives
\[ \sum_v a_v J_v(kr) \left. \frac{dF_v^l}{d\theta} \right|_{\theta=\theta_0} + \sum_\mu b_\mu J_\mu(kr) \frac{F_\mu^l(\cos \theta_0)}{\sin \theta_0} = 0 . \] \quad (25)

Since \( F_\mu^l(\cos \theta_0) = 0 \) by (24), we set
\[ \left. \frac{dF_v^l}{d\theta} \right|_{\theta=\theta_0} = 0 \] \quad (26)
and thus the values of \( v \) are determined.
Next, the boundary conditions (22b) and (22c) are applied to
determine the four sets of unknown expansion co-efficients \( a, b, c_n, \) and \( d_n \). For the \( \theta \) component of (22b) there results

\[
\sum_{n=1}^{\infty} \left[ \gamma_n J_n(k) \frac{P_n^l(\cos \theta)}{\sin \theta} + \Gamma_n J_n'(k) \frac{dP_n^l}{d\theta} \right] \cos \theta = 0, \quad 0 \leq \theta < \theta_0,
\]

and for the \( \phi \) component,

\[
\sum_{n=1}^{\infty} \left[ \gamma_n J_n(k) \frac{dP_n^l}{d\theta} + \Gamma_n J_n'(k) \frac{P_n^l(\cos \theta)}{\sin \theta} \right] \sin \theta = 0, \quad \theta_0 \leq \theta < \pi.
\]

Similarly, for the \( \theta \)-component of (22c) there results

\[
\sum_{n=1}^{\infty} \left[ c_n J_n(k) \frac{dP_n^l}{d\theta} + d_n J_n'(k) \frac{P_n^l(\cos \theta)}{\sin \theta} \right] \sin \theta = 0, \quad 0 \leq \theta < \theta_0,
\]

and for the \( \phi \) component,

\[
\sum_{n=1}^{\infty} \left[ c_n J_n(k) \frac{P_n^l(\cos \theta)}{\sin \theta} + d_n J_n'(k) \frac{dP_n^l}{d\theta} \right] \cos \theta = 0, \quad \theta_0 \leq \theta < \pi.
\]
\[ \frac{1}{n} \sum_{n=1}^{\infty} \left[ \frac{\gamma_n^J \mu_n(kb)}{\sin \theta} - \frac{\gamma_n \mu_n(kb)}{\sin \theta} \right] \sin \theta \]

\[ + \frac{1}{n} \sum_{n=1}^{\infty} \left[ \frac{c_n h_n(kb)}{\sin \theta} - \frac{d_n h_n(kb)}{\sin \theta} \right] \sin \theta \]

\[ = \frac{1}{n} \sum_{\nu} \left[ a_{\nu} \chi_{\nu}(kb) \frac{dP_n^{1/2}}{d\theta} \right] - \sum_{\mu} b_{\mu} \chi_{\mu}(kb) \frac{P_n^{1/2}(\cos \theta)}{\sin \theta} \]

\[ \sin \theta, 0 \leq \theta \leq \theta_0, \]  

(29)

and for the \( \phi \)-component,

\[ \frac{1}{n} \sum_{n=1}^{\infty} \left[ \frac{\gamma_n^J \mu_n(kb)}{\sin \theta} - \frac{\gamma_n \mu_n(kb)}{\sin \theta} \right] \cos \phi \]

\[ + \frac{1}{n} \sum_{n=1}^{\infty} \left[ \frac{c_n h_n(kb)}{\sin \theta} - \frac{d_n h_n(kb)}{\sin \theta} \right] \cos \phi \]

\[ = \frac{1}{n} \sum_{\nu} \left[ a_{\nu} \chi_{\nu}(kb) \frac{dP_n^{1/2}}{d\theta} \right] - \sum_{\mu} b_{\mu} \chi_{\mu}(kb) \frac{P_n^{1/2}(\cos \theta)}{\sin \theta} \]

\[ \cos \phi, 0 \leq \theta \leq \theta_0. \]  

(30)

These four equations, (27), (28), (29), and (30), are functions of \( \theta \), (27) and (28) over the interval \( 0 \leq \theta \leq \pi \) and (29) and (30) over the interval \( 0 \leq \theta \leq \theta_0 \). In the solution of Rogers and Schultz (1960) these four equations were manipulated in a process that involved differentiation with respect to \( \theta \). It is well known that an infinite series can be integrated term-by-term with non-stringent requirements on the nature of convergence, whereas term-by-term differentiation of
an infinite series is valid only with strict requirements on the convergence of the series. Since the exact nature of the convergence of the infinite series expansions in (27) through (30) is unknown, we here use an integration process, in order to avoid the problems encountered with differentiation.

First we multiply (27) by \( P_m^1(\cos \theta) \), multiply (28) by \( \sin \theta \frac{dP_m^1}{d\theta} \), and subtract the two results. We then integrate the resulting equation with respect to \( \theta \) over the interval \( 0 \) to \( \pi \). It is necessary to evaluate two integrals with limits of \( 0 \) to \( \pi \) and two integrals with limits \( 0 \) to \( \theta_0 \). The integrals are common to boundary value problems of this type and can be evaluated by using the associated Legendre differential equation, and (24) and (26). The integral that appears as a factor in the \( c_n \) summation fortunately involves the Kronecker delta, \( \delta_{mn} \), enabling the coefficient \( c_m \) to be separated.

The coefficient \( d_m \) is separated in exactly the same manner except that (27) is multiplied by \( \sin \theta \frac{dP_m^1}{d\theta} \) and (28) by \( P_m^1(\cos \theta) \).

To separate \( a_\alpha \), (29) is multiplied by \( \sin \theta \frac{dP_\alpha^1}{d\theta} \) and (30) by \( P_\alpha^1 \) and the results added. The subscript \( \alpha \) denotes a particular value of the infinite set \( v \). This equation is then integrated with respect to \( \theta \) over the interval \( 0 \) to \( \theta_0 \). Again the integrals can be evaluated by using the associated Legendre differential equation, and (24) and (26). Here the integral associated with the \( a_v \) summation involves the Kronecker delta, \( \delta_{v\alpha} \), enabling the coefficient \( a_\alpha \) to be separated.
The coefficient \( b_\beta \) is separated in the same manner as is \( a_\alpha \), except that (29) is multiplied by \( P_\beta^1 \) and (30) by \( \sin \theta \frac{dP_\beta^1}{d\theta} \).

The subscript \( \beta \) denotes a particular value of the infinite set \( \mu \).

If the values of \( \gamma_n \) and \( \Gamma_n \) given by (8) are then substituted in the four separated equations, there result:

\[
c_m = \frac{-i^m (2m+1) J_m'(kb)}{m(m+1) h_m(kb)}
\]

\[
c_m + \frac{(2m+1) \sin \theta}{2[m(m+1)]^2 h_m(kb)} \frac{dP_m^1}{d\theta} \bigg|_{\theta=\theta_0} \sum_\nu a_\nu \gamma_\nu(kb) P_\nu^1(\cos \theta_0)
\]

\[
d_n = \frac{i^{n+1} (2m+1) J_n'(kb)}{m(m+1) h_n'(kb)}
\]

\[
d_n + \frac{(2m+1) P_m^1(\cos \theta_0)}{2m(m+1) h_n'(kb)} \sum_\nu a_\nu \gamma_\nu(kb) P_\nu^1(\cos \theta_0)
\]

\[
d_n + \frac{(2m+1) \sin \theta_0 P_m^1(\cos \theta_0)}{2m(m+1) h_n'(kb)} \sum_\mu \frac{b_{\mu \mu} J_{\mu}'(kb)}{m(m+1) - \mu(\mu+1)} \left. \frac{dP_\mu^1}{d\theta} \right|_{\theta=\theta_0}
\]

\[
a_\alpha = \frac{\sin \theta_0 P_\alpha^1(\cos \theta_0)}{B_\alpha J_\alpha'(kb)} \sum_{n=1}^{\infty} \left\{ \left[ c_n h_n'(kb) + \frac{i^n (2n+1) J_n'(kb)}{n(n+1)} \right] \frac{dP_n^1}{d\theta} \bigg|_{\theta=\theta_0} \right\}
\]

\[
a_\alpha - \frac{P_\alpha^1(\cos \theta_0)}{a(\alpha+1) \beta_\alpha J_\alpha'(kb)} \sum_{n=1}^{\infty} \left\{ \left[ d_n h_n(kb) - \frac{i^{n+1} (2n+1) J_n(kb)}{n(n+1)} \right] P_n^1(\cos \theta_0) \right\}
\]
 where the quantities $B_\alpha$ and $B_\beta$ are defined by

$$
B_\gamma = \int_0^{\theta_0} \sin \theta \left( P_\gamma^1 \right)^2 d\theta .
$$

The reader may wish to refer to Appendix A for the analytic details of the derivation of (31) thru (35).

Equations (31) through (34) could be manipulated into four equations with each set of coefficients appearing in only one equation, but the form of the end result would be less convenient for numerical computation. Therefore, (31) through (34), together with (16) through (21), (24), (26), and (35) represent the formal solution of the problem.

We have completed the solution without the necessity of satisfying boundary condition (22d), the finite energy condition at the edge of the cone. Rogers and Schultz (1960) used numerical results to show that this finite energy condition appears to be satisfied at the edge of the cone.

One of the primary objectives of the solution of this problem is to investigate the radar cross-section of the cone. The radar cross-section, $\sigma$, is defined to be

$$
\sigma = \lim_{r \to \infty} 4\pi r^2 \left| \frac{\bar{S}_{III}}{S_{II}} \right|^2 , \tag{36}
$$

where $\bar{S} = \frac{1}{2} \text{Re} \left\{ \vec{E} \times \vec{H}^* \right\}$, the average Poynting vector. For our coordinate system, the radar cross-section evaluated at $\theta = 0$ is more
precisely termed the back scattering radar cross-section, \( \sigma_{BS} \). By using some simple algebra, \( \sigma_{BS} \) can be shown to be expressed by

\[
\sigma_{BS} = \frac{\lambda^2}{4\pi} \left| \sum_{n} i^n n(n+1) \left( c_n - id_n \right) \right|^2,
\]

where \( \lambda \) is the wavelength of the incident plane wave. In order to determine the back-scattering radar cross-section, then, we must first determine the sets of \( c_n \) and \( d_n \).

7. The Numerical Solution

Equations (31) through (34) represent an infinite number of equations in an infinite number of unknown expansion coefficients. The expansion coefficients, therefore, do not enjoy the property of finality. It is important, then, to calculate as many of the coefficients as possible in order to insure that the values of the lowest order coefficients are reasonably accurate. The number of coefficients calculated in each set is designated by \( n_0 \). All numerical work was done for \( \theta_0 = 165^\circ \) (a cone apex angle of 30°). The calculations have been carried out for a rather large number of values of \( ka \) in order to determine rather well the details of the graph of \( \sigma_{BS} \) vs. \( ka \), \( a \) being the radius of the base of the cone.

An examination of (31) through (34) indicates that the following sets of constants need to be determined: \( \mu, \nu, B_\mu, B_\nu, P^1_n(\cos 165^\circ), \quad P^1_\nu(\cos 165^\circ), \quad \frac{dP^1_n}{d\sigma}_{\theta=165^\circ}, \quad \frac{dP^1_\mu}{d\sigma}_{\theta=165^\circ}, \quad J_n(kb), \quad J_\mu(kb), \quad J_\nu(kb), \quad \).
$J_n'(kb)$, $J_n''(kb)$, $J_n''(kb)$, $h_n(kb)$, and $h_n'(kb)$. The first thirty values, each, of $u$ and $v$, as determined from (24) and (26), were taken from Waterman's paper (1963), and these have seven-place precision. With the exception of the radial functions, the remaining sets of constants were calculated by Schultz, Bolle, and Schindler (1963), using a Burroughs Datatron 205 computer. The reader may wish to refer to their work for a detailed presentation of the methods used in calculating these constants. Appendix B herein lists the first fifty-one values in each set of these constants along with the given sets $u$ and $v$. The radial function constants were calculated by using the infinite series representations for the spherical Bessel functions.

When these constants are substituted into (31) through (34), $4n_o$ equations in $4n_o$ unknowns result. These $4n_o$ equations must then be solved for the four sets of $n_o$ expansion coefficients. Two different methods were used to accomplish this. First, an iterative method was used for values of $ka$ in the Rayleigh region. Secondly, a more complicated method, but one that is usable for any value of $ka$ (a standard matrix solution), was used for the higher values of $ka$.

For the iterative solution we first assume initial values of the $c_n$ and $d_n$ in (33) and (34) and obtain initial values of the $a_v$ and $b_u$. These values of $a_v$ and $b_u$ are then substituted into (31) and (32) and new values for the $c_n$ and $d_n$ are obtained. The process is then repeated, and continued until the values of the coefficients approach a final value. This method did not converge for values of $ka$ greater than 0.518.
The second method is more complicated but more useful since it is applicable for higher values of $k\alpha$. This method involves a straightforward matrix multiplication. If $n_a$ coefficients are to be calculated, (31) through (34) can be written in the form

\begin{align}
\vec{c} &= E_1 + E_2 \vec{a} \\
\vec{d} &= F_1 + F_2 \vec{a} + F_3 \vec{b} \\
\vec{b} &= G_1 \vec{c} + G_2 + G_3 \vec{a} + G_4 + G_5 \\
\vec{a} &= H_1 \vec{d} + H_2,
\end{align}

where

\begin{align}
\vec{c} &= \begin{bmatrix}
c_1 \\
c_2 \\
\vdots \\
c_{n_a}
\end{bmatrix}, \\
\vec{d} &= \begin{bmatrix}
d_1 \\
d_2 \\
\vdots \\
d_{n_a}
\end{bmatrix}, \\
\vec{a} &= \begin{bmatrix}
a_{v_1} \\
a_{v_2} \\
\vdots \\
a_{v_{n_a}}
\end{bmatrix}, \\
\vec{b} &= \begin{bmatrix}
b_{v_1} \\
b_{v_2} \\
\vdots \\
b_{v_{n_a}}
\end{bmatrix}.
\end{align}
$E_1, F_1, G_2, G_4$, and $H_2$ are $n_0$-by-1 matrices, and $E_2, F_2, F_3, G_1, G_3$, and $H_1$ are $n_0$-by-$n_0$ matrices. Since only $\vec{c}$ and $\vec{d}$ are needed to calculate $\sigma_{BS}$, we substitute (40) and (41) into (38) and (39) to eliminate $\vec{a}$ and $\vec{b}$. The two resulting relations can then be written in the form

$$
\begin{align*}
\begin{bmatrix}
I - E_2 G_1
\end{bmatrix} \vec{c} + \begin{bmatrix}
- E_2 G_3
\end{bmatrix} \vec{d} &= \begin{bmatrix}
E_1 + E_2 (G_2 + G_4)
\end{bmatrix} \\
\begin{bmatrix}
- F_2 G_1
\end{bmatrix} \vec{c} + \begin{bmatrix}
I - F_2 G_3 - F_3 H_1
\end{bmatrix} \vec{d} &= \begin{bmatrix}
F_1 + F_2 (G_2 + G_4) + F_3 H_2
\end{bmatrix},
\end{align*}
$$

(44)

where $I$ is the identity matrix. If we define

$$
\vec{x} = \begin{bmatrix}
\vec{c} \\
\vec{d}
\end{bmatrix},
$$

(46)

then (44) and (45) can be expressed as

$$
A \vec{x} = \vec{b},
$$

(47)

and so the desired solution is

$$
\vec{x} = A^{-1} \vec{b}.
$$

(48)

We then have the values of $c_1, c_2, \ldots, c_{n_0}, d_1, d_2, \ldots, d_{n_0}$, enabling $\sigma_{BS}$ to be calculated by using (37).

All calculations were accomplished by using an IBM-7090 digital computer programmed in FORTRAN. The quantities $k_b$ and $n_0$ were input parameters which could be changed at will. The sets of constants $\mu$, $\nu$, $B_\mu$, $B_\nu$, and the values of the associated Legendre functions were read into the machine as input data, whereas the values of the radial
functions were calculated at the beginning of the program, since the latter are dependent upon \( k_b \). In the case of the iterative method of solution, the values of \( a_v, b_u, c_n, \) and \( d_n \) were printed out, either after every iteration or after every fifth iteration, depending upon the speed of convergence to the final values. A subroutine for \( \sigma_{BS} \) was included at the end of the program, and the value of \( \sigma_{BS} \) was also printed. In the case of the matrix method, only the values of \( c_n, d_n, \) and \( \sigma_{BS} \) were printed, as the \( a_v \) and \( b_u \) were not computed in this latter method. Also, in the matrix method an additional input parameter \( s_o \) was used, \( s_o \) being the number of terms retained in the summations in the elements of the matrices \( G, G^*, \) and \( H \) (see Appendix C).

8. The Numerical Results

Table 1 lists 60 calculated values of \( \sigma_{BS}/\pi a^2 \) for 50 different values of \( k_a \). For some values of \( k_a \), \( \sigma_{BS}/\pi a^2 \) was calculated for several values of \( n_o \) with \( k_a \) fixed, in order to determine the sensitivity of \( \sigma_{BS}/\pi a^2 \) to \( n_o \), \( n_o \) being the number of expansion coefficients calculated in each set. Nine values of \( \sigma_{BS}/\pi a^2 \) were calculated by using the iteration method. As mentioned previously, it was found that the iterations would not converge for values of \( k_a \) above 0.518. The matrix method was then used for all calculations for \( k_a > 0.518 \).

Two different programs were written using the matrix method. A maximum of only 30 coefficients \( (n_o = 30) \) could be calculated by using the first program, and the maximum value of \( s_o \) was also limited to 30.
<table>
<thead>
<tr>
<th>$k_b$ (1)</th>
<th>$k_a$ (2)</th>
<th>$n_o$ (3)</th>
<th>Iteration method</th>
<th>Matrix method</th>
</tr>
</thead>
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(1) $b$ is the slant height of the cone
(2) $a$ is the base radius of the cone
(3) $n_o$ is the number of expansion coefficients calculated in each of the sets $a_v$, $b_w$, $c_n$, and $d_n$
(4) $s_o$ is the number of terms retained in the summations in the elements of the matrices $G_2$, $G_4$, and $H_2$
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<th>$k_b(1)$</th>
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</table>
The second program offered greater flexibility, the maximum values of 
\(n_o\) and \(s_o\) being 45 and 50, respectively. These latter maximum values 
of \(n_o\) and \(s_o\) were limited by the memory capacity of the IBM-7090 digital 
computer. Both of the matrix programs were used to calculate \(\sigma_{BS}/na^2\) 
for \(ka = 0.933\), and the values obtained were the same, 1.530, for the 
same number of terms, \(n_o\). This was reassuring, considering the 
completely dissimilar nature of the sequence of calculations in the 
two matrix programs.

Fig. 3 shows a graph of \(\sigma_{BS}/na^2\) vs. \(ka\), where \(a\) is the radius of 
the base of the cone (\(a = b \sin 15^\circ\)). For those values of \(\sigma_{BS}/na^2\) 
calculated for several different values of \(n_o\) with \(ka\) fixed, the value 
of \(\sigma_{BS}/na^2\) corresponding to the largest \(n_o\) is used in Fig. 3.

The values of \(\sigma_{BS}/na^2\) in the Rayleigh region (\(ka < 0.4\)) are not 
shown in Fig. 3, but it is important to note that \(\sigma_{BS}/na^2\) obeys 
extremely well the \(\lambda^{-4}\) law predicted for this region. Furthermore, by 
using an approximate method of Siegel (Siegel, 1959), applicable in the 
Rayleigh region, it can be shown that at \(ka = 0.0259\) the normalized 
back-scattering radar cross-section is approximately \(3.56 \times 10^{-6}\). The 
value for \(ka = 0.0259\) from Table 1 is \(3.93 \times 10^{-6}\), which agrees with 
the approximate value of Siegel rather well.

For higher values of \(ka\) the graph shows unexpectedly rapid 
fluctuations. It is believed that these are caused by convergence 
difficulties, especially since a curve obtained by using \(n_o = 30\) instead 
of \(n_o = 45\) showed even wilder fluctuations for \(ka > 3.2\).
Fig. 3. Values of the Normalized Back-Scattering Radar Cross Section ($\frac{\sigma_{BS}}{\pi a^2}$) for a 30-Degree Perfectly Conducting Cone of Slant Height $b$. 

$k = \frac{2 \pi}{\lambda}$

$a = b \sin 15^\circ$

Keys (Experimental)

Keller

Siegel (Rayleigh)

Schultz, et al
Also shown in Fig. 3 are values of $\sigma_{RS}/\pi a^2$ calculated by using Keller's modified geometrical optics theory (Keller, 1960). Double diffraction effects are included.

Mr. John E. Keys of the Defence Research Telecommunications Establishment, Ottawa, Ontario, very kindly supplied the present authors with detailed data from measurements similar to those on which he and R. I. Primich reported in the Canadian Journal of Physics (1959). Mr. Keys has given his permission for the inclusion of these measurements in the present report, and they are plotted in Fig. 3. These measurements were made on flat-based cones but Mr. Keys has informed the present authors that he has made measurements on spherically-capped cones, of the type analyzed in the present work, and these measurements are indistinguishable from those made on flat-based cones.

Likewise included in Fig. 3 is the straight-line graph representing the results obtained by using Siegel's modified Rayleigh theory (Siegel, 1959).

The conclusions to be drawn from Fig. 3 are rather obvious and will not be discussed. It is of interest, however, to point out that the irregularities appearing at ka values of about 1.0, 2.2, and 2.8 in the curve illustrating the present work, occur in a region where the calculated results are believed to be accurate, so it is considered that these are bona fide irregularities. It was thought that they might be caused by diffraction from the tip of the cone, but, when this effect was included in the Keller-theory calculations, the changes in the $\sigma$-values were too minute to be noticeable.
In order to look into possible resonance effects as the cause of these irregularities, the following table was made up.

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<th>2.2</th>
<th>2.8</th>
</tr>
</thead>
<tbody>
<tr>
<td>( k_b )</td>
<td>3.86</td>
<td>8.49</td>
<td>10.82</td>
</tr>
<tr>
<td>( \frac{2a}{\lambda} )</td>
<td>0.318</td>
<td>0.700</td>
<td>0.892</td>
</tr>
<tr>
<td>( \frac{b}{\lambda} )</td>
<td>0.614</td>
<td>1.352</td>
<td>1.723</td>
</tr>
<tr>
<td>( \frac{\pi a}{\lambda} )</td>
<td>0.590</td>
<td>1.10</td>
<td>1.40</td>
</tr>
<tr>
<td>( \frac{b+a}{\lambda} )</td>
<td>0.773</td>
<td>1.702</td>
<td>2.17</td>
</tr>
</tbody>
</table>

The values of \( \pi a/\lambda \) make it appear that these irregularities may be resonances in response caused by current paths from top to bottom of the cone along the edge of the base. In view of the fact that the base edge of the cone is very important in determining the scattering characteristics, it is not surprising that these resonance effects occur.
ACKNOWLEDGMENTS

One of the authors (F.V.S.) wishes to acknowledge the helpfulness of discussions with Dr. R. E. Kleinman and Dr. T. B. A. Senior of the Radiation Laboratory of the University of Michigan. The assistance of Dr. I. Marx and Mr. M. R. Halsey of Purdue University in the early phases of the work was also very valuable.

In addition, Dr. C. C. Rogers, now of Rose Polytechnic Institute, and Dr. J. K. Schindler, presently at the Air Force Cambridge Research Laboratories, made major contributions which are indispensable parts of the present report.
REFERENCES


Kleinman, R. E. and T. F. A. Senior (1963), "Diffraction and Scattering by Regular Bodies - II: the Cone", The University of Michigan Radiation Laboratory Report No. 348-2-T.


APPENDIX A

ANALYTIC SOLUTION FOR EXPANSION COEFFICIENTS

The solution for the expansion coefficients will now be presented in its analytic detail.

First we multiply (27) by \( P_m^l(\cos \theta) \), multiply (28) by \( \sin \theta \frac{dP_m^l}{d\theta} \), and subtract the two results, obtaining,

\[
\begin{align*}
\sum_{n=1}^{\infty} & \gamma_n h_n(kb) \left[ \frac{P_m^l P_n^l}{\sin \theta} + \sin \theta \frac{dP_m^l}{d\theta} \frac{dP_n^l}{d\theta} \right] \\
+ & \sum_{n=1}^{\infty} \gamma_n h'_n(kb) \left[ P_m^l \frac{dP_n^l}{d\theta} + \frac{dP_m^l}{d\theta} P_n^l \right] \\
+ & \sum_{n=1}^{\infty} c_n h_n(kb) \left[ \frac{P_m^l P_n^l}{\sin \theta} + \sin \theta \frac{dP_m^l}{d\theta} \frac{dP_n^l}{d\theta} \right] \\
+ & \sum_{n=1}^{\infty} d_n h'_n(kb) \left[ P_m^l \frac{dP_n^l}{d\theta} + \frac{dP_m^l}{d\theta} P_n^l \right] \\
+ & \sum_{\nu} \alpha_{\nu} \gamma_{\nu}(kb) \left[ \frac{P_m^l P_{\nu}^l}{\sin \theta} + \sin \theta \frac{dP_m^l}{d\theta} \frac{dP_{\nu}^l}{d\theta} \right] \\
= & \left\{ \begin{array}{l}
+ \sum_{\mu} b_{\mu} \gamma_{\mu}(kb) \left[ P_m^l \frac{dP_{\mu}^l}{d\theta} + \frac{dP_m^l}{d\theta} P_{\mu}^l \right], 0 \leq \theta < \theta_0 \\
0, \quad \theta_0 \leq \theta \leq \pi \end{array} \right\}
\]}

By integrating both sides of (1-1) with respect to \( \theta \) over the interval 0 to \( \pi \) and combining terms, we obtain,
\[ 
\int_{\theta=0}^{\theta=\pi} \left[ \frac{F_m^l P_n^l}{\sin \theta} + \sin \theta \frac{dF_m^l}{d\theta} \frac{dP_n^l}{d\theta} \right] d\theta = \int_{\theta=0}^{\theta=\pi} \left[ \frac{P_m^l P_n^l}{\sin \theta} + \sin \theta \frac{dP_m^l}{d\theta} \frac{dP_n^l}{d\theta} \right] d\theta = \text{constant}\]

The third integral is evaluated by utilizing the associated Legendre equation,
\[
\frac{d}{d\theta} \left( \sin \theta \frac{dP_1}{d\theta} \right) + n(n+1) \sin \theta \frac{P_1}{\sin \theta} - \frac{P_1}{n} = 0. 
\]  
(1-6)

Multiplying by \( \frac{P_1}{m} \) and using the product differentiation rule, we obtain

\[
\frac{P_1}{m} \frac{P_1}{n} \sin \theta + \sin \theta \frac{dP_1}{m \frac{m}{d\theta}} \frac{dP_1}{n \frac{d\theta}} = \frac{d}{d\theta} \left( \sin \theta \frac{P_1}{m} \frac{dP_1}{n \frac{d\theta}} \right) \cdot n(n+1) \sin \theta \frac{P_1}{m} \frac{P_1}{n} .
\]

(1-7)

Integrating from 0 to \( \theta_0 \), there results

\[
\int_0^{\theta_0} \left[ \frac{P_1}{m} \frac{P_1}{n} + \sin \theta \frac{dP_1}{m \frac{d\theta}} \frac{dP_1}{n \frac{d\theta}} \right] d\theta = \left[ \sin \theta \frac{P_1}{m} \frac{dP_1}{n \frac{d\theta}} \right]_0^{\theta_0} + \int_0^{\theta_0} n(n+1) \sin \theta \frac{P_1}{m} \frac{P_1}{n} d\theta .
\]

(1-8)

The integral in (1-8) can easily be evaluated by using (1-7). If \( m \) and \( n \) are interchanged in (1-7) and the result subtracted from (1-7), the ensuing equation is

\[
\left[ m(m+1) - n(n+1) \right] \sin \theta \frac{P_1}{m} \frac{P_1}{n} = \frac{d}{d\theta} \left[ \sin \theta \frac{P_1}{m} \frac{dP_1}{n \frac{d\theta}} - \sin \theta \frac{P_1}{n} \frac{dP_1}{m \frac{d\theta}} \right] .
\]

(1-9)

The integral in (1-8) may now be evaluated.

\[
\int_0^{\theta_0} n(n+1) \sin \theta \frac{P_1}{m} \frac{P_1}{n} d\theta = \frac{n(n+1)}{m(m+1) - n(n+1)} \left[ \sin \theta \left( \frac{P_1}{m} \frac{dP_1}{n \frac{d\theta}} - \frac{P_1}{n} \frac{dP_1}{m \frac{d\theta}} \right) \right]_0^{\theta_0} .
\]

(1-10)

Substituting (1-10) in (1-8) and combining terms, there results

\[
\int_0^{\theta_0} \left[ \frac{P_1}{m} \frac{P_1}{n} + \sin \theta \frac{dP_1}{m \frac{d\theta}} \frac{dP_1}{n \frac{d\theta}} \right] d\theta = \frac{n(n+1)}{m(m+1) - n(n+1)} \left[ \sin \theta \frac{P_1}{m} \frac{dP_1}{n \frac{d\theta}} \right]_0^{\theta_0} - \frac{n(n+1)}{m(m+1) - n(n+1)} \left[ \sin \theta \frac{P_1}{n} \frac{dP_1}{m \frac{d\theta}} \right]_0^{\theta_0} .
\]

(1-11)

The third integral in (1-2) is then obtained by letting \( n = \nu \) in (1-11) and making use of the condition (26).
The separation of $d_m$ is accomplished in a similar manner by multiplying (27) by $\sin \theta \frac{dF^l}{d\theta}$, multiplying (28) by $F^l_m$, and subtracting the two results, obtaining,
Integrating both sides of (1-15) with respect to $\theta$ over the interval 0 to $\pi$ and combining terms, we obtain,

$$
\int_{n=1}^{\infty} \left[ \gamma_n \jmath'_n(kb) + c_n h'_n(kb) \right] \int_0^{\pi} \left[ \frac{P_n^l}{P_n^m} + \frac{P_n^l}{P_n^m} \right] d\theta
$$

$$
+ \sum_{n=1}^{\infty} \left[ \Gamma_n \jmath'_n(kb) + \Delta h'_n(kb) \right] \int_0^{\pi} \left[ \frac{P_n^l}{P_n^m} + \sin \theta \frac{P_n^l}{P_n^m} \right] d\theta
$$

$$= \sum_{\nu} \gamma_{\nu} \jmath_{\nu}(kb) \int_0^{\theta_o} \left[ \frac{dP_n^l}{d\theta} \frac{P_n^l}{P_n^m} + \frac{P_n^l}{P_n^m} \frac{dP_n^l}{d\theta} \right] d\theta
$$

$$+ \sum_{\mu} \beta_{\mu} \jmath_{\mu}(kb) \int_0^{\theta_o} \left[ \frac{dP_n^l}{d\theta} \frac{P_n^l}{P_n^m} + \sin \theta \frac{P_n^l}{P_n^m} \frac{dP_n^l}{d\theta} \right] d\theta \ . \ \ (1-16)
$$

The values of the first and second integrals in (1-16) are given by (1-4) and (1-3), respectively. The third integral is easily evaluated.

$$
\int_0^{\theta_o} \left[ \frac{dP_n^l}{d\theta} \frac{P_n^l}{P_n^m} + \frac{P_n^l}{P_n^m} \frac{dP_n^l}{d\theta} \right] d\theta = \int_0^{\theta_o} d \left[ \frac{P_n^l}{P_n^m} \right] = P_n^l (\cos \theta_o) P_n^l (\cos \theta_o) \ \ (1-17)
$$

The fourth integral is evaluated by using (1-11) with $n$ replaced by $\mu$.

$$
\int_0^{\theta_o} \left[ \frac{P_n^l}{P_n^m} \frac{P_n^l}{P_n^m} + \sin \theta \frac{dP_n^l}{d\theta} \frac{dP_n^l}{d\theta} \right] d\theta = \frac{m(m+1)}{m(m+1)-\mu(\mu+1)} \sin \theta_o \left[ \frac{P_n^l}{P_n^m} \right] \ \ (1-18)
$$
By substituting these results into (1-16), we obtain

\[
\sum_{n=1}^{\infty} \left[ \Gamma_n j_n'(kb) + d_n h_n'(kb) \right] \frac{2[m(m+1)]^2}{2m+1} \delta_{mn}
\]

\[
= \sum_{\nu} a_{\nu} j_{\nu}(kb) \frac{P_{\nu}^1(\cos \theta_o)}{P_{\nu}^1(\cos \theta_o)}
+ \sum_{\mu} b_{\mu} j_{\mu}'(kb) \frac{m(m+1) \sin \theta}{m(m+1)-\mu(\mu+1)} \left[ P_{\mu}^1 \frac{dP_{\mu}^1}{d\theta} \right]_{\theta=\theta_o},
\]

(1-19)

which can be solved for \(d_m\):

\[
d_m = -\Gamma_m j_m'(kb) + \frac{(2m+1) P_m^1(\cos \theta_o)}{2m(m+1) h_m'(kb)} \sum_{\nu} a_{\nu} j_{\nu}(kb) \frac{P_{\nu}^1(\cos \theta_o)}{P_{\nu}^1(\cos \theta_o)}
+ \frac{(2m+1) \sin \theta_o}{2m(m+1) h_m'(kb)} \sum_{\mu} b_{\mu} j_{\mu}'(kb) \frac{dP_{\mu}^1}{d\theta} \bigg|_{\theta=\theta_o}.
\]

(1-20)

The separation of \(a_{\nu}\) is accomplished by multiplying (29) by

\[
\sin \theta \frac{dP_{\alpha}^1}{d\theta},
\]

multiplying (30) by \(P_{\alpha}^1\) and adding the results, obtaining,

\[
\sum_{n=1}^{\infty} \Gamma_n j_n'(kb) \left[ \frac{P_{\alpha}^1 P_n^1}{\sin \theta} + \sin \theta \frac{dP_{\alpha}^1}{d\theta} \frac{dP_n^1}{d\theta} \right]
\]

\[
- \sum_{n=1}^{\infty} \Gamma_n j_n'(kb) \left[ \frac{P_{\alpha}^1 dP_n^1}{d\theta} + \frac{dP_{\alpha}^1}{d\theta} P_n^1 \right]
\]

\[
+ \sum_{n=1}^{\infty} c_n h_n'(kb) \left[ \frac{P_{\alpha}^1 P_n^1}{\sin \theta} + \sin \theta \frac{dP_{\alpha}^1}{d\theta} \frac{dP_n^1}{d\theta} \right]
\]
\[
\sum_{n=1}^{\infty} d_n h_n(k\beta) \left[ \frac{d^4}{d\alpha^4} p_n + \frac{d^2}{d\alpha^2} p_n \right] \\
= \sum_{n} a_n j_n(k\beta) \left[ \frac{d^4}{d\alpha^4} p_n + \frac{d^2}{d\alpha^2} p_n \right]
\]

where the subscript \( \alpha \) denotes a particular value of the infinite set of \( \nu \). Integrating both sides of (1-21) with respect to \( \theta \) over the interval \( 0 \) to \( \theta_o \) and combining terms, we obtain,

\[
\sum_{n=1}^{\infty} \left[ w_n j_n(k\beta) + c_n h_n(k\beta) \right] \int_0^{\theta} \left[ \frac{d^4}{d\alpha^4} p_n + \sin \theta \frac{d^2}{d\alpha^2} p_n \right] d\theta \\
- \sum_{n=1}^{\infty} \left[ \frac{d^4}{d\alpha^4} h_n(k\beta) + d_n h_n(k\beta) \right] \int_0^{\theta} \left[ \frac{d^4}{d\alpha^4} p_n + \frac{d^2}{d\alpha^2} p_n \right] d\theta
\]

\[
= \sum_{n} a_n j_n(k\beta) \int_0^{\theta} \left[ \frac{d^4}{d\alpha^4} p_n + \sin \theta \frac{d^2}{d\alpha^2} p_n \right] d\theta \\
- \sum_{n} b_n j_n(k\beta) \int_0^{\theta} \left[ \frac{d^4}{d\alpha^4} p_n + \frac{d^2}{d\alpha^2} p_n \right] d\theta.
\]

(1-22)

The first integral in (1-22) is evaluated by using (1-11), replacing \( m \) by \( \alpha \) and noting that \( \alpha \) is a particular value of the set of \( \nu \).

\[
\int_0^{\theta} \left[ \frac{d^4}{d\alpha^4} p_n + \sin \theta \frac{d^2}{d\alpha^2} p_n \right] d\theta = \frac{\alpha(\alpha+1) \sin \theta \alpha}{\alpha(\alpha+1)-n(n+1)} \left[ \frac{d^4}{d\alpha^4} p_n \right]_{\theta=\theta_o}
\]

(1-23)
The second and fourth integrals are easily evaluated.

\[
\int_0^\theta \left[ P_\alpha^1 \frac{dP_\mu^1}{d\theta} + \frac{dP_\alpha^1}{d\theta} P_\mu^1 \right] d\theta = \int_0^\theta d \left[ P_\alpha^1 P_\mu^1 \right] = P_\alpha^1 (\cos \theta) P_\mu^1 (\cos \theta)
\]

\[ (1-24) \]

\[
\int_0^\theta \left[ P_\alpha^1 \frac{dP_\mu^1}{d\theta} + \frac{dP_\alpha^1}{d\theta} P_\mu^1 \right] d\theta = \int_0^\theta d \left[ P_\alpha^1 P_\mu^1 \right] = \left[ P_\alpha^1 P_\mu^1 \right] \bigg|_0^\theta = 0.
\]

\[ (1-25) \]

The third integral is evaluated by using (1-11),

\[
\int_0^\theta \left[ \frac{P_\alpha^1}{\sin \theta} P_\mu^1 + \sin \theta \frac{dP_\alpha^1}{d\theta} \frac{dP_\mu^1}{d\theta} \right] d\theta = \delta_{\alpha\nu} \int_0^\theta \left[ \frac{(P_\alpha^1)^2}{\sin \theta} + \sin \theta \left( \frac{dP_\alpha^1}{d\theta} \right)^2 \right] d\theta,
\]

\[ (1-26) \]

where \( \delta_{\alpha\nu} \) is the Kronecker delta. The integral in (1-26) can be evaluated by the use of (1-7). If (1-7) is re-written with \( m \) and \( n \) replaced by \( \alpha \), there results

\[
\frac{(P_\alpha^1)^2}{\sin \theta} + \sin \theta \left( \frac{dP_\alpha^1}{d\theta} \right)^2 = \frac{d}{d\theta} \left( \sin \theta P_\alpha^1 \frac{dP_\alpha^1}{d\theta} \right) + \alpha(\alpha+1) \sin \theta (P_\alpha^1)^2.
\]

\[ (1-27) \]

The integral in (1-26) is then

\[
\int_0^\theta \left[ \frac{(P_\alpha^1)^2}{\sin \theta} + \sin \theta \left( \frac{dP_\alpha^1}{d\theta} \right)^2 \right] d\theta = \left( \sin \theta P_\alpha^1 \frac{dP_\alpha^1}{d\theta} \right) \bigg|_0^\theta + \alpha(\alpha+1) \int_0^\theta \sin \theta (P_\alpha^1)^2 \, d\theta.
\]

\[ (1-28) \]

The first term on the right side of (1-28) is seen to vanish, and then, if we define

\[
B_\alpha = \int_0^\theta \sin \theta (P_\alpha^1)^2 \, d\theta,
\]

\[ (1-29) \]
for (1-26) there results

\[ \int_{0}^{\theta_0} \left[ \frac{\sin \theta}{\sin \theta} \frac{dP^1_{\alpha}}{d\theta} + \sin \theta \frac{dP^1_{\nu}}{d\theta} \right] d\theta \equiv \sigma(\alpha+1) B_{\alpha} \delta_{\alpha\nu} . \]  

(1-30)

Upon substituting these results in (1-22), we obtain

\[ \sum_{n=1}^{\infty} \left[ \gamma_{n-1}^\prime(kb) + c_n h_n^\prime(kb) \right] \sigma(\alpha+1) \sin \theta_0 \left[ \frac{\sin \theta}{\sin(\alpha+1)-n(n+1)} \frac{dP^1_{\alpha}}{d\theta} \right] \theta=\theta_0 \]

\[ = - \sum_{n=1}^{\infty} \left[ \gamma_{n-1}^\prime(kb) + d_n h_n^\prime(kb) \right] P^1_{\alpha} (\cos \theta_0) P^1_n (\cos \theta_0) \]

\[ = \sum_{n=1}^{\infty} a_{\nu} j_{n-1}^\prime(kb) \sigma(\alpha+1) B_{\alpha} \delta_{\alpha\nu} , \]  

(1-31)

and solving for \( a_{\alpha} \),

\[ a_{\alpha} = \frac{\sin \theta_0}{B_{\alpha} j_{\alpha}^\prime(kb)} \sum_{n=1}^{\infty} \frac{\gamma_{n-1}^\prime(kb) + c_n h_n^\prime(kb)}{\sigma(\alpha+1)-n(n+1)} \frac{dP^1_n}{d\theta} |_{\theta=\theta_0} \]

\[ = - \frac{\sigma(\alpha+1) B_{\alpha} j_{\alpha}^\prime(kb)}{P^1_{\alpha}(\cos \theta_0)} \sum_{n=1}^{\infty} \left[ \gamma_{n-1}^\prime(kb) + d_n h_n^\prime(kb) \right] P^1_n (\cos \theta_0) . \]  

(1-32)

The last coefficient to be separated is \( b, \). This is accomplished by multiplying (29) by \( P^1_{\beta} \), multiplying (30) by \( \sin \theta \frac{dP^1_{\beta}}{d\theta} \), and adding the results, obtaining,

\[ \sum_{n=1}^{\infty} \gamma_{n-1}^\prime(kb) \left[ \frac{P^1_{\beta} dP^1_{n}}{d\theta} + \frac{dP^1_{\beta}}{d\theta} P^1_{n} \right] \]

\[ - \sum_{n=1}^{\infty} \gamma_{n-1}^\prime(kb) \left[ \frac{P^1_{\beta} P^1_{n} + \sin \theta \frac{dP^1_{\beta}}{d\theta} P^1_{n}}{\sin \theta} \frac{dP^1_{n}}{d\theta} \right] \]
where the subscript $\beta$ denotes a particular value of the infinite set $\mu$. Upon integrating both sides of (1-33) with respect to $\theta$ over the interval $0$ to $\theta_0$ and combining terms, we obtain,

$$
\sum_{n=1}^{\infty} \left[ \gamma_n \gamma_n' + c_n h_n' \right] \int_0^{\theta_0} \left[ \frac{P_1}{\beta} \frac{dP_1}{d\theta} + \frac{\partial P_1}{\partial \theta} P_1 \right] d\theta 
$$

$$
- \sum_{n=1}^{\infty} \left[ \Gamma_n \Gamma_n' + d_n h_n' \right] \int_0^{\theta_0} \left[ \frac{P_1}{\beta} \frac{dP_1}{d\theta} + \frac{\partial P_1}{\partial \theta} P_1 \right] d\theta 
$$

$$
= \sum_{\nu=1}^{\infty} s_{\nu} s_{\nu}' \int_0^{\theta_0} \left[ \frac{P_1}{\beta} \frac{dP_1}{d\theta} + \frac{\partial P_1}{\partial \theta} P_1 \right] d\theta 
$$

$$
- \sum_{\mu=1}^{\infty} b_{\mu} b_{\mu}' \int_0^{\theta_0} \left[ \frac{P_1}{\beta} \frac{dP_1}{d\theta} + \frac{\partial P_1}{\partial \theta} P_1 \right] d\theta . \quad (1-34)
$$

The first integral in (1-34) is easily evaluated, noting that $\beta$ is a particular value of the set $\mu$. 
The second integral is evaluated by using (1-11) and replacing \( \mu \) by \( \beta \), obtaining,

\[
\int_0^{\theta_0} \left[ \frac{\partial \mathbf{r}^\perp}{\partial \theta} \mathbf{F}^\perp_n + \sin \theta \frac{\partial \mathbf{r}^\perp}{\partial \theta} \mathbf{F}^\perp_\nu \right] d\theta = \int_0^{\theta_0} d\left[ \frac{\partial \mathbf{r}^\perp_n}{\partial \theta} \mathbf{F}^\perp_n \right] d\theta = \left[ \frac{\partial \mathbf{r}^\perp_n}{\partial \theta} \mathbf{F}^\perp_n \right]_0^{\theta_0} = 0
\]  

(1-35)

The third integral in (1-34) also is elementary.

\[
\int_0^{\theta_0} \left[ \frac{\partial \mathbf{r}^\perp}{\partial \theta} \mathbf{F}^\perp_\nu \right] d\theta = \int_0^{\theta_0} d\left[ \frac{\partial \mathbf{r}^\perp_\nu}{\partial \theta} \mathbf{F}^\perp_\nu \right] d\theta = \left[ \frac{\partial \mathbf{r}^\perp_\nu}{\partial \theta} \mathbf{F}^\perp_\nu \right]_0^{\theta_0} = 0
\]  

(1-36)

The fourth integral is evaluated by using (1-11),

\[
\int_0^{\theta_0} \left[ \frac{\partial \mathbf{r}^\perp}{\partial \theta} \mathbf{F}^\perp_n + \sin \theta \frac{\partial \mathbf{r}^\perp}{\partial \theta} \mathbf{F}^\perp_\nu \right] d\theta = \delta_{\beta \mu} \int_0^{\theta_0} \left[ \frac{(\mathbf{r}^\perp_\beta)^2}{\sin \theta} + \sin \theta \left( \frac{\partial \mathbf{r}^\perp_\beta}{\partial \theta} \right)^2 \right] d\theta ,
\]  

(1-37)

where \( \delta_{\beta \mu} \) is the Kronecker delta. The integral in (1-38) is evaluated by using (1-28) with \( \mu \) replaced by \( \beta \).

\[
\int_0^{\theta_0} \left[ \frac{(\mathbf{r}^\perp_\beta)^2}{\sin \theta} + \sin \theta \left( \frac{\partial \mathbf{r}^\perp_\beta}{\partial \theta} \right)^2 \right] d\theta = \left( \frac{\partial \mathbf{r}^\perp_\beta}{\partial \theta} \right)^2 \left[ \frac{(\mathbf{r}^\perp_\beta)^2}{\sin \theta} \right]_0^{\theta_0} + \beta \cdot (\mathbf{r}^\perp_\beta)^2 \int_0^{\theta_0} \sin \theta \left( \frac{\partial \mathbf{r}^\perp_\beta}{\partial \theta} \right)^2 d\theta
\]  

(1-39)

The first term on the right side of (1-39) is seen to vanish, and then, if we define

\[
B_\beta = \int_0^{\theta_0} \sin \theta \left( \frac{\partial \mathbf{r}^\perp_\beta}{\partial \theta} \right)^2 d\theta ,
\]  

(1-40)
for (1-38) there results

\[ \int_0^{\theta_0} \left[ \frac{B_1}{\sin \theta} + \sin \theta \frac{dP_1}{d\theta} \frac{dP_2}{d\theta} \right] d\theta = \beta(\beta+1) B_{\beta} \delta_{\beta, \mu}. \]  

(1-41)

Upon substituting the results of these integrations into (1-34), we obtain

\[ -\sum_{\mu} b_{\mu} \left( k_b \right) \beta(\beta+1) B_{\beta} \delta_{\beta, \mu}, \]  

(1-42)

and solving for \( b_{\beta} \),

\[ b_{\beta} = \frac{\sin \theta_0}{\beta(\beta+1) B_{\beta}} \left| \frac{dP_1}{d\theta} \right|_{\theta=\theta_0} \sum_{n=1}^{\infty} \frac{\left( n \right)^n \left( k_b \right) + d_n \left( k_b \right) }{n(n+1)-\beta(\beta+1)} \left| \frac{dP_1}{d\theta} \right|_{\theta=\theta_0} P_n^1 \cos \theta_0. \]  

(1-43)

Equations (1-14), (1-20), (1-32), and (1-43) represent the formal solution for the expansion co-efficients, and, when the values of \( \gamma_n \) and \( \Gamma_n \) are substituted by using (8), are equivalent to (31) through (34).
APPENDIX B

LEGENDRE FUNCTION CONSTANTS

<table>
<thead>
<tr>
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(1) Donated by Dr. F. C. Waterman of AVCO.

(2) Determined from \(P_u^1(\cos 165^\circ) = 0\).

(3) Determined from \(\left. \frac{dP_v^1}{d\theta} \right|_{\theta=165^\circ} = 0\).

(4) Defined by \(B_\gamma = \int_{165^\circ}^{\theta} \sin \theta \ (P_\gamma^1)^2 \ d\theta\).
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APPENDIX C

DEFINITIONS OF MATRICE ELEMENTS

The elements of the matrices $F$ (36) through (41) are given by (31) through (34). If we let $[A]_{ij}$ denote the $j$th element of the $i$th row of a matrix $A$, then the matrix elements are defined by

$$\begin{align*}
\left[ E_{1} \right]_{m} &= \frac{-i^{m} (2m+1) J_{m}^{\prime}(kt)}{m(m+1) n_{m}(kt)} \\
\left[ F_{0} \right]_{mn} &= \frac{(2m+1) \sin \theta_{o} \frac{dF_{m}^{\prime}}{d\theta} \bigg|_{\theta=\theta_{o}} \left[ J_{m}^{\prime}(kt) - J_{m}^{\prime}(\nu+1) \right] F_{\nu}^{\prime}(\cos \theta_{o})}{2^{m(m+1)} n_{m}(kt) \left[ \nu_{m}(\nu+1) - m(m+1) \right]} \\
\left[ F_{1} \right]_{m} &= \frac{i^{m+1} (2m+1) J_{m+1}^{\prime}(kt)}{m(m+1) n_{m}(kt)} \\
\left[ F_{2} \right]_{mn} &= \frac{(2m+1) F_{m}^{\prime}(\cos \nu_{c}) J_{\nu_{c}}^{\prime}(kt) F_{\nu}^{\prime}(\cos \theta_{o})}{2^{m(m+1)} n_{m}(kt)} \\
\left[ F_{3} \right]_{mn} &= \frac{(2m+1) \sin \theta_{o} F_{m}^{\prime}(\cos \theta_{o}) J_{m}^{\prime}(kt) \frac{dF_{m}^{\prime}}{d\theta} \bigg|_{\theta=\theta_{o}}}{2^{m(m+1)} n_{m}(kt) \left[ \nu_{m}(\nu+1) - n(n+1) \right]} \\
\left[ G_{1} \right]_{mn} &= \frac{\sin \nu_{c} F_{m}^{\prime}(\cos \theta_{o}) h_{m}^{\prime}(kt) \frac{dF_{m}^{\prime}}{d\theta} \bigg|_{\theta=\theta_{o}}}{n_{m}(kt) \left[ \nu_{m}(\nu+1) - n(n+1) \right]} \\
\left[ G_{2} \right]_{mn} &= \frac{\sin \theta_{o} F_{m}^{\prime}(\cos \theta_{o})}{J_{\nu_{c}}^{\prime}(kt) F_{\nu}^{\prime}(\nu)} \sum_{n=1}^{s_{c}} F_{n}^{\prime}(2n+1) \frac{dF_{n}^{\prime}}{d\theta} \bigg|_{\theta=\theta_{o}} \frac{dF_{1}^{\prime}}{d\theta} \bigg|_{\theta=\theta_{o}}}{n(n+1) \left[ \nu_{m}(\nu+1) - n(n+1) \right]} 
\end{align*}$$

(3-1) (3-2) (3-3) (3-4) (3-5) (3-6) (3-7)
\[
\begin{align*}
\{G_3\}_{mn} &= - \frac{P^l_m (\cos \theta_o) J^m_n (kb) P^l_{n}(\cos \theta_o)}{v_m (v + 1) B_{v_m} J^m_{v_m} (kb)} \\
\{G_4\}_{mn} &= \frac{P^l_m (\cos \theta_o)}{v_m (v + 1) B_{v_m} J^m_{v_m} (kb)} \sum_{n=1}^{s_o} \frac{I_{n+1}^{(2n+1)} J_n (kb) P^l_{n}(\cos \theta_o)}{n(n+1)} \\
\{H_1\}_{mn} &= \frac{\sin \theta_o}{\cos \theta_o} \frac{dP^l_m}{d\theta} \left. \right|_{\theta = \theta_o} h_m (kb) n(n+1) P^l_r (\cos \theta_o) \\
\{H_2\}_{mn} &= - \frac{\sin \theta_o}{\cos \theta_o} \frac{dP^l_m}{d\theta} \left. \right|_{\theta = \theta_o} h_m (kb) \sum_{n=1}^{s_o} \frac{I_{n+1}^{(2n+1)} J_n (kb) P^l_{n}(\cos \theta_o)}{n(n+1) - u_m (\mu_m + 1)} \\
\end{align*}
\]

\( u_m \) and \( v_m \) represent the \( m \)th value of the sets \( u \) and \( v \), respectively.